## Chaos: Classical and Quantum Part I: Deterministic Chaos



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No man but a blockhead ever wrote except for money Samuel Johnson

This book is a result of collaborative labors of many people over a span of several decades. Coauthors of a chapter or a section are indicated in the byline to the chapter/section title. If you are referring to a specific coauthored section rather than the entire book, cite it as (for example):
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## Acknowledgements

I feel I never want to write another book. What's the good! I can eke living on stories and little articles, that don't cost a tithe of the output a book costs. Why write novels any more!
D.H. Lawrence

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Fritz Haake's heartfelt lament on page 239 was uttered at the end of the first conference presentation of cycle expansions, in 1988. Joseph Ford introduced himself to the authors of this book by the email quoted on page 457. G.P. Morriss advice to students as how to read the introduction to this book, page 4, was offerred during a 2002 graduate course in Dresden. Kerson Huang's interview of C.N. Yang quoted on page 128 is available on ChaosBook.org/extras.

Who is the 3-legged dog reappearing throughout the book? Long ago, when we were innocent and knew not Borel measurable $\alpha$ to $\Omega$ sets, P. Cvitanović asked V. Baladi a question about dynamical zeta functions, who then asked J.-P. Eckmann, who then asked D. Ruelle. The answer was transmitted back: "The master says: 'It is holomorphic in a strip' ". Hence His Master's Voice logo, and the 3-legged dog is us, still eager to fetch the bone. The answer has made it to the book, though not precisely in His Master's voice. As a matter of fact, the answer is the book. We are still chewing on it.

Profound thanks to all the unsung heroes - students and colleagues, too
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## Chapter 1

## Overture

If I have seen less far than other men it is because I have stood behind giants.
Edoardo Specchio
Rereading classic theoretical physics textbooks leaves a sense that there are holes large enough to steam a Eurostar train through them. Here we learn about harmonic oscillators and Keplerian ellipses - but where is the chapter on chaotic oscillators, the tumbling Hyperion? We have just quantized hydrogen, where is the chapter on the classical 3-body problem and its implications for quantization of helium? We have learned that an instanton is a solution of field-theoretic equations of motion, but shouldn't a strongly nonlinear field theory have turbulent solutions? How are we to think about systems where things fall apart; the center cannot hold; every trajectory is unstable?

This chapter offers a quick survey of the main topics covered in the book. We start out by making promises - we will right wrongs, no longer shall you suffer the slings and arrows of outrageous Science of Perplexity. We relegate a historical overview of the development of chaotic dynamics to appendix A, and head straight to the starting line: A pinball game is used to motivate and illustrate most of the concepts to be developed in this book.

Throughout the book
indicates that the section requires a hearty stomach and is probably best skipped on first reading

fast track points you where to skip to
tells you where to go for more depth on a particular topic
indicates an exercise that might clarify a point in the text
indicates that a figure is still missing - you are urged to fetch it

This is a textbook, not a research monograph, and you should be able to follow the thread of the argument without constant excursions to sources. Hence there are no literature references in the text proper, all learned remarks and bibliographical pointers are relegated to the "Commentary" section at the end of each chapter.

### 1.1 Why this book?

> It seems sometimes that through a preoccupation with science, we acquire a firmer hold over the vicissitudes of life and meet them with greater calm, but in reality we have done no more than to find a way to escape from our sorrows.
> Hermann Minkowski in a letter to David Hilbert

The problem has been with us since Newton's first frustrating (and unsuccessful) crack at the 3-body problem, lunar dynamics. Nature is rich in systems governed by simple deterministic laws whose asymptotic dynamics are complex beyond belief, systems which are locally unstable (almost) everywhere but globally recurrent. How do we describe their long term dynamics?

The answer turns out to be that we have to evaluate a determinant, take a logarithm. It would hardly merit a learned treatise, were it not for the fact that this determinant that we are to compute is fashioned out of infinitely many infinitely small pieces. The feel is of statistical mechanics, and that is how the problem was solved; in the 1960's the pieces were counted, and in the 1970's they were weighted and assembled in a fashion that in beauty and in depth ranks along with thermodynamics, partition functions and path integrals amongst the crown jewels of theoretical physics.

Then something happened that might be without parallel; this is an area of science where the advent of cheap computation had actually subtracted from our collective understanding. The computer pictures and numerical plots of fractal science of the 1980's have overshadowed the deep insights of the 1970's, and these pictures have since migrated into textbooks. Fractal science posits that certain quantities (Lyapunov exponents, generalized dimensions, ...) can be estimated on a computer. While some of the numbers so obtained are indeed mathematically sensible characterizations of fractals, they are in no sense observable and measurable on the length-scales and time-scales dominated by chaotic dynamics.

Even though the experimental evidence for the fractal geometry of nature is circumstantial, in studies of probabilistically assembled fractal aggregates we know of nothing better than contemplating such quantities.

In deterministic systems we can do much better. Chaotic dynamics is generated by the interplay of locally unstable motions, and the interweaving of their global stable and unstable manifolds. These features are robust and accessible in systems as noisy as slices of rat brains. Poincaré, the first to understand deterministic chaos, already said as much (modulo rat brains). Once the topology of chaotic dynamics is understood, a powerful theory yields the macroscopically measurable consequences of chaotic dynamics, such as atomic spectra, transport coefficients, gas pressures.

That is what we will focus on in this book. This book is a self-contained graduate textbook on classical and quantum chaos. We teach you how to evaluate a determinant, take a logarithm - stuff like that. Ideally, this should take 100 pages or so. Well, we fail - so far we have not found a way to traverse this material in less than a semester, or 200-300 page subset of this text. Nothing can be done about that.

### 1.2 Chaos ahead

Things fall apart; the centre cannot hold.
W.B. Yeats: The Second Coming

The study of chaotic dynamical systems is no recent fashion. It did not start with the widespread use of the personal computer. Chaotic systems have been studied for over 200 years. During this time many have contributed, and the field followed no single line of development; rather one sees many interwoven strands of progress.

In retrospect many triumphs of both classical and quantum physics seem a stroke of luck: a few integrable problems, such as the harmonic oscillator and the Kepler problem, though "non-generic", have gotten us very far. The success has lulled us into a habit of expecting simple solutions to simple equations - an expectation tempered for many by the recently acquired ability to numerically scan the phase space of non-integrable dynamical systems. The initial impression might be that all of our analytic tools have failed us, and that the chaotic systems are amenable only to numerical and statistical investigations. Nevertheless, a beautiful theory of deterministic chaos, of predictive quality comparable to that of the traditional perturbation expansions for nearly integrable systems, already exists.

In the traditional approach the integrable motions are used as zerothorder approximations to physical systems, and weak nonlinearities are then accounted for perturbatively. For strongly nonlinear, non-integrable systems such expansions fail completely; at asymptotic times the dynamics exhibits amazingly rich structure which is not at all apparent in the integrable approximations. However, hidden in this apparent chaos is a rigid skeleton, a self-similar tree of cycles (periodic orbits) of increasing lengths. The insight of the modern dynamical systems theory is that the zeroth-order approximations to the harshly chaotic dynamics should be very different

Figure 1.1: A physicist's bare bones game of pinball.

from those for the nearly integrable systems: a good starting approximation here is the linear stretching and folding of a baker's map, rather than the periodic motion of a harmonic oscillator.

So, what is chaos, and what is to be done about it? To get some feeling for how and why unstable cycles come about, we start by playing a game of pinball. The reminder of the chapter is a quick tour through the material covered in this book. Do not worry if you do not understand every detail at the first reading - the intention is to give you a feeling for the main themes of the book. Details will be filled out later. If you want to get a particular point clarified right now, $\mathbb{l} 8$ on the margin points at the appropriate section.

### 1.3 The future as in a mirror

All you need to know about chaos is contained in the introduction of the [Cvitanović et al. "Chaos: Classical and Quantum"] book. However, in order to understand the introduction you will first have to read the rest of the book.
Gary Morriss
That deterministic dynamics leads to chaos is no surprise to anyone who has tried pool, billiards or snooker - the game is about beating chaos so we start our story about what chaos is, and what to do about it, with a game of pinball. This might seem a trifle, but the game of pinball is to chaotic dynamics what a pendulum is to integrable systems: thinking clearly about what "chaos" in a game of pinball is will help us tackle more difficult problems, such as computing diffusion constants in deterministic gases, or computing the helium spectrum.

We all have an intuitive feeling for what a ball does as it bounces among the pinball machine's disks, and only high-school level Euclidean geometry is needed to describe its trajectory. A physicist's pinball game is the game of pinball stripped to its bare essentials: three equidistantly placed reflecting disks in a plane, figure 1.1. A physicist's pinball is free, frictionless, pointlike, spin-less, perfectly elastic, and noiseless. Point-like pinballs are shot at the disks from random starting positions and angles; they spend some time bouncing between the disks and then escape.

At the beginning of the 18th century Baron Gottfried Wilhelm Leibniz was confident that given the initial conditions one knew everything a deterministic system would do far into the future. He wrote [1.1], anticipating by a century and a half the oft-quoted Laplace's "Given for one instant an intelligence which could comprehend all the forces by which nature is animated...":

That everything is brought forth through an established destiny is just as certain as that three times three is nine. [...] If, for example, one sphere meets another sphere in free space and if their sizes and their paths and directions before collision are known, we can then foretell and calculate how they will rebound and what course they will take after the impact. Very simple laws are followed which also apply, no matter how many spheres are taken or whether objects are taken other than spheres. From this one sees then that everything proceeds mathematically - that is, infallibly - in the whole wide world, so that if someone could have a sufficient insight into the inner parts of things, and in addition had remembrance and intelligence enough to consider all the circumstances and to take them into account, he would be a prophet and would see the future in the present as in a mirror.

Leibniz chose to illustrate his faith in determinism precisely with the type of physical system that we shall use here as a paradigm of "chaos". His claim is wrong in a deep and subtle way: a state of a physical system can never be specified to infinite precision, there is no way to take all the circumstances into account, and a single trajectory cannot be tracked, only a ball of nearby initial points makes physical sense.

### 1.3.1 What is "chaos"?

I accept chaos. I am not sure that it accepts me. Bob Dylan, Bringing It All Back Home

A deterministic system is a system whose present state is in principle fully determined by its initial conditions, in contrast to a stochastic system, for which the initial conditions determine the present state only partially, due to noise, or other external circumstances beyond our control. For a stochastic system, the present state reflects the past initial conditions plus the particular realization of the noise encountered along the way.

A deterministic system with sufficiently complicated dynamics can fool us into regarding it as a stochastic one; disentangling the deterministic from the stochastic is the main challenge in many real-life settings, from stock markets to palpitations of chicken hearts. So, what is "chaos"?

In a game of pinball, any two trajectories that start out very close to each other separate exponentially with time, and in a finite (and in practice, a very small) number of bounces their separation $\delta \mathbf{x}(t)$ attains the magnitude of $L$, the characteristic linear extent of the whole system, figure 1.2.

Figure 1.2: Sensitivity to initial conditions: two pinballs that start out very close to each other separate exponentially with time.


This property of sensitivity to initial conditions can be quantified as

$$
|\delta \mathbf{x}(t)| \approx e^{\lambda t}|\delta \mathbf{x}(0)|
$$

where $\lambda$, the mean rate of separation of trajectories of the system, is called the Lyapunov exponent. For any finite accuracy $\delta x=|\delta \mathbf{x}(0)|$ of the initial data, the dynamics is predictable only up to a finite Lyapunov time

$$
\begin{equation*}
T_{\text {Lyap }} \approx-\frac{1}{\lambda} \ln |\delta x / L|, \tag{1.1}
\end{equation*}
$$

despite the deterministic and, for Baron Leibniz, infallible simple laws that rule the pinball motion.

A positive Lyapunov exponent does not in itself lead to chaos. One could try to play 1- or 2-disk pinball game, but it would not be much of a game; trajectories would only separate, never to meet again. What is also needed is mixing, the coming together again and again of trajectories. While locally the nearby trajectories separate, the interesting dynamics is confined to a globally finite region of the phase space and thus the separated trajectories are necessarily folded back and can re-approach each other arbitrarily closely, infinitely many times. For the case at hand there are $2^{n}$ topologically distinct $n$ bounce trajectories that originate from a given disk. More generally, the number of distinct trajectories with $n$ bounces can be quantified as

$$
N(n) \approx e^{h n}
$$

where the topological entropy $h(h=\ln 2$ in the case at hand $)$ is the growth rate of the number of topologically distinct trajectories.

The appellation "chaos" is a confusing misnomer, as in deterministic dynamics there is no chaos in the everyday sense of the word; everything proceeds mathematically - that is, as Baron Leibniz would have it, infallibly. When a physicist says that a certain system exhibits "chaos", he means that the system obeys deterministic laws of evolution, but that the outcome


Figure 1.3: Dynamics of a chaotic dynamical system is (a) everywhere locally unstable (positive Lyapunov exponent) and (b) globally mixing (positive entropy). (A. Johansen)
is highly sensitive to small uncertainties in the specification of the initial state. The word "chaos" has in this context taken on a narrow technical meaning. If a deterministic system is locally unstable (positive Lyapunov exponent) and globally mixing (positive entropy) - figure 1.3 - it is said to be chaotic.

While mathematically correct, the definition of chaos as "positive Lyapunov + positive entropy" is useless in practice, as a measurement of these quantities is intrinsically asymptotic and beyond reach for systems observed in nature. More powerful is Poincare's vision of chaos as the interplay of local instability (unstable periodic orbits) and global mixing (intertwining of their stable and unstable manifolds). In a chaotic system any open ball of initial conditions, no matter how small, will in finite time overlap with any other finite region and in this sense spread over the extent of the entire asymptotically accessible phase space. Once this is grasped, the focus of theory shifts from attempting to predict individual trajectories (which is impossible) to a description of the geometry of the space of possible outcomes, and evaluation of averages over this space. How this is accomplished is what this book is about.

A definition of "turbulence" is even harder to come by. Intuitively, the word refers to irregular behavior of an infinite-dimensional dynamical system described by deterministic equations of motion - say, a bucket of boiling water described by the Navier-Stokes equations. But in practice the word "turbulence" tends to refer to messy dynamics which we understand poorly. As soon as a phenomenon is understood better, it is reclaimed and renamed: "a route to chaos", "spatiotemporal chaos", and so on.

In this book we shall develop a theory of chaotic dynamics for low dimensional attractors visualized as a succession of nearly periodic but unstable motions. In the same spirit, we shall think of turbulence in spatially extended systems in terms of recurrent spatiotemporal patterns. Pictorially, dynamics drives a given spatially extended system through a repertoire of unstable patterns; as we watch a turbulent system evolve, every so often we catch a glimpse of a familiar pattern:


For any finite spatial resolution, the system follows approximately for a finite time a pattern belonging to a finite alphabet of admissible patterns, and the long term dynamics can be thought of as a walk through the space of such patterns. In this book we recast this image into mathematics.

### 1.3.2 When does "chaos" matter?

> Whether 'tis nobler in the mind to suffer The slings and arrows of outrageous fortune, Or to take arms against a sea of troubles, And by opposing end them? W. Shakespeare, Hamlet

When should we be mindful of chaos? The solar system is "chaotic", yet we have no trouble keeping track of the annual motions of planets. The rule of thumb is this; if the Lyapunov time (1.1) (the time by which a phase space region initially comparable in size to the observational accuracy extends across the entire accessible phase space) is significantly shorter than the observational time, you need to master the theory that will be developed here. That is why the main successes of the theory are in statistical mechanics, quantum mechanics, and questions of long term stability in celestial mechanics.

In science popularizations too much has been made of the impact of "chaos theory", so a number of caveats are already needed at this point.

At present the theory is in practice applicable only to systems with a low intrinsic dimension - the minimum number of coordinates necessary to capture its essential dynamics. If the system is very turbulent (a description of its long time dynamics requires a space of high intrinsic dimension) we are out of luck. Hence insights that the theory offers in elucidating problems of fully developed turbulence, quantum field theory of strong interactions and early cosmology have been modest at best. Even that is a caveat with qualifications. There are applications - such as spatially extended (nonequilibrium) systems and statistical mechanics applications where the few important degrees of freedom can be isolated and studied profitably by methods to be described here.

Thus far the theory has had limited practical success when applied to the very noisy systems so important in the life sciences and in economics. Even though we are often interested in phenomena taking place on time scales much longer than the intrinsic time scale (neuronal interburst intervals, cardiac pulses, etc.), disentangling "chaotic" motions from the environmental noise has been very hard.

### 1.4 A game of pinball

Formulas hamper the understanding.
S. Smale

We are now going to get down to the brasstacks. But first, a disclaimer: If you understand most of the rest of this chapter on the first reading, you either do not need this book, or you are delusional. If you do not understand it, is not because the people who wrote it are so much smarter than you: the most one can hope for at this stage is to give you a flavor of what lies ahead. If a statement in this chapter mystifies/intrigues, fast forward to a section indicated by feel you need. Of course, we think that you need to learn ALL of it, or otherwise we would not have written it in the first place.

Confronted with a potentially chaotic dynamical system, we analyze it through a sequence of three distinct stages; I. diagnose, II. count, III. measure. First we determine the intrinsic dimension of the system - the minimum number of coordinates necessary to capture its essential dynamics. If the system is very turbulent we are, at present, out of luck. We know only how to deal with the transitional regime between regular motions and chaotic dynamics in a few dimensions. That is still something; even an infinite-dimensional system such as a burning flame front can turn out to have a very few chaotic degrees of freedom. In this regime the chaotic dynamics is restricted to a space of low dimension, the number of relevant parameters is small, and we can proceed to step II; we count and classify all possible topologically distinct trajectories of the system into a hierarchy whose successive layers require increased precision and patience on the part of the observer. This we shall do in sect. 1.4.1. If successful, we can proceed with step III: investigate the weights of the different pieces of the system.

We commence our analysis of the pinball game with steps I, II: diagnose, count. We shall return to step III - measure - in sect. 1.5.

With the game of pinball we are in luck - it is a low dimensional system, free motion in a plane. The motion of a point particle is such that after a collision with one disk it either continues to another disk or it escapes. If we label the three disks by 1,2 and 3 , we can associate every trajectory with an itinerary, a sequence of labels indicating the order in which the disks are visited; for example, the two trajectories in figure 1.2 have itineraries _2313_, _23132321_ respectively. The itinerary is finite for a scattering trajectory, coming in from infinity and escaping after a finite number of collisions, infinite for a trapped trajectory, and infinitely repeating for a periodic orbit. Parenthetically, in this subject the words "orbit" and "trajectory" refer to one and the same thing.
[-8
[18 chapter 11
[1-2 chapter 13
[-2


Such labeling is the simplest example of symbolic dynamics. As the particle cannot collide two times in succession with the same disk, any two consecutive symbols must differ. This is an example of pruning, a rule that forbids certain subsequences of symbols. Deriving pruning rules is in

Figure 1.4: Binary labeling of the 3-disk pinball trajectories; a bounce in which the trajectory returns to the preceding disk is labeled 0 , and a bounce which results in continuation to the third disk is labeled 1.

general a difficult problem, but with the game of pinball we are lucky there are no further pruning rules.

The choice of symbols is in no sense unique. For example, as at each bounce we can either proceed to the next disk or return to the previous disk, the above 3 -letter alphabet can be replaced by a binary $\{0,1\}$ alphabet, figure 1.4. A clever choice of an alphabet will incorporate important features of the dynamics, such as its symmetries.

Suppose you wanted to play a good game of pinball, that is, get the pinball to bounce as many times as you possibly can - what would be a winning strategy? The simplest thing would be to try to aim the pinball so it bounces many times between a pair of disks - if you managed to shoot it so it starts out in the periodic orbit bouncing along the line connecting two disk centers, it would stay there forever. Your game would be just as good if you managed to get it to keep bouncing between the three disks forever, or place it on any periodic orbit. The only rub is that any such orbit is unstable, so you have to aim very accurately in order to stay close to it for a while. So it is pretty clear that if one is interested in playing well, unstable periodic orbits are important - they form the skeleton onto which all trajectories trapped for long times cling.

### 1.4.1 Partitioning with periodic orbits

A trajectory is periodic if it returns to its starting position and momentum. We shall refer to the set of periodic points that belong to a given periodic orbit as a cycle.

Short periodic orbits are easily drawn and enumerated - some examples are drawn in figure 1.5 - but it is rather hard to perceive the systematics of orbits from their shapes. In mechanics a trajectory is fully and uniquely specified by its position and momentum at a given instant, and no two distinct phase space trajectories can intersect. Their projections onto arbitrary subspaces, however, can and do intersect, in rather unilluminating ways. In the pinball example the problem is that we are looking at the projections of a 4-dimensional phase space trajectories onto a 2-dimensional subspace, the configuration space. A clearer picture of the dynamics is obtained by constructing a phase space Poincaré section.

Suppose that the pinball has just bounced off disk 1. Depending on its position and outgoing angle, it could proceed to either disk 2 or 3 . Not much happens in between the bounces - the ball just travels at constant velocity along a straight line - so we can reduce the four-dimensional flow to a two-

Figure 1.5: Some examples of 3-disk cycles: (a) $\overline{12123}$ and $\overline{13132}$ are mapped into each other by the flip across 1 axis. Similarly (b) $\overline{123}$ and $\overline{132}$ are related by flips, and (c) $\overline{1213}$, $\overline{1232}$ and $\overline{1323}$ by rotations. (d) The cycles $\overline{121212313}$ and $\overline{121212323}$ are related by rotaion and time reversal. These symmetries are discussed in more detail in chapter 22. (from ref. [1.2])

a)

b)

c)

d)

(b)


Figure 1.6: (a) The Poincaré section coordinates for the 3-disk game of pinball. (b) Collision sequence $\left(s_{1}, p_{1}\right) \mapsto\left(s_{2}, p_{2}\right) \mapsto\left(s_{3}, p_{3}\right)$ from the boundary of a disk to the boundary of the next disk presented in the Poincaré section coordinates.
dimensional map $f$ that takes the coordinates of the pinball from one disk edge to another disk edge. Let us state this more precisely: the trajectory just after the moment of impact is defined by marking $s_{n}$, the arc-length position of the $n$th bounce along the billiard wall, and $p_{n}=p \sin \phi_{n}$ the momentum component parallel to the billiard wall at the point of impact, figure 1.6. Such a section of a flow is called a Poincaré section, and the particular choice of coordinates (due to Birkhoff) is particularly smart, as it conserves the phase-space volume. In terms of the Poincaré section, the dynamics is reduced to the return map $P:\left(s_{n}, p_{n}\right) \mapsto\left(s_{n+1}, p_{n+1}\right)$ from the boundary of a disk to the boundary of the next disk. The explicit form of this map is easily written down, but it is of no importance right now.

Next, we mark in the Poincaré section those initial conditions which do not escape in one bounce. There are two strips of survivors, as the trajectories originating from one disk can hit either of the other two disks, or escape without further ado. We label the two strips $\mathcal{M}_{0}, \mathcal{M}_{1}$. Embedded

Figure 1.7: (a) A trajectory starting out from disk 1 can either hit another disk or escape. (b) Hitting two disks in a sequence requires a much sharper aim. The cones of initial conditions that hit more and more consecutive disks are nested within each other, as in figure 1.8.

(a)

(b)


Figure 1.8: The 3-disk game of pinball Poincaré section, trajectories emanating from the disk 1 with $x_{0}=$ (arclength, parallel momentum $)=\left(s_{0}, p_{0}\right)$, disk radius : center separation ratio $a: R=1: 2.5$. (a) Strips of initial points $\mathcal{M}_{12}, \mathcal{M}_{13}$ which reach disks 2, 3 in one bounce, respectively. (b) Strips of initial points $\mathcal{M}_{121}, \mathcal{M}_{131} \mathcal{M}_{132}$ and $\mathcal{M}_{123}$ which reach disks $1,2,3$ in two bounces, respectively. The Poincaré sections for trajectories originating on the other two disks are obtained by the appropriate relabeling of the strips. (Y. Lan)
within them there are four strips $\mathcal{M}_{00}, \mathcal{M}_{10}, \mathcal{M}_{01}, \mathcal{M}_{11}$ of initial conditions that survive for two bounces, and so forth, see figures 1.7 and 1.8. Provided that the disks are sufficiently separated, after $n$ bounces the survivors are divided into $2^{n}$ distinct strips: the $\mathcal{M}_{i}$ th strip consists of all points with itinerary $i=s_{1} s_{2} s_{3} \ldots s_{n}, s=\{0,1\}$. The unstable cycles as a skeleton of chaos are almost visible here: each such patch contains a periodic point $\overline{s_{1} s_{2} s_{3} \ldots s_{n}}$ with the basic block infinitely repeated. Periodic points are skeletal in the sense that as we look further and further, the strips shrink but the periodic points stay put forever.

We see now why it pays to utilize a symbolic dynamics; it provides a navigation chart through chaotic phase space. There exists a unique trajectory for every admissible infinite length itinerary, and a unique itinerary labels every trapped trajectory. For example, the only trajectory labeled by $\overline{12}$ is the 2-cycle bouncing along the line connecting the centers of disks 1 and 2 ; any other trajectory starting out as $12 \ldots$ either eventually escapes or hits the 3rd disk.

### 1.4.2 Escape rate

What is a good physical quantity to compute for the game of pinball? Such system, for which almost any trajectory eventually leaves a finite region (the pinball table) never to return, is said to be open, or a repeller. The repeller escape rate is an eminently measurable quantity. An example of such a
measurement would be an unstable molecular or nuclear state which can be well approximated by a classical potential with the possibility of escape in certain directions. In an experiment many projectiles are injected into such a non-confining potential and their mean escape rate is measured, as in figure 1.1. The numerical experiment might consist of injecting the pinball between the disks in some random direction and asking how many times the pinball bounces on the average before it escapes the region between the disks.

[^0]For a theorist a good game of pinball consists in predicting accurately the asymptotic lifetime (or the escape rate) of the pinball. We now show how periodic orbit theory accomplishes this for us. Each step will be so simple that you can follow even at the cursory pace of this overview, and still the result is surprisingly elegant.

Consider figure 1.8 again. In each bounce the initial conditions get thinned out, yielding twice as many thin strips as at the previous bounce. The total area that remains at a given time is the sum of the areas of the strips, so that the fraction of survivors after $n$ bounces, or the survival probability is given by

$$
\begin{align*}
& \hat{\Gamma}_{1}=\frac{\left|\mathcal{M}_{0}\right|}{|\mathcal{M}|}+\frac{\left|\mathcal{M}_{1}\right|}{|\mathcal{M}|}, \quad \hat{\Gamma}_{2}=\frac{\left|\mathcal{M}_{00}\right|}{|\mathcal{M}|}+\frac{\left|\mathcal{M}_{10}\right|}{|\mathcal{M}|}+\frac{\left|\mathcal{M}_{01}\right|}{|\mathcal{M}|}+\frac{\left|\mathcal{M}_{11}\right|}{|\mathcal{M}|} \\
& \hat{\Gamma}_{n}=\frac{1}{|\mathcal{M}|} \sum_{i}^{(n)}\left|\mathcal{M}_{i}\right| \tag{1.2}
\end{align*}
$$

where $i$ is a label of the $i$ th strip, $|\mathcal{M}|$ is the initial area, and $\left|\mathcal{M}_{i}\right|$ is the area of the $i$ th strip of survivors. $i=01,10,11, \ldots$ is a label, not a binary number. Since at each bounce one routinely loses about the same fraction of trajectories, one expects the sum (1.2) to fall off exponentially with $n$ and tend to the limit

$$
\begin{equation*}
\hat{\Gamma}_{n+1} / \hat{\Gamma}_{n}=e^{-\gamma_{n}} \rightarrow e^{-\gamma} . \tag{1.3}
\end{equation*}
$$

The quantity $\gamma$ is called the escape rate from the repeller.

### 1.5 Chaos for cyclists

Étant données des équations ... et une solution particuliére quelconque de ces équations, on peut toujours trouver une solution périodique (dont la période peut, il est vrai, étre trés longue), telle que la différence entre les deux solutions soit aussi petite qu'on le veut, pendant un temps aussi long qu'on le veut. D'ailleurs, ce qui nous rend ces solutions périodiques si précieuses, c'est qu'elles sont, pour ansi dire, la seule bréche par où nous puissions esseyer de pénétrer dans une place jusqu'ici réputée inabordable.
H. Poincaré, Les méthodes nouvelles de la méchanique céleste

We shall now show that the escape rate $\gamma$ can be extracted from a highly convergent exact expansion by reformulating the sum (1.2) in terms of unstable periodic orbits.

If, when asked what the 3 -disk escape rate is for a disk of radius 1 , center-center separation 6 , velocity 1 , you answer that the continuous time escape rate is roughly $\gamma=0.4103384077693464893384613078192 \ldots$, you do not need this book. If you have no clue, hang on.

### 1.5.1 Size of a partition

Not only do the periodic points keep track of topological ordering of the strips, but, as we shall now show, they also determine their size.

As a trajectory evolves, it carries along and distorts its infinitesimal neighborhood. Let

$$
x(t)=f^{t}\left(x_{0}\right)
$$

denote the trajectory of an initial point $x_{0}=x(0)$. Expanding $f^{t}\left(x_{0}+\delta x_{0}\right)$ tolinear order, the evolution of the distance to a neighboring trajectory $x_{i}(t)+\delta x_{i}(t)$ is given by the Jacobian matrix:

$$
\delta x_{i}(t)=\sum_{j=1}^{d} \mathbf{J}^{t}\left(x_{0}\right)_{i j} \delta x_{0 j}, \quad \mathbf{J}^{t}\left(x_{0}\right)_{i j}=\frac{\partial x_{i}(t)}{\partial x_{0 j}} .
$$

A trajectory of a pinball moving on a flat surface is specified by two position coordinates and the direction of motion, so in this case $d=3$. Evaluation of of $x(t)$ along the flow; its eigenvectors and eigenvalues give the directions and the corresponding rates of expansion or contraction. The trajectories that start out in an infinitesimal neighborhood are separated along
the unstable directions (those whose eigenvalues are greater than unity in magnitude), approach each other along the stable directions (those whose eigenvalues are less than unity in magnitude), and maintain their distance along the marginal directions (those whose eigenvalues equal unity in magnitude). In our game of pinball the beam of neighboring trajectories is defocused along the unstable eigendirection of the Jacobian matrix $\mathbf{J}$.

As the heights of the strips in figure 1.8 are effectively constant, we can concentrate on their thickness. If the height is $\approx L$, then the area of the $i$ th strip is $\mathcal{M}_{i} \approx L l_{i}$ for a strip of width $l_{i}$.

Each strip $i$ in figure 1.8 contains a periodic point $x_{i}$. The finer the intervals, the smaller the variation in flow across them, so the contribution from the strip of width $l_{i}$ is well-approximated by the contraction around the periodic point $x_{i}$ within the interval,

$$
\begin{equation*}
l_{i}=a_{i} /\left|\Lambda_{i}\right| \tag{1.4}
\end{equation*}
$$

where $\Lambda_{i}$ is the unstable eigenvalue of the Jacobian matrix $\mathbf{J}^{t}\left(x_{i}\right)$ evaluated at the $i$ th periodic point for $t=T_{p}$, the full period (due to the low dimensionality, the Jacobian can have at most one unstable eigenvalue). Only the magnitude of this eigenvalue matters, we can disregard its sign. The prefactors $a_{i}$ reflect the overall size of the system and the particular distribution of starting values of $x$. As the asymptotic trajectories are strongly mixed by bouncing chaotically around the repeller, we expect their distribution to be insensitive to smooth variations in the distribution of initial points.

To proceed with the derivation we need the hyperbolicity assumption: for large $n$ the prefactors $a_{i} \approx O(1)$ are overwhelmed by the exponential growth of $\Lambda_{i}$, so we neglect them. If the hyperbolicity assumption is justified, we can replace $\left|\mathcal{M}_{i}\right| \approx L l_{i}$ in (1.2) by $1 /\left|\Lambda_{i}\right|$ and consider the sum

$$
\Gamma_{n}=\sum_{i}^{(n)} 1 /\left|\Lambda_{i}\right|,
$$

where the sum goes over all periodic points of period $n$. We now define a generating function for sums over all periodic orbits of all lengths:

$$
\begin{equation*}
\Gamma(z)=\sum_{n=1}^{\infty} \Gamma_{n} z^{n} . \tag{1.5}
\end{equation*}
$$

Recall that for large $n$ the $n$th level sum (1.2) tends to the limit $\Gamma_{n} \rightarrow e^{-n \gamma}$, so the escape rate $\gamma$ is determined by the smallest $z=e^{\gamma}$ for which (1.5) diverges:

$$
\begin{equation*}
\Gamma(z) \approx \sum_{n=1}^{\infty}\left(z e^{-\gamma}\right)^{n}=\frac{z e^{-\gamma}}{1-z e^{-\gamma}} . \tag{1.6}
\end{equation*}
$$

This is the property of $\Gamma(z)$ that motivated its definition. Next, we devise a formula for (1.5) expressing the escape rate in terms of periodic orbits:

$$
\begin{align*}
\Gamma(z)= & \sum_{n=1}^{\infty} z^{n} \sum_{i}^{(n)}\left|\Lambda_{i}\right|^{-1} \\
= & \frac{z}{\left|\Lambda_{0}\right|}+\frac{z}{\left|\Lambda_{1}\right|}+\frac{z^{2}}{\left|\Lambda_{00}\right|}+\frac{z^{2}}{\left|\Lambda_{01}\right|}+\frac{z^{2}}{\left|\Lambda_{10}\right|}+\frac{z^{2}}{\left|\Lambda_{11}\right|} \\
& +\frac{z^{3}}{\left|\Lambda_{000}\right|}+\frac{z^{3}}{\left|\Lambda_{001}\right|}+\frac{z^{3}}{\left|\Lambda_{010}\right|}+\frac{z^{3}}{\left|\Lambda_{100}\right|}+\ldots \tag{1.7}
\end{align*}
$$

For sufficiently small $z$ this sum is convergent. The escape rate $\gamma$ is now given by the leading pole of (1.6), rather than by a numerical extrapolation of a sequence of $\gamma_{n}$ extracted from (1.3). As any finite truncation $n<$ $n_{\text {trunc }}$ of (1.7) is a polynomial in $z$, convergent for any $z$, finding this pole requires that we know something about $\Gamma_{n}$ for any $n$, and that might be a tall order.

We could now proceed to estimate the location of the leading singularity of $\Gamma(z)$ from finite truncations of (1.7) by methods such as Padé approximants. However, as we shall now show, it pays to first perform a simple resummation that converts this divergence into a zero of a related function.

### 1.5.2 Dynamical zeta function

If a trajectory retraces a prime cycle $r$ times, its expanding eigenvalue is $\Lambda_{p}^{r}$. A prime cycle $p$ is a single traversal of the orbit; its label is a non-repeating symbol string of $n_{p}$ symbols. There is only one prime cycle for each cyclic permutation class. For example, $p=\overline{0011}=\overline{1001}=\overline{1100}=\overline{0110}$ is prime, but $\overline{0101}=\overline{01}$ is not. $\quad$ By the chain rule for derivatives the stability of a cycle is the same everywhere along the orbit, so each prime cycle of length $n_{p}$ contributes $n_{p}$ terms to the sum (1.7). Hence (1.7) can be rewritten as

$$
\begin{equation*}
\Gamma(z)=\sum_{p} n_{p} \sum_{r=1}^{\infty}\left(\frac{z^{n_{p}}}{\left|\Lambda_{p}\right|}\right)^{r}=\sum_{p} \frac{n_{p} t_{p}}{1-t_{p}}, \quad t_{p}=\frac{z^{n_{p}}}{\left|\Lambda_{p}\right|} \tag{1.8}
\end{equation*}
$$

where the index $p$ runs through all distinct prime cycles. Note that we have resummed the contribution of the cycle $p$ to all times, so truncating the summation up to given $p$ is not a finite time $n \leq n_{p}$ approximation, but an asymptotic, infinite time estimate based by approximating stabilities of all cycles by a finite number of the shortest cycles and their repeats. The $n_{p} z^{n_{p}}$ factors in (1.8) suggest rewriting the sum as a derivative

$$
\Gamma(z)=-z \frac{d}{d z} \sum_{p} \ln \left(1-t_{p}\right) .
$$

Hence $\Gamma(z)$ is a logarithmic derivative of the infinite product

$$
\begin{equation*}
1 / \zeta(z)=\prod_{p}\left(1-t_{p}\right), \quad t_{p}=\frac{z^{n_{p}}}{\left|\Lambda_{p}\right|} \tag{1.9}
\end{equation*}
$$

This function is called the dynamical zeta function, in analogy to the Riemann zeta function, which motivates the choice of "zeta" in its definition as $1 / \zeta(z)$. This is the prototype formula of periodic orbit theory. The zero of $1 / \zeta(z)$ is a pole of $\Gamma(z)$, and the problem of estimating the asymptotic escape rates from finite $n$ sums such as (1.2) is now reduced to a study of the zeros of the dynamical zeta function (1.6.2). The escape rate is related by (1.6) to a divergence of $\Gamma(z)$, and $\Gamma(z)$ diverges whenever $1 / \zeta(z)$ has a
zero.
[108 sect. 19.1
[10

$$
z_{p}=\left|\Lambda_{p}\right|^{1 / n_{p}}
$$

for each term in the product. What's the problem?" Dead wrong!

### 1.5.3 Cycle expansions

How are formulas such as (1.6.2) used? We start by computing the lengths and eigenvalues of the shortest cycles. This usually requires some numerical work, such as the Newton's method searches for periodic solutions; we shall assume that the numerics are under control, and that all short cycles up to given length have been found. In our pinball example this can be done by elementary geometrical optics. It is very important not to miss any short cycles, as the calculation is as accurate as the shortest cycle dropped - including cycles longer than the shortest omitted does not improve the accuracy (unless exponentially many more cycles are included). The result of such numerics is a table of the shortest cycles, their periods and their stabilities.

Now expand the infinite product (1.6.2), grouping together the terms of the same total symbol string length

$$
\begin{align*}
1 / \zeta= & \left(1-t_{0}\right)\left(1-t_{1}\right)\left(1-t_{10}\right)\left(1-t_{100}\right) \cdots \\
= & 1-t_{0}-t_{1}-\left[t_{10}-t_{1} t_{0}\right]-\left[\left(t_{100}-t_{10} t_{0}\right)+\left(t_{101}-t_{10} t_{1}\right)\right] \\
& -\left[\left(t_{1000}-t_{0} t_{100}\right)+\left(t_{1110}-t_{1} t_{110}\right)\right. \\
& \left.+\left(t_{1001}-t_{1} t_{001}-t_{101} t_{0}+t_{10} t_{0} t_{1}\right)\right]-\ldots \tag{1.10}
\end{align*}
$$

The virtue of the expansion is that the sum of all terms of the same total length $n$ (grouped in brackets above) is a number that is exponentially smaller than a typical term in the sum, for geometrical reasons we explain in the next section.

The calculation is now straightforward. We substitute a finite set of the eigenvalues and lengths of the shortest prime cycles into the cycle expansion (1.10), and obtain a polynomial approximation to $1 / \zeta$. We then vary $z$ in (1.6.2) and determine the escape rate $\gamma$ by finding the smallest $z=e^{\gamma}$ for which (1.10) vanishes.

### 1.5.4 Shadowing

When you actually start computing this escape rate, you will find out that the convergence is very impressive: only three input numbers (the two fixed points $\overline{0}, \overline{1}$ and the 2-cycle $\overline{10}$ ) already yield the pinball escape rate to $3-4$ significant digits! We have omitted an infinity of unstable cycles; so why does approximating the dynamics by a finite number of the shortest cycle eigenvalues work so well?

The convergence of cycle expansions of dynamical zeta functions is a consequence of the smoothness and analyticity of the underlying flow. Intuitively, one can understand the convergence in terms of the geometrical picture sketched in figure 1.9; the key observation is that the long orbits are shadowed by sequences of shorter orbits.

A typical term in (1.10) is a difference of a long cycle $\{a b\}$ minus its shadowing approximation by shorter cycles $\{a\}$ and $\{b\}$

$$
\begin{equation*}
t_{a b}-t_{a} t_{b}=t_{a b}\left(1-t_{a} t_{b} / t_{a b}\right)=t_{a b}\left(1-\left|\frac{\Lambda_{a b}}{\Lambda_{a} \Lambda_{b}}\right|\right), \tag{1.11}
\end{equation*}
$$

where $a$ and $b$ are symbol sequences of the two shorter cycles. If all orbits are weighted equally ( $t_{p}=z^{n_{p}}$ ), such combinations cancel exactly; if orbits of similar symbolic dynamics have similar weights, the weights in such combinations almost cancel.

This can be understood in the context of the pinball game as follows. Consider orbits $\overline{0}, \overline{1}$ and $\overline{01}$. The first corresponds to bouncing between any two disks while the second corresponds to bouncing successively around all three, tracing out an equilateral triangle. The cycle $\overline{01}$ starts at one disk, say disk 2. It then bounces from disk 3 back to disk 2 then bounces from disk 1 back to disk 2 and so on, so its itinerary is $\overline{2321}$. In terms of the bounce types shown in figure 1.4, the trajectory is alternating between 0 and 1 . The incoming and outgoing angles when it executes these bounces are very close to the corresponding angles for 0 and 1 cycles. Also the distances traversed between bounces are similar so that the 2 -cycle expanding eigenvalue $\Lambda_{01}$ is close in magnitude to the product of the 1-cycle eigenvalues $\Lambda_{0} \Lambda_{1}$.

To understand this on a more general level, try to visualize the partition of a chaotic dynamical system's phase space in terms of cycle neighborhoods as a tessellation of the dynamical system, with smooth flow approximated by its periodic orbit skeleton, each "face" centered on a periodic point, and


Figure 1.9: Approximation to (a) a smooth dynamics by (b) the skeleton of periodic points, together with their linearized neighborhoods. Indicated are segments of two 1-cycles and a 2-cycle that alternates between the neighborhoods of the two 1-cycles, shadowing first one of the two 1-cycles, and then the other.
the scale of the "face" determined by the linearization of the flow around the periodic point, figure 1.9.

The orbits that follow the same symbolic dynamics, such as $\{a b\}$ and a "pseudo orbit" $\{a\}\{b\}$, lie close to each other in phase space; long shadowing pairs have to start out exponentially close to beat the exponential growth in separation with time. If the weights associated with the orbits are multiplicative along the flow (for example, by the chain rule for products of derivatives) and the flow is smooth, the term in parenthesis in (1.11) falls off exponentially with the cycle length, and therefore the curvature expansions are expected to be highly convergent.

### 1.6 Evolution

In physics, when we do not understand something, we give it a name.
Matthias Neubert
The above derivation of the dynamical zeta function formula for the escape rate has one shortcoming; it estimates the fraction of survivors as a function of the number of pinball bounces, but the physically interesting quantity is the escape rate measured in units of continuous time. For continuous time flows, the escape rate (1.2) is generalized as follows. Define a finite phase space region $\mathcal{M}$ such that a trajectory that exits $\mathcal{M}$ never reenters. For example, any pinball that falls of the edge of a pinball table in figure 1.1 is gone forever. Start with a uniform distribution of initial points. The fraction of initial $x$ whose trajectories remain within $\mathcal{M}$ at time $t$ is
expected to decay exponentially

$$
\Gamma(t)=\frac{\int_{\mathcal{M}} d x d y \delta\left(y-f^{t}(x)\right)}{\int_{\mathcal{M}} d x} \rightarrow e^{-\gamma t}
$$

The integral over $x$ starts a trajectory at every $x \in \mathcal{M}$. The integral over $y$ tests whether this trajectory is still in $\mathcal{M}$ at time $t$. The kernel of this integral

$$
\begin{equation*}
\mathcal{L}^{t}(y, x)=\delta\left(y-f^{t}(x)\right) \tag{1.12}
\end{equation*}
$$

is the Dirac delta function, as for a deterministic flow the initial point $x$ maps into a unique point $y$ at time $t$. For discrete time, $f^{n}(x)$ is the $n$th iterate of the map $f$. For continuous flows, $f^{t}(x)$ is the trajectory of the initial point $x$, and it is appropriate to express the finite time kernel $\mathcal{L}^{t}$ in terms of a generator of infinitesimal time translations

$$
\mathcal{L}^{t}=e^{t \mathcal{A}}
$$

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very much in the way the quantum evolution is generated by the Hamiltonian $H$, the generator of infinitesimal time quantum transformations.

As the kernel $\mathcal{L}$ is the key to everything that follows, we shall give it a name, and refer to it and its generalizations as the evolution operator for a $d$-dimensional map or a $d$-dimensional flow.

The number of periodic points increases exponentially with the cycle length (in the case at hand, as $2^{n}$ ). As we have already seen, this exponential proliferation of cycles is not as dangerous as it might seem; as a matter of fact, all our computations will be carried out in the $n \rightarrow \infty$ limit. Though a quick look at chaotic dynamics might reveal it to be complex beyond belief, it is still generated by a simple deterministic law, and with some luck and insight, our labeling of possible motions will reflect this simplicity. If the rule that gets us from one level of the classification hierarchy to the next does not depend strongly on the level, the resulting hierarchy is approximately self-similar. We now turn such approximate self-similarity to our advantage, by turning it into an operation, the action of the evolution operator, whose iteration encodes the self-similarity.

### 1.6.1 Trace formula

Recasting dynamics in terms of evolution operators changes everything. So far our formulation has been heuristic, but in the evolution operator formalism the escape rate and any other dynamical average are given by exact formulas, extracted from the spectra of evolution operators. The key tools are trace formulas and spectral determinants.


Figure 1.10: The trace of an evolution operator is concentrated in tubes around prime cycles, of length $T_{p}$ and thickness $1 /\left|\Lambda_{p}\right|^{r}$ for the $r$ th repetition of the prime cycle $p$.

The trace of an operator is given by the sum of its eigenvalues. The explicit expression (1.12) for $\mathcal{L}^{t}(x, y)$ enables us to evaluate the trace. Identify $y$ with $x$ and integrate $x$ over the whole phase space. The result is an expression for $\operatorname{tr} \mathcal{L}^{t}$ as a sum over neighborhoods of prime cycles $p$ and their repetitions

$$
\begin{equation*}
\operatorname{tr} \mathcal{L}^{t}=\sum_{p} T_{p} \sum_{r=1}^{\infty} \frac{\delta\left(t-r T_{p}\right)}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|} \tag{1.13}
\end{equation*}
$$

This formula has a simple geometrical interpretation sketched in figure 1.10. After the $r$ th return to a Poincaré section, the initial tube $\mathcal{M}_{p}$ has been stretched out along the expanding eigendirections, with the overlap with the initial volume given by $1 /\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right| \rightarrow 1 /\left|\Lambda_{p}\right|$, the same weight we obtained heuristically in sect. 1.5.1.

The "spiky" sum (1.13) is disquieting in the way reminiscent of the Poisson resummation formulas of Fourier analysis; the left-hand side is the smooth eigenvalue sum $\operatorname{tr} e^{\mathcal{A} t}=\sum e^{s_{\alpha} t}$, while the right-hand side equals zero everywhere except for the set $t=r T_{p}$. A Laplace transform smooths the sum over Dirac delta functions in cycle periods and yields the trace formula for the eigenspectrum $s_{0}, s_{1}, \cdots$ of the classical evolution operator:

$$
\begin{align*}
\int_{0_{+}}^{\infty} d t e^{-s t} \operatorname{tr} \mathcal{L}^{t} & =\operatorname{tr} \frac{1}{s-\mathcal{A}}= \\
\sum_{\alpha=0}^{\infty} \frac{1}{s-s_{\alpha}} & =\sum_{p} T_{p} \sum_{r=1}^{\infty} \frac{e^{r\left(\beta \cdot A_{p}-s T_{p}\right)}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|} . \tag{1.14}
\end{align*}
$$

The beauty of trace formulas lies in the fact that everything on the right-hand-side - prime cycles $p$, their periods $T_{p}$ and the stability eigenvalues of $\mathbf{J}_{p}$ - is an invariant property of the flow, independent of any coordinate choice.

### 1.6.2 Spectral determinant

The eigenvalues of a linear operator are given by the zeros of the appropriate determinant. One way to evaluate determinants is to expand them in terms of traces, using the identities

$$
\begin{aligned}
\ln \operatorname{det}(s-\mathcal{A}) & =\operatorname{tr} \ln (s-\mathcal{A}) \\
\frac{d}{d s} \ln \operatorname{det}(s-\mathcal{A}) & =\operatorname{tr} \frac{1}{s-\mathcal{A}}
\end{aligned}
$$

and integrating over $s$. In this way the spectral determinant of an evolution operator becomes related to the traces that we have just computed:

$$
\begin{equation*}
\operatorname{det}(s-\mathcal{A})=\exp \left(-\sum_{p} \sum_{r=1}^{\infty} \frac{1}{r} \frac{e^{-s T_{p} r}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|}\right) . \tag{1.15}
\end{equation*}
$$

The $1 / r$ factor is due to the $s$ integration, leading to the replacement $T_{p} \rightarrow$ $T_{p} / r T_{p}$ in the periodic orbit expansion (1.14).

The motivation for recasting the eigenvalue problem in this form is sketched in figure 1.11; exponentiation improves analyticity and trades in a divergence of the trace sum for a zero of the spectral determinant. In this way we have retaced the heuristic derivation of the divergent sum (1.6) and the dynamical zeta function, but this time with no approximations: formula (1.15) is exact. The computation of the zeros of $\operatorname{det}(s-\mathcal{A})$ proceeds very much like the computations of sect. 1.5.3.

### 1.7 From chaos to statistical mechanics

While the above replacement of dynamics of individual trajectories by evolution operators which propagate densities might feel like just another bit of mathematical voodoo, actually something very radical has taken place. Consider a chaotic flow, such as the stirring of red and white paint by some deterministic machine. If we were able to track individual trajectories, the fluid would forever remain a striated combination of pure white and pure red; there would be no pink. What is more, if we reversed the stirring, we would return to the perfect white/red separation. However, we know that this cannot be true - in a very few turns of the stirring stick the thickness of the layers goes from centimeters to Ångströms, and the result is irreversibly pink.

Understanding the distinction between evolution of individual trajectories and the evolution of the densities of trajectories is key to understanding statistical mechanics - this is the conceptual basis of the second law of thermodynamics, and the origin of irreversibility of the arrow of time for

Figure 1.11: Spectral determinant is preferable to the trace as it vanishes smoothly at the leading eigenvalue, while the trace formula diverges.

deterministic systems with time-reversible equations of motion: reversibility is attainable for distributions whose measure in the space of density functions goes exponentially to zero with time.

By going to a description in terms of the asymptotic time evolution operators we give up tracking individual trajectories for long times, but instead gain a very effective description of the asymptotic trajectory densities. This will enable us, for example, to give exact formulas for transport coefficients such as the diffusion constants without any probabilistic assumptions (such as the stosszahlansatz of Boltzmann).

A century ago it seemed reasonable to assume that statistical mechanics applies only to systems with very many degrees of freedom. More recent is the realization that much of statistical mechanics follows from chaotic dynamics, and already at the level of a few degrees of freedom the evolution of densities is irreversible. Furthermore, the theory that we shall develop here generalizes notions of "measure" and "averaging" to systems far from equilibrium, and transports us into regions hitherto inaccessible with the tools of equilibrium statistical mechanics.

The concepts of equilibrium statistical mechanics do help us, however, to understand the ways in which the simple-minded periodic orbit theory falters. A non-hyperbolicity of the dynamics manifests itself in power-law correlations and even "phase transitions".

### 1.8 A guide to the literature

But the power of instruction is seldom of much efficacy, except in those happy dispositions where it is almost superfluous.
Gibbon
This text aims to bridge the gap between the physics and mathematics dynamical systems literature. The intended audience is Henri Roux, the perfect physics graduate student with a theoretical bent who does not believe anything he does not understand. As a complementary presentation we recommend Gaspard's monograph [1.4] which covers much of the same ground in a highly readable and scholarly manner.

As far as the prerequisites are concerned - this book is not an introduction to nonlinear dynamics. Nonlinear science requires a one semester basic course (advanced undergraduate or first year graduate). A good start is the textbook by Strogatz [1.5], an introduction to flows, fixed points,
manifolds, bifurcations. It is the most accessible introduction to nonlinear dynamics - it starts out with differential equations, and its broadly chosen examples and many exercises make it a favorite with students. It is not strong on chaos. There the textbook of Alligood, Sauer and Yorke [1.6] is preferable: an elegant introduction to maps, chaos, period doubling, symbolic dynamics, fractals, dimensions - a good companion to this book. An introduction more comfortable to physicists is the textbook by Ott [1.7], with the baker's map used to illustrate many key techniques in analysis of chaotic systems. It is perhaps harder than the above two as the first book on nonlinear dynamics.

An introductory course should give students skills in qualitative and numerical analysis of dynamical systems for short times (trajectories, fixed points, bifurcations) and familiarize them with Cantor sets and symbolic dynamics for chaotic systems. A good introduction to numerical experimentation with physically realistic systems is Tufillaro, Abbott, and Reilly [1.8]. Korsch and Jodl [1.9] and Nusse and Yorke [1.10] also emphasize hands-on approach to dynamics. With this, and a graduate level-exposure to statistical mechanics, partial differential equations and quantum mechanics, the stage is set for any of the one-semester advanced courses based on this book. The courses taught so far start out with the introductory chapters on qualitative dynamics, symbolic dynamics and flows, and then continue in different directions:

Deterministic chaos. Chaotic averaging, evolution operators, trace formulas, zeta functions, cycle expansions, Lyapunov exponents, billiards, transport coefficients, thermodynamic formalism, period doubling, renormalization operators.

Spatiotemporal dynamical systems. Partial differential equations for dissipative systems, weak amplitude expansions, normal forms, symmetries and bifurcations, pseudospectral methods, spatiotemporal chaos.

Quantum chaos. Semiclassical propagators, density of states, trace formulas, semiclassical spectral determinants, billiards, semiclassical helium, diffraction, creeping, tunneling, higher-order $\hbar$ corrections.

This book concentrates on periodic orbit theory. The role of unstable periodic orbits was already fully appreciated by Poincaré [1.11, 1.12], who noted that hidden in the apparent chaos is a rigid skeleton, a tree of $c y$ cles (periodic orbits) of increasing lengths and self-similar structure, and suggested that the cycles should be the key to chaotic dynamics. Periodic orbits have been at core of much of the mathematical work on the theory of the classical and quantum dynamical systems ever since. We refer the reader to the reprint selection [1.13] for an overview of some of that literature. This book offers a breach into a domain hitherto reputed unreachable, a domain traditionally traversed only by mathematical physicists and pure mathematicians. What distinguishes it from pure mathematics is the emphasis on computation and numerical convergence of methods offered. A rigorous proof, the end of the story as far as a mathematician is concerned, might state that in a given setting, for times in excess of $10^{32}$ years, turbu-
lent dynamics settles onto an attractor of dimension less than 600. Such a theorem is of a little use for a physicist, especially if a numerical experiment indicates that within the span of the best simulation the dynamics seems to have settled on a (transient?) attractor of dimension less than 3.

If you find this book not rigorous enough, you should turn to the mathematics literature. The most extensive reference is the treatise by Katok and Hasselblatt [1.14], an impressive compendium of modern dynamical systems theory. The fundamental papers in this field, all still valuable reading, are Smale [1.15], Bowen [1.16] and Sinai [1.17]. Sinai's paper is prescient and offers a vision and a program that ties together dynamical systems and statistical mechanics. It is written for readers versed in statistical mechanics. Markov partitions were introduced by Sinai in ref. [1.18]. The classical text (though certainly not an easy read) on the subject of dynamical zeta functions is Ruelle's Statistical Mechanics, Thermodynamic Formalism [1.19]. In Ruelle's monograph transfer operator technique (or the "Perron-Frobenius theory") and Smale's theory of hyperbolic flows are applied to zeta functions and correlation functions. The status of the theory from Ruelle's point of view is compactly summarized in his 1995 Pisa lectures [1.20]. Further excellent mathematical references on thermodynamic formalism are Parry and Pollicott's monograph [1.21] with emphasis on the symbolic dynamics aspects of the formalism, and Baladi's clear and compact reviews of the theory dynamical zeta functions [1.22, 1.23].

A graduate level introduction to statistical mechanics from the dynamical point view is given by Dorfman [1.24]; the Gaspard monograph [1.4] covers the same ground in more depth. Driebe monograph [1.25] offers a nice introduction to the problem of irreversibility in dynamics. The role of "chaos" in statistical mechanics is critically dissected by Bricmont in his highly readable essay "Science of Chaos or Chaos in Science?" [1.26].

If you were wandering while reading this introduction "what's up with rat brains?", the answer is yes indeed, there is a line of research in neuronal dynamics that focuses on possible unstable periodic states, described for example in ref. [1.28, 1.29, 1.30, 1.31].

## A guide to exercises

God can afford to make mistakes. So can Dada!
Dadaist Manifesto
The essence of this subject is incommunicable in print; the only way to develop intuition about chaotic dynamics is by computing, and the reader is urged to try to work through the essential exercises. As not to fragment the text, the exercises are indicated by text margin boxes such as the one on this margin, and collected at the end of each chapter. The problems that you should do have underlined titles. The rest (smaller type) are optional. Difficult problems are marked by any number of $* * *$ stars. If you solve
one of those, it is probably worth a publication. By the end of a (twosemester) course you should have completed at least three small projects: (a) compute everything for a one-dimensional repeller, (b) compute escape rate for a 3 -disk game of pinball, (c) compute a part of the quantum 3-disk game of pinball, or the helium spectrum, or if you are interested in statistical rather than the quantum mechanics, compute a transport coefficient. The essential steps are:

## - Dynamics

1. count prime cycles, exercise 1.1
2. pinball simulator, exercise 6.1, exercise 17.4
3. pinball stability, exercise 8.1, exercise 17.4
4. pinball periodic orbits, exercise 17.6 , exercise 17.5
5. helium integrator, exercise 2.10 , exercise 17.8
6. helium periodic orbits, exercise 34.4, exercise 17.9

## - Averaging, numerical

1. pinball escape rate, exercise 10.3
2. Lyapunov exponent, exercise 20.2

## - Averaging, periodic orbits

1. cycle expansions, exercise 18.1, exercise 18.2
2. pinball escape rate, exercise 18.4 , exercise 18.5
3. cycle expansions for averages, exercise 18.1, exercise 19.3
4. cycle expansions for diffusion, exercise 23.1
5. desymmetrization exercise 22.1
6. semiclassical quantization exercise 32.3
7. ortho-, para-helium, lowest eigenenergies exercise 34.7

Solutions for some of the problems are given in appendix N. Often going through a solution is more instructive than reading the chapter that problem is supposed to illustrate.

## Résumé

The goal of this text is an exposition of the best of all possible theories of deterministic chaos, and the strategy is: 1) count, 2) weigh, 3) add up.

In a chaotic system any open ball of initial conditions, no matter how small, will spread over the entire accessible phase space. Hence the theory focuses on describing the geometry of the space of possible outcomes, and evaluating averages over this space, rather than attempting the impossible:
precise prediction of individual trajectories. The dynamics of distributions of trajectories is described in terms of evolution operators. In the evolution operator formalism the dynamical averages are given by exact formulas, extracted from the spectra of evolution operators. The key tools are trace formulas and spectral determinants.

The theory of evaluation of the spectra of evolution operators presented here is based on the observation that the motion in dynamical systems of few degrees of freedom is often organized around a few fundamental cycles. These short cycles capture the skeletal topology of the motion on a strange attractor/repeller in the sense that any long orbit can approximately be pieced together from the nearby periodic orbits of finite length. This notion is made precise by approximating orbits by prime cycles, and evaluating the associated curvatures. A curvature measures the deviation of a longer cycle from its approximation by shorter cycles; smoothness and the local instability of the flow implies exponential (or faster) fall-off for (almost) all curvatures. Cycle expansions offer an efficient method for evaluating classical and quantum observables.

The critical step in the derivation of the dynamical zeta function was the hyperbolicity assumption, that is, the assumption of exponential shrinkage of all strips of the pinball repeller. By dropping the $a_{i}$ prefactors in (1.4), we have given up on any possibility of recovering the precise distribution of starting $x$ (which should anyhow be impossible due to the exponential growth of errors), but in exchange we gain an effective description of the asymptotic behavior of the system. The pleasant surprise of cycle expansions (1.6.2) is that the infinite time behavior of an unstable system is as easy to determine as the short time behavior.

To keep the exposition simple we have here illustrated the utility of cycles and their curvatures by a pinball game, but topics covered in this book - unstable flows, Poincaré sections, Smale horseshoes, symbolic dynamics, pruning, discrete symmetries, periodic orbits, averaging over chaotic sets, evolution operators, dynamical zeta functions, spectral determinants, cycle expansions, quantum trace formulas and zeta functions, and so on to the semiclassical quantization of helium - should give the reader some confidence in the general applicability of the theory. The formalism should work for any average over any chaotic set which satisfies two conditions:

1. the weight associated with the observable under consideration is multiplicative along the trajectory,
2. the set is organized in such a way that the nearby points in the symbolic dynamics have nearby weights.

The theory is applicable to evaluation of a broad class of quantities characterizing chaotic systems, such as the escape rates, Lyapunov exponents, transport coefficients and quantum eigenvalues. One of the surprises is that the quantum mechanics of classically chaotic systems is very much like the classical mechanics of chaotic systems; both are described by nearly the same zeta functions and cycle expansions, with the same dependence on the topology of the classical flow.

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## Exercises

Exercise 1.1 3-disk symbolic dynamics. As periodic trajectories will turn out to be our main tool to breach deep into the realm of chaos, it pays to start familiarizing oneself with them now by sketching and counting the few shortest prime cycles (we return to this in sect. 13.4). Show that the 3-disk pinball has $3 \cdot 2^{n}$ itineraries of length $n$. List periodic orbits of lengths 2, 3, 4, $5, \cdots$. Verify that the shortest 3 -disk prime cycles are 12, 13, 23, 123, 132, 1213, 1232, 1323, 12123, … Try to sketch them.

Exercise 1.2 Sensitivity to initial conditions. Assume that two pinball trajectories start out parallel, but separated by 1 Ångström, and the disks are of radius $a=1 \mathrm{~cm}$ and center-to-center separation $R=6 \mathrm{~cm}$. Try to estimate in how many bounces the separation will grow to the size of system (assuming that the trajectories have been picked so they remain trapped for at least that long). Estimate the Who's Pinball Wizard's typical score (number of bounces) in a game without cheating, by hook or crook (by the end of chapter 18 you should be in position to make very accurate estimates).

## Chapter 2

## Flows

Poetry is what is lost in translation.
Robert Frost
(R. Mainieri, P. Cvitanović and E.A. Spiegel)

We start out with a recapitulation of the basic notions of dynamics. Our aim is narrow; we keep the exposition focused on prerequisites to the applications to be developed in this text. We assume that the reader is familiar with dynamics on the level of the introductory texts mentioned in sect. 1.8, and concentrate here on developing intuition about what a dynamical system can do. It will be a coarse brush sketch - a full description of all possible behaviors of dynamical systems is beyond human ken. Anyway, for a novice there is no shortcut through this lengthy detour; a sophisticated traveler might prefer to skip this well-trodden territory and embark upon the journey at chapter 9 .


### 2.1 Dynamical systems

In a dynamical system we observe the world as a function of time. We express our observations as numbers and record how they change with time; given sufficiently detailed information and understanding of the underlying natural laws, we see the future in the present as in a mirror. The motion of the planets against the celestial firmament provides an example. Against the daily motion of the stars from East to West, the planets distinguish themselves by moving among the fixed stars. Ancients discovered that by knowing a sequence of planet's positions - latitudes and longitudes - its future position could be predicted.

For the solar system, tracking the latitude and longitude in the celestial sphere suffices to completely specify the planet's apparent motion. All
possible values for positions and velocities of the planets form the phase space of the system. More generally, a state of a physical system, at a given instant in time, can be represented by a single point in an abstract space called state space or phase space $\mathcal{M}$. As the system changes, so does the representative point in phase space. We refer to the evolution of such points as dynamics, and the function $f^{t}$ which specifies where the representative point is at time $t$ as the evolution rule.

If there is a definite rule $f$ that tells us how this representative point moves in $\mathcal{M}$, the system is said to be deterministic. For a deterministic dynamical system, the evolution rule takes one point of the phase space and maps it into exactly one point. However, this is not always possible. For example, knowing the temperature today is not enough to predict the temperature tomorrow; knowing the value of a stock today will not determine its value tomorrow. The phase space can be enlarged, in the hope that in a sufficiently large phase space it is possible to determine an evolution rule, so we imagine that knowing the state of the atmosphere, measured over many points over the entire planet should be sufficient to determine the temperature tomorrow. Even that is not quite true, and we are less hopeful when it comes to stocks.

For a deterministic system almost every point has a unique future, so trajectories cannot intersect. We say "almost" because there might exist a set of measure zero (tips of wedges, cusps, etc.) for which a trajectory is [習 chapter 12 not defined. We may think such sets a nuisance, but it is quite the contrary - they will enable us to partition phase space, so that the dynamics can be better understood.

Locally, the phase space $\mathcal{M}$ looks like $\mathbb{R}^{d}$, meaning that $d$ numbers are sufficient to determine what will happen next. Globally, it may be a more complicated manifold formed by patching together several pieces of $\mathbb{R}^{d}$, forming a torus, a cylinder, or some other geometric object. When we need to stress that the dimension $d$ of $\mathcal{M}$ is greater than one, we may refer to the point $x \in \mathcal{M}$ as $x_{i}$ where $i=1,2,3, \ldots, d$. The evolution rule $f^{t}: \mathcal{M} \rightarrow \mathcal{M}$ tells us where a point $x$ is in $\mathcal{M}$ after a time interval $t$. The pair $(\mathcal{M}, f)$ constitute a dynamical system.

The dynamical systems we will be studying are smooth. This is expressed mathematically by saying that the evolution rule $f^{t}$ can be differentiated as many times as needed. Its action on a point $x$ is sometimes indicated by $f(x, t)$ to remind us that $f$ is really a function of two variables: the time and a point in phase space. Note that time is relative rather than absolute, so only the time interval is necessary. This follows from the fact that a point in phase space completely determines all future evolution, and it is not necessary to know anything else. The time parameter can be a real variable ( $t \in \mathbb{R}$ ), in which case the evolution is called a flow, or an integer $(t \in \mathbb{Z})$, in which case the evolution advances in discrete steps in time, given by iteration of a map. Actually, the evolution parameter need not be the physical time; for example, a time-stationary solution of a partial differential equation is parametrized by spatial variables. In such situations one talks of a "spatial profile" rather than a "flow".


Figure 2.1: (a) A trajectory traced out by the evolution rule $f^{t}$. Starting from the phase space point $x$, after a time $t$, the point is at $f^{t}(x)$. (b) The evolution rule $f^{t}$ can be used to map a region $\mathcal{M}_{i}$ of the phase space into the region $f^{t}\left(\mathcal{M}_{i}\right)$.

Nature provides us with innumerable dynamical systems. They manifest themselves through their trajectories: given an initial point $x_{0}$, the evolution rule traces out a sequence of points $x(t)=f^{t}\left(x_{0}\right)$, the trajectory through the point $x_{0}=x(0)$. A trajectory is parameterized by the time $t$ and thus belongs to $\left(f^{t}\left(x_{0}\right), t\right) \in \mathcal{M} \times \mathbb{R}$. By extension, we can also talk of the evolution of a region $\mathcal{M}_{i}$ of the phase space: just apply $f^{t}$ to every point in $\mathcal{M}_{i}$ to obtain a new region $f^{t}\left(\mathcal{M}_{i}\right)$, as in figure 2.1.

Because $f^{t}$ is a single-valued function, any point of the trajectory can be used to label the trajectory. If we mark the trajectory by its inital point $x_{0}$, we are describing it in the Lagrangian coordinates. We can regard the transport of the material point at $t=0$ to its current point $x(t)=f^{t}\left(x_{0}\right)$ as a coordinate transformation from the Lagrangian coordinates to the Eulerian coordinates.

The subset of points in $\mathcal{M}$ that belong to the (possibly infinite) trajectory of a given point $x_{0}$ is called the orbit of $x_{0}$; we shall talk about forward orbits, backward orbits, periodic orbits, etc.. For a flow, an orbit is a continuous curve; for a map, it is a sequence of points.

What are the possible trajectories? This is a grand question, and there are many answers, chapters to follow offering some. Here we shall classify possible trajectories as:

$$
\begin{array}{rll}
\text { stationary: } & f^{t}(x)=x & \text { for all } t \\
\text { periodic: } & f^{t}(x)=f^{t+T_{p}}(x) & \text { for a given minimum period } T_{p} \\
\text { aperiodic: } & f^{t}(x) \neq f^{t^{\prime}}(x) & \text { for all } t \neq t^{\prime}
\end{array}
$$

The ancient tried to make sense of all dynamics in terms of periodic motions; epicycles, integrable systems. The embarassing truth is that for a generic dynamical systems most motions are aperiodic. We will classify aperiodic motions into two subtypes: those that wander off, and those that keep coming back.

A point $x \in \mathcal{M}$ is called a wandering point, if there exists an open neighborhood $\mathcal{M}_{0}$ of $x$ to which the trajectory never returns

$$
\begin{equation*}
f^{t}(x) \notin \mathcal{M}_{0} \quad \text { for all } \quad t>t_{\min } . \tag{2.1}
\end{equation*}
$$

In physics literature, the dynamics of such state is often referred to as transient.

A periodic orbit corresponds to a trajectory that returns exactly to the initial point in a finite time. Periodic orbits form a very small subset of the phase space, in the same sense that rational numbers are a set of zero measure on the unit interval. For times much longer than a typical "turnover" time, it makes sense to relax the notion of exact periodicity, and replace it by the notion of recurrence. A point is recurrent or nonwandering if for any open neighborhood $\mathcal{M}_{0}$ of $x$ and any time $t_{\text {min }}$ there exists a later time $t$, such that

$$
\begin{equation*}
f^{t}(x) \in \mathcal{M}_{0} . \tag{2.2}
\end{equation*}
$$

In other words, the trajectory of a non-wandering point reenters the neighborhood $\mathcal{M}_{0}$ infinitely often. We shall denote by $\Omega$ the non-wandering set of $f$, that is, the union of all the non-wandering points of $\mathcal{M}$. The set $\Omega$, the non-wandering set of $f$, is the key to understanding the long-time behavior of a dynamical system; all calculations undertaken here will be carried out on non-wandering sets.

So much about individual trajectories. What about clouds of initial points? If there exists a connected phase space volume that maps into itself under forward evolution (and you can prove that by the method of Lyapunov functionals, or several other methods available in the literature), the flow is globally contracting onto a subset of $\mathcal{M}$ which we shall refer to as the attractor. The attractor may be unique, or there can coexist any number of distinct attracting sets, each with its own basin of attraction, the set of all points that fall into the attractor under foward evolution. The attractor can be a fixed point, a periodic orbit, aperiodic, or any combination of the above. The most interesting case is that of an aperiodic recurrent attractor, to which we shall refer loosely as a strange attractor. We say 'loosely', as will soon become apparent that diagnosing and proving existence of a genuine, card-carrying strange attractor is a highly nontrivial undertaking.

Conversely, if we can enclose the non-wandering set $\Omega$ by a connected phase space volume $\mathcal{M}_{0}$ and then show that almost all points within $\mathcal{M}_{0}$, but not in $\Omega$, eventually exit $\mathcal{M}_{0}$, we refer to the non-wandering set $\Omega$ as a repeller. An example of a repeller is not hard to come by - the pinball game of sect. 1.3 is a simple chaotic repeller.

It would seem, having said that the periodic points are so exceptional that almost all non-wandering points are aperiodic, that we have given up the ancients' fixation on periodic motions. Nothing could be further from truth. As longer and longer cycles approximate more and more accurately finite segments of aperiodic trajectories, we shall establish control over nonwandering sets by defining them as the closures of the union of all periodic points.

Before we can work out an example of a non-wandering set and get a better grip on what chaotic motion might look like, we need to ponder flows in a little more depth.

### 2.2 Flows

> There is no beauty without some strangeness.
> William Blake

A flow is a continuous-time dynamical system. The evolution rule $f^{t}$ is a family of mappings of $\mathcal{M} \rightarrow \mathcal{M}$ parameterized by $t \in \mathbb{R}$. Because $t$ represents a time interval, any family of mappings that forms an evolution rule must satisfy:
(a) $f^{0}(x)=x \quad$ (in 0 time there is no motion)
(b) $f^{t}\left(f^{t^{\prime}}(x)\right)=f^{t+t^{\prime}}(x) \quad$ (the evolution law is the same at all times)
(c) the mapping $(x, t) \mapsto f^{t}(x)$ from $\mathcal{M} \times \mathbb{R}$ into $\mathcal{M}$ is continuous.

The family of mappings $f^{t}(x)$ thus forms a continuous (forward semi-) group. Why "semi-" group? It may fail to form a group if the dynamics is not reversible, and the rule $f^{t}(x)$ cannot be used to rerun the dynamics backwards in time, with negative $t$; with no reversibility, we cannot define the inverse $f^{-t}\left(f^{t}(x)\right)=f^{0}(x)=x$, in which case the family of mappings $f^{t}(x)$ does not form a group. In exceedingly many situations of interest for times beyond the Lyapunov time, for asymptotic attractors, for dissipative partial differential equations, for systems with noise, for non-invertible maps - the dynamics cannot be run backwards in time, hence, the circumspect emphasis on semigroups. On the other hand, there are many settings of physical interest, where dynamics is reversible (such as finite-dimensional Hamiltonian flows), and where the family of evolution maps $f^{t}$ does form a group.

For infinitesimal times, flows can be defined by differential equations. We write a trajectory as

$$
\begin{equation*}
x(t+\tau)=f^{t+\tau}\left(x_{0}\right)=f\left(f\left(x_{0}, t\right), \tau\right) \tag{2.3}
\end{equation*}
$$

and express the time derivative of a trajectory at point $x(t)$,

$$
\begin{equation*}
\left.\frac{d x}{d \tau}\right|_{\tau=0}=\left.\partial_{\tau} f\left(f\left(x_{0}, t\right), \tau\right)\right|_{\tau=0}=\dot{x}(t) \tag{2.4}
\end{equation*}
$$



Figure 2.2: (a) The two-dimensional vector field for the Duffing system (2.6), together with a short trajectory segment. (b) The flow lines. Each "comet" represents the same time interval of a trajectory, starting at the tail and ending at the head. The longer the comet, the faster the flow in that region.
as the time derivative of the evolution rule, a vector evaluated at the same point. By considering all possible trajectories, we obtain the vector $\dot{x}(t)$ at any point $x \in \mathcal{M}$. This vector field is a (generalized) velocity field:

$$
\begin{equation*}
v(x)=\dot{x}(t) . \tag{2.5}
\end{equation*}
$$

Newton's laws, Lagrange's method, or Hamilton's method are all familiar procedures for obtaining a set of differential equations for the vector field $v(x)$ that describes the evolution of a mechanical system. Equations of mechanics may appear different in form from (2.5), as they are often involve higher time derivatives, but an equation that is second or higher order in time can always be rewritten as a set of first order equations.

We are concerned here with a much larger world of general flows, mechanical or not, defined by a time-independent vector field (2.5). At each point of the phase space a vector indicates the local direction in which the orbit evolves. The length of the vector $|v(x)|$ is proportional to the speed at the point $x$, and the direction and length of $v(x)$ changes from point to point. When the phase space is a complicated manifold embedded in $\mathbb{R}^{d}$, one can no longer think of the vector field as being embedded in the phase space. Instead, we have to imagine that each point $x$ of phase space has a different tangent plane $T \mathcal{M}_{x}$ attached to it. The vector field lives in the union of all these tangent planes, a space called the tangent bundle $\mathbf{T} \mathcal{M}$.

Example 2.1 A two-dimensional vector field $v(x)$ : A simple example of a flow is afforded by the unforced Duffing system

$$
\begin{align*}
\dot{x}(t) & =y(t) \\
\dot{y}(t) & =-0.15 y(t)+x(t)-x(t)^{3} \tag{2.6}
\end{align*}
$$

plotted in figure 2.2. The velocity vectors are drawn superimposed over the configuration coordinates $(x(t), y(t))$ of phase space $\mathcal{M}$, but they belong to a different space, the tangent bundle $\mathbf{T} \mathcal{M}$.

$$
\begin{equation*}
\text { If } \quad v\left(x_{q}\right)=0, \tag{2.7}
\end{equation*}
$$

Figure 2.3: A trajectory of the Rössler flow
 at time $t=250$. (G. Simon)
$x_{q}$ is an equilibrium point (often referred to as a stationary, fixed, or stagnation point), and the trajectory remains forever stuck at $x_{q}$. Otherwise the trajectory passing though $x_{0}$ at time $t=0$ can be obtained by integrating the equations (2.5):

$$
\begin{equation*}
x(t)=f^{t}\left(x_{0}\right)=x_{0}+\int_{0}^{t} d \tau v(x(\tau)), \quad x(0)=x_{0} . \tag{2.8}
\end{equation*}
$$

We shall consider here only autonomous flows, that is, flows for which the velocity field $v_{i}$ is stationary, not explicitly dependent on time. A nonautonomous system

$$
\begin{equation*}
\frac{d y}{d \tau}=w(y, \tau), \tag{2.9}
\end{equation*}
$$

can always be converted into a system where time does not appear explicitly. To do so, extend ("suspend") phase space to be ( $d+1$ )-dimensional by defining $x=\{y, \tau\}$, with a stationary vector field

$$
v(x)=\left[\begin{array}{c}
w(y, \tau)  \tag{2.10}\\
1
\end{array}\right]
$$

The new flow $\dot{x}=v(x)$ is autonomous, and the trajectory $y(\tau)$ can be read page 47 off $x(t)$ by ignoring the last component of $x$.

Example 2.2 A flow with a strange attractor: The Duffing flow of figure 2.2 is bit of a bore - every trajectory ends up in one of the two attractive equilibrium points. Let's construct a flow that does not die out, but exhibits a recurrent dynamics. Start with a harmonic oscillator

$$
\begin{equation*}
\dot{x}=-y, \quad \dot{y}=x \tag{2.11}
\end{equation*}
$$

The solutions are $A e^{i t}, A e^{-i t}$, and the whole $x-y$ plane rotates with constant angular velocity $\theta=1$, period $T=2 \pi$. Now make the system unstable by adding

$$
\begin{equation*}
\dot{x}=-y, \quad \dot{y}=x+a y, \quad a>0 . \tag{2.12}
\end{equation*}
$$

The plane is still rotating with constant angular velocity, but trajectories are now spiraling out. In general, any flow in the plane either escapes, falls into an attracting equilibrium point, or converges to a limit cycle - richer dynamics requires at least one more dimension. In order to prevent the trajectory from escaping to $\infty$, kick it into 3rd dimension when $x$ reaches some value $c$ by adding

$$
\begin{equation*}
\dot{z}=b+z(x-c), \quad c>0 . \tag{2.13}
\end{equation*}
$$

Now $z$ shoots upwards exponentially, $z \simeq e^{(x-c) t}$. In order to bring it back, start decreasing $x$ by modifing its evolution equation to

$$
\dot{x}=-y-z .
$$

Large $z$ drives the trajectory toward $x=0$; there the exponential contraction by $e^{-c t}$ kicks in, and the trajectory drops back toward the $x-y$ plane. This frequently studied example of an autonomous flow is called the Rössler system (for definitiveness we fix the parameters $a, b, c$ in what follows):

$$
\begin{align*}
\dot{x} & =-y-z \\
\dot{y} & =x+a y \\
\dot{z} & =b+z(x-c), \quad a=b=0.2, \quad c=5.7 \tag{2.14}
\end{align*}
$$

The system is as simple as they get - it would be linear, were it not for the sole bilinear term $z x$. Even for so "simple" a system the nature of long-time solutions is far from obvious.

There are two repelling equilibrium points:

$$
\begin{align*}
& \left(x^{-}, y^{-}, z^{-}\right)=(0.0070,-0.0351,0.0351) \\
& \left(x^{+}, y^{+}, z^{+}\right)=(5.6929,-28.464,28.464) \tag{2.15}
\end{align*}
$$

One is close to the origin by construction - the other, some distance away, must exist because the equilibrium has a $2 n d$-order nonlinearity.

To see what other solutions look like we need to resort to numerical integration. A typical numerically integrated long-time trajectory is sketched in figure 2.3. As we shall show in sect. 4.1, for this flow any finite volume of initial conditions shrinks with time, so the flow is contracting. Trajectories that start out sufficiently close to the origin seem to converge to a strange attractor. We say "seem", as there exists no proof that such an attractor is asymptotically aperiodic - it might well be that what we see is but a long transient on a way to an attractive periodic orbit. For now, accept that figure 2.3 and similar figures in what follows are examples of "strange attractors". (continued in exercise 2.8)
(Rytis Paškauskas)

fast track:
chapter 3, p. 51

### 2.3 Computing trajectories

On two occasions I have been asked [by members of Parliament], 'Pray, Mr. Babbage, if you put into the machine wrong figures, will the right answers come out?' I am not able rightly to apprehend the kind of confusion of ideas that could provoke such a question. Charles Babbage

You have not learned dynamics unless you know how to integrate numerically whatever dynamical equations you face. Sooner or later, you need to implement some finite time-step prescription for integration of the equations of motion (2.5). The simplest is the Euler integrator which advances the trajectory by $\delta \tau \times$ velocity at each time step:

$$
\begin{equation*}
x_{i} \rightarrow x_{i}+v_{i}(x) \delta \tau \tag{2.16}
\end{equation*}
$$

This might suffice to get you started, but as soon as you need higher numerical accuracy, you will need something better. There are many excellent reference texts and computer programs that can help you learn how to solve differential equations numerically using sophisticated numerical tools, such as pseudo-spectral methods or implicit methods. If a "sophisticated" integration routine takes days and gobbles up terabits of memory, you are using brain-damaged high level software. Try writing a few lines of your own Runge-Kuta code in some mundane everyday language. While you absolutely need to master the requisite numerical methods, this is neither the time nor the place to expound upon them; how you learn them is your business. And if you have developed some nice routines for solving problems in this text or can point another student to some, let us know.

In the next section we dispose of the fear of "infinite-dimensional" dynamical systems - you might prefer to skip the section on first reading.

fast track:
chapter 3, p. 51

### 2.4 Infinite-dimensional flows

> Romeo: 'Misshapen chaos of well seeming forms!'
> W. Shakespeare, Romeo and Julliet, Act I, Scene I
-
There is only one honorable cause that would justify sweating through so much formalism - this is but the sharpening of a pencil in order that we may attack the Navier-Stokes equation,

$$
\begin{equation*}
\rho\left(\frac{\partial \mathbf{u}}{\partial t}+\mathbf{u} \cdot \nabla \mathbf{u}\right)=-\nabla p+\eta\left(\nabla^{2} \mathbf{u}+\frac{1}{3} \nabla(\operatorname{div} \mathbf{u})\right)+\mathbf{f} \tag{2.17}
\end{equation*}
$$

and solve the problem of turbulence. Being realistic, we are not so foolhardy to immediately plunge into the problem - there are too many dimensions and indices. Instead, we start small, in one spatial dimension, $\mathbf{u} \rightarrow u$, $\mathbf{u} \cdot \nabla \mathbf{u} \rightarrow \frac{1}{2} \partial_{x} u^{2}$, assume constant $\rho$, forget about the pressure $p$, and so on. This line of reasoning, as well as many other equally sensible threads of thought, such as the amplitude equations obtained via weakly nonlinear

Figure 2.4: Spatiotemporally periodic solution $u_{0}(x, t)$. We have divided $x$ by $\pi$ and plotted only the $x>0$ part, since we work in the subspace of the odd solutions, $u(x, t)=$ $-u(-x, t) . \quad N=16$ Fourier modes truncation
 with $\nu=0.029910$. (From ref. [2.10])
stability analysis of steady flows, lead to the essentially same nonlinear PDEs, like the one that we turn to in the next section.

Flows described by partial differential equations are considered infinite dimensional because if one writes them down as a set of ordinary differential equations (ODEs), one needs infinitely many of them to represent the dynamics of one partial differential equation (PDE). Even though their phase space is infinite dimensional, the dynamics of many systems of physical interest converges toward finite-dimensional global attractors. We illustrate how this works with a concrete example, the Kuramoto-Sivashinsky system.


### 2.4.1 Fluttering flame front

The Kuramoto-Sivashinsky system, arising in the description of the flame front flutter of gas burning in a cylindrically symmetric burner on your kitchen stove and many other problems of greater import, is one of the simplest partial differential equations that exhibit chaos. It is a dynamical system extended in one spatial dimension, defined by

$$
\begin{equation*}
u_{t}=\left(u^{2}\right)_{x}-u_{x x}-\nu u_{x x x x} \tag{2.18}
\end{equation*}
$$

In this equation $t \geq 0$ is the time and $x \in[0,2 \pi]$ is the spatial coordinate. The subscripts $x$ and $t$ denote partial derivatives with respect to $x$ and $t$; $u_{t}=d u / d t, u_{x x x x}$ stands for the 4th spatial derivative of the "height of the flame front" (more correctly the "velocity of the flame front") $u=u(x, t)$ at position $x$ and time $t$. The "viscosity" parameter $\nu$ controls the suppression of solutions with fast spatial variations. We take note, as in the NavierStokes equation (2.17), of the $u \partial_{x} u$ "inertial" term, the $\partial_{x}^{2} u$ "diffusive" term (both with a "wrong" sign), etc.

The term $\left(u^{2}\right)_{x}$ makes this a nonlinear system. It is one of the simplest conceivable nonlinear PDE, playing the role in the theory of spatially extended systems analogous to the role that the $x^{2}$ nonlinearity plays in the dynamics of iterated mappings. The time evolution of a solution of the Kuramoto-Sivashinsky system is illustrated by figure 2.4. How are such solutions computed? The salient feature of such partial differential equations
is a theorem saying that for any finite value of the phase-space contraction parameter $\nu$, the asymptotic dynamics is describable by a finite set of "inertial manifold" ordinary differential equations.

We are studying a "flame front" $u(x, t)=u(x+2 \pi, t)$ periodic on the $x \in[0,2 \pi]$ interval, so a reasonable strategy (but by no means the only one) is to expand it in a discrete spatial Fourier series:

$$
\begin{equation*}
u(x, t)=\sum_{k=-\infty}^{+\infty} b_{k}(t) e^{i k x} \tag{2.19}
\end{equation*}
$$

Since $u(x, t)$ is real, $b_{k}=b_{-k}^{*}$. Substituting (2.19) into (2.18) yields the infinite ladder of evolution equations for the Fourier coefficients $b_{k}$ :

$$
\begin{equation*}
\dot{b}_{k}=\left(k^{2}-\nu k^{4}\right) b_{k}+i k \sum_{m=-\infty}^{\infty} b_{m} b_{k-m} . \tag{2.20}
\end{equation*}
$$

As it follows from this equation that $\dot{b}_{0}=0$, the solution integrated over space is constant in time. We shall consider only the case of this average the mean value of $u$ - equal to zero, $b_{0}=\int d x u(x, t)=0$.

The coefficients $b_{k}$ are in general complex functions of time $t$. We can isolate a smaller subspace of the system (2.20) further by considering the case of $b_{k}$ pure imaginary, $b_{k}=i a_{k}$, where $a_{k}$ are real, with the evolution equations

$$
\begin{equation*}
\dot{a}_{k}=\left(k^{2}-\nu k^{4}\right) a_{k}-k \sum_{m=-\infty}^{\infty} a_{m} a_{k-m} . \tag{2.21}
\end{equation*}
$$

This picks out the subspace of odd solutions $u(x, t)=-u(-x, t)$, so $a_{-k}=$ $-a_{k}$. By picking this subspace we eliminate the continuous translational symmetry from our consideration; that is probably not an option for an experimentalist, but will do for our purposes.

That is the infinite set of ordinary differential equations promised at the beginning of the section.

The trivial solution $u(x, t)=0$ is an equilibrium point of (2.18), but that is basically all we know as far as analytical solutions are concerned. You can integrate numerically the Fourier modes (2.21), truncating the ladder of equations to a finite number of modes $N$, that is, set $a_{k}=0$ for $k>N$. In the applied mathematics literature such a truncation is called a Gälerkin truncation, or a Gälerkin projection. For the parameter values explored below, $N \leq 16$ truncations were deemed sufficiently accurate. In other words, even though our starting point (2.18) is an infinite-dimensional dynamical system, the asymptotic dynamics unfolds on a finite-dimensional attracting manifold, and so we are back on the familiar territory of sect. 2.2:


Figure 2.5: Projections of a typical 16-dimensional trajectory onto different 3dimensional subspaces, coordinates (a) $\left\{a_{1}, a_{2}, a_{3}\right\}$, (b) $\left\{a_{1}, a_{2}, a_{4}\right\}$. $N=16$ Fourier modes truncation with $\nu=0.029910$. (From ref. [2.10].)
the theory of a finite number of ODEs applies to this infinite-dimensional PDE as well.

Once the trajectory is computed in Fourier space, we can recover and plot the corresponding spatiotemporal pattern $u(x, t)$ over the configuration space using (2.19), as in figure 2.4.

### 2.4.2 Fourier modes truncations

Thinking is extra price.
Julio Cesar Solla
Consider now the case of initial $a_{k}$ sufficiently small that the bilinear $a_{m} a_{k-m}$ terms in (2.21) can be neglected. Then we have a set of decoupled linear equations for $a_{k}$ whose solutions are exponentials, at most a finite number for which $k^{2}>\nu k^{4}$ is growing with time, and infinitely many with $\nu k^{4}>k^{2}$ decaying in time. The growth of the unstable long wavelengths (low $|k|$ ) excites the short wavelengths through the nonlinear term in (2.21). The excitations thus transferred are dissipated by the strongly damped short wavelengths, and a "chaotic equilibrium" can emerge. The very short wavelengths $|k| \gg 1 / \sqrt{\nu}$ remain small for all times, but the intermediate wavelengths of order $|k| \sim 1 / \sqrt{\nu}$ play an important role in maintaining the dynamical equilibrium. As the damping parameter decreases, the solutions increasingly take on shock front character poorly represented by the Fourier basis, and many higher harmonics may need to be kept in truncations of (2.21).

Hence, while one may truncate the high modes in the expansion (2.21), care has to be exercised to ensure that no modes essential to the dynamics are chopped away. In practice one does this by repeating the same calculation at different truncation cutoffs $N$, and making sure that the inclusion of additional modes has no effect within the desired accuracy. For the figures given here, the numerical calculations were performed taking $N=16$ and the damping parameter value $\nu=0.029910$, for which the system is chaotic (as far as can be determined numerically).

The problem with such high-dimensional truncations of the infinite tower of equations (2.21) is that the dynamics is difficult to visualize. The best we can do without much programming is to examine the trajectory's projections onto any three axes $a_{i}, a_{j}, a_{k}$, as in figure 2.5.

Examination of numerical plots such as figure 2.5 suggests that a more thoughtful approach would be to find a coordinate transformation $y=h(x)$ to a "center manifold", such that in the new, curvilinear coordinates largescale dynamics takes place in $\left(y_{1}, y_{2}\right)$ coordinates, with exponentially small dynamics in $y_{3}, y_{4} \cdots$. But - thinking is extra price - we do not know how to actually accomplish this.


We can now start to understand the remark on page 35 that for infinite dimensional systems time reversability is not an option: evolution forward in time strongly damps the higher Fourier modes. There is no turning back: if we reverse the time, the infinity of high modes that contract strongly forward in time now explodes, instantly rendering evolution backward in time meaningless. As everything in dynamics, this claim is also wrong, in a subtle way: if the initial $u(x, 0)$ is in the non-wandering set (2.2), the trajectory is well defined both forward and backward in time. For practical purposes, this subtlety is not of much use, as any time-reversed numerical trajectory in a finite-mode truncation will explode very quickly, unless special precautions are taken.

## Commentary

Remark 2.1 Dynamical systems. The defintinition of dynamical systems given in sect. 2.1 is ascribed to M.W. Hirsch [2.19] .

Remark 2.2 Model ODE and PDE systems. The Duffing system (2.6) arises in the study of electronic circuits. Rössler system was introduced in ref. [2.3] as a simplified set of equations describing no particular physical system, but capturing the essence of chaos in a simplest imaginable smooth flow. Otto Rössler, a man of classical education, was inspired in this quest by that rarely cited grandfather of chaos, Anaxagoras (456 B.C.). This, and references to earlier work can be found in refs. [2.4, 2.6]. We recommend in particular the inimitable Abraham and Shaw illustrated classic [2.5] for its beautiful sketches of the Rössler and many other flows.

The theorem on finite dimenionality of inertial manifolds of phase-space contracting PDE flows is proven in ref. [2.7]. The Kuramoto-Sivashinsky equation was introduced in refs. [2.8, 2.9]; sect. 2.4 is based on V. Putkaradze's term project (see ChaosBook.org/extras). and on the Christiansen et al. article [2.10]. Holmes, Lumley and Berkooz [2.11] offer a delightful discussion of why this system deserves study as a staging ground for studying turbulence in full-fledged Navier-Stokes equation. How good a description of a flame front this equation is
need not concern us here; suffice it to say that such model amplitude equations for interfacial instabilities arise in a variety of contexts - see e.g. ref. [2.12] - and this one is perhaps the simplest physically interesting spatially extended nonlinear system.

Remark 2.3 Diagnosing chaos. In sect. 1.3.1 we have stated that a deterministic system exhibits "chaos" if its dynamics is locally unstable (positive Lyapunov exponent) and globally mixing (positive entropy). In sect. 10.3 we shall define Lyapunov exponents, and discuss their evaluation, but already at this point it would be handy to have a few quick numerical methods to diagnose chaotic dynamics. Laskar's frequency analysis method [2.16] is useful for extracting quasi-periodic and weakly chaotic regions of phase space in Hamiltonian dynamics with many degrees of freedom. For references to several other numerical methods, see ref. [2.17].

## Résumé

A dynamical system - a flow, or an iterated map - is defined by specifying a pair $(\mathcal{M}, f)$, where $\mathcal{M}$ is the phase space and $f: \mathcal{M} \rightarrow \mathcal{M}$. The key concepts in exploration of the long time dynamics are the notions of recurrence and of the non-wandering set of $f$, the union of all the non-wandering points of $\mathcal{M}$.

Chaotic dynamics with a low-dimensional attractor can be visualized as a succession of nearly periodic but unstable motions. In the same spirit, turbulence in spatially extended systems can be described in terms of recurrent spatiotemporal patterns. Pictorially, dynamics drives a given spatially extended system through a repertoire of unstable patterns; as we watch a turbulent system evolve, every so often we catch a glimpse of a familiar pattern. For any finite spatial resolution and finite time the system follows approximately a pattern belonging to a finite repertoire of possible patterns, and the long-term dynamics can be thought of as a walk through the space of such patterns. Recasting this image into mathematics is the subject of this book.

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## Exercises

The problems that you should do have underlined titles. The rest (smaller type) are optional. Difficult problems are marked by any number of ${ }^{* * *}$ stars.

Exercise 2.1 Trajectories do not intersect. A trajectory in the phase space $\mathcal{M}$ is the set of points one gets by evolving $x \in \mathcal{M}$ forwards and backwards in time:

$$
C_{x}=\left\{y \in \mathcal{M}: f^{t}(x)=y \text { for } t \in \mathbb{R}\right\}
$$

Show that if two trajectories intersect, then they are the same curve.

Exercise 2.2 Evolution as a group. The trajectory evolution $f^{t}$ is a oneparameter (semi-)group where

$$
f^{t+s}=f^{t} \circ f^{s}
$$

Show that it is a commutative (semi-)group.

In this case, the commutative character of the (semi-)group of evolution functions comes from the commutative character of the time parameter under addition. Can you think of any other (semi-)group replacing time?

Exercise 2.3 Almost ODE's.
(a) Consider the point $x$ on $\mathbb{R}$ evolving according $\dot{x}=e^{\dot{x}}$. Is this an ordinary differential equation?
(b) Is $\dot{x}=x(x(t))$ an ordinary differential equation?
(c) What about $\dot{x}=x(t+1)$ ?

Exercise 2.4 All equilibrium points are fixed points. Show that a point of a vector field $v$ where the velocity is zero is a fixed point of the dynamics $f^{t}$.

Exercise 2.5 Gradient systems. Gradient systems (or "potential problems") are a simple class of dynamical systems for which the velocity field is given by the gradient of an auxiliary function, the "potential" $\phi$

$$
\dot{x}=-\nabla \phi(x)
$$

where $x \in \mathbb{R}^{d}$, and $\phi$ is a function from that space to the reals $\mathbb{R}$.
(a) Show that the velocity of the particle is in the direction of most rapid decrease of the function $\phi$.
(b) Show that all extrema of $\phi$ are fixed points of the flow.
(c) Show that it takes an infinite amount of time for the system to reach an equilibrium point.
(d) Show that there are no periodic orbits in gradient systems.

Exercise 2.6 Runge-Kutta integration. Implement the fourth-order Runge-Kutta integration formula (see, for example, ref. [2.13]) for $\dot{x}=v(x)$ :

$$
\begin{align*}
x_{n+1} & =x_{n}+\frac{k_{1}}{6}+\frac{k_{2}}{3}+\frac{k_{3}}{3}+\frac{k_{4}}{6}+O\left(\delta \tau^{5}\right) \\
k_{1} & =\delta \tau v\left(x_{n}\right), \quad k_{2}=\delta \tau v\left(x_{n}+k_{1} / 2\right) \\
k_{3} & =\delta \tau v\left(x_{n}+k_{2} / 2\right), \quad k_{4}=\delta \tau v\left(x_{n}+k_{3}\right) . \tag{2.22}
\end{align*}
$$

If you already know your Runge-Kutta, program what you believe to be a better numerical integration routine, and explain what is better about it.

Exercise 2.7 Rössler system. Use the result of exercise 2.6 or some other integration routine to integrate numerically the Rössler system (2.14). Does the result look like a "strange attractor"? If you happen to already know what fractal dimensions are, argue (possibly on basis of numerical integration) that this attractor is of dimension smaller than $\mathbb{R}^{3}$.

## Exercise 2.8 Equilibria of the Rössler system.

(a) Find all equilibrium points $\left(x^{q}, y^{q}, z^{q}\right)$ of the Rössler system (2.14). How many are there?
(b) Assume that $b=a$ in (2.14). Define parameters

$$
\begin{align*}
& \epsilon=a / c \\
& D=1-4 \epsilon^{2}  \tag{2.23}\\
& p^{ \pm}=(1 \pm \sqrt{D}) / 2
\end{align*}
$$

Express all the equilibria in terms of $\left(p^{ \pm}, c\right)$. Expand equilibria to the first subleading order in $\epsilon$. Note that expansions in $\epsilon$ make sense because for $a=b=0.2, c=5.7$ in (2.14), $\epsilon \approx 0.035$.
(continued as exercise 3.1)

## (Rytis Paškauskas)

Exercise 2.9 Can you integrate me? Integrating equations numerically is not for the faint of heart. It is not always possible to establish that a set of nonlinear ordinary differential equations has a solution for all times and there are many cases were the solution only exists for a limited time interval, as, for example, for the equation $\dot{x}=x^{2}, \quad x(0)=1$.
(a) For what times do solutions of

$$
\dot{x}=x(x(t))
$$

exist? Do you need a numerical routine to answer this question?
(b) Let's test the integrator you wrote in exercise 2.6. The equation $\ddot{x}=-x$ with initial conditions $x(0)=2$ and $\dot{x}=0$ has as solution $x(t)=e^{-t}\left(1+e^{2 t}\right)$. Can your integrator reproduce this solution for the interval $t \in[0,10]$ ? Check you solution by plotting the error as compared to the exact result.
(c) Now we will try something a little harder. The equation is going to be third order

$$
\dddot{x}+0.6 \ddot{x}+\dot{x}-|x|+1=0,
$$

which can be checked - numerically - to be chaotic. As initial conditions we will always use $\ddot{x}(0)=\dot{x}(0)=x(0)=0$. Can you reproduce the result $x(12)=0.8462071873$ (all digits are significant)? Even though the equation being integrated is chaotic, the time intervals are not long enough for the exponential separation of trajectories to be noticeble (the exponential growth factor is $\approx 2.4$ ).
(d) Determine the time interval for which the solution of $\dot{x}=x^{2}, x(0)=1$ exists.

Exercise 2.10 Classical collinear helium dynamics. In order to apply periodic orbit theory to quantization of helium we shall need to compute classical periodic orbits of the helium system. In this exercise we commence their evaluation for the collinear helium atom (5.6)

$$
H=\frac{1}{2} p_{1}^{2}+\frac{1}{2} p_{2}^{2}-\frac{Z}{r_{1}}-\frac{Z}{r_{2}}+\frac{1}{r_{1}+r_{2}}
$$

The nuclear charge for helium is $Z=2$. Colinear helium has only 3 degrees of freedom and the dynamics can be visualized as a motion in the $\left(r_{1}, r_{2}\right), r_{i} \geq 0$ quadrant. In $\left(r_{1}, r_{2}\right)$-coordinates the potential is singular for $r_{i} \rightarrow 0$ nucleuselectron collisions. These 2-body collisions can be regularized by rescaling the coordinates, with details given in sect. 7.3. In the transformed coordinates $\left(x_{1}, x_{2}, p_{1}, p_{2}\right)$ the Hamiltonian equations of motion take the form

$$
\begin{array}{ll}
\dot{P}_{1}=2 Q_{1}\left[2-\frac{P_{2}^{2}}{8}-Q_{2}^{2}\left(1+\frac{Q_{2}^{2}}{R^{4}}\right)\right] ; & \dot{Q}_{1}=\frac{1}{4} P_{1} Q_{2}^{2} \\
\dot{P}_{2}=2 Q_{2}\left[2-\frac{P_{1}^{2}}{8}-Q_{1}^{2}\left(1+\frac{Q_{1}^{2}}{R^{4}}\right)\right] ; & \dot{Q}_{2}=\frac{1}{4} P_{2} Q_{1}^{2} . \tag{2.24}
\end{array}
$$

where $R=\left(Q_{1}^{2}+Q_{2}^{2}\right)^{1 / 2}$.
(a) Integrate the equations of motion by the fourth order Runge-Kutta computer routine of exercise 2.6 (or whatever integration routine you like). A convenient way to visualize the 3-d phase space orbit is by projecting it onto the 2-dimensional $\left(r_{1}(t), r_{2}(t)\right)$ plane.

## Exercise 2.11 Infinite dimensional dynamical systems are not smooth.

Many of the operations we consider natural for finite dimensional systems do not have smooth behavior in infinite dimensional vector spaces. Consider, as an example, a concentration $\phi$ diffusing on $\mathbb{R}$ according to the diffusion equation

$$
\partial_{t} \phi=\frac{1}{2} \nabla^{2} \phi .
$$

(a) Interpret the partial differential equation as an infinite dimensional dynamical system. That is, write it as $\dot{x}=F(x)$ and find the velocity field.
(b) Show by examining the norm

$$
\|\phi\|^{2}=\int_{\mathbb{R}} d x \phi^{2}(x)
$$

that the vector field $F$ is not continuous.
(c) Try the norm

$$
\|\phi\|=\sup _{x \in \mathbb{R}}|\phi(x)| .
$$

Is F continuous?
(d) Argue that the semi-flow nature of the problem is not the cause of our difficulties.
(e) Do you see a way of generalizing these results?

## Chapter 3

## Maps

(R. Mainieri and P. Cvitanović)

The time parameter in the sect. 2.1 definition of a dynamical system can be either continuous or discrete. Discrete time dynamical systems arise naturally from flows; one can observe the flow at fixed time intervals (by strobing it), or one can record the coordinates of the flow when a special event happens (the Poincaré section method). This triggering event can be as simple as vanishing of one of the coordinates, or as complicated as the flow cutting through a curved hypersurface.

### 3.1 Poincaré sections

Successive trajectory intersections with a Poincaré section, a $d$-dimensional hypersurface or a set of hypersurfaces $\mathcal{P}$ embedded in the ( $d+1$ )-dimensional phase space $\mathcal{M}$, define the Poincaré return map $P(x)$, a $d$-dimensional map of form

$$
\begin{equation*}
x^{\prime}=P(x)=f^{\tau(x)}(x), \quad x^{\prime}, x \in \mathcal{P} . \tag{3.1}
\end{equation*}
$$

(For economy of notation, the maps of this chapter will be taken to be $d$ dimensional, the associated flows $(d+1)$-dimensional). Here the first return function $\tau(x)$ - sometimes referred to as the ceiling function - is the time of flight to the next section for a trajectory starting at $x$. The choice of the section hypersurface $\mathcal{P}$ is altogether arbitrary. The hypersurface can be specified implicitly through a function $U(x)$ that is zero whenever a point $x$ is on the Poincaré section.

The gradient of $U(x)$ evaluated at $x \in \mathcal{P}$ serves a two-fold function. First, the flow should pierce the hypersurface $\mathcal{P}$, rather than being tangent to it. A nearby point $x+\delta x$ is in the hypersurface $\mathcal{P}$ if $U(x+\delta x)=0$.

A nearby point on the trajectory is given by $\delta x=v \delta t$, so a traversal is ensured by the transversality condition

$$
\begin{equation*}
(v \cdot U)=v_{j}(x) U_{j}(x) \neq 0, \quad U_{j}(x)=\frac{d}{d x_{j}} U(x), \quad x \in \mathcal{P} . \tag{3.2}
\end{equation*}
$$

Second, the gradient $U_{j}$ defines the orientation of the hypersurface $\mathcal{P}$. The flow is oriented as well, and even the shortest periodic orbit can pierce $\mathcal{P}$ twice, traversing it in either direction. Hence a more natural definition of Poincaré return map $P(x)$ is obtained by supplementing (3.1) with the orientation condition

$$
\begin{array}{ll}
x_{n+1}=P\left(x_{n}\right), \quad & U\left(x_{n+1}\right)=U\left(x_{n}\right)=0, \quad n \in \mathbb{Z}^{+} \\
& \sum_{j=1}^{d+1} v_{j}\left(x_{n}\right) U_{j}\left(x_{n}\right)>0 . \tag{3.3}
\end{array}
$$

In this way the continus time $t$ flow $f^{t}(x)$ is reduced to a discrete time $n$ sequence $x_{n}$ of successive oriented trajectory traversals of $\mathcal{P}$.

With a sufficiently clever choice of a Poincaré section or a set of sections, any orbit of interest intersects a section. Depending on the application, one might need to convert the discrete time $n$ back to the continuous flow time. This is accomplished by adding up the first return function times $\tau\left(x_{n}\right)$, with the accumulated flight time given by

$$
\begin{equation*}
t_{n+1}=t_{n}+\tau\left(x_{n}\right), \quad t_{0}=0, \quad x_{n} \in \mathcal{P} . \tag{3.4}
\end{equation*}
$$

Other quantities integrated along the trajectory can be defined in a similar manner, and will need to be evaluated in the process of evaluating dynamical averages.

A few examples may help visualize this.

Example 3.1 Hyperplane $\mathcal{P}$ : The simplest choice of a Poincaré section is a plane specified by a point (located at the tip of the vector $r_{0}$ ) and a direction vector a perpendicular to the plane. A point $x$ is in this plane if it satisfies the condition

$$
\begin{equation*}
U(x)=\left(x-r_{0}\right) \cdot a=0 . \tag{3.5}
\end{equation*}
$$

Consider a circular periodic orbit centered at $r_{0}$, but not lying in $\mathcal{P}$. It pierces the hyperplane twice; the $(v \cdot a)>0$ traversal orientation condition (3.3) ensures that the first return time is the full period of the cycle.


Figure 3.1: (Right:) a sequence of Poincaré sections of the Rössler strange attractor, defined by planes through the $z$ axis, oriented at angles (a) $-60^{\circ}$ (b) $0^{\circ}$, (c) $60^{\circ}$, (d) $120^{\circ}$, in the $x-y$ plane. (Left:) side and $x-y$ plane view of trajectory with Poincaré sections superimposed. (Rytis Paškauskas)

Example 3.2 Pendulum: The phase space of a simple pendulum is 2-dimensional: momentum on the vertical axis and position on the horizontal axis. We choose the Poincaré section to be the positive horizontal axis. Now imagine what happens as a point traces a trajectory through this phase space. As long as the motion is oscillatory, in the pendulum all orbits are loops, so any trajectory will periodically intersect the line, that is the Poincaré section, at one point.

Consider next a pendulum with friction, such as the unforced Duffing system plotted in figure 2.2. Now every trajectory is an inward spiral, and the trajectory will intersect the Poincaré section $y=0$ at a series of points that get closer and closer to either of the equilibrium points; the Duffing oscillator at rest.

Motion of a pendulum is so simple that you can sketch it yourself on a piece of paper. The next two examples offer a better illustration of the utility of visualization of dynamics by means of Poincaré sections.

Example 3.3 Rössler attractor: Consider figure 2.3, a typical trajectory of the 3-dimensional Rössler flow (2.14). It wraps around the $z$ axis, so a good choice for a Poincaré section is a plane passing through the $z$ axis. A sequence of such Poincaré sections placed radially at increasing angles with respect to the $x$ axis, figure 3.1, illustrates the "stretch \& fold" action of the Rössler flow. To orient yourself, compare this with figure 2.3, and note the different $z$-axis scales. Figure 3.1 assembles these sections into a series of snapshots of the flow. A line segment $[A, B]$, traversing the width of the attractor, starts out close to the $x-y$ plane, and after the stretching $(a) \rightarrow$ (b) followed by the folding (c) $\rightarrow$ (d), the folded segment returns close to the $x-y$ plane strongly compressed. In one Poincaré return the $[A, B]$ interval is stretched, folded and


Figure 3.2: Return maps for the $R_{n} \rightarrow R_{n+1}$ radial distance Poincaré sections of figure 3.1. (Rytis Paškauskas)
mapped onto itself, so the flow is expanding. It is also mixing, as in one Poincaré return the point $C$ from the interior of the attractor is mapped into the outer edge, while the edge point $B$ lands in the interior.

Once a particular Poincaré section is picked, we can also exhibit the return map (3.1), as in figure 3.2. Cases (a) and (d) are examples of nice 1-to-1 return maps. However, (b) and (c) appear multimodal and non-invertible, artifacts of projections of a 2-dimensional return map $\left(R_{n}, z_{n}\right) \rightarrow\left(R_{n+1}, z_{n+1}\right)$ onto a 1-dimensional subspace $R_{n} \rightarrow R_{n+1}$. (continued in exercise 3.1)

The dynamics of high-dimensional flows, such as a truncation of the infinite tower of the Kuramoto-Sivashinsky modes (2.21) can be difficult to visualize. The question is: how is one to look at such a flow? It is not clear that restricting the dynamics to a Poincaré section necessarily helps - after all, a section reduces a $(d+1)$-dimensional flow to a $d$-dimensional map, and how much is gained by replacing a continuous flow in 16 dimensionas by a set of points in 15 dimensions? Nevertheless, as we now show, much can be gleaned by examining a trajectory's projection onto a subspace spanned by two or three coordinate axes.

Example 3.4 Kuramoto-Sivashinsky attractor: Consider an N-mode truncation to the infinite tower of the Kuramoto-Sivashinsky coupled Fourier modes (2.21). We fix (arbitrarily) the Poincaré section to be the hyperplane $a_{1}=0$, and integrate (2.21) with the initial conditions $a_{1}=0$, and arbitrary values of the coordinates $a_{2}, \ldots, a_{N}$ (where $N$ is the truncation order). When $a_{1}$ becomes 0 the next time and the flow crosses the $a_{1}=0$ hyperplane in the same direction as initially, the coordinates $a_{2}, \ldots, a_{N}$ are mapped into $\left(a_{2}^{\prime}, \ldots a_{N}^{\prime}\right)=P\left(a_{2}, \ldots, a_{N}\right)$, where $P$ is the Poincaré return map (3.1) of the $(N-1)$-dimensional $a_{1}=0$ hyperplane into itself. Figure 3.3 is an example of a typical result. We have picked - arbitrarily - a subspace such as $a_{6}(n+1)$ vs. $a_{6}(n)$ in order to visualize the dynamics. While the topology of the attractor is still obscure, one thing is clear: even though the flow is infinite dimensional, the attractor is finite and thin, barely thicker than a line.

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The above examples illustrate why a Poincaré section gives a more informative snapshot of the flow than the full flow portrait. For example, while the full flow portrait of the Rössler flow figure 2.3 gives us no sense of the thickness of the attractor, we see clearly in the Rössler Poincaré

Figure 3.3: The attractor of the KuramotoSivashinsky system (2.21), plotted as the $a_{6}$ component of the $a_{1}=0$ Poincaré section return map. Here 10,000 Poincaré section returns of a typical trajectory are plotted. Also indicated are the periodic points $0,1,01$ and 10. $N=16$ Fourier modes truncation with $\nu=0.029910$. (From ref. [2.10].)

sections figure 3.1 that even though the return map is $2-d \rightarrow 2-d$, the flow contraction is so strong that for all practical purposes it renders the return map 1-dimensional. Similarly, while no fine structure is discernable in the full flow portraits of the Kuramoto-Sivashinsky dynamics figure 2.5, the Poincaré return map figure 3.3 reveals the fractal structure in the asymptotic attractor.

### 3.2 Constructing a Poincaré section

- ${ }^{3}$
For almost any flow of physical interest a Poincaré section is not available in analytic form. We describe here a numerical method for determining a Poincaré section.

Consider the system (2.5) of ordinary differential equations in the vector variable $x=\left(x_{1}, x_{2}, \ldots, x_{d}\right)$

$$
\begin{equation*}
\frac{d x_{i}}{d t}=v_{i}(x, t) \tag{3.6}
\end{equation*}
$$

where the flow velocity $v$ is a vector function of the position in phase space $x$ and the time $t$. In general $v$ cannot be integrated analytically and we will have to resort to numerical integration to determine the trajectories of the system. Our task is to determine the points at which the numerically integrated trajectory traverses a given hypersurface. The hypersurface will be specified implicitly through a function $U(x)$ that is zero whenever a point $x$ is on the Poincaré section, such as the hyperplane (3.5).

If we use a tiny step size in our numerical integrator, we can observe the value of $U$ as we integrate; its sign will change as the trajectory crosses the hypersurface. The problem with this method is that we have to use a very small integration time step. In order to actually land on the Poincaré section one might try to interpolate the intersection point from the two trajectory points on either side of the hypersurface. However, there is a better way.

Let $t_{a}$ be the time just before $U$ changes sign, and $t_{b}$ the time just after it changes sign. The method for landing exactly on the Poincaré section will be to convert one of the space coordinates into an integration variable for the part of the trajectory between $t_{a}$ and $t_{b}$. Using

$$
\begin{equation*}
\frac{d x_{k}}{d x_{1}} \frac{d x_{1}}{d t}=\frac{d x_{k}}{d x_{1}} v_{1}(x, t)=v_{k}(x, t) \tag{3.7}
\end{equation*}
$$

we can rewrite the equations of motion (3.6) as

$$
\begin{equation*}
\frac{d t}{d x_{1}}=\frac{1}{v_{1}}, \cdots, \frac{d x_{d}}{d x_{1}}=\frac{v_{d}}{v_{1}} . \tag{3.8}
\end{equation*}
$$

Now we use $x_{1}$ as the "time" in the integration routine and integrate it from $x_{1}\left(t_{a}\right)$ to the value of $x_{1}$ on the hypersurface, which can be found from the hypersurface intersection condition (3.5). The quantity $x_{1}$ need not be perpendicular to the Poincaré section; any $x_{i}$ can be chosen as the integration variable, privided the $x_{i}$-axis is not parallel to the Poincaré section at the trajectory intersection point. If the section crossing is transverse (see (3.2)), $v_{1}$ cannot vanish in the short segment bracketed by the integration step preceeding the section, and the point on the Poincaré section.

Of course, once you have determined the Poincaré section point, you use it as the initial point of the next segment of the trajectory, integrated again in the time variable, not in $x_{1}$, as globally $v_{1}$ vanishes at least twice for each complete Poincaré section return.

### 3.3 Do it again

Though we have motivated discrete time dynamics by considering sections of a continuous flow, there are many settings in which dynamics is inherently discrete, and naturally described by repeated iterations of the same map

$$
f: \mathcal{M} \rightarrow \mathcal{M}
$$

or sequences of consecutive applications of a finite set of maps,

$$
\begin{equation*}
\left\{f_{A}, f_{B}, \ldots f_{Z}\right\}: \mathcal{M} \rightarrow \mathcal{M} \tag{3.9}
\end{equation*}
$$

for example maps relating different sections among a set of Poincaré sections. The discrete "time" is then an integer, the number of applications of a map. As explicitly writing out formulas involving repeated applications of a set of maps can be awkward, we streamline the notation by denoting a map composition by "o"

$$
\begin{equation*}
\left.f_{Z}\left(\cdots f_{B}\left(f_{A}(x)\right)\right) \cdots\right)=f_{Z} \circ \cdots f_{B} \circ f_{A}(x), \tag{3.10}
\end{equation*}
$$

and the $n$th iterate of map $f$ by

$$
f^{n}(x)=f \circ f^{n-1}(x)=f\left(f^{n-1}(x)\right), \quad f^{0}(x)=x
$$

The trajectory of $x$ is the set of points

$$
\left\{x, f(x), f^{2}(x), \ldots, f^{n}(x)\right\}
$$

and the orbit of $x$ is the subset of all points of $\mathcal{M}$ that can be reached by iterations of $f$.

The functional form of such Poincaré return maps $P$ as figures 3.2 and 3.3 can be approximated by tabulating the results of integration of the flow from $x$ to the first Poincaré section return for many $x \in \mathcal{P}$, and constructing a function that interpolates through these points. If we find a good approximation to $P(x)$, we can get rid of numerical integration altogether, by replacing the continuous time trajectory $f^{t}(x)$ by iteration of the Poincaré return map $P(x)$. Constructing accurate $P(x)$ for a given flow can be tricky, but we can already learn much from approximate Poincaré return maps. Multinomial approximations

$$
\begin{equation*}
P_{k}(x)=a_{k}+\sum_{j=1}^{d+1} b_{k j} x_{j}+\sum_{i, j=1}^{d+1} c_{k i j} x_{i} x_{j}+\ldots, \quad x \in \mathcal{P} \tag{3.11}
\end{equation*}
$$

to Poincaré return maps

$$
\left(\begin{array}{c}
x_{1, n+1} \\
x_{2, n+1} \\
\ldots \\
x_{d, n+1}
\end{array}\right)=\left(\begin{array}{c}
P_{1}\left(x_{n}\right) \\
P_{2}\left(x_{n}\right) \\
\ldots \\
P_{d}\left(x_{n}\right)
\end{array}\right), \quad x_{n}, x_{n+1} \in \mathcal{P}
$$

motivate the study of model mappings of the plane, such as the Hénon map.

Example 3.5 Hénon map: The map

$$
\begin{align*}
x_{n+1} & =1-a x_{n}^{2}+b y_{n} \\
y_{n+1} & =x_{n} \tag{3.12}
\end{align*}
$$

is a nonlinear 2-dimensional map most frequently employed in testing various hunches about chaotic dynamics. The Hénon map is sometimes written as a 2 -step recurrence relation

$$
\begin{equation*}
x_{n+1}=1-a x_{n}^{2}+b x_{n-1} . \tag{3.13}
\end{equation*}
$$

An $n$-step recurrence relation is the discrete-time analogue of an nth order differential equation, and it can always be replaced by a set of $n 1$-step recurrence relations.

Figure 3.4: The strange attractor and an unstable period 7 cycle of the Hénon map (3.12) with $a=1.4, b=0.3$. The periodic points
 in the cycle are connected to guide the eye. (K.T. Hansen [1.3])

The Hénon map is the simplest map that captures the "stretch \& fold" dynamics of return maps such as Rössler's, figure 3.1. It can be obtained by a truncation of a polynomial approximation (3.11) to a Poincaré return map to second order.

A quick sketch of the long-time dynamics of such a mapping (an example is depicted in figure 3.4), is obtained by picking an arbitrary starting point and iterating (3.12) on a computer. We plot here the dynamics in the $\left(x_{n}, x_{n+1}\right)$ plane, rather than in the $\left(x_{n}, y_{n}\right)$ plane, because we think of the Hénon map as a model return map $x_{n} \rightarrow x_{n+1}$. As we shall soon see, periodic orbits will be key to understanding the long-time dynamics, so we also plot a typical periodic orbit of such a system, in this case an unstable period 7 cycle. Numerical determination of such cycles will be explained in sect. 31.1, and the cycle point labels 0111010, 1110100, $\cdots$ in sect. 12.2.

Example 3.6 Lozi map: Another example frequently employed is the Lozi map, a linear, "tent map" version of the Hénon map given by

$$
\begin{align*}
x_{n+1} & =1-a\left|x_{n}\right|+b y_{n} \\
y_{n+1} & =x_{n} . \tag{3.14}
\end{align*}
$$

Though not realistic as an approximation to a smooth flow, the Lozi map is a very helpful tool for developing intuition about the topology of a large class of maps of the "stretch \& fold" type.

What we get by iterating such maps is - at least qualitatively - not unlike what we get from Poincaré section of flows (see figures 3.2 and 3.3). For an arbitrary initial point this process might converge to a stable limit cycle, to a strange attractor, to a false attractor (due to roundoff errors), or diverge. In other words, mindless iteration is essentially uncontrollable, and we will need to resort to more thoughtful explorations. As we shall explain in due course below, strategies for systematic exploration rely on stable/unstable manifolds, periodic points, saddle-straddle methods and so on.

Example 3.7 Parabola: One iteration of the Hénon map stretches and folds a region of the $(x, y)$ plane centered around the origin. The parameter a controls the amount of stretching, while the parameter $b$ controls the thickness of the folded image through the "1-step memory" term $b x_{n-1}$ in (3.13). In figure 3.4 the parameter $b$ is rather large, $b=0.3$, so the attractor is rather thick, with the transverse fractal structure
clearly visible. For vanishingly small $b$ the Hénon map reduces to the 1-dimensional logisticm

$$
\begin{equation*}
x_{n+1}=1-a x_{n}^{2} \tag{3.15}
\end{equation*}
$$

By setting $b=0$ we lose determinism ("surjectivity", if mathematise is your native 2 dialect), as on reals the inverse of map (3.15) has two preimages $\left\{x_{n-1}^{+}, x_{n-1}^{-}\right\}^{-}$for most $x_{n}$. Still, this 1-dimensional approximation is very instructive.

As we shall see in sect. 11.3.1, an understanding of 1-dimensional dynamics is indeed the essential prerequisite to unravelling the qualitative dynamics of many higher-dimensional dynamical systems. For this reason many expositions of the theory of dynamical systems commence with a study of 1-dimensional maps. We prefer to stick to flows, as that is where the physics is.

## Commentary

Remark 3.1 Determining a Poincaré section. The idea of changing the integration variable from time to one of the coordinates, although simple, avoids the alternative of having to interpolate the numerical solution to determine the intersection. The trick described in sect. 3.2 is due to Hénon [3.4, 11.13, 1.8].

Remark 3.2 Hénon, Lozi maps. The Hénon map is of no particular physical import in and of itself - its significace lies in the fact that it is a minimal normal form for modeling flows near a saddle-node bifurcation, and that it is a prototype of the stretching and folding dynamics that leads to deterministic chaos. It is generic in the sense that it can exhibit arbitrarily complicated symbolic dynamics and mixtures of hyperbolic and non-hyperbolic behaviors. Its construction was motivated by the best known early example of "deterministic chaos", the Lorenz equation [2.1]. Y. Pomeau's studies of the Lorenz attractor on an analog computer, and his insights into its stretching and folding dynamics motivated Hénon [3.1] to introduce the Hénon map in 1976. Hénon's and Lorenz's original papers can be found in reprint collections refs. [3.2, 3.3]. They are a pleasure to read, and are still the best introduction to the physics motivating such models. A detailed description of the dynamics of the Hénon map is given by Mira and coworkers [3.5], as well as very many other authors.

The Lozi map [3.7] is particularly convenient in investigating the symbolic dynamics of 2- $d$ mappings. Both the Lorenz and Lozi systems are uniformly smooth systems with singularities. The continuity of measure for the Lozi map was proven by M. Misiurewicz [3.8], and the existence of the SRB measure was established by L.-S. Young.

[^1]
## Résumé

In recurrent dynamics a trajectory exits a region in phase space and then reenters it infinitely often, with a mean return time. If the orbit is periodic,
it returns after a full period. So, on average, nothing much really happens along the trajectory - what is important is behavior of neighboring trajectories transverse to the flow. This observation motivates a replacament of the continuous time flow by iterative mapping, the Poincaré return map.

The visualization of strange attractors is greatly facilitated by a felicitous choice of Poincaré sections, and the reduction of flows to Poincaré return maps. This observation motivates in turn the study of discrete-time dynamical systems generated by iterations of maps.

A particularly natural application of the Poincaré section method is the reduction of a billiard flow to a boundary-to-boundary return map, described in chapter 6 below. As we shall show in chapter 7, further simplification of a Poincaré return map, or any nonlinear map, can be attained through rectifying these maps locally by means of smooth conjugacies.

## References

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## Exercises

Exercise 3.1 Poincaré sections of the Rössler flow. (continuation of exercise 2.8) Calculate numerically a Poincaré section (or several Poincaré sections) of the Rössler flow. As the Rössler flow phase space is 3-dimensional, the flow maps onto a 2-dimensional Poincaré section. Do you see that in your numerical results? How good an approximation would a replacement of the return map for this section by a 1-dimensional map be? More precisely, estimate the thickness of the strange attractor. (continued as exercise 4.3)

## (Rytis Paškauskas)

Exercise 3.2 Arbitrary Poincaré sections. We will generalize the construction of Poincaré sections so that they can have any shape, as specified by the equation $U(x)=0$.
(a) Start by modifying your integrator so that you can change the coordinates once you get near the Poincaré section. You can do this easily by writing the equations as

$$
\begin{equation*}
\frac{d x_{k}}{d s}=\kappa f_{k} \tag{3.16}
\end{equation*}
$$

with $d t / d s=\kappa$, and choosing $\kappa$ to be 1 or $1 / f_{1}$. This allows one to switch between $t$ and $x_{1}$ as the integration "time."
(b) Introduce an extra dimension $x_{n+1}$ into your system and set

$$
\begin{equation*}
x_{n+1}=U(x) . \tag{3.17}
\end{equation*}
$$

How can this be used to find a Poincaré section?

Exercise 3.3 Classical collinear helium dynamics. (continuation of exercise 2.10)

Make a Poincaré surface of section by plotting $\left(r_{1}, p_{1}\right)$ whenever $r_{2}=0$ : Note that for $r_{2}=0, p_{2}$ is already determined by (5.6). Compare your results with figure 34.3(b).
(Gregor Tanner, Per Rosenqvist)

Exercise 3.4 Hénon map fixed points. Show that the two fixed points $\left(x_{0}, x_{0}\right),\left(x_{1}, x_{1}\right)$ of the Hénon map (3.12) are given by

$$
\begin{align*}
& x_{0}=\frac{-(1-b)-\sqrt{(1-b)^{2}+4 a}}{2 a}, \\
& x_{1}=\frac{-(1-b)+\sqrt{(1-b)^{2}+4 a}}{2 a} . \tag{3.18}
\end{align*}
$$

## Exercise 3.5 How strange is the Hénon attractor?

(a) Iterate numerically some 100,000 times or so the Hénon map

$$
\left[\begin{array}{l}
x^{\prime} \\
y^{\prime}
\end{array}\right]=\left[\begin{array}{l}
1-a x^{2}+y \\
b x
\end{array}\right]
$$

for $a=1.4, b=0.3$. Would you describe the result as a "strange attractor"? Why?
(b) Now check how robust the Hénon attractor is by iterating a slightly different Hénon map, with $a=1.39945219, b=0.3$. Keep at it until the "strange" attracttor vanishes like the smile of the Chesire cat. What replaces it? Would you describe the result as a "strange attractor"? Do you still have confidence in your own claim for the part (a) of this exercise?

Exercise 3.6 Fixed points of maps. A continuous function $F$ is a contraction of the unit interval if it maps the interval inside itself.
(a) Use the continuity of $F$ to show that a one-dimensional contraction $F$ of the interval $[0,1]$ has at least one fixed point.
(b) In a uniform (hyperbolic) contraction the slope of $F$ is always smaller than one, $\left|F^{\prime}\right|<1$. Is the composition of uniform contractions a contraction? Is it uniform?

## Chapter 4

## Local stability

(R. Mainieri and P. Cvitanović)

So far we have concentrated on description of the trajectory of a single initial point. Our next task is to define and determine the size of a neighborhood of $x(t)$. We shall do this by assuming that the flow is locally smooth, and describe the local geometry of the neighborood by studying the flow linearized around $x(t)$. Nearby points aligned along the stable (contracting) directions remain in the neighborhood of the trajectory $x(t)=f^{t}\left(x_{0}\right)$; the ones to keep an eye on are the points which leave the neighborhood along the unstable directions. As we shall demonstrate in chapter 14, in hyperbolic systems what matters are the expanding directions. The reprecussion are far-reaching: As long as the number of unstable directions is finite, the same theory applies to finite-dimensional ODEs, phase-space volume preserving Hamiltonian flows, and dissipative, volume contracting infinite-dimensional PDEs.

### 4.1 Flows transport neighborhoods

As a swarm of representative points moves along, it carries along and distorts neighborhoods, as sketched in figure 2.1(b). The deformation of an infinitesimal neighborhood is best understood by considering a trajectory originating near $x_{0}=x(0)$ with an initial infinitesimal displacement $\delta x(0)$, and letting the flow transport the displacement $\delta x(t)$ along the trajectory $x\left(x_{0}, t\right)=f^{t}\left(x_{0}\right)$. The system of linear equations of variations for the displacement of the infinitesimally close neighbor $x+\delta x$ follows from the flow equations (2.5) by Taylor expanding to linear order

$$
\dot{x}_{i}+\dot{\delta} x_{i}=v_{i}(x+\delta x) \approx v_{i}(x)+\sum_{j} \frac{\partial v_{i}}{\partial x_{j}} \delta x_{j} .
$$

The infinitesimal displacement $\delta x$ is thus transported along the trajectory $x\left(x_{0}, t\right)$, with time variation given by

$$
\begin{equation*}
\frac{d}{d t} \delta x_{i}\left(x_{0}, t\right)=\left.\sum_{j} \frac{\partial v_{i}(x)}{\partial x_{j}}\right|_{x=x\left(x_{0}, t\right)} \delta x_{j}\left(x_{0}, t\right) \tag{4.1}
\end{equation*}
$$

As both the displacement and the trajectory depend on the initial point $x_{0}$ and the time $t$, we shall often abbreviate the notation to $x\left(x_{0}, t\right) \rightarrow x(t) \rightarrow$ $x, \delta x_{i}\left(x_{0}, t\right) \rightarrow \delta x_{i}(t) \rightarrow \delta x$ in what follows. Taken together, the set of equations

$$
\begin{equation*}
\dot{x}_{i}=v_{i}(x), \quad \dot{\delta} x_{i}=\sum_{j} A_{i j}(x) \delta x_{j} \tag{4.2}
\end{equation*}
$$

governs the dynamics in the tangent bundle $(x, \delta x) \in \mathbf{T} \mathcal{M}$ obtained by adjoining the $d$-dimensional tangent space $\delta x \in \mathbf{T}_{x} \mathcal{M}$ to every point $x \in \mathcal{M}$ in the $d$-dimensional phase space $\mathcal{M} \subset \mathbb{R}^{d}$. The matrix of variations

$$
\begin{equation*}
A_{i j}(x)=\frac{\partial v_{i}(x)}{\partial x_{j}} \tag{4.3}
\end{equation*}
$$

describes the instantaneous rate of shearing of the infinitesimal neighborhood of $x(t)$ by the flow.

Taylor expanding a finite time flow to linear order,

$$
\begin{equation*}
f_{i}^{t}\left(x_{0}+\delta x\right)=f_{i}^{t}\left(x_{0}\right)+\sum_{j} \frac{\partial f_{i}^{t}\left(x_{0}\right)}{\partial x_{0 j}} \delta x_{j}+\cdots \tag{4.4}
\end{equation*}
$$

one finds that the linearized neighborhood is transported by the Jacobian (or fundamental) matrix

$$
\begin{equation*}
\delta x(t)=\mathbf{J}^{t}\left(x_{0}\right) \delta x(0), \quad \mathbf{J}_{i j}^{t}\left(x_{0}\right)=\left.\frac{\partial x_{i}(t)}{\partial x_{j}}\right|_{x=x_{0}} \tag{4.5}
\end{equation*}
$$

which describes the deformation of an infinitesimal neighborhood at finite time $t$ in the co-moving frame of $x(t)$, that is, transformation of the initial, Lagrangian coordinate frame into the current, Eulerian coordinate frame.

As this is a deformation in the linear approximation, one can think of it as a linear deformation of an infinitesimal sphere enveloping $x_{0}$ into an ellipsoid around $x(t)$, described by the eigenvectors and eigenvalues of the Jacobian matrix of the linearized flow, figure 4.1 Nearby trajectories separate along the unstable directions, approach each other along the stable directions, and change their distance along the marginal directions at a

Figure 4.1: Jacobian matrix maps an infinitesimal spherical neighborhood of $x_{0}$ into an ellipsoidal neighborhood time $t$ later.

rate slower than exponential. In the literature adjectives neutral or indifferent are often used instead of "marginal", (attracting) stable directions are sometimes called "asymptotically stable", the (neutrally) stable directions "stable", and so on: but all one is saying is that the eigenvalues of the Jacobian matrix have maginitude smaller, equal, or larger than 1.

One of the eigendirections is what one might expect, the direction of the flow itself. To see that, consider two points along a trajectory separated by infinitesimal flight time $\delta t: \delta x(0)=f^{\delta t}\left(x_{0}\right)-x_{0}=v\left(x_{0}\right) \delta t$. At time $t$,

$$
\delta x(t)=f^{t+\delta t}\left(x_{0}\right)-f^{t}\left(x_{0}\right)=f^{\delta t}(x(t))-x(t)=v(x(t)) \delta t,
$$

while from (4.5) $\delta x(t)=\mathbf{J}^{t}\left(x_{0}\right) \delta x(0)$. Dividing both equations by $\delta t$ we observe that $\mathbf{J}^{t}\left(x_{0}\right)$ transports the velocity vector at $x_{0}$ to the velocity vector at $x(t)$ at time $t$ :

$$
\begin{equation*}
v(x(t))=\mathbf{J}^{t}\left(x_{0}\right) v\left(x_{0}\right) . \tag{4.6}
\end{equation*}
$$

As noted above, the Jacobian matrix (4.5) maps the initial, Lagrangian coordinate frame into the current, Eulerian coordinate frame.

The velocity at point $x(t)$ in general does not point in the same direction as the velocity at point $x_{0}$, so this is not an eigenvalue condition for $\mathbf{J}^{t}$; the Jacobian matrix computed for an arbitrary segment of an arbitrary trajectory has no invariant meaning.

As the eigenvalues of finite time $\mathbf{J}^{t}$ have invariant meaning only for periodic orbits, we postpone their interpretation to chapter 8. However, already at this stage we see that if the orbit is periodic, $x\left(T_{p}\right)=x(0)$, at any point along cycle $p$ the velocity $v$ is an eigenvector of the Jacobian matrix $\mathbf{J}_{p}=\mathbf{J}^{T_{p}}$ with a unit eigenvalue,

$$
\begin{equation*}
\mathbf{J}^{T_{p}}(x) v(x)=v(x), \quad x \in p . \tag{4.7}
\end{equation*}
$$

Two successive points along the cycle separated by $\delta x(0)$ have exactly the same separation after a completed period $\delta x\left(T_{p}\right)=\delta x(0)$, hence eigenvalue 1.

As we started by assuming that we know the equations of motion, from (4.3) we also know $\mathbf{A}$, the instantaneous rate of shear of an infinitesimal neighborhood $\delta x_{i}(t)$ of the trajectory $x(t)$. What we do not know is the finite time deformation matrix $\mathbf{J}^{t}$. Our next task is to relate the instantaneous shear rate $\mathbf{A}$ to $\mathbf{J}^{t}$.

We are interested in smooth, differentiable flows. If a flow is smooth, in a sufficiently small neighborhood it is essentially linear. Hence the next section, which might seem an embarrassment (what is a section on linear flows doing in a book on nonlinear dynamics?), offers a firm stepping stone on the way to understanding nonlinear flows. If you know your eigenvalues and eigenvectors, you may prefer to fast forward here.


### 4.2 Linear flows

Linear fields are the simplest vector fields. Described by linear differential equations which can be solved explicitly, with solutions that are good for all times. The phase space for linear differential equations is $\mathcal{M}=\mathbb{R}^{d}$, and the equations of motion (2.5) are written in terms of a vector $x$ and a constant matrix of variations $\mathbf{A}$ as

$$
\begin{equation*}
\dot{x}=v(x)=\mathbf{A} x . \tag{4.8}
\end{equation*}
$$

Solving this equation means finding the phase space trajectory

$$
x(t)=\left(x_{1}(t), x_{2}(t), \ldots, x_{d}(t)\right)
$$

passing through the point $x_{0}$.
If $x(t)$ is a solution with $x(0)=x_{0}$ and $y(t)$ another solution with $y(0)=y_{0}$, then the linear combination $a x(t)+b y(t)$ with $a, b \in \mathbb{R}$ is also a solution, but now starting at the point $a x_{0}+b y_{0}$. At any instant in time, the space of solutions is a $d$-dimensional vector space, which means that one can find a basis of $d$ linearly independent solutions. How do we solve the linear differential equation (4.8)? If instead of a matrix equation we have a scalar one, $\dot{x}=a x$, with $a$ a real number, then the solution is

$$
\begin{equation*}
x(t)=e^{t a} x(0) \tag{4.9}
\end{equation*}
$$

In order to solve the matrix case, it is helpful to rederive the solution (4.9) by studying what happens for a short time step $\delta t$. If at time $t=0$ the position is $x(0)$, then

$$
\begin{equation*}
\frac{x(\delta t)-x(0)}{\delta t}=a x(0) \tag{4.10}
\end{equation*}
$$

which we iterate $m$ times to obtain the Euler's formula for compounding interest

$$
\begin{equation*}
x(t) \approx\left(1+\frac{t}{m} a\right)^{m} x(0) \tag{4.11}
\end{equation*}
$$

The term in parentheses acts on the initial condition $x(0)$ and evolves it to $x(t)$ by taking $m$ small time steps $\delta t=t / m$. As $m \rightarrow \infty$, the term in parentheses converges to $e^{t a}$. Consider now the matrix version of equation (4.10):

$$
\begin{equation*}
\frac{x(\delta t)-x(0)}{\delta t}=\mathbf{A} x(0) \tag{4.12}
\end{equation*}
$$

A representative point $x$ is now a vector in $\mathbb{R}^{d}$ acted on by the matrix $\mathbf{A}$, as in (4.8). Denoting by $\mathbf{1}$ the identity matrix, and repeating the steps (4.10) and (4.11) we obtain Euler's formula for the exponential of a matrix

$$
\begin{equation*}
x(t)=\lim _{m \rightarrow \infty}\left(\mathbf{1}+\frac{t}{m} \mathbf{A}\right)^{m} x(0)=e^{t \mathbf{A}} x(0) . \tag{4.13}
\end{equation*}
$$

We will use this expression as the definition of the exponential of a matrix.


How do we compute the exponential (4.13)?

Example 4.1 Stability eigenvalues, diagonal case: Should we be so lucky that A happens to be a diagonal matrix $\mathbf{A}_{D}$ with eigenvalues real $\left(\lambda_{1}, \lambda_{2}, \ldots, \lambda_{d}\right)$, the exponential is simply

$$
\mathbf{J}^{t}=e^{t \mathbf{A}_{D}}=\left(\begin{array}{ccc}
e^{t \lambda_{1}} & \cdots & 0  \tag{4.14}\\
& \ddots & \\
0 & \cdots & e^{t \lambda_{d}}
\end{array}\right) .
$$

Throughout this text the symbol $\Lambda_{k}$ will always denote the $k$ th stability eigenvalue (in literature sometimes referred to as the stability multiplier) of the finite time Jacobian matrix $\mathbf{J}^{t}$. Symbol $\lambda_{k}$ will be reserved for the $k$ th stability exponent (in cases of interest, often the same thing as the Lyapunov exponent)

$$
\begin{equation*}
\left|\Lambda_{k}\right|=e^{t \lambda_{k}}, \quad \Lambda_{k}=\Lambda_{k}\left(x_{0}, t\right), \lambda_{k}=\lambda_{k}\left(x_{0}, t\right) . \tag{4.15}
\end{equation*}
$$

Example 4.2 Stability eigenvalues, diagonalizable case: Suppose next that $\mathbf{A}$ is diagonalizable and that $\mathbf{U}$ is the matrix that brings it to its diagonal form $\mathbf{A}_{D}=$ $\mathbf{U}^{-1} \mathbf{A U}$. The transformation $\mathbf{U}$ is a linear coordinate transformation which rotates, skews, and possibly flips the coordinate axis of the vector space. Then J can also be brought to a diagonal form by inserting factors $\mathbf{1}=\mathbf{U U}^{-1}$ between the steps of the product (4.13):

$$
\mathbf{J}^{t}=e^{t \mathbf{A}}=\mathbf{U} e^{t \mathbf{A}_{D}} \mathbf{U}^{-1}
$$

In either example, the action of both $\mathbf{A}$ and $\mathbf{J}$ is very simple; the axes of orthogonal coordinate system where $\mathbf{A}$ is diagonal are also the eigendirections of both $\mathbf{A}$ and $\mathbf{J}^{t}$, and under the flow the neighborhood is deformed by a multiplication by an eigenvalue factor for each coordinate axis.

As $\mathbf{A}$ has only real entries, it will in general have either real eigenvalues, or complex conjugate pairs of eigenvalues. That is not surprising, but also the corresponding eigenvectors can be either real or complex. All coordinates used in defining the flow are real numbers, so what is the meaning of a complex eigenvector?

Example 4.3 Complex stability eigenvalues: To develop some intuition about that, let us work out the behavior for the simplest nontrivial case, the case where

$$
\mathbf{A}=\left(\begin{array}{ll}
A_{11} & A_{12}  \tag{4.17}\\
A_{21} & A_{22}
\end{array}\right)
$$

The eigenvalues $\lambda_{1}, \lambda_{2}$ of $\mathbf{A}$ are the roots

$$
\begin{equation*}
\lambda_{1,2}=\frac{1}{2}\left(\operatorname{tr} \mathbf{A} \pm \sqrt{(\operatorname{tr} \mathbf{A})^{2}-4 \operatorname{det} \mathbf{A}}\right) \tag{4.18}
\end{equation*}
$$

of the characteristic equation

$$
\begin{align*}
\operatorname{det}(\mathbf{A}-z \mathbf{1}) & =\left(\lambda_{1}-z\right)\left(\lambda_{2}-z\right)=0  \tag{4.19}\\
\left\|\begin{array}{cc}
A_{11}-z & A_{12} \\
A_{21} & A_{22}-z
\end{array}\right\| & =z^{2}-\left(A_{11}+A_{22}\right) z+\left(A_{11} A_{22}-A_{12} A_{21}\right)
\end{align*}
$$

The qualitative behavior of $e^{\mathbf{A}}$ for real eigenvalues $\lambda_{1}, \lambda_{2} \in \mathbb{R}$ will differ from the case that they form a complex conjugate pair,

$$
\lambda_{1}=\lambda+i \theta, \quad \lambda_{2}=\lambda_{1}^{*}=\lambda-i \theta
$$

These two possibilities are refined further into sub-cases depending on the signs of the real part. The matrix might have only one eigenvector, or two linearly independent eigenvectors, which may or may not be orthogonal. Along each of these directions the motion is of the form $x_{k} \exp \left(t \lambda_{k}\right)$. If the exponent $\lambda_{k}$ is positive, then the component $x_{k}$ will grow; if the exponent $\lambda_{k}$ is negative, it will shrink.

We sketch the full set of possibilities in figure $4.2(\mathrm{a})$, and work out in detail the case when $\mathbf{A}$ can be brought to the diagonal form, and the case of degenerate eigenvalues.

Example 4.4 Complex stability eigenvalues, diagonal: If A can be brought to the diagonal form, the solution (4.13) to the differential equation (4.8) can be written either as

$$
\binom{x_{1}(t)}{x_{2}(t)}=\left(\begin{array}{cc}
e^{t \lambda_{1}} & 0  \tag{4.20}\\
0 & e^{t \lambda_{2}}
\end{array}\right)\binom{x_{1}(0)}{x_{2}(0)}
$$

or

$$
\binom{x_{1}(t)}{x_{2}(t)}=e^{t \lambda}\left(\begin{array}{cc}
e^{i t \theta} & 0  \tag{4.21}\\
0 & e^{-i t \theta}
\end{array}\right)\binom{x_{1}(0)}{x_{2}(0)} .
$$

In the case $\lambda_{1}>0, \lambda_{2}<0, x_{1}$ grows exponentially with time, and $x_{2}$ contracts exponentially. This behavior, called a saddle, is sketched in figure 4.2(b), as are the remaining possibilities: in/out nodes, inward/outward spirals, and the center. Spirals arise from taking a real part of the action of $\mathbf{J}^{t}$ on a complex eigenvector. The magnitude of $|x(t)|$ diverges exponentially when $\lambda>0$, and contracts toward 0 when the $\lambda<0$, whereas the imaginary phase $\theta$ controls its oscillations.

In general $\mathbf{J}^{t}$ is neither diagonal, nor diagonalizable, nor constant along the trajectory. Still, any matrix, including $\mathbf{J}^{t}$, can be expressed in the singular value decomposition form

$$
\mathbf{J}=\mathbf{U D V}^{T}
$$

where $\mathbf{D}$ is diagonal, and $\mathbf{U}, \mathbf{V}$ are orthogonal matrices. The diagonal elements $\Lambda_{1}, \Lambda_{2}, \ldots, \Lambda_{d}$ of $\mathbf{D}$ are called the stability eigenvalues.

Under the action of the flow an infinitesimally small ball of initial points is deformed into an ellipsoid: $\Lambda_{i}$ is the relative stretching of the $i$ th principal axis of the ellipsoid, the columns of the matrix $\mathbf{V}$ are the principal axes $\mathbf{e}_{i}$ of stretching in the Lagrangian coordinate frame, and the orthogonal matrix $\mathbf{U}$ gives the orientation of the ellipse in the Eulerian coordinates.

Now that we have some feeling for the qualitative behavior of eigenvectors and eigenvalues, we are ready to return to the general case: nonlinear flows.

### 4.3 Stability of flows

How do you determine the eigenvalues of the finite time local deformation $\mathbf{J}^{t}$ for a general nonlinear smooth flow? The Jacobian matrix is computed by integrating the equations of variations (4.2)

$$
\begin{equation*}
x(t)=f^{t}\left(x_{0}\right), \quad \delta x\left(x_{0}, t\right)=A^{t}\left(x_{0}\right) \delta x\left(x_{0}, 0\right) \tag{4.22}
\end{equation*}
$$

The equations of variations are linear, so we should be able to integrate them - but in order to make sense of the answer, we derive it step by step.
(a)

(b)


Figure 4.2: (a) Qualitatively distinct types of stability exponents of a [ $2 \times 2$ ] stability matrix. (b) Streamlines for several typical 2-dimensional flows: saddle (hyperbolic), in node (attracting), center (elliptic), in spiral.

### 4.3.1 Stability of equilibria

For a start, consider the case where $x_{q}$ is an equilibrium point (2.7). Expanding around the equilibrium point $x_{q}$, using the fact that the matrix $\mathbf{A}=\mathbf{A}\left(x_{q}\right)$ in (4.2) is constant, and integrating,

$$
\begin{equation*}
f^{t}(x)=x_{q}+e^{\mathbf{A} t}\left(x-x_{q}\right)+\cdots \tag{4.23}
\end{equation*}
$$

we verify that the simple formula (4.13) applies also to the Jacobian matrix of an equilibrium point,

$$
\begin{equation*}
\mathbf{J}^{t}\left(x_{q}\right)=e^{\mathbf{A} t} \quad \mathbf{A}=\mathbf{A}\left(x_{q}\right) . \tag{4.24}
\end{equation*}
$$

Example 4.5 Equilibria of the Rössler flow. The Rösler system (2.14) has two equilibrium points

$$
\begin{aligned}
& \left(x^{-}, y^{-}, z^{-}\right)=(0.0070,-0.0351,0.0351) \\
& \left(x^{+}, y^{+}, z^{+}\right)=(5.6929,-28.4648,28.4648)
\end{aligned}
$$

The two equilibria together their stability exponents now yield quite detailed information about the flow. Figure 4.3 shows two trajectories which start in the neighborhood of the " + " equilibrium point. Trajectories to the right of the " + " equilibrium point escape, and those to the left spiral toward the "-" equilibrium point, where they seem to wander chaoticaly for all times. The stable manifold of " + " equilibrium point thus serves as a attraction basin boundary. Consider now the linearized stability exponents of the two equilibria

$$
\left.\begin{array}{rlr}
\left(\lambda_{1}^{-}, \lambda_{2}^{-} \pm i \theta_{2}^{-}\right) & =\left(\begin{array}{rr}
-5.686, & 0.0970 \pm i 0.9951
\end{array}\right) \\
\left(\lambda_{1}^{+}, \lambda_{2}^{+} \pm i \theta_{2}^{+}\right)= & (0.1929, & -4.596 \times 10^{-6} \pm i 5.428 \tag{4.25}
\end{array}\right)
$$

The $\lambda_{2}^{+} \pm i \theta_{2}^{+}$complex eigenvalue pair implies that that neighborhood of the outer equilibrium point rotates with angular period $T_{+} \approx\left|2 \pi / \theta_{2}^{+}\right|=1.1575$. The stability multiplier by which a trajectory that starts near the " + " equilibrium point contracts in the stable manifold plane is the excrutiatingly slow $\Lambda_{2}^{+} \approx \exp \left(\lambda_{2}^{+} T_{+}\right)=0.9999947$ per rotation. For each period the point of the stable manifold moves away along the unstable

Figure 4.3: Two trajectories of the Rössler flow initiated in the neighborhood of the
 " + " equilibrium point (2.15). (R. Paškauskas)
eigendirection by factor $\Lambda_{1}^{+} \approx \exp \left(\lambda_{1}^{+} T_{+}\right)=1.2497$. Hence the slow spiraling on both sides of the " + " equilibrium point.

The $\lambda_{2}^{-} \pm i \theta_{2}^{-}$complex eigenvalue pair tells us that neighborhood of the "-" equilibrium point rotates with angular period $T_{-} \approx\left|2 \pi / \theta_{2}^{-}\right|=6.313$, slightly faster than the harmonic oscillator estimate in (2.11). The stability multiplier by which a trajectory that starts near the "-" equilibrium point spirals away per one rotation is $\Lambda_{\text {radial }} \approx \exp \left(\lambda_{2}^{-} T_{-}\right)=1.84$. The $\lambda^{-}$eigenvalue is essentially the $z$ expansion correcting parameter c introduced in (2.13). For each Poincaré section return, the trajectory is contracted into the stable manifold by the amazing factor of $\Lambda_{1} \approx \exp \left(\lambda_{1}^{-} T_{-}\right)=10^{-15.6}(!)$.

Suppose you start with a 1 mm interval pointing in the $\Lambda_{1}$ eigendirection. After one Poincaré return the interval is of order of $10^{-4}$ fermi, the furthest we will get into subnuclear structure in this book. Of course, from the mathematical point of view, the flow is reversible, and the Poincaré return map is invertible. (Rytis Paškauskas)

## Example 4.6 Why does a flame front flutter? The Kuramoto-

Sivashinsky flat flame front $u(x, t)=0$ is an equilibrium point of (2.18). The matrix of variations (4.3) follows from (2.21)

$$
\begin{equation*}
A_{k j}(a)=\frac{\partial v_{k}(x)}{\partial a_{j}}=\left(k^{2}-\nu k^{4}\right) \delta_{k j}-2 k a_{k-j} . \tag{4.26}
\end{equation*}
$$

For the $u(x, t)=0$ equilibrium solution the matrix of variations is diagonal, and - as in (4.14) - so is the Jacobian matrix $\mathbf{J}_{k j}^{t}(0)=\delta_{k j} e^{\left(k^{2}-\nu k^{4}\right) t}$.

For $\nu>1, u(x, t)=0$ is the globally attractive stable equilibrium. As the "viscosity" decreases below $\nu \leq 1$, the dynamics goes through a rich sequence of bifurcations on which we shall not dwell here. The $|k|<1 / \sqrt{\nu}$ long wavelength perturbations of the flat-front equilibrium are linearly unstable, while all $|k|>1 / \sqrt{\nu}$ short wavelength perturbations are strongly contractive. The high $k$ eigenvalues, corresponding to rapid variations of the flame front, decay so fast that the corresponding eigendirections are physically irrelevant. To illustrate the rapid contraction in the non-leading eigendirections we plot in figure 4.4 the eigenvalues of the equilibrium in the unstable regime, for relatively low viscosity $\nu$, and compare them with the stability eigenvalues of the least unstable cycle for the same value of $\nu$. The equilibrium solution is very unstable, in

Figure 4.4: Lyapunov exponents $\lambda_{\overline{1}, k}$ versus $k$ for the least unstable spatio-temporally periodic orbit $\overline{1}$ of the Kuramoto-Sivashinsky system, compared with the stability exponents of the $u(x, t)=0$ stationary solution, $\lambda_{k}=k^{2}-\nu k^{4}$. The eigenvalue $\lambda_{\overline{1}, k}$ for $k \geq 8$ falls below the numerical accuracy of integration and are not meaningful. $N=16$ Fourier mode truncation, $\nu=0.029924$. The cycle $\overline{1}$ was computed using methods of chapter 17. (From ref. [2.10])


> 5 eigendirections, the least unstable cycle only in one. Note that for $k>7$ the rate of contraction is so strong that higher eigendirections are numerically meaningless for either solution; even though the flow is infinite-dimensional, the attracting set must be rather thin.

### 4.3.2 Stability of trajectories

Next, consider the case of a general, non-stationary trajectory $x(t)$. The exponential of a constant matrix can be defined either by its Taylor series expansion, or in terms of the Euler limit (4.13):

$$
\begin{align*}
e^{t \mathbf{A}} & =\sum_{k=0}^{\infty} \frac{t^{k}}{k!} \mathbf{A}^{k}  \tag{4.27}\\
& =\lim _{m \rightarrow \infty}\left(\mathbf{1}+\frac{t}{m} \mathbf{A}\right)^{m} \tag{4.28}
\end{align*}
$$

Taylor expanding is fine if $\mathbf{A}$ is a constant matrix. However, only the second, tax-accountant's discrete step definition of an exponential is appropriate for the task at hand, as for a dynamical system the local rate of neighborhood distortion $\mathbf{A}(x)$ depends on where we are along the trajectory. The linearized neighborhood is multiplicatively deformed along the flow, and the $m$ discrete time step approximation to $\mathbf{J}^{t}$ is therefore given by a generalization of the Euler product (4.28):

$$
\begin{align*}
\mathbf{J}^{t} & =\lim _{m \rightarrow \infty} \prod_{n=m}^{1}\left(\mathbf{1}+\delta t \mathbf{A}\left(x_{n}\right)\right)=\lim _{m \rightarrow \infty} \prod_{n=m}^{1} e^{\delta t \mathbf{A}\left(x_{n}\right)}  \tag{4.29}\\
& =\lim _{m \rightarrow \infty} e^{\delta t \mathbf{A}\left(x_{n}\right)} e^{\delta t \mathbf{A}\left(x_{m-1}\right)} \cdots e^{\delta t \mathbf{A}\left(x_{2}\right)} e^{\delta t \mathbf{A}\left(x_{1}\right)}
\end{align*}
$$

where $\delta t=\left(t-t_{0}\right) / m$, and $x_{n}=x\left(t_{0}+n \delta t\right)$. Slightly perverse indexing of the products indicates that in our convention the successive infinitesimal deformation are applied by multiplying from the left. The two formulas for $\mathbf{J}^{t}$ agree to leading order in $\delta t$, and the $m \rightarrow \infty$ limit of this procedure is the integral

$$
\begin{equation*}
\mathbf{J}_{i j}^{t}\left(x_{0}\right)=\left[\mathbf{T} e^{\int_{0}^{t} d \tau \mathbf{A}(x(\tau))}\right]_{i j}, \tag{4.30}
\end{equation*}
$$

where $\mathbf{T}$ stands for time-ordered integration, defined as the continuum limit of the successive left multiplications (4.29). This formula for $\mathbf{J}$ is the main result of this chapter.

It is evident from the time-ordered product structure (4.29) that the Jacobian matrices are multiplicative along the flow,

$$
\begin{equation*}
\mathbf{J}^{t+t^{\prime}}(x)=\mathbf{J}^{t^{\prime}}\left(x^{\prime}\right) \mathbf{J}^{t}(x), \quad \text { where } x^{\prime}=f^{t}(x) \tag{4.31}
\end{equation*}
$$

The formula (4.29) is a matrix generalization of the crude Euler integrator (2.16) and is neither smart not accurate. Much better numerical accuracy is obtained by the following observation: To linear order in $\delta t$,

$$
\mathbf{J}^{t+\delta t}-\mathbf{J}^{t}=\delta t \mathbf{A} \mathbf{J}^{t}+O\left((\delta t)^{2}\right),
$$

so the Jacobian matrix itself satisfies the linearized equation (4.1)

$$
\begin{equation*}
\frac{d}{d t} \mathbf{J}^{t}(x)=\mathbf{A}(x) \mathbf{J}^{t}(x), \quad \text { initial condition } \mathbf{J}^{0}(x)=\mathbf{1} \tag{4.32}
\end{equation*}
$$

Given a numerical routine for integrating the equations of motion, evaluation of the Jacobian matrix requires minimal additional programming effort; one simply extends the $d$-dimensional integration routine and integrates concurrently with $f^{t}(x)$ the $d^{2}$ elements of $\mathbf{J}^{t}(x)$.

The qualifier "simply" is perhaps too glib. Integration will work for short finite times, but for exponentially unstable flows one quickly runs into numerical over- and/or underflow problems, so further thought will have to go into implementating this calculation.


### 4.4 Stability of maps

The transformation of an infinitesimal neighborhood of a trajectory under the iteration of a map follows from Taylor expanding the iterated mapping at discrete time $n$ to linear order, as in (4.4). The linearized neighborhood is transported by the Jacobian matrix evaluated at a discrete set of times $n=1,2, \ldots$,

$$
\begin{equation*}
\mathbf{J}_{i j}^{n}\left(x_{0}\right)=\left.\frac{\partial f_{i}^{n}(x)}{\partial x_{j}}\right|_{x=x_{0}} \tag{4.33}
\end{equation*}
$$

This matrix is in the literature sometimes called the fundamental matrix. As the simplest example, consider a 1-dimensional map. The chain rule yields the stability of the $n$th iterate

$$
\begin{equation*}
\Lambda\left(x_{0}\right)_{n}=\frac{d}{d x} f^{n}\left(x_{0}\right)=\prod_{m=0}^{n-1} f^{\prime}\left(x_{m}\right), \quad x_{m}=f^{m}\left(x_{0}\right) . \tag{4.34}
\end{equation*}
$$

The 1-step product formula for the stability of the $n$th iterate of a $d$ dimensional map

$$
\begin{align*}
\mathbf{J}^{n}\left(x_{0}\right) & =\mathbf{J}\left(x_{n-1}\right) \cdots \mathbf{J}\left(x_{1}\right) \mathbf{J}\left(x_{x_{0}}\right), \\
\mathbf{J}(x)_{k l} & =\frac{\partial}{\partial x_{l}} f_{k}(x), \quad x_{m}=f^{m}\left(x_{0}\right) \tag{4.35}
\end{align*}
$$

follows from the chain rule for matrix derivatives

$$
\frac{\partial}{\partial x_{i}} f_{j}(f(x))=\left.\sum_{k=1}^{d} \frac{\partial}{\partial y_{k}} f_{j}(y)\right|_{y=f(x)} \frac{\partial}{\partial x_{i}} f_{k}(x)
$$

If you prefer to think of a discrete time dynamics as a sequence of Poincaré section returns, then (4.35) follows from (4.31), Jacobian matrices are multiplicative along the flow.

Example 4.7 Hénon map Jacobian matrix: For the Hénon map (3.12) the Jacobian matrix for the nth iterate of the map is

$$
\mathbf{J}^{n}\left(x_{0}\right)=\prod_{m=n}^{1}\left(\begin{array}{cc}
-2 a x_{m} & b  \tag{4.36}\\
1 & 0
\end{array}\right), \quad x_{m}=f_{1}^{m}\left(x_{0}, y_{0}\right) .
$$

The determinant of the Hénon one time step Jacobian matrix (4.36) is constant,

$$
\begin{equation*}
\operatorname{det} \mathbf{J}=\Lambda_{1} \Lambda_{2}=-b \tag{4.37}
\end{equation*}
$$

so in this case only one eigenvalue $\Lambda_{1}=-b / \Lambda_{2}$ needs to be determined. This is not an accident; a constant Jacobian was one of desiderata that led Hénon to construct a map of this particular form.


### 4.4.1 Stability of Poincaré return maps

> (R. Paškauskas and P. Cvitanović)

Figure 4.5: If $x(t)$ intersects the Poincaré section $\mathcal{P}$ at time $\tau$, the nearby $x(t)+\delta x(t)$ trajectory intersects it time $\tau+\delta t$ later. As ( $U^{\prime}$. $\left.v^{\prime} \delta t\right)=-\left(U^{\prime} \cdot \mathbf{J} \delta x\right)$, the difference in arrival times is given by $\delta t=-\left(U^{\prime} \cdot \mathbf{J} \delta x\right) /\left(U^{\prime} \cdot v^{\prime}\right)$.


We now relate the linear stability of the Poincaré return map $P$ : $\mathcal{P} \rightarrow \mathcal{P}$ defined in sect. 3.1 to the stability of the continuous time flow in the full phase space.

The hypersurface $\mathcal{P}$ can be specified implicitly through a function $U(x)$ that is zero whenever a point $x$ is on the Poincaré section. A nearby point $x+\delta x$ is in the hypersurface $\mathcal{P}$ if $U(x+\delta x)=0$, and the same is true for variations around the first return point $x^{\prime}=x(\tau)$, so expanding $U\left(x^{\prime}\right)$ to linear order in $\delta x$ leads to the condition

$$
\begin{equation*}
\left.U_{i}\left(x^{\prime}\right) \frac{d x_{i}^{\prime}}{d x_{j}}\right|_{\mathcal{P}}=0 \tag{4.38}
\end{equation*}
$$

In what follows $U_{i}$ is the gradient of $U$ defined in (3.2), unprimed quantities refer to the starting point $x=x_{0} \in \mathcal{P}, v=v\left(x_{0}\right)$, and the primed quantities to the first return: $x^{\prime}=x(\tau), v^{\prime}=v\left(x^{\prime}\right), U^{\prime}=U\left(x^{\prime}\right)$. For brevity we shall also denote the full phase space Jacobian matrix at the first return by $\mathbf{J}=\mathbf{J}^{\tau}\left(x_{0}\right)$. Both the first return $x^{\prime}$ and the time of flight to the next Poincaré section $\tau(x)$ depend on the starting point $x$, so the Jacobian matrix

$$
\begin{equation*}
\hat{\mathbf{J}}(x)_{i j}=\left.\frac{d x_{i}^{\prime}}{d x_{j}}\right|_{\mathcal{P}} \tag{4.39}
\end{equation*}
$$

with both initial and the final variation constrained to the Poincaré section hypersurface $\mathcal{P}$ is related to the continuous flow Jacobian matrix by

$$
\left.\frac{d x_{i}^{\prime}}{d x_{j}}\right|_{\mathcal{P}}=\frac{\partial x_{i}^{\prime}}{\partial x_{j}}+\frac{d x_{i}^{\prime}}{d t} \frac{d t}{d x_{j}}=\mathbf{J}_{i j}+v_{i}^{\prime} \frac{d t}{d x_{j}}
$$

The return time variation $d t / d x$, figure 4.5 , is eliminated by substituting this expression into the constraint (4.38),

$$
0=U_{i}^{\prime} \mathbf{J}_{i j}+\left(v^{\prime} \cdot U^{\prime}\right) \frac{d t}{d x_{j}}
$$

yielding the projection of the full space $(d+1)$-dimensional Jacobian matrix to the Poincaré map $d$-dimensional Jacobian matrix:

$$
\begin{equation*}
\hat{\mathbf{J}}_{i j}=\left(\delta_{i k}-\frac{v_{i}^{\prime} U_{k}^{\prime}}{\left(v^{\prime} \cdot U^{\prime}\right)}\right) \mathbf{J}_{k j} \tag{4.40}
\end{equation*}
$$

Substituting (4.6) we verify that the initial velocity $v(x)$ is a zero-eigenvector of $\hat{\mathbf{J}}$

$$
\begin{equation*}
\hat{\mathbf{J}} v=0, \tag{4.41}
\end{equation*}
$$

so the Poincaré section eliminates variations parallel to $v$, and $\hat{\mathbf{J}}$ is a rank $d$ matrix, that is, one less than the dimension of the continuus time flow.

## Résumé

A neighborhood of a trajectory deforms as it is transported by a flow. In the linear approximation, the matrix of variations $\mathbf{A}$ describes the shearing/compression/expansion of an infinitesimal neighborhood in an infinitesimal time step. The deformation after a finite time $t$ is described by the Jacobian matrix

$$
\mathbf{J}^{t}\left(x_{0}\right)=\mathbf{T} e^{\int_{0}^{t} d \tau \mathbf{A}(x(\tau))}
$$

where $\mathbf{T}$ stands for the time-ordered integration, or, for discrete time maps, by multiplication along the $n$ points $x_{0}, x_{1}, x_{2}, \ldots, x_{n-1}$ on the trajectory of $x_{0}$,

$$
\mathbf{J}^{n}\left(x_{0}\right)=\prod_{m=n-1}^{0} \mathbf{J}\left(x_{n}\right) \mathbf{J}\left(x_{n-1}\right) \cdots \mathbf{J}\left(x_{2}\right) \mathbf{J}\left(x_{1}\right)
$$

with $\mathbf{J}(x)$ the single time step Jacobian matrix. Here symbol $\Lambda_{k}$ will always denote the $k$ th stability eigenvalue of the finite time Jacobian matrix $\mathbf{J}^{t}$, and symbol $\lambda_{k}$ the the $k$ th stability exponent

$$
|\Lambda|=e^{t \lambda}, \quad \Lambda_{ \pm}=e^{t(\lambda \pm i \theta)} .
$$

For a complex stability eigenvalue the conjugate eigenvalue pair describes rotational motion in the plane defined by the corresponding pair of eigenvectors, with period $T=2 \pi / \theta$.

The eigenvalues and eigendirections of the Jacobian matrix describe the deformation of an initial infinitesimal sphere of neighboring trajectories into an ellipsoid a finite time $t$ later. Nearby trajectories separate exponentially along unstable directions, approach each other along stable directions, and change slowly (algebraically) their distance along marginal directions. The Jacobian matrix $\mathbf{J}^{t}$ is in general neither symmetric, nor diagonalizable by a rotation, nor do its (left or right) eigenvectors define an orthonormal coordinate frame. Furthermore, although the stability matrices are multiplicative along the flow, in dimensions higher than one their eigenvalues in general are not. This lack of multiplicativity has important repercussions for both classical and quantum dynamics.

## References

[4.1] J.-L. Thiffeault, "Derivatives and constraints in chaotic flows: Asymptotic behaviour and a numerical method", preprint (2001).

## Exercises

## Exercise 4.1 Trace-log of a matrix. Prove that

$\operatorname{det} M=e^{\operatorname{tr} \ln M}$.
for an arbitrary finite dimensional matrix M.
Exercise 4.2 Stability, diagonal case. Verify the relation (4.16)

$$
\mathbf{J}^{t}=e^{t \mathbf{A}}=\mathbf{U}^{-1} e^{t \mathbf{A}_{D}} \mathbf{U}, \quad \text { where } \quad \mathbf{A}_{D}=\mathbf{U A U}^{-1}
$$

Exercise 4.3 Topology of the Rössler flow. (continuation of exercise 3.1)
(a) Show that equation $|\operatorname{det}(A-\lambda \mathbf{1})|=0$ for Rössler system in the notation of exercise 2.23 can be written as

$$
\begin{equation*}
\lambda^{3}+\lambda^{2} c\left(p^{\mp}-\epsilon\right)+\lambda\left(p^{ \pm} / \epsilon+1-c^{2} \epsilon p^{\mp}\right) \mp c \sqrt{D}=0 \tag{4.42}
\end{equation*}
$$

(b) Solve (4.42) for eigenvalues $\lambda^{ \pm}$for each equilibrium point as an expansion in powers of $\epsilon$. Derive

$$
\begin{align*}
& \lambda_{1}^{-}=-c+\epsilon c /\left(c^{2}+1\right)+o(\epsilon) \\
& \lambda_{2}^{-}=\epsilon c^{3} /\left[2\left(c^{2}+1\right)\right]+o\left(\epsilon^{2}\right) \\
& \theta_{2}^{-}=1+\epsilon /\left[2\left(c^{2}+1\right)\right]+o(\epsilon) \\
& \lambda_{1}^{+}=c \epsilon(1-\epsilon)+o\left(\epsilon^{3}\right)  \tag{4.43}\\
& \lambda_{2}^{+}=-\epsilon^{5} c^{2} / 2+o\left(\epsilon^{6}\right) \\
& \theta_{2}^{+}=\sqrt{1+1 / \epsilon}(1+o(\epsilon))
\end{align*}
$$

Compare with exact eigenvalues. What are dynamical implications of the extravagant value of $\lambda_{1}^{-}$?
(continued as exercise 4.3)
Exercise 4.4 A contracting baker's map. Consider a contracting (or "dissipative") baker's map, acting on a unit square $[0,1]^{2}=[0,1] \times[0,1]$, defined by

$$
\begin{aligned}
& \binom{x_{n+1}}{y_{n+1}}=\binom{x_{n} / 3}{2 y_{n}} \quad y_{n} \leq 1 / 2 \\
& \binom{x_{n+1}}{y_{n+1}}=\binom{x_{n} / 3+1 / 2}{2 y_{n}-1} \quad y_{n}>1 / 2
\end{aligned}
$$

This map shrinks strips by a factor of $1 / 3$ in the $x$-direction, and then stretches (and folds) them by a factor of 2 in the $y$-direction.

By how much does the phase space volume contract fo one iteration of the map?

## Chapter 5

## Newtonian dynamics

You might think that the strangeness of contracting flows, flows such as the Rössler flow of figure 2.3 is of concern only to chemists; real physicists do Hamiltonians, right? Not at all - while it is easier to visualize aperiodic dynamics when a flow is contracting onto a lower-dimensional attracting set, there are plenty examples of chaotic flows that do preserve the full symplectic invariance of Hamiltonian dynamics. The truth is, the whole story started with Poincaré's restricted 3-body problem, a realization that chaos rules also in general (non-Hamiltonian) flows came much later.

Here we briefly review parts of classical dynamics that we will need later on; symplectic invariance, canononical transformations, and stability of Hamiltonian flows. We discuss billiard dynamics in some detail in chapter 6 .

### 5.1 Hamiltonian flows

(P. Cvitanović and L.V. Vela-Arevalo)

An important class of flows are Hamiltonian flows, given by a Hamiltonian $H(q, p)$ together with the Hamilton's equations of motion

$$
\begin{equation*}
\dot{q}_{i}=\frac{\partial H}{\partial p_{i}}, \quad \dot{p}_{i}=-\frac{\partial H}{\partial q_{i}}, \tag{5.1}
\end{equation*}
$$

with the $2 D$ phase space coordinates $x$ split into the configuration space coordinates and the conjugate momenta of a Hamiltonian system with $D$ degrees of freedom (dof):

$$
\begin{equation*}
x=(\mathbf{q}, \mathbf{p}), \quad \mathbf{q}=\left(q_{1}, q_{2}, \ldots, q_{D}\right), \quad \mathbf{p}=\left(p_{1}, p_{2}, \ldots, p_{D}\right) . \tag{5.2}
\end{equation*}
$$

The energy, or the value of the Hamiltonian function at the phase space

Figure 5.1: Phase plane of the unforced, undamped Duffing oscillator. The trajectories lie on level sets of the Hamiltonian (5.4).

point $x=(\mathbf{q}, \mathbf{p})$ is constant along the trajectory $x(t)$,

$$
\begin{align*}
\frac{d}{d t} H(\mathbf{q}(t), \mathbf{p}(t)) & =\frac{\partial H}{\partial q_{i}} \dot{q}_{i}(t)+\frac{\partial H}{\partial p_{i}} \dot{p}_{i}(t) \\
& =\frac{\partial H}{\partial q_{i}} \frac{\partial H}{\partial p_{i}}-\frac{\partial H}{\partial p_{i}} \frac{\partial H}{\partial q_{i}}=0 \tag{5.3}
\end{align*}
$$

so the trajectories lie on surfaces of constant energy, or level sets of the Hamiltonian $\{(q, p): H(q, p)=E\}$. For 1-dof Hamiltonian systems this is basically the whole story.

Example 5.1 Unforced undamped Duffing oscillator: When the damping term is removed from the Duffing oscillator (2.6), the system can be written in Hamiltonian form with the Hamiltonian

$$
\begin{equation*}
H(q, p)=\frac{p^{2}}{2}-\frac{q^{2}}{2}+\frac{q^{4}}{4} \tag{5.4}
\end{equation*}
$$

This is a 1-dof Hamiltonian system, with a 2-dimensional phase space, the plane ( $q, p$ ). The Hamilton's equations (5.1) are

$$
\begin{equation*}
\dot{q}=p, \quad \dot{p}=q-q^{3} \tag{5.5}
\end{equation*}
$$

For 1-dof systems, the "surfaces" of constant energy (5.3) are simply curves in the phase plane $(q, p)$, and the dynamics is very simple: the curves of constant energy are the trajectories, as shown in figure 5.1.

Thus all 1-dof systems are integrable, in the sense that the entire phase plane is foliated by curves of constant energy, either periodic - as is the case for the harmonic oscillator (a "bound state") - or open (a "scattering trajectory"). Add one more degree of freedom, and chaos breaks loose.

Example 5.2 Collinear helium: In chapter 34, we shall apply the periodic orbit theory to the quantization of helium. In particular, we will study collinear helium, a doubly charged nucleus with two electrons arranged on a line, an electron on each side of the nucleus. The Hamiltonian for this system is

$$
\begin{equation*}
H=\frac{1}{2} p_{1}^{2}+\frac{1}{2} p_{2}^{2}-\frac{2}{r_{1}}-\frac{2}{r_{2}}+\frac{1}{r_{1}+r_{2}} . \tag{5.6}
\end{equation*}
$$

Figure 5.2: A typical collinear helium trajectory in the $r_{1}-r_{2}$ plane; the trajectory enters along the $r_{1}$-axis and then, like almost every other trajectory, after a few bounces escapes to infinity, in this case along the $r_{2}$-axis.

$r_{1}$

Collinear helium has 2 dof, and thus a 4-dimensional phase space $\mathcal{M}$, which energy conservation reduces to 3 dimensions. The dynamics can be projected onto the 2 dimensional configuration plane, the $\left(r_{1}, r_{2}\right), r_{i} \geq 0$ quadrant, figure 5.2. It looks messy, and, indeed, it will turn out to be no less chaotic than a pinball bouncing between three disks. As always, a Poincaré section will be more informative than this rather arbitrary projection of the flow.

### 5.2 Stability of Hamiltonian flows

Hamiltonian flows offer an illustration of the ways in which an invariance of equations of motion can affect the dynamics. In the case at hand, the symplectic invariance will reduce the number of independent stability eigenvalues by a factor of 2 or 4 .

### 5.2.1 Canonical transformations

The equations of motion for a time-independent, $D$-dof Hamiltonian (5.1) can be written

$$
\dot{x}_{i}=\omega_{i j} H_{j}(x), \quad \omega=\left(\begin{array}{cc}
0 & \mathbf{I}  \tag{5.7}\\
-\mathbf{I} & 0
\end{array}\right), \quad H_{j}(x)=\frac{d}{d x_{j}} H(x),
$$

where $x=(\mathbf{q}, \mathbf{p}) \in \mathcal{M}$ is a phase space point, $H_{k}=\partial_{k} H$ is the column vector of partial derivatives of $H, \mathbf{I}$ is the $[D \times D]$ unit matrix, and $\omega$ the $[2 D \times 2 D]$ symplectic form

$$
\begin{equation*}
\omega^{T}=-\omega, \quad \omega^{2}=-\mathbf{1} \tag{5.8}
\end{equation*}
$$

The linearized motion in the vicinity $x+\delta x$ of a phase space trajectory $x(t)=(\mathbf{q}(t), \mathbf{p}(t))$ is described by the Jacobian matrix (4.22). For Hamiltonian flows the matrix of variations (4.32) takes form

$$
\begin{equation*}
\frac{d}{d t} \mathbf{J}^{t}(x)=A(x) \mathbf{J}^{t}(x), \quad A_{i j}(x)=\omega_{i k} H_{k j}(x) \tag{5.9}
\end{equation*}
$$

where the matrix of second derivatives $H_{k n}=\partial_{k} \partial_{n} H$ is called the Hessian matrix. From the symmetry of $H_{k n}$ it follows that

$$
\begin{equation*}
A^{T} \omega+\omega A=0 . \tag{5.10}
\end{equation*}
$$

This is the defining property for infinitesimal generators of symplectic (or canonical) transformations, transformations which leave the symplectic form $\omega$ invariant.

More explicitely: just as the rotation group $O(d)$ is the Lie group of all matrix transformations which preserve a bilinear symmetric invariant $x^{2}=x_{i} \delta_{i j} x_{j}$, that is, the length squared, the symplectic group $S p(2 D)$ is defined as the Lie group of all matrix transformations $x^{\prime}=\mathbf{J} x$ which preserve a bilinear antisymmetric invariant $x_{i} \omega_{i j} y_{j}$. The symplectic Lie algebra $s p(2 D)$ follows by writing $\mathbf{J}=\exp (\delta t A)$ and linearizing $\mathbf{J}=1+$ $\delta t A$. This yields (5.10) as the defining property of infinitesimal symplectic transformations.

Consider now a smooth nonlinear change of variables of form $y=h(x)$, and the Hamiltonian as the function of the new phase space variables, $K(x)=H(h(x))$. Rewriting (5.7) as a function of $y$ and employing the chain rule we find

$$
\begin{align*}
\partial K(x) & =(\partial h)^{T} \partial H(h(x))=(\partial h)^{T}(-\omega \dot{y}) \\
& =-(\partial h)^{T} \omega(\partial h) \dot{x}, \tag{5.11}
\end{align*}
$$

where $(\partial h)_{k n}=\partial h_{k} / \partial x_{n}$. Consider the case that $\partial h$, the linearization of $h$, is a symplectic transformation, that is, a map that preserves the symplectic form,

$$
\begin{equation*}
(\partial h)^{T} \omega(\partial h)=\omega . \tag{5.12}
\end{equation*}
$$

In that case $K(x)$ also induces a Hamiltonian flow $\dot{x}_{i}=\omega_{i j} K_{j}(x)$, and $h$ is called a canonical transformation. We care about canonical transformations for two reasons. First (and this is a dark art), if the canonical transformation $h$ is very cleverly chosen, the flow in new coordinates might be considerably simpler than the original flow. Second,

Example 5.3 Hamiltonian flows are canonical: For Hamiltonian flows it follows from (5.10) that $\frac{d}{d t}\left(\mathbf{J}^{T} \omega \mathbf{J}\right)=0$, and since at the initial time $\mathbf{J}^{0}\left(x_{0}\right)=\mathbf{1}$, $\mathbf{J}$ is a symplectic transformation $\mathbf{J}^{T} \omega \mathbf{J}=\omega$. This equality is valid for all times, so a Hamiltonian flow $f^{t}(x)$ is a canonical transformation, with the linearization $\partial_{x} f^{t}(x)$ a symplectic transformation (5.12):

$$
\begin{equation*}
\mathbf{J}^{T} \omega \mathbf{J}=\omega . \tag{5.13}
\end{equation*}
$$

For notational brevity here we have suppressed the dependence on time and the initial point, $\mathbf{J}=\mathbf{J}^{t}\left(x_{0}\right)$.


Figure 5.3: (a) Stability exponents of a Hamiltonian equilibrium point, 2-dof. (b) Stability of a symplectic map in $\mathbb{R}^{4}$.

### 5.2.2 Stability of equilibria of Hamiltonian flows

For an equilibrium point $x_{q}$ the matrix of variations $A$ is constant. Its eigenvalues describe the linear stability of the equilibrium point. In the case of Hamiltonian flows, from (5.10) it follows that the characteristic polynomial of $A$ for an equilibrium $x_{q}$ satisfies

$$
\begin{align*}
\operatorname{det}(A-\lambda \mathbf{1}) & =\operatorname{det}\left(\omega^{-1}(A-\lambda \mathbf{1}) \omega\right)=\operatorname{det}(-\omega A \omega-\lambda \mathbf{1}) \\
& =\operatorname{det}\left(A^{T}+\lambda \mathbf{1}\right)=\operatorname{det}(A+\lambda \mathbf{1}) \tag{5.14}
\end{align*}
$$

$A$ is the matrix (5.10) with real matrix elements, so its eigenvalues are either real or come in complex pairs. Symplectic invariance implies in addition that if $\lambda$ is an eigenvalue, then $-\lambda, \lambda^{*}$ and $-\lambda^{*}$ are also eigenvalues. Distinct symmetry classes of the stability exponents of an equilibrium point in a 2-dof system are displayed in figure 5.3(a). It is worth noting that while the linear stability of equilibria in a Hamiltonian system always respects this symmetry, the nonlinear stability can be completely different.

### 5.3 Symplectic maps

A stability eigenvalue $\Lambda=\Lambda\left(x_{0}, t\right)$ associated to a trajectory is an eigenvalue of the Jacobian matrix $\mathbf{J}$. The transpose $\mathbf{J}^{T}$ and the inverse $\mathbf{J}^{-1}$ are related by

$$
\begin{equation*}
\mathbf{J}^{-1}=-\omega \mathbf{J}^{T} \omega, \tag{5.15}
\end{equation*}
$$

so the characteristic polynomial satisfies

$$
\begin{align*}
\operatorname{det}(\mathbf{J}-\Lambda \mathbf{1}) & =\operatorname{det}\left(\omega(\mathbf{J}-\Lambda \mathbf{1}) \omega^{-1}\right)=\operatorname{det}(\omega \mathbf{J} \omega-\Lambda \mathbf{1}) \\
& =\operatorname{det}\left(\left(\mathbf{J}^{T}\right)^{-1}-\Lambda \mathbf{1}\right)=\operatorname{det}\left(\mathbf{J}^{-1}\right) \operatorname{det}(\mathbf{1}-\Lambda \mathbf{J}) \\
& =\Lambda^{2 D} \operatorname{det}\left(\mathbf{J}-\Lambda^{-1} \mathbf{1}\right) . \tag{5.16}
\end{align*}
$$

Hence if $\Lambda$ is an eigenvalue of $\mathbf{J}$, so are $1 / \Lambda, \Lambda^{*}$ and $1 / \Lambda^{*}$. Real (nonmarginal, $|\Lambda| \neq 1$ ) eigenvalues always come paired as $\Lambda, 1 / \Lambda$. An immediate consequence of symplectic invariance is

$$
\begin{equation*}
\operatorname{det} \mathbf{J}=1 \tag{5.17}
\end{equation*}
$$

The complex eigenvalues come in pairs $\Lambda, \Lambda^{*},|\Lambda|=1$, or in loxodromic quartets $\Lambda, 1 / \Lambda, \Lambda^{*}$ and $1 / \Lambda^{*}$. These possibilities are illustrated in figure 5.3(b).

Example 5.4 2-dimensional symplectic maps: In the 2-dimensional case the eigenvalues (8.2) depend only on $\operatorname{tr} \mathbf{J}^{t}$

$$
\begin{equation*}
\Lambda_{1,2}=\frac{1}{2}\left(\operatorname{tr} \mathbf{J}^{t} \pm \sqrt{\left(\operatorname{tr} \mathbf{J}^{t}-2\right)\left(\operatorname{tr} \mathbf{J}^{t}+2\right)}\right) . \tag{5.18}
\end{equation*}
$$

The trajectory is elliptic if the residue $\left|\operatorname{tr} \mathbf{J}^{t}\right|-2 \leq 0$, with complex eigenvalues $\Lambda_{1}=$ $e^{i \theta t}, \Lambda_{2}=\Lambda_{1}^{*}=e^{-i \theta t}$. If $\left|\operatorname{tr} \mathbf{J}^{t}\right|-2>0, \lambda$ is real, and the trajectory is either

$$
\begin{align*}
\text { hyperbolic } & \Lambda_{1}=e^{\lambda t}, \quad \Lambda_{2}=e^{-\lambda t}, \text { or }  \tag{5.19}\\
\text { inverse hyperbolic } & \Lambda_{1}=-e^{\lambda t}, \quad \Lambda_{2}=-e^{-\lambda t} . \tag{5.20}
\end{align*}
$$

Example 5.5 Hamiltonian Hénon map, reversibility: By (4.37) the Hénon map (3.12) for $b=-1$ value is the simplest $2-d$ orientation preserving area-preserving map, often studied to better understand topology and symmetries of Poincaré sections of 2 dof Hamiltonian flows. We find it convenient to multiply (3.13) by a and absorb the a factor into $x$ in order to bring the Hénon map for the $b=-1$ parameter value into the form

$$
\begin{equation*}
x_{i+1}+x_{i-1}=a-x_{i}^{2}, \quad i=1, \ldots, n_{p} \tag{5.21}
\end{equation*}
$$

The 2-dimensional Hénon map for $b=-1$ parameter value

$$
\begin{align*}
x_{n+1} & =a-x_{n}^{2}-y_{n} \\
y_{n+1} & =x_{n} . \tag{5.22}
\end{align*}
$$

is Hamiltonian (symplectic) in the sense that it preserves area in the $[x, y]$ plane.
For definitiveness, in numerical calculations we shall fix here (arbitrarily) the stretching parameter value to $a=6$, a value large enough to guarantee that all roots of $0=f^{n}(x)-x$ (periodic points) are real.

### 5.3.1 Poincaré invariants

Let $C$ a region in the phase space and $V(0)$ its volume. Denoting the flow of the Hamiltonian system by $f^{t}(x)$, the volume of $C$ after a time $t$ is $V(t)=f^{t}(C)$, and using (5.17) we derive the Liouville theorem:

$$
\begin{align*}
V(t)= & \int_{f^{t}(C)} d x=\int_{C}\left\|\frac{\partial f^{t}\left(x^{\prime}\right)}{\partial x}\right\| d x^{\prime} \\
& \int_{C} \operatorname{det}(\mathbf{J}) d x^{\prime}=\int_{C} d x^{\prime}=V(0), \tag{5.23}
\end{align*}
$$

Hamiltonian flows preserve phase-space volumes.
The symplectic structure of Hamilton's equations buys us much more than the "incompressibility," or the phase space volume conservation. Consider the symplectic product of two infinitesimal vectors

$$
\begin{align*}
(\delta x, \delta \hat{x}) & =\delta x^{T} \omega \delta \hat{x}=\delta p_{i} \delta \hat{q}_{i}-\delta q_{i} \delta \hat{p}_{i} \\
& =\sum_{i=1}^{D}\left\{\text { oriented area in the }\left(q_{i}, p_{i}\right) \text { plane }\right\} \tag{5.24}
\end{align*}
$$

Time $t$ later we have

$$
\left(\delta x^{\prime}, \delta \hat{x}^{\prime}\right)=\delta x^{T} \mathbf{J}^{T} \omega \mathbf{J} \delta \hat{x}=\delta x^{T} \omega \delta \hat{x} .
$$

This has the following geometrical meaning. We imagine there is a reference phase space point. We then define two other points infinitesimally close so that the vectors $\delta x$ and $\delta \hat{x}$ describe their displacements relative to the reference point. Under the dynamics, the three points are mapped to three new points which are still infinitesimally close to one another. The meaning of the above expression is that the area of the parallelopiped spanned by the three final points is the same as that spanned by the inital points. The integral (Stokes theorem) version of this infinitesimal area invariance states that for Hamiltonian flows the $D$ oriented areas $\mathcal{V}_{i}$ bounded by $D$ loops $\Omega \mathcal{V}_{i}$, one per each ( $q_{i}, p_{i}$ ) plane, are separately conserved:

$$
\begin{equation*}
\int_{\mathcal{V}} d p \wedge d q=\oint_{\Omega \mathcal{V}} p \cdot d q=\text { invariant. } \tag{5.25}
\end{equation*}
$$

Morally a Hamiltonian flow is really $D$-dimensional, even though its phase space is $2 D$-dimensional. Hence for Hamiltonian flows one emphasizes $D$, the number of the degrees of freedom.

## Commentary

Remark 5.1 Symplectic. The term symplectic - Greek for twining or plaiting together - was introduced into mathematics by Hermann Weyl. "Canonical" lineage is church-doctrinal: Greek "kanon", referring to a reed used for measurement, came to mean in Latin a rule or a standard.

Remark 5.2 The sign convention of $\omega$. The overall sign of $\omega$, the symplectic invariant in (5.7), is set by the convention that the Hamilton's principal function (for energy conserving flows) is given by $R\left(q, q^{\prime}, t\right)=\int_{q}^{q^{\prime}} p_{i} d q_{i}-E t$. With this sign convention the action along a classical path is minimal, and the kinetic energy of a free particle is positive.

Remark 5.3 Symmetries of the symbol square. For a more detailed discussion of symmetry lines see refs. [17.4, 3.5, 12.46]. In discussion of example ??, we have followed refs. [5.3, 5.4, 5.5].

## References

[5.1] J.E. Marsden and T.S. Ratiu, Introduction to Mechanics and Symmetry (Springer, New York, 1994).
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[5.3] D.G. Sterling, H.R. Dullin and J.D. Meiss, "Homoclinic Bifurcations for the Hénon Map," Physica D 134, 153 (1999); chao-dyn/9904019.
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[5.5] A. Gómez and J.D. Meiss, "Reversible Polynomial Automorphisms of the Plane: the Involutory Case," Phys. Lett. A 312, 49 (2003); nlin. CD/0209055.
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## Exercises

Exercise 5.1 Complex nonlinear Schrödinger equation. Consider the complex nonlinear Schrödinger equation in one spatial dimension [5.1]:
$i \frac{\partial \phi}{\partial t}+\frac{\partial^{2} \phi}{\partial x^{2}}+\beta \phi|\phi|^{2}=0, \quad \beta \neq 0$.
(a) Show that the function $\psi: \mathbb{R} \rightarrow \mathbf{C}$ defining the traveling wave solution $\phi(x, t)=\psi(x-c t)$ for $c>0$ satisfies a second-order complex differential equation equivalent to a Hamiltonian system in $\mathbb{R}^{4}$ relative to the noncanonical symplectic form whose matrix is given by

$$
w_{c}=\left[\begin{array}{cccc}
0 & 0 & 1 & 0 \\
0 & 0 & 0 & 1 \\
-1 & 0 & 0 & -c \\
0 & -1 & c & 0
\end{array}\right] .
$$

(b) Analyze the equilibria of the resulting Hamiltonian system in $\mathbb{R}^{4}$ and determine their linear stability properties.
(c) Let $\psi(s)=e^{i c s / 2} a(s)$ for a real function $a(s)$ and determine a second order equation for $a(s)$. Show that the resulting equation is Hamiltonian and has heteroclinic orbits for $\beta<0$. Find them.
(d) Find "soliton" solutions for the complex nonlinear Schrödinger equation.

> (Luz V. Vela-Arevalo)

## Exercise 5.2 When is a linear transformation canonical?

(a) Let $A$ be a $n \times n$ invertible matrix. Show that the map $\phi: \mathbb{R}^{2 n} \rightarrow \mathbb{R}^{2 n}$ given by $(\mathbf{q}, \mathbf{p}) \mapsto\left(A \mathbf{q},\left(A^{-1}\right)^{T} \mathbf{p}\right)$ is a canonical transformation.
(b) If $\mathbf{R}$ is a rotation in $\mathbb{R}^{3}$, show that the map $(\mathbf{q}, \mathbf{p}) \mapsto(\mathbf{R} \mathbf{q}, \mathbf{R} \mathbf{p})$ is a canonical transformation.
(Luz V. Vela-Arevalo)

Exercise 5.3 Cherry's example. What follows [5.2] is mostly a reading exercise, about a Hamiltonian system that is linearly stable but nonlinearly unstable. Consider the Hamiltonian system on $\mathbb{R}^{4}$ given by

$$
H=\frac{1}{2}\left(q_{1}^{2}+p_{1}^{2}\right)-\left(q_{2}^{2}+p_{2}^{2}\right)+\frac{1}{2} p_{2}\left(p_{1}^{2}-q_{1}^{2}\right)-q_{1} q_{2} p_{1} .
$$

(a) Show that this system has an equilibrium at the origin, which is linearly stable. (The linearized system consists of two uncoupled oscillators with frequencies in ratios 2:1).
(b) Convince yourself that the following is a family of solutions parametrized by a constant $\tau$ :

$$
\begin{array}{ll}
q_{1}=-\sqrt{2} \frac{\cos (t-\tau)}{t-\tau}, & q_{2}=\frac{\cos 2(t-\tau)}{t-\tau} \\
p_{1}=\sqrt{2} \frac{\sin (t-\tau)}{t-\tau}, & p_{2}=\frac{\sin 2(t-\tau)}{t-\tau}
\end{array}
$$

These solutions clearly blow up in finite time; however they start at $t=0$ at a distance $\sqrt{3} / \tau$ from the origin, so by choosing $\tau$ large, we can find solutions starting arbitrarily close to the origin, yet going to infinity in a finite time, so the origin is nonlinearly unstable.
(Luz V. Vela-Arevalo)

## Chapter 6

## Billiards

> We owe it to a book to withhold judgment until we reach page 100 .
> Henrietta McNutt, George Johnson's seventh-grade English teacher

The dynamics that we have the best intuitive grasp on, and find easiest to grapple with both numerically and conceptually, is the dynamics of billiards. For billiards, discrete time is altogether natural; a particle moving through a billiard suffers a sequence of instantaneous kicks, and executes simple motion inbetween, so there is no need to contrive a Poincaré section. We have already used this system in sect. 1.3 as the intuitively most accessible example of chaos. Here we define billiard dynamics more precisely, anticipating the applications to come.

### 6.1 Billiard dynamics

A billiard is defined by a connected region $Q \subset \mathbb{R}^{D}$, with boundary $\partial Q \subset$ $\mathbb{R}^{D-1}$ separating $Q$ from its complement $\mathbb{R}^{D} \backslash Q$. The region $Q$ can consist of one compact, finite volume component (in which case the billiard phase space is bounded, as for the stadium billiard figure 6.1 ), or can be infinite in extent, with its complement $\mathbb{R}^{D} \backslash Q$ consisting of one or several finite or infinite volume components (in which case the phase space is open, as for the 3 -disk pinball game figure 1.1). In what follows we shall most often restrict our attention to planar billiards. A point particle of mass $m$ and momentum $p_{n}=m v_{n}$ moves freely within the billiard, along a straight line, until it encounters the boundary. There it reflects specularly (specular = mirrorlike), with no change in the tangential component of momentum, and instantaneous reversal of the momentum component normal to the boundary,

$$
\begin{equation*}
p^{\prime}=p-2(p \cdot \hat{n}) \hat{n} \tag{6.1}
\end{equation*}
$$



Figure 6.1: The stadium billiard is a 2-dimensional domain bounded by two semicircles of radius $d=1$ connected by two straight walls of length $2 a$. At the points where the straight walls meet the semi-circles, the curvature of the border changes discontinuously; these are the only singular points of the flow. The length $a$ is the only parameter.


Figure 6.2: (a) A planar billiard trajectory is fixed by specifying the perimeter length parametrized by $s$ and the outgoing trajectory angle $\phi$, both measured counterclockwise with respect to the outward normal $\hat{n}$. (b) The Birkhoff phase space coordinate pair $(s, p)$ fully specifies the trajectory, where $p=|p| \sin \phi$ is the momentum component tangential to the boundary (and we set $|p|=1$ whenever convenient).
with $\hat{n}$ the unit vector normal to the boundary $\partial Q$ at the collision point. The angle of incidence equals the angle of reflection, as illustrated in figure 6.2. A billiard is a Hamiltonian system with a $2 D$-dimensional phase space $x=(q, p)$ and potential $V(q)=0$ for $q \in Q, V(q)=\infty$ for $q \in \partial Q$.

A billiard flow has a natural Poincaré section defined by Birkhoff coordinates $s_{n}$, the arc length position of the $n$th bounce measured along the billiard boundary, and $p_{n}=|p| \sin \phi_{n}$, the momentum component parallel to the boundary, where $\phi_{n}$ is the angle between the outgoing trajectory and the normal to the boundary. We measure both the arc length $s$, and the parallel momentum $p$ counterclockwise relative to the outward normal (see figure 6.2 as well as figure 1.6 (a)). In $D=2$, the Poincaré section is a cylinder (topologically an annulus), figure 6.3 , where the parallel momentum $p$ ranges for $-|p|$ to $|p|$, and the $s$ coordinate is cyclic along each connected component of $\partial Q$. The volume in the full phase space is preserved by the Liouville theorem (5.23). The Birkhoff coordinates $x=(s, p) \in \mathcal{P}$, are

Figure 6.3: In $D=2$ the billiard Poincaré section is a cylinder, with the parallel momentum $p$ ranging over $p \in\{-1,1\}$, and with the $s$ coordinate is cyclic along each connected component of $\partial Q$. The rectangle figure 6.2 (b) is such cylinder unfolded, with periodic boundary conditions glueing together the left and the right edge of the rectangle.

the natural choice, because with them the the Poincaré return map preserves the phase space volume in the $(s, p)$ parametrized Poincaré section (a perfectly good coordinate set $(s, \phi)$ does not do that).

Without loss of generality we set $m=|v|=|p|=1$. Poincaré section condition eliminates one dimension, and the energy conservation $|p|=1$ eliminates another, so the Poincaré section return map P is $(2 D-2)$ dimensional.

The dynamics is given by the Poincaré return map

$$
\begin{equation*}
P:\left(s_{n}, p_{n}\right) \mapsto\left(s_{n+1}, p_{n+1}\right) \tag{6.2}
\end{equation*}
$$

from the $n$th collision to the $(n+1)$ st collision. The discrete time dynamics map $P$ is equivalent to the Hamiltonian flow (5.1) in the sense that both describe the same full trajectory. Let $t_{n}$ denate the instant of $n$th collision. Then the position of the pinball $\in Q$ at time $t_{n}+\tau \leq t_{n+1}$ is given by $2 D-2$ Poincaré section coordinates $\left(s_{n}, p_{n}\right) \in \mathcal{P}$ together with $\tau$, the distance reached by the pinball along the $n$th section of its trajectory.

Example 6.1 3-disk game of pinball: In case of bounces off a circular disk, the position coordinate $s=r \theta$ is given by angle $\theta \in[0,2 \pi]$. For example, for the 3 -disk game of pinball of figure 1.4 and figure 1.6 we have two types of collisions:

$$
\begin{align*}
& P_{0}: \begin{cases}\phi^{\prime}=-\phi+2 \arcsin p & \text { back-reflection } \\
p^{\prime}=-p+\frac{a}{R} \sin \phi^{\prime}\end{cases}  \tag{6.3}\\
& P_{1}: \begin{cases}\phi^{\prime}=\phi-2 \arcsin p+2 \pi / 3 \\
p^{\prime}=p-\frac{a}{R} \sin \phi^{\prime} & \text { reflect to 3rd disk. }\end{cases} \tag{6.4}
\end{align*}
$$

Here $a=$ radius of a disk, and $R=$ center-to-center separation. Actually, as in this example we are computing intersections of circles and straight lines, nothing more than high-school geometry is required. There is no need to compute arcsin's either - one only needs to compute a square root per each reflection, and the simulations can be very fast.

Trajectory of the pinball in the 3-disk billiard is generated by a series of Poe 96 and $P_{1}$ 's. At each step on has to check whether the trajectory intersects the desired disk (and no disk inbetween). With minor modifications, the above formulas are valid for any smooth billiard as long as we replace $a$ by the local curvature of the boundary at the point of collision.

### 6.2 Stability of billiards

We turn next to the question of local stability of discrete time billiard systems. Infinitesimal equations of variations (4.2) do not apply, but the multiplicative structure (4.31) of the finite-time Jacobian matrices does. As they are more physical than most maps studied by dynamicists, let us work out the billiard stability in some detail.

On the face of it, a plane billiard phase space is 4 -dimensional. However, one dimension can be eliminated by energy conservation, and the other by the fact that the magnitude of the velocity is constant. We shall now show how going to a local frame of motion leads to a $[2 \times 2]$ Jacobian matrix.

Consider a 2-dimensional billiard with phase space coordinates $x=$ $\left(q_{1}, q_{2}, p_{1}, p_{2}\right)$. Let $t_{k}$ be the instant of the $k$ th collision of the pinball with the billiard boundary, and $t_{k}^{ \pm}=t_{k} \pm \epsilon, \epsilon$ positive and infinitesimal. With the mass and the velocity equal to 1 , the momentum direction can be specified by angle $\theta: x=\left(q_{1}, q_{2}, \sin \theta, \cos \theta\right)$. Now parametrize the $2-d$ neighborhood of a trajectory segment by $\delta x=(\delta z, \delta \theta)$, where

$$
\begin{equation*}
\delta z=\delta q_{1} \cos \theta-\delta q_{2} \sin \theta, \tag{6.5}
\end{equation*}
$$

$\delta \theta$ is the variation in the direction of the pinball motion. Due to energy conservation, there is no need to keep track of $\delta q_{\|}$, variation along the flow, as that remains constant. $\left(\delta q_{1}, \delta q_{2}\right)$ is the coordinate variation transverse to the $k$ th segment of the flow. From the Hamilton's equations of motion for a free particle, $d q_{i} / d t=p_{i}, d p_{i} / d t=0$, we obtain the equations of motion (4.1) for the linearized neighborhood

$$
\begin{equation*}
\frac{d}{d t} \delta \theta=0, \quad \frac{d}{d t} \delta z=\delta \theta \tag{6.6}
\end{equation*}
$$

Let $\delta \theta_{k}=\delta \theta\left(t_{k}^{+}\right)$and $\delta z_{k}=\delta z\left(t_{k}^{+}\right)$be the local coordinates immediately after the $k$ th collision, and $\delta \theta_{k}^{-}=\delta \theta\left(t_{k}^{-}\right), \delta z_{k}^{-}=\delta z\left(t_{k}^{-}\right)$immediately before. Integrating the free flight from $t_{k-1}^{+}$to $t_{k}^{-}$we obtain

$$
\begin{align*}
& \delta z_{k}^{-}=\delta z_{k-1}+\tau_{k} \delta \theta_{k-1}, \quad \tau_{k}=t_{k}-t_{k-1} \\
& \delta \theta_{k}^{-}=\delta \theta_{k-1}, \tag{6.7}
\end{align*}
$$

and the stability matrix (4.30) for the $k$ th free flight segment is

$$
\mathbf{J}_{T}\left(x_{k}\right)=\left(\begin{array}{cc}
1 & \tau_{k}  \tag{6.8}\\
0 & 1
\end{array}\right) .
$$

At incidence angle $\phi_{k}$ (the angle between the outgoing particle and the outgoing normal to the billiard edge), the incoming transverse variation $\delta z_{k}^{-}$

Figure 6.4: Defocusing of a beam of nearby trajectories at a billiard collision. (A. Wirzba)

projects onto an arc on the billiard boundary of length $\delta z_{k}^{-} / \cos \phi_{k}$. The corresponding incidence angle variation $\delta \phi_{k}=\delta z_{k}^{-} / \rho_{k} \cos \phi_{k}, \rho_{k}=$ local radius of curvature, increases the angular spread to

$$
\begin{align*}
\delta z_{k} & =-\delta z_{k}^{-} \\
\delta \theta_{k} & =-\delta \theta_{k}^{-}-\frac{2}{\rho_{k} \cos \phi_{k}} \delta z_{k}^{-}, \tag{6.9}
\end{align*}
$$

so the Jacobian matrix associated with the reflection is

$$
\mathbf{J}_{R}\left(x_{k}\right)=-\left(\begin{array}{rr}
1 & 0  \tag{6.10}\\
r_{k} & 1
\end{array}\right), \quad r_{k}=\frac{2}{\rho_{k} \cos \phi_{k}} .
$$

The full Jacobian matrix for $n_{p}$ consecutive bounces describes a beam of trajectories defocused by $\mathbf{J}_{T}$ along the free flight (the $\tau_{k}$ terms below) and defocused/refocused at reflections by $\mathbf{J}_{R}$ (the $r_{k}$ terms below)

$$
\mathbf{J}_{p}=(-1)^{n_{p}} \prod_{k=n_{p}}^{1}\left(\begin{array}{cc}
1 & \tau_{k}  \tag{6.11}\\
0 & 1
\end{array}\right)\left(\begin{array}{cc}
1 & 0 \\
r_{k} & 1
\end{array}\right)
$$

where $\tau_{k}$ is the flight time of the $k$ th free-flight segment of the orbit, $r_{k}=$ $2 / \rho_{k} \cos \phi_{k}$ is the defocusing due to the $k$ th reflection, and $\rho_{k}$ is the radius of curvature of the billiard boundary at the $k$ th scattering point (for our 3 -disk game of pinball, $\rho=1$ ). As the billiard dynamics is phase space volume preserving, $\operatorname{det} \mathbf{J}=1$, and the eigenvalues are given by (5.18).

This is still another example of the Jacobian matrix chain rule (4.35) for discrete time systems, rather similar to the Hénon map stability (4.36). Stability of every flight segment or reflection taken alone is a shear with two unit eigenvalues,

$$
\operatorname{det} \mathbf{J}_{T}=\operatorname{det}\left(\begin{array}{cc}
1 & \tau_{k}  \tag{6.12}\\
0 & 1
\end{array}\right), \quad \operatorname{det} \mathbf{J}_{R}=\operatorname{det}\left(\begin{array}{cc}
1 & 0 \\
r_{k} & 1
\end{array}\right)
$$

but acting in concert in the intervowen sequence (6.11) they can lead to a hyperbolic deformation of the infinitesimal neighborhood of a billiard trajectory.

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As a concrete application, consider the 3-disk pinball system of sect. 1.3. Analytic expressions for the lengths and eigenvalues of $\overline{0}, \overline{1}$ and $\overline{10}$ cycles follow from elementary geometrical considerations. Longer cycles require numerical evaluation by methods such as those described in chapter 17.

## Commentary

Remark 6.1 Billiards. The 3-disk game of pinball is to chaotic dynamics what a pendulum is to integrable systems; the simplest physical example that captures the essence of chaos. Another contender for the title of the "harmonic oscillator of chaos" is the baker's map which is used as the red thread through Ott's introduction to chaotic dynamics [1.7]. The baker's map is the simplest reversible dynamical system which is hyperbolic and has positive entropy. We will not have much use for the baker's map here, as due to its piecewise linearity it is so nongeneric that it misses all of the subtleties of cycle expansions curvature corrections that will be central to this treatise.

That the 3 -disk game of pinball is a quintessential example of deterministic chaos appears to have been first noted by B. Eckhardt [6.1]. The model was studied in depth classically, semiclassically and quantum mechanically by P. Gaspard and S.A. Rice [6.2], and used by P. Cvitanović and B. Eckhardt [6.3] to demonstrate applicability of cycle expansions to quantum mechanical problems. It has been used to study the higher order $\hbar$ corrections to the Gutzwiller quantization by P. Gaspard and D. Alonso Ramirez [6.4], construct semiclassical evolution operators and entire spectral determinants by P. Cvitanović and G. Vattay [6.5], and incorporate the diffraction effects into the periodic orbit theory by G. Vattay, A. Wirzba and P.E. Rosenqvist [6.6]. The full quantum mechanics and semiclassics of scattering systems is developed here in the 3-disk scattering context in chapter 32. Gaspard's monograph [1.4], which we warmly recommend, utilizies the 3 -disk system in much more depth than will be attained here. For further links check ChaosBook.org.

A pinball game does miss a number of important aspects of chaotic dynamics: generic bifurcations in smooth flows, the interplay between regions of stability and regions of chaos, intermittency phenomena, and the renormalization theory of the "border of order" between these regions. To study these we shall have to face up to much harder challenge, dynamics of smooth flows.

Nevertheless, pinball scattering is relevant to smooth potentials. The game of pinball may be thought of as the infinite potential wall limit of a smooth potential, and pinball symbolic dynamics can serve as a covering symbolic dynamics in smooth potentials. One may start with the infinite wall limit and adiabatically relax an unstable cycle onto the corresponding one for the potential under investigation. If things go well, the cycle will remain unstable and isolated, no new orbits (unaccounted for by the pinball symbolic dynamics) will be born, and the lost orbits will be accounted for by a set of pruning rules. The validity of this adiabatic approach has to be checked carefully in each application, as things can easily go wrong; for example, near a bifurcation the same naive symbol string assignments can refer to a whole island of distinct periodic orbits.

Remark 6.2 Further reading. The chapter 1 of Gaspard monograph [1.4] is recommended reading if you are interested in Hamiltonian flows, and billiards in particular. A. Wirzba has generalized the stability analysis of sect. 6.2 to scattering off 3-dimensional spheres (follow the links in ChaosBook.org/extras).

A clear discussion of linear stability for the general $d$-dimensional case is given in Gaspard [1.4], sect. 1.4.

## Résumé

A particulary natural application of the Poincaré section method is the reduction of a billiard flow to a boundary-to-boundary return map.

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## Exercises

Exercise 6.1 A pinball simulator. Implement the disk $\rightarrow$ disk maps to compute a trajectory of a pinball for a given starting point, and a given $R: a=$ (center-to-center distance):(disk radius) ratio for a 3-disk system. As this requires only computation of intersections of lines and circles together with specular reflections, implementation should be within reach of a highschool student. Please start working on this program now; it will be continually expanded in chapters to come, incorporating the Jacobian calculations, Newton root-finding, and so on.

Fast code will use elementary geometry (only one $\sqrt{\cdots}$ per iteration, rest are multiplications) and eschew trigonometric functions. Provide a graphic display of the trajectories and of the Poincaré section iterates. To be able to compare with the numerical results of coming chapters, work with R:a $=6$ and/or 2.5 values. Draw the correct versions of figure 1.8 or figure 11.4 for $R: a=2.5$ and $/$ or 6 .

Exercise 6.2 Trapped orbits. Shoot 100,000 trajectories from one of the disks, and trace out the strips of figure 1.8 for various R:a by color coding the initial points in the Poincaré section by the number of bounces preceeding their escape. Try also $R: a=6: 1$, though that might be too thin and require some magnification. The initial conditions can be randomly chosen, but need not - actually a clearer picture is obtained by systematic scan through regions of interest.

Exercise 6.3 Pinball stability. Add to your exercise 6.1 pinball simulator a routine that computes the the [ $2 \times 2$ ] Jacobian matrix. To be able to compare with the numerical results of coming chapters, work with $R: a=6$ and/or 2.5 values.

Exercise 6.4 Stadium billiard. Consider the Bunimovich stadium [6.8, 6.9] defined in figure 6.1. The Jacobian matrix associated with the reflection is given by (6.10). Here we take $\rho_{k}=-1$ for the semicircle sections of the boundary, and $\cos \phi_{k}$ remains constant for all bounces in a rotation sequence. The time of flight between two semicircle bounces is $\tau_{k}=2 \cos \phi_{k}$. The Jacobian matrix of one semicircle reflection folowed by the flight to the next bounce is

$$
\mathbf{J}=(-1)\left(\begin{array}{cc}
1 & 2 \cos \phi_{k} \\
0 & 1
\end{array}\right)\left(\begin{array}{cc}
1 & 0 \\
-2 / \cos \phi_{k} & 1
\end{array}\right)=(-1)\left(\begin{array}{cc}
-3 & 2 \cos \phi_{k} \\
2 / \cos \phi_{k} & 1
\end{array}\right) .
$$

A shift must always be followed by $k=1,2,3, \cdots$ bounces along a semicircle, hence the natural symbolic dynamics for this problem is $n$-ary, with the corresponding Jacobian matrix given by shear (ie. the eigenvalues remain equal to 1 throughout the whole rotation), and $k$ bounces inside a circle lead to

$$
\mathbf{J}^{k}=(-1)^{k}\left(\begin{array}{cc}
-2 k-1 & 2 k \cos \phi  \tag{6.13}\\
2 k / \cos \phi & 2 k-1
\end{array}\right) .
$$

The Jacobian matrix of a cycle $p$ of length $n_{p}$ is given by

$$
\mathbf{J}_{p}=(-1)^{\sum n_{k}} \prod_{k=1}^{n_{p}}\left(\begin{array}{cc}
1 & \tau_{k}  \tag{6.14}\\
0 & 1
\end{array}\right)\left(\begin{array}{cc}
1 & 0 \\
n_{k} r_{k} & 1
\end{array}\right) .
$$

Adopt your pinball simulator to the Bunimovich stadium.

Exercise 6.5 A test of your pinball simulator. Test your exercise 6.3 pinball simulator by computing numerically cycle stabilities by tracking distances to nearby orbits. Compare your result with the exact analytic formulas of exercise 8.1 and 8.2.

Exercise 6.6 Birkhoff coordinates. Prove that the Birkhoff coordinates are phase-space volume preserving. Hint: compute the determinant of (6.11).

## Chapter 7

## Get straight

A Hamiltonian system is said to be "integrable" if one can find a change of coordinates to an action-angle coordinate frame where the phase space dynamics is described by motion on circles, one circle for each degree of freedom. In the same spirit, a natural description of a hyperbolic, unstable flow would be attained if one found a change of coordinates into a frame where the stable/unstable manifolds are straight lines, and the flow is along hyperbolas. Achieving this globally for anything but a handful of contrived examples is too much to hope for. Still, as we shall now show, we can make some headway on straightening out the flow locally.

Even though such nonlinear coordinate transformations are very important, especially in celestial mechanics, we shall not necessarily use them much in what follows, so you can safely skip this chapter on the first reading. Except, perhaps, you might want to convince yourself that cycle stabilities are indeed metric invariants of flows (sect. 8.5), and you might like transformations that turn a Keplerian ellipse into a harmonic oscillator (example 7.2) and regularize the 2-body Coulomb collisions (sect. 7.3) in classical helium.

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chapter 9, p. 123

### 7.1 Changing coordinates

Problems are handed down to us in many shapes and forms, and they are not always expressed in the most convenient way. In order to simplify a given problem, one may stretch, rotate, bend and mix the coordinates, but in doing so, the vector field will also change. The vector field lives in a (hyper)plane tangent to phase space and changing the coordinates of phase space affects the coordinates of the tangent space as well, in a way that we will now describe.

Denote by $h$ the conjugation function which maps the coordinates of the
initial phase space $\mathcal{M}$ into the reparametrized phase space $\mathcal{M}^{\prime}=h(\mathcal{M})$, with a point $x \in \mathcal{M}$ related to a point $y \in \mathcal{M}^{\prime}$ by $y=h(x)$. The change of coordinates must be one-to-one and span both $\mathcal{M}$ and $\mathcal{M}^{\prime}$, so given any point $y$ we can go back to $x=h^{-1}(y)$. For smooth flows the reparametrized dynamics should support the same number of derivatives as the initial one. If $h$ is a (piecewise) analytic function, we refer to $h$ as a smooth conjugacy.

The evolution rule $g^{t}\left(y_{0}\right)$ on $\mathcal{M}^{\prime}$ can be computed from the evolution rule $f^{t}\left(x_{0}\right)$ on $\mathcal{M}$ by taking the initial point $y_{0} \in \mathcal{M}^{\prime}$, going back to $\mathcal{M}$, evolving, and then mapping the final point $x(t)$ back to $\mathcal{M}^{\prime}$ :

$$
\begin{equation*}
y(t)=g^{t}\left(y_{0}\right)=h \circ f^{t} \circ h^{-1}\left(y_{0}\right) . \tag{7.1}
\end{equation*}
$$

Here "०" stands for functional composition $h \circ f(x)=h(f(x))$, so (7.1) is a shorthand for $y(t)=h\left(f^{t}\left(h^{-1}\left(y_{0}\right)\right)\right)$.

The vector field $\dot{x}=v(x)$ in $\mathcal{M}$, locally tangent to the flow $f^{t}$, is related to the flow by differentiation (2.4) along the trajectory. The vector field $\dot{y}=w(y)$ in $\mathcal{M}^{\prime}$, locally tangent to $g^{t}$ follows by the chain rule:

$$
\begin{align*}
w(y) & =\left.\frac{d g^{t}}{d t}(y)\right|_{t=0}=\left.\frac{d}{d t}\left(h \circ f^{t} \circ h^{-1}(y)\right)\right|_{t=0} \\
& =h^{\prime}\left(h^{-1}(y)\right) v\left(h^{-1}(y)\right)=h^{\prime}(x) v(x) \tag{7.2}
\end{align*}
$$

With the indices reinstated, this stands for

$$
\begin{equation*}
w_{i}(y)=\frac{\partial h_{i}(x)}{\partial x_{j}} v_{j}(x), \quad y_{i}=h_{i}(x) . \tag{7.3}
\end{equation*}
$$

Imagine that the phase space is a rubber sheet with the flow lines drawn on it. A coordinate change $h$ corresponds to pulling and tugging on the rubber sheet smoothly, without cutting, glueing, or self-intersections of the distorted rubber sheet. Trajectories that are closed loops in $\mathcal{M}$ will remain closed loops in the new manifold $\mathcal{M}^{\prime}$, but their shapes will change. Globally $h$ deforms the rubber sheet in a highly nonlinear manner, but locally it simply rescales and shears the tangent field by $\partial_{j} h_{i}$, hence the simple transformation law (7.2) for the velocity fields.

The time itself is a parametrization of points along flow lines, and it can also be reparametrized, $s=s(t)$, with the attendent modification of (7.2). An example is the 2-body collision regularization of the helium Hamiltonian (5.6), to be undertaken in sect. 7.3 below.

### 7.2 Rectification of flows

A profitable way to exploit invariance of dynamics under smooth conjugacies is to use it to pick out the simplest possible representative of an
equivalence class. In general and globally these are just words, as we have no clue how to pick such "canonical" representative, but for smooth flows we can always do it localy and for sufficiently short time, by appealing to the rectification theorem, a fundamental theorem of ordinary differential equations. The theorem assures us that there exists a solution (at least for a short time interval) and what the solution looks like. The rectification theorem holds in the neighborhood of points of the vector field $v(x)$ that are not singular, that is, everywhere except for the equilibrium points (2.7), and points at which $v$ is infinite. According to the theorem, in a small neighborhood of a non-singular point there exists a change of coordinates $y=h(x)$ such that $\dot{x}=v(x)$ in the new coordinates takes the standard form

$$
\begin{align*}
& \dot{y}_{1}=1 \\
& \dot{y}_{2}=\dot{y}_{3}=\cdots=\dot{y}_{d}=0, \tag{7.4}
\end{align*}
$$

with unit velocity flow along $y_{1}$, and no flow along any of the remaining directions.

Example 7.1 Harmonic oscillator, rectified: As a simple example of global rectification of a flow consider the harmonic oscillator

$$
\begin{equation*}
\dot{q}=p, \quad \dot{p}=-q . \tag{7.5}
\end{equation*}
$$

The trajectories $x(t)=(q(t), p(t))$ just go around the origin, so a fair guess is that the system would have a simpler representation in polar coordinates $y=(r, \theta)$ :

$$
h^{-1}:\left\{\begin{array}{l}
q=h_{1}^{-1}(r, \theta)=r \cos \theta  \tag{7.6}\\
p=h_{2}^{-1}(r, \theta)=r \sin \theta
\end{array}\right. \text {. }
$$

The Jacobian matrix of the transformation is

$$
h^{\prime}=\left[\begin{array}{cc}
\cos \theta & \sin \theta  \tag{7.7}\\
-\frac{\sin \theta}{r} & -\frac{\cos \theta}{r}
\end{array}\right]
$$

resulting in (7.2) of rectified form

$$
\begin{equation*}
\dot{r}=0, \quad \dot{\theta}=-1 . \tag{7.8}
\end{equation*}
$$

In the new coordinates the radial coordinate $r$ is constant, and the angular coordinate $\theta$ wraps around a cylinder with constant angular velocity. There is a subtle point in this change of coordinates: the domain of the map $h^{-1}$ is not the plane $\mathbb{R}^{2}$, but rather the plane minus the origin. We had mapped a plane into a cylinder, and coordinate transformations should not change the topology of the space in which the dynamics takes place; the coordinate transformation is not defined on the equilibrium point $x=(0,0)$, or $r=0$.

Figure 7.1: Coordinates for the helium three body problem in the plane.


### 7.3 Classical dynamics of collinear helium

(G. Tanner)

So far much has been said about 1-dimensional maps, game of pinball and other curious but rather idealized dynamical systems. If you have become impatient and started wondering what good are the methods learned so far in solving real life physical problems, good news are here. We will apply here concepts of symbolic dynamics to nothing less than the helium, a dreaded three-body Coulomb problem.

Can we really jump from three static disks directly to three charged particles moving under the influence of their mutually attracting or repelling forces? It turns out, we can, but we have to do it with care. The full problem is indeed not accessible in all its detail, but we are able to analyze a somewhat simpler subsystem - collinear helium. This system plays an important role in the classical dynamics of the full three-body problem.

The full classical helium system consists of two electrons of mass $m_{e}$ and charge $-e$ moving about a positively charged nucleus of mass $m_{h e}$ and charge $+2 e$.

The helium electron-nucleus mass ratio $m_{h e} / m_{e}=1836$ is so large that we may work in the infinite nucleus mass approximation $m_{h e}=\infty$, fixing the nucleus at the origin. Finite nucleus mass effects can be taken into account without any substantial difficulty. We are now left with two electrons moving in three spatial dimensions around the origin. The total angular momentum of the combined electron system is still conserved. In the special case of angular momentum $L=0$, the electrons move in a fixed plane containing the nucleus. The three body problem can then be written in terms of three independent coordinates only, the electron-nucleus distances $r_{1}$ and $r_{2}$ and the inter-electron angle $\Theta$, see figure 7.1.

This looks like something we can lay our hands on; the problem has been reduced to three degrees of freedom, six phase space coordinates in all, and the total energy is conserved. But let us go one step further; the electrons are attracted by the nucleus but repelled by each other. They will tend to stay as far away from each other as possible, preferably on opposite sides of the nucleus. It is thus worth having a closer look at the situation where the three particles are all on a line with the nucleus being somewhere between the two electrons. If we, in addition, let the electrons

Figure 7.2: Collinear helium, with the two electrons on opposite sides of the nucleus.

have momenta pointing towards the nucleus as in figure 7.2 , then there is no force acting on the electrons perpendicular to the common interparticle axis. That is, if we start the classical system on the dynamical subspace $\Theta=\pi$, $\frac{d}{d t} \Theta=0$, the three particles will remain in this collinear configuration for all times.

### 7.3.1 Scaling

In what follows we will restrict the dynamics to this collinear subspace. It is a system of two degrees of freedom with the Hamiltonian

$$
\begin{equation*}
H=\frac{1}{2 m_{e}}\left(p_{1}^{2}+p_{2}^{2}\right)-\frac{2 e^{2}}{r_{1}}-\frac{2 e^{2}}{r_{2}}+\frac{e^{2}}{r_{1}+r_{2}}=E, \tag{7.9}
\end{equation*}
$$

where $E$ is the total energy. As the dynamics is restricted to the fixed energy shell, the four phase space coordinates are not independent; the energy shell dependence can be made explicit by writing $\left(r_{1}, r_{2}, p_{1}, p_{2}\right) \rightarrow$ $\left(r_{1}(E), r_{2}(E), p_{1}(E), p_{2}(E)\right)$.

We will first consider the dependence of the dynamics on the energy $E$. A simple analysis of potential versus kinetic energy tells us that if the energy is positive both electrons can escape to $r_{i} \rightarrow \infty, i=1,2$. More interestingly, a single electron can still escape even if $E$ is negative, carrying away an unlimited amount of kinetic energy, as the total energy of the remaining inner electron has no lower bound. Not only that, but one electron will escape eventually for almost all starting conditions. The overall dynamics thus depends critically on whether $E>0$ or $E<0$. But how does the dynamics change otherwise with varying energy? Fortunately, not at all. Helium dynamics remains invariant under a change of energy up to a simple scaling transformation; a solution of the equations of motion at a fixed energy $E_{0}=-1$ can be transformed into a solution at an arbitrary energy $E<0$ by scaling the coordinates as

$$
r_{i}(E)=\frac{e^{2}}{(-E)} r_{i}, \quad p_{i}(E)=\sqrt{-m_{e} E} p_{i}, \quad i=1,2,
$$

together with a time transformation $t(E)=e^{2} m_{e}^{1 / 2}(-E)^{-3 / 2} t$. We include the electron mass and charge in the scaling transformation in order to obtain a non-dimensionalized Hamiltonian of the form

$$
\begin{equation*}
H=\frac{p_{1}^{2}}{2}+\frac{p_{2}^{2}}{2}-\frac{2}{r_{1}}-\frac{2}{r_{2}}+\frac{1}{r_{1}+r_{2}}=-1 . \tag{7.10}
\end{equation*}
$$

The case of negative energies chosen here is the most interesting one for us. It exhibits chaos, unstable periodic orbits and is responsible for the bound states and resonances of the quantum problem treated in sect. 34.5.

### 7.3.2 Regularization of two-body collisions

Next, we have a closer look at the singularities in the Hamiltonian (7.10). Whenever two bodies come close to each other, accelerations become large, numerical routines require lots of small steps, and numerical precision suffers. No numerical routine will get us through the singularity itself, and in collinear helium electrons have no option but to collide with the nucleus. Hence a regularization of the differential equations of motions is a necessary prerequisite to any numerical work on such problems, both in celestial mechanics (where a spaceship executes close approaches both at the start and its destiantion) and in quantum mechanics (where much of semiclassical physics is dominated by returning classical orbits that probe the quantum wave function at the nucleus).

There is a fundamental difference between two-body collisions $r_{1}=0$ or $r_{2}=0$, and the triple collision $r_{1}=r_{2}=0$. Two-body collisions can be regularized, with the singularities in equations of motion removed by a suitable coordinate transformation together with a time transformation preserving the Hamiltonian structure of the equations. Such regularization is not possible for the triple collision, and solutions of the differential equations can not be continued through the singularity at the origin. As we shall see, the chaos in collinear helium originates from this singularity of triple collisions.

A regularization of the two-body collisions is achieved by means of the Kustaanheimo-Stiefel (KS) transformation, which consists of a coordinate dependent time transformation which stretches the time scale near the origin, and a canonical transformation of the phase space coordinates. In order to motivate the method, we apply it first to the 1-dimensional Kepler problem

$$
\begin{equation*}
H=\frac{1}{2} p^{2}-\frac{2}{x}=E \tag{7.11}
\end{equation*}
$$

Example 7.2 Keplerian ellipse, rectified: To warm up, consider the $E=0$ case, starting at $x=0$ at $t=0$. Even though the equations of motion are singular at the intial point, we can immediately integrate

$$
\frac{1}{2} \dot{x}^{2}-\frac{2}{x}=0
$$

by means of separation of variables

$$
\begin{equation*}
\sqrt{x} d x=\sqrt{2} d t, \quad x=(3 t)^{\frac{2}{3}}, \tag{7.12}
\end{equation*}
$$

and observe that the solution is not singular. The aim of regularization is to compensate for the infinite acceleration at the origin by introducing a fictitious time, in terms of which the passage through the origin is smooth.

A time transformation $d t=f(q, p) d \tau$ for a system described by a Hamiltonian $H(q, p)=E$ leaves the Hamiltonian structure of the equations of motion unaltered, if the Hamiltonian itself is transformed into $\mathcal{H}(q, p)=f(q, p)(H(q, p)-E)$. For the 1dimensional Coulomb problem with (7.11) we choose the time transformation $d t=x d \tau$ which lifts the $|x| \rightarrow 0$ singularity in (7.11) and leads to a new Hamiltonian

$$
\begin{equation*}
\mathcal{H}=\frac{1}{2} x p^{2}-2-E x=0 \tag{7.13}
\end{equation*}
$$

The solution (7.12) is now parametrized by the fictitous time $d \tau$ through a pair of equations

$$
x=\tau^{2}, \quad t=\frac{1}{3} \tau^{3}
$$

The equations of motion are, however, still singular as $x \rightarrow 0$ :

$$
\frac{d^{2} x}{d \tau^{2}}=-\frac{1}{2 x} \frac{d x}{d \tau}+x E
$$

Appearance of the square root in (7.12) now suggests a canonical transformation of form

$$
\begin{equation*}
x=Q^{2}, \quad p=\frac{P}{2 Q} \tag{7.14}
\end{equation*}
$$

which maps the Kepler problem into that of a harmonic oscillator with Hamiltonian

$$
\begin{equation*}
H(Q, P)=\frac{1}{8} P^{2}-E Q^{2}=2 \tag{7.15}
\end{equation*}
$$

with all singularities completely removed.

We now apply this method to collinear helium. The basic idea is that one seeks a higher-dimensional generalization of the "square root removal" trick (7.14), by introducing a new vector $Q$ with property $r=|Q|^{2}$. In this simple 1-dimensional example the KS transformation can be implemented by

$$
\begin{equation*}
r_{1}=Q_{1}^{2}, \quad r_{2}=Q_{2}^{2}, \quad p_{1}=\frac{P_{1}}{2 Q_{1}}, \quad p_{2}=\frac{P_{2}}{2 Q_{2}} \tag{7.16}
\end{equation*}
$$

and reparametrization of time by $d \tau=d t / r_{1} r_{2}$. The singular behavior in the original momenta at $r_{1}$ or $r_{2}=0$ is again compensated by stretching the time scale at these points. The Hamiltonian structure of the equations of motions with respect to the new time $\tau$ is conserved, if we consider the Hamiltonian

$$
\begin{equation*}
H_{k o}=\frac{1}{8}\left(Q_{2}^{2} P_{1}^{2}+Q_{1}^{2} P_{2}^{2}\right)-2 R_{12}^{2}+Q_{1}^{2} Q_{2}^{2}\left(-E+1 / R_{12}^{2}\right)=0 \tag{7.17}
\end{equation*}
$$



Figure 7.3: (a) A typical trajectory in the $r_{1}-r_{2}$ plane; the trajectory enters here along the $r_{1}$ axis and escapes to infinity along the $r_{2}$ axis; (b) Poincaré map ( $r_{2}=0$ ) for collinear helium. Strong chaos prevails for small $r_{1}$ near the nucleus.
with $R_{12}=\left(Q_{1}^{2}+Q_{2}^{2}\right)^{1 / 2}$, and we will take $E=-1$ in what follows. The equations of motion now have the form

$$
\begin{array}{ll}
\dot{P}_{1}=2 Q_{1}\left[2-\frac{P_{2}^{2}}{8}-Q_{2}^{2}\left(1+\frac{Q_{2}^{2}}{R_{12}^{4}}\right)\right] ; & \dot{Q}_{1}=\frac{1}{4} P_{1} Q_{2}^{2}  \tag{7.18}\\
\dot{P}_{2}=2 Q_{2}\left[2-\frac{P_{1}^{2}}{8}-Q_{1}^{2}\left(1+\frac{Q_{1}^{2}}{R_{12}^{4}}\right)\right] ; & \dot{Q}_{2}=\frac{1}{4} P_{2} Q_{1}^{2} .
\end{array}
$$

Individual electron-nucleus collisions at $r_{1}=Q_{1}^{2}=0$ or $r_{2}=Q_{2}^{2}=0$ no longer pose a problem to a numerical integration routine. The equations (7.18) are singular only at the triple collision $R_{12}=0$, that is, when both electrons hit the nucleus at the same time.

The new coordinates and the Hamiltonian (7.17) are very useful when calculating trajectories for collinear helium; they are, however, less intuitive as a visualization of the three-body dynamics. We will therefore refer to the old coordinates $r_{1}, r_{2}$ when discussing the dynamics and the periodic orbits.

To summarize, we have brought a 3-body problem into a form where the 2 -body collisions have been transformed away, and the phase space trajectories computable numerically. To appreciate the full beauty of what has been attained, you have to fast-forward to chapter 34; we are already "almost" ready to quantize helium by semiclassical methods.

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chapter 8, p. 111

### 7.4 Rectification of maps



In sect. 7.2 we had argued that nonlinear coordinate transformations can be profitably employed to simplify the representation of a flow. We
shall now apply the same idea to nonlinear maps, and determine a smooth nonlinear change of coordinates that flattens out the vicinity of a fixed point and makes the map linear in an open neighborhood. In its simplest form the idea can be implemented only for an isolated nondegenerate fixed point (otherwise are needed in the normal form expansion around the point), and only in a finite neigborhood of a point, as the conjugating function in general has a finite radius of convergence. In sect. 8.4 we will extend the method to periodic orbits.

### 7.4.1 Rectification of a fixed point in one dimension

Consider a 1-dimensional map $x_{n+1}=f\left(x_{n}\right)$ with a fixed point at $x=0$, with stability $\Lambda=f^{\prime}(0)$. If $|\Lambda| \neq 1$, one can determine term-by-term the power series for a smooth conjugation $h(x)$ centered at the fixed point, $h(0)=0$, that flattens out the neighborhood of the fixed point

$$
\begin{equation*}
f(x)=h^{-1}(\Lambda h(x)) \tag{7.19}
\end{equation*}
$$

and replaces the nonlinear map $f(x)$ by a linear map $y_{n+1}=\Lambda y_{n}$.
To compute the conjugation $h$ we use the functional equation $h^{-1}(\Lambda x)=$ $f\left(h^{-1}(x)\right)$ and the expansions

$$
\begin{align*}
f(x) & =\Lambda x+x^{2} f_{2}+x^{3} f_{3}+\ldots \\
h^{-1}(x) & =x+x^{2} h_{2}+x^{3} h_{3}+\ldots \tag{7.20}
\end{align*}
$$

If $h(x)$ is a conjugation, so is any scaling $h(b x)$ of the function for a real number $b$. Hence the value of $h^{\prime}(0)$ is not determined by the functional equation (7.19); it is convenient to set $h^{\prime}(0)=1$.

The algebra is not particularly illuminating and best left to computers. In any case, for the time being we will not use much beyond the first, linear term in these expansions.

Here we assume $\Lambda \neq 1$. If the fixed point has first $k-1$ derivatives vanishing, the conjugacy is to the $k$ th normal form.

In several dimensions, $\Lambda$ is replaced by the Jacobian matrix, and one has to check that the eigenvalues $\mathbf{J}$ are non-resonant, that is, there is no integer linear relation between the stability exponents (8.4).

## Commentary

Remark 7.1 Rectification of maps. The methods outlined above are standard in the analysis of fixed points and construction of normal forms for bifurcations,
see for example ref. [1.14, 7.2, 7.4, 7.5, 7.6, 7.7, 7.8, 7.9, 3.6]. The geometry underlying such methods is pretty, and we enjoyed reading, for example, Percival and Richards [7.10], chaps. 2 and 4 of Ozorio de Almeida's monograph [7.11], and, as always, Arnol'd [7.1].

Recursive formulas for evaluation of derivatives needed to evaluate (7.20) are given, for example, in Appendix A of ref. [9.5]).

## Résumé

Dynamics $(\mathcal{M}, f)$ is invariant under the group of all smooth conjugacies

$$
(\mathcal{M}, f) \rightarrow\left(\mathcal{M}^{\prime}, g\right)=\left(h(\mathcal{M}), h \circ f \circ h^{-1}\right) .
$$

This invariance can be used to (i) find a simplified representation for the flow and (ii) identify a set of invariants, numbers computed within a particular choice of $(\mathcal{M}, f)$, but invariant under all $\mathcal{M} \rightarrow h(\mathcal{M})$ smooth conjugacies.

The $2 D$-dimensional phase space of an integrable Hamiltonian system of $D$ degrees of freedom is fully foliated by $D$-tori. In the same spirit, for a uniformly hyperbolic, chaotic dynamical system one would like to change into a coordinate frame where the stable/unstable manifolds form a set of transversally interesecting hyper-planes, with the flow everywhere locally hyperbolic. That cannot be achieved in general: Fully globally integrable and fully globally chaotic flows are a very small subset of all possible flows.

What we really care about is developping invariant notions of what a given dynamical system is. The totality of smooth one-to-one nonlinear coordinate transformations $h$ which map all trajectories of a given dynamical system $\left(\mathcal{M}, f^{t}\right)$ onto all trajectories of dynamical systems $\left(\mathcal{M}^{\prime}, g^{t}\right)$ gives us a huge equivalence class, much larger than the equivalence classes familiar from the theory of linear transformations, such as the rotation group $O(d)$ or the Galilean group of all rotations and translations in $\mathbb{R}^{d}$. In the theory of Lie groups, the full invariant specification of an object is given by a finite set of Casimir invariants. What a good full set of invariants for a group of general nonlinear smooth conjugacies might be is not known, but the set of all periodic orbits and their stability eigenvalues will turn out to be a good start.

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## Exercises

Exercise 7.1 Coordinate transformations. Changing coordinates is conceptually simple, but can become confusing when carried out in detail. The difficulty arises from confusing functional relationships, such as $x(t)=h^{-1}(y(t))$ with numerical relationships, such as $w(y)=h^{\prime}(x) v(x)$. Working through an example will clear this up.
(a) The differential equation in the $\mathcal{M}$ space is $\dot{x}=\left\{2 x_{1}, x_{2}\right\}$ and the change of coordinates from $\mathcal{M}$ to $\mathcal{M}^{\prime}$ is $h\left(x_{1}, x_{2}\right)=\left\{2 x_{1}+x_{2}, x_{1}-x_{2}\right\}$. Solve for $x(t)$. Find $h^{-1}$.
(b) Show that in the transformed space $\mathcal{M}^{\prime}$, the differential equation is

$$
\frac{d}{d t}\left[\begin{array}{l}
y_{1}  \tag{7.21}\\
y_{2}
\end{array}\right]=\frac{1}{3}\left[\begin{array}{c}
5 y_{1}+2 y_{2} \\
y_{1}+4 y_{2}
\end{array}\right] .
$$

Solve this system. Does it match the solution in the $\mathcal{M}$ space?

Exercise 7.2 Linearization for maps. Let $f: C \rightarrow C$ be a map from the complex numbers into themselves, with a fixed point at the origin and analytic there. By manipulating power series, find the first few terms of the map $h$ that conjugates $f$ to $\alpha z$, that is,

$$
f(z)=h^{-1}(\alpha h(z)) .
$$

There are conditions on the derivative of $f$ at the origin to assure that the conjugation is always possible. Can you formulate these conditions by examining the series?
(difficulty: medium)

## Chapter 8

## Cycle stability

Topological features of a dynamical system - singularities, periodic orbits, and the ways in which the orbits intertwine - are invariant under a general continuous change of coordinates. Surprisingly, there exist quantities that depend on the notion of metric distance between points, but nevertheless do not change value under a smooth change of coordinates. Local quantities such as the stability eigenvalues of equilibria and periodic orbits, and global quantities such as Lyapunov exponents, metric entropy, and fractal dimensions are examples of properties of dynamical systems independent of coordinate choice.

We now turn to the first, local class of such invariants, linear stability of periodic orbits of flows and maps. This will give us metric information about local dynamics. The material covered here is aimed at unbelieving Thomases - if you already know that the stability eigenvalues of periodic orbits are invariants of a flow, you can skip this chapter.


### 8.1 Stability of periodic orbits

As noted on page 33, a trajectory can be stationary, periodic or aperiodic. For chaotic systems almost all trajectories are aperiodic - nevertheless, stationary and periodic orbits will turn out to be the key to unraveling chaotic dynamics. Here we note a few of the properties that makes them so precious to a theorist.

An obvious virtue of periodic orbits is that they are topological invariants: a fixed point remains a fixed point for any choice of coordinates, and similarly a periodic orbit remains periodic in any representation of the dynamics. Any re-parametrization of a dynamical system that preserves its topology has to preserve topological relations between periodic orbits, such as their relative inter-windings and knots. So the mere existence of periodic

Figure 8.1: For a prime cycle $p$, Jacobian matrix $\mathbf{J}_{p}$ returns an infinitesimal spherical neighborhood of $x_{0} \in p$ stretched into an ellipsoid, with overlap ratio along the eigenvector $\mathbf{e}_{i}$ of $\mathbf{J}_{p}(x)$ given by the eigenvalue $\Lambda_{p, i}$. These ratios are invariant under smooth nonlinear reparametrizations of phase space coordinates, and are intrinsic property of cycle $p$.

orbits suffices to partially organize the spatial layout of a non-wandering set. No less important, as we shall now show, is the fact that cycle stability eigenvalues are metric invariants: they determine the relative sizes of neighborhoods in a non-wandering set.

To prove this, we start by noting that due to the multiplicative structure (4.31) of Jacobian matrices, the stability of the $r$ th repeat of a prime cycle $p$ of period $T_{p}$ is

$$
\begin{equation*}
\mathbf{J}^{r T_{p}}(x)=\mathbf{J}^{T_{p}}\left(f^{(r-1) T_{p}}(x)\right) \cdots \mathbf{J}^{T_{p}}\left(f^{T_{p}}(x)\right) \mathbf{J}^{T_{p}}(x)=\left(\mathbf{J}_{p}(x)\right)^{r} \tag{8.1}
\end{equation*}
$$

where $\mathbf{J}_{p}(x)=\mathbf{J}^{T_{p}}(x)$ is the stability matrix for a single traversal of the prime cycle $p, x \in p$ is any point on the cycle, and $f^{r T_{p}}(x)=x$ as $f^{t}(x)$ returns to $x$ every multiple of the period $T_{p}$. Hence, it suffices to restrict our considerations to the stability of prime cycles.

### 8.1.1 Stability eigenvalues and exponents

We sort the stability eigenvalues $\Lambda_{p, 1}, \Lambda_{p, 2}, \ldots, \Lambda_{p, d}$ of the $[d \times d]$ Jacobian matrix $\mathbf{J}_{p}$ evaluated on the $p$-cycle into sets $\{e, m, c\}$

$$
\begin{array}{rll}
\text { expanding: } & \left\{\Lambda_{p}\right\}_{e}=\left\{\Lambda_{p, j}:\left|\Lambda_{p, j}\right|>1\right\} \\
\text { marginal: } & \left\{\Lambda_{p}\right\}_{m}=\left\{\Lambda_{p, j}:\left|\Lambda_{p, j}\right|=1\right\}  \tag{8.2}\\
\text { contracting: } & \left\{\Lambda_{p}\right\}_{c}=\left\{\Lambda_{p, j}:\left|\Lambda_{p, j}\right|<1\right\}
\end{array}
$$

and denote by $\Lambda_{p}$ (no $j$ th eigenvalue index) the product of expanding eigenvalues

$$
\begin{equation*}
\Lambda_{p}=\prod_{e} \Lambda_{p, e} \tag{8.3}
\end{equation*}
$$

As $\mathbf{J}_{p}$ is a real matrix, complex eigenvalues always come in complex conjugate pairs, $\Lambda_{p, i+1}=\Lambda_{p, i}^{*}$, so the product of expanding eigenvalues $\Lambda_{p}$ is always real.

Cycle stability exponents are the stretching/contraction rates per unit time

$$
\begin{equation*}
\lambda_{p, i}=\frac{1}{T_{p}} \ln \left|\Lambda_{p, i}\right| \tag{8.4}
\end{equation*}
$$

This definition is motivated by the form of the stability exponents for the linear dynamical systems, for example (4.14), as well as the fact that exponents so defined can be interpreted as Lyapunov exponents (10.32) evaluated on the prime cycle $p$. As in the three cases of (8.2), we sort the stability exponents into three sets

$$
\begin{array}{rll}
\text { expanding: } & \left\{\lambda_{p}\right\}_{e}=\left\{\lambda_{p, i}: \lambda_{p, i}>0\right\} \\
\text { marginal: } & \left\{\lambda_{p}\right\}_{m}=\left\{\lambda_{p, i}: \lambda_{p, i}=0\right\} \\
\text { contracting: } & \left\{\lambda_{p}\right\}_{c}=\left\{\lambda_{p, i}: \lambda_{p, i}<0\right\} . \tag{8.5}
\end{array}
$$

A periodic orbit $p$ of a $d$-dimensional flow or a map is stable if all its stability exponents (other than the vanishing longitudinal stability exponent, to be explained in sect. 8.2.1 below) are strictly negative, $\left|\lambda_{p, i}\right|<0$. The region of system parameter values for which a periodic orbit $p$ is stable is called the stability window of $p$. The set $\mathcal{M}_{p}$ of initial points that are asymptotically attracted to $p$ as $t \rightarrow+\infty$ (for a fixed set of system parameter values) is called the basin of attraction of $p$.

If all stability exponents (other than the vanishing longitudinal stability exponent) of all periodic orbits of a flow are strictly bounded away from zero, $\left|\lambda_{i}\right| \geq \lambda_{\text {min }}>0$, the flow is said to be hyperbolic. Otherwise the flow is said to be nonhyperbolic. In particular, if all $\left|\lambda_{i}\right|=0$, the orbit is said to be elliptic. Such orbits proliferate in Hamiltonian flows.

We often do care about the sign of $\Lambda_{p, i}$ and, if $\Lambda_{p, i}$ is complex, its phase

$$
\begin{equation*}
\Lambda_{p, j}= \pm e^{T_{p}\left(\lambda_{p, j} \pm i \theta_{p, j}\right)} . \tag{8.6}
\end{equation*}
$$

[188 sect. 5.2
Keeping track of this by case-by-case enumeration is a self-inflicted, unnecessary nuisance, followed in much of the literature. To avoid this, almost all of our formulas will be stated in terms of the stability eigenvalues $\Lambda_{j}$ rather than in the terms of the overall signs, stability exponents $\lambda_{j}$ and phases $\theta_{j}$.

Example 8.1 1-dimensional maps: The simplest example of cycle stability is afforded by 1-dimensional maps. The stability of a prime cycle $p$ follows from the chain rule (4.34) for stability of the $n_{p}$ th iterate of the map

$$
\begin{equation*}
\Lambda_{p}=\frac{d}{d x_{0}} f^{n_{p}}\left(x_{0}\right)=\prod_{m=0}^{n_{p}-1} f^{\prime}\left(x_{m}\right), \quad x_{m}=f^{m}\left(x_{0}\right) . \tag{8.7}
\end{equation*}
$$

$\Lambda_{p}$ is a property of the cycle, not the initial point, as taking any periodic point in the $p$ cycle as the initial point yields the same result.

A critical point $x_{c}$ is a value of $x$ for which the mapping $f(x)$ has vanishing derivative, $f^{\prime}\left(x_{c}\right)=0$. For future reference we note that a periodic orbit of a 1dimensional map is stable if

$$
\left|\Lambda_{p}\right|=\left|f^{\prime}\left(x_{n_{p}}\right) f^{\prime}\left(x_{n_{p}-1}\right) \cdots f^{\prime}\left(x_{2}\right) f^{\prime}\left(x_{1}\right)\right|<1,
$$

and superstable if the orbit includes a critical point, so that the above product vanishes. For a stable periodic orbit of period $n$ the slope of the nth iterate $f^{n}(x)$ evaluated on a periodic point $x$ (fixed point of the $n$th iterate) lies between -1 and 1. If $\left|\Lambda_{p}\right|>1$, p-cycle is unstable.

Example 8.2 Stability of cycles for maps: No matter what method we had used to determine the unstable cycles, the theory to be developed here requires that their stability eigenvalues be evaluated as well. For maps a Jacobian matrix is easily evaluated by picking any cycle point as a starting point, running once around a prime cycle, and multiplying the individual cycle point stability matrices according to (4.35). For example, the Jacobian matrix $\mathbf{J}_{p}$ for a Hénon map (3.12) prime cycle $p$ of length $n_{p}$ is given by (4.36),

$$
\mathbf{J}_{p}\left(x_{0}\right)=\prod_{m=n}^{1}\left(\begin{array}{cc}
-2 a x_{m} & b \\
1 & 0
\end{array}\right), \quad x_{m} \in p
$$

and the Jacobian matrix $\mathbf{J}_{p}$ for a 2-dimensional billiard prime cycle $p$ of length $n_{p}$

$$
\mathbf{J}_{p}=(-1)^{n_{p}} \prod_{k=n_{p}}^{1}\left(\begin{array}{cc}
1 & \tau_{k} \\
0 & 1
\end{array}\right)\left(\begin{array}{cc}
1 & 0 \\
r_{k} & 1
\end{array}\right)
$$

follows from (6.11).

### 8.2 Cycle stabilities are cycle invariants

The 1-dimensional map cycle stability $\Lambda_{p}$ is a product of derivatives over all points around the cycle, and is therefore independent of which periodic point is chosen as the initial one. In higher dimensions the form of the Jacobian matrix $\mathbf{J}_{p}\left(x_{0}\right)$ in (8.1) does depend on the choice of coordinates and the initial point $x_{0} \in p$. Nevertheless, as we shall now show, the cycle stability eigenvalues are intrinsic property of a cycle also for multidimensional flows. Consider the $i$ th eigenvalue, eigenvector pair $\left(\Lambda_{p, i}, \mathbf{e}_{i}\right)$ computed from $\mathbf{J}_{p}$ evaluated at a cycle point,

$$
\begin{equation*}
\mathbf{J}_{p}(x) \mathbf{e}_{i}(x)=\Lambda_{p, i} \mathbf{e}_{i}(x), \quad x \in p \tag{8.8}
\end{equation*}
$$

Consider another point on the cycle at time $t$ later, $x^{\prime}=f^{t}(x)$ whose Jacobian matrix is $\mathbf{J}_{p}\left(x^{\prime}\right)$. By the group property (4.31), $\mathbf{J}^{T_{p}+t}=\mathbf{J}^{t+T_{p}}$, and the Jacobian matrix at $x^{\prime}$ can be written either as

$$
\mathbf{J}^{T_{p}+t}(x)=\mathbf{J}^{T_{p}}\left(x^{\prime}\right) \mathbf{J}^{t}(x)=\mathbf{J}_{p}\left(x^{\prime}\right) \mathbf{J}^{t}(x), \quad \text { or } \quad \mathbf{J}^{t+T_{p}}(x)=\mathbf{J}^{t}(x) \mathbf{J}_{p}(x) .
$$

Multiplying (8.8) by $\mathbf{J}^{t}(x)$, we find that the Jacobian matrix evaluated at $x^{\prime}$ has the same eigenvalue,

$$
\begin{equation*}
\mathbf{J}_{p}\left(x^{\prime}\right) \mathbf{e}_{i}\left(x^{\prime}\right)=\Lambda_{p, i} \mathbf{e}_{i}\left(x^{\prime}\right), \quad \mathbf{e}_{i}\left(x^{\prime}\right)=\mathbf{J}^{t}(x) \mathbf{e}_{i}(x), \tag{8.9}
\end{equation*}
$$

but with the eigenvector $\mathbf{e}_{i}$ transported along the flow $x \rightarrow x^{\prime}$ to $\mathbf{e}_{i}\left(x^{\prime}\right)=$ $\mathbf{J}^{t}(x) \mathbf{e}_{i}(x)$. Hence, $\mathbf{J}_{p}$ evaluated anywhere along the cycle has the same set of stability eigenvalues $\left\{\Lambda_{p, 1}, \Lambda_{p, 2}, \cdots \Lambda_{p, d}\right\}$. As quantities such as $\operatorname{tr} \mathbf{J}_{p}(x)$, $\operatorname{det} \mathbf{J}_{p}(x)$ depend only on the eigenvalues of $\mathbf{J}_{p}(x)$ and not on the starting point $x$, in expressions such as $\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}(x)\right)$ we may omit reference to any particular cycle point $x$ :

$$
\begin{equation*}
\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)=\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}(x)\right) \quad \text { for any } x \in p . \tag{8.10}
\end{equation*}
$$

We postpone the proof that the cycle stability eigenvalues are smooth conjugacy invariants of the flow to sect. 8.5.

### 8.2.1 Marginal eigenvalues

The presence of marginal eigenvalues signals either an invariance of the flow (which one should immediately exploit to simplify the problem), or a non-hyperbolicity of a flow (a source of much pain, hard to avoid).

Example 8.3 A periodic orbit of a continuous flow has a marginal eigenvalue: As $\mathbf{J}^{t}(x)$ transports the velocity field $v(x)$ by (4.6), after a complete period

$$
\begin{equation*}
\mathbf{J}_{p}(x) v(x)=v(x), \tag{8.11}
\end{equation*}
$$

so a periodic orbit of a flow always has an eigenvector $\mathbf{e}_{\|}(x)=v(x)$ parallel to the local velocity field with the unit eigenvalue

$$
\begin{equation*}
\Lambda_{p, \|}=1 . \tag{8.12}
\end{equation*}
$$

The continuous invariance that gives rise to this marginal eigenvalues is the invariance 7.2 of a cycle under a translation of its points along the cycle: two points on the payctle initially distance $\delta x$ apart, $x^{\prime}(0)-x(0)=\delta x(0)$, are separated by the exactly same $\delta x$ after a full period $T_{p}$. As we shall see in sect. 8.3, this marginal stability direction can be eliminated by cutting the cycle by a Poincaré section and eliminating the continuous flow Jacobian matrix in favor of the Jacobian matrix of the Poincaré return map.

If the flow is governed by a time-independent Hamiltonian, the energy is conserved, and that leads to an additional marginal eigenvalue (remember, by symplectic invariance (5.16) real eigenvalues come in pairs).

### 8.3 Stability of Poincaré map cycles

(R. Paškauskas and P. Cvitanović)

If a continuous flow periodic orbit $p$ pierces the Poincaré section $\mathcal{P}$ once, the section point is a fixed point of the Poincaré return map $P$ with stability (4.40)

$$
\begin{equation*}
\hat{\mathbf{J}}_{i j}=\left(\delta_{i k}-\frac{v_{i} U_{k}}{(v \cdot U)}\right) \mathbf{J}_{k j} \tag{8.13}
\end{equation*}
$$

with all primes dropped, as the initial and the final points coincide, $x^{\prime}=$ $f^{T_{p}}(x)=x$. If the periodic orbit $p$ pierces the set of Poincaré sections $\mathcal{P} n$ times, the same observation applies to the $n$th iterate of $P$.

We have already established in (4.41) that the velocity $v(x)$ is a zeroeigenvector of the Poincaré section Jacobian matrix, $\hat{\mathbf{J}} v=0$. Consider next ( $\Lambda_{p, \alpha}, \mathbf{e}_{\alpha}$ ), the full phase space $\alpha$ th (eigenvalue, eigenvector) pair (8.8), evaluated at a cycle point on a Poincaré section,

$$
\begin{equation*}
\mathbf{J}(x) \mathbf{e}_{\alpha}(x)=\Lambda_{\alpha} \mathbf{e}_{\alpha}(x), \quad x \in \mathcal{P} . \tag{8.14}
\end{equation*}
$$

Multiplying (8.13) by $\mathbf{e}_{\alpha}$ and inserting (8.14), we find that the Poincaré section Jacobian matrix $\hat{\mathbf{J}}$ has the same eigenvalue as the full phase space Jacobian matrix,

$$
\begin{equation*}
\hat{\mathbf{J}}(x) \hat{\mathbf{e}}_{\alpha}(x)=\Lambda_{\alpha} \hat{\mathbf{e}}_{\alpha}(x), \quad x \in \mathcal{P} . \tag{8.15}
\end{equation*}
$$

where $\hat{\mathbf{e}}_{\alpha}$ is a projection of the full phase space eigenvector onto the Poincaré section:

$$
\begin{equation*}
\left(\hat{\mathbf{e}}_{\alpha}\right)_{i}=\left(\delta_{i k}-\frac{v_{i} U_{k}}{(v \cdot U)}\right)\left(\mathbf{e}_{\alpha}\right)_{k} \tag{8.16}
\end{equation*}
$$

Hence, $\hat{\mathbf{J}}_{p}$ evaluated on any Poincaré section point along the cycle $p$ has the same set of stability eigenvalues $\left\{\Lambda_{p, 1}, \Lambda_{p, 2}, \cdots \Lambda_{p, d}\right\}$ as the full phase space Jacobian matrix $\mathbf{J}_{p}$.

### 8.4 Rectification of a 1-dimensional periodic orbit



In sect. 7.4.1 we have constructed the conjugation function for a fixed point. Here we turn to the problem of constructing it for periodic orbits. Each point around the cycle has a differently distorted neighborhood, with differing second and higher order derivatives, so we need to compute a different conjugation function $h_{a}$ at each cycle point $x_{a}$. We expand the map $f$ around each cycle point along the cycle,

$$
y_{a}(\phi)=f_{a}(\phi)-x_{a+1}=\phi f_{a, 1}+\phi^{2} f_{a, 2}+\ldots
$$

where $x_{a}$ is a point on the cycle, $f_{a}(\phi)=f\left(x_{a}+\phi\right)$ is centered on the periodic orbit, and the index $k$ in $f_{a, k}$ refers to the $k$ th order in the expansion (7.20).

For a periodic orbit the conjugation formula (7.19) generalizes to

$$
f_{a}(\phi)=h_{a+1}^{-1}\left(f_{a}^{\prime}(0) h_{a}(\phi)\right), \quad a=1,2, \cdots, n
$$

point by point. The conjugationg functions $h_{a}$ are obtained in the same way as before, by equating coefficients of the expansion (7.20), and assuming that the cycle stability $\Lambda=\prod_{a=0}^{n-1} f^{\prime}\left(x_{a}\right)$ is not marginal, $|\Lambda| \neq 1$. The explicit expressions for $h_{a}$ in terms of $f$ are obtained by iterating around the whole cycle,

$$
\begin{equation*}
f^{n}\left(x_{a}+\phi\right)=h_{a}^{-1}\left(\Lambda h_{a}(\phi)\right)+x_{a} . \tag{8.17}
\end{equation*}
$$

evaluated at each cycle point $a$. Again we have the freedom to set $h_{a}^{\prime}(0)=1$ for all $a$.

### 8.4.1 Repeats of cycles

We have traded in our initial nonlinear map $f$ for a (locally) linear map $\Lambda y$ and an equally complicated conjugation function $h$. What is gained by rewriting the map $f$ in terms of the conjugacy function $h$ ? Once the neighborhood of a fixed point is linearized, the repeats of it are trivialized; from the conjugation formula (7.20) one can compute the derivatives of a function composed with itself $r$ times:

$$
f^{r}(x)=h^{-1}\left(\Lambda^{r} h(x)\right) .
$$

One can already discern the form of the expansion for arbitrary repeats; the answer will depend on the conjugacy function $h(x)$ computed for a single repeat, and all the dependence on the repeat number will be carried by factors polynomial in $\Lambda^{r}$, a considerable simplification. The beauty of the idea is difficult to gauge at this stage - an appreciation only sets in when one starts computing perturbative corrections, be it in celestial mechanics (where the method was born), be it the quantum or stochastic corrections to "semiclassical" approximations.

### 8.5 Smooth conjugacies and cycle stability

In sect. 8.2 we have established that for a given flow the cycle stability eigenvalues are intrinsic to a given cycle, independent of the staring point along the cycle. Now we can prove a much stronger statement; cycle stability eigenvalues are metric invariants of the flow, the same in any representation of the dynamical system.

That the cycle stability eigenvalues are an invariant property of the given dynamical system follows from elementary considerations of sect. 7.1: If the same dynamics is given by a map $f$ in $x$ coordinates, and a map $g$ in the $y=h(x)$ coordinates, then $f$ and $g$ (or any other good representation) are related by (7.2), a reparametrization and a coordinate transformation $g=h \circ f \circ h^{-1}$. As both $f$ and $g$ are arbitrary representations of the
dynamical system, the explicit form of the conjugacy $h$ is of no interest, only the properties invariant under any transformation $h$ are of general import. Furthermore, a good representation should not mutilate the data; $h$ must be a smooth conjugacy which maps nearby cycle points of $f$ into nearby cycle points of $g$. This smoothness guarantees that the cycles are not only topological invariants, but that their linearized neighborhoods are also metrically invariant. For a fixed point $f(x)=x$ of a 1-dimensional map this follows from the chain rule for derivatives,

$$
\begin{align*}
g^{\prime}(y) & =h^{\prime}\left(f \circ h^{-1}(y)\right) f^{\prime}\left(h^{-1}(y)\right) \frac{1}{h^{\prime}(x)} \\
& =h^{\prime}(x) f^{\prime}(x) \frac{1}{h^{\prime}(x)}=f^{\prime}(x), \tag{8.18}
\end{align*}
$$

and the generalization to the stability eigenvalues of periodic orbits of $d$ dimensional flows is immediate.

As stability of a flow can always be rewritten as stability of a Poincaré section return map, we find that the stability eigenvalues of any cycle, for a flow or a map in arbitrary dimension, is a metric invariant of the dynamical
(2) 7.2
page 110 system.

### 8.6 Neighborhood of a cycle

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[帠 remark 10.3
(2) 4.1
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The Jacobian of the flow (or the sum of stability exponents) is easily evaluated.

Consider $\operatorname{det} \mathbf{J}^{t}\left(x_{0}\right)=\prod_{i=1}^{d} \Lambda_{i}\left(x_{0}, t\right)$, the product of the stability eigenvalues. We shall refer to this determinant as the Jacobian of the flow. By means of the time-ordered product (4.29) and the identity $\ln \operatorname{det} M=$ $\operatorname{tr} \ln M$ the Jacobian is given by

$$
\begin{equation*}
\operatorname{det} \mathbf{J}^{t}\left(x_{0}\right)=\exp \left[\int_{0}^{t} d \tau \operatorname{tr} \mathbf{A}(x(\tau))\right]=\exp \left[\int_{0}^{t} d \tau \partial_{i} v_{i}(x(\tau))\right] . \tag{8.19}
\end{equation*}
$$

(Here, as elsewhere in this book, a repeated index implies summation.) As the divergence $\partial_{i} v_{i}$ is a scalar quantity, the integral in the exponent needs no time ordering. All we need to do is evaluate the time average

$$
\begin{align*}
\left\langle\partial_{i} v_{i}\right\rangle_{t} & =\frac{1}{t} \int_{0}^{t} d \tau \sum_{i=1}^{d} A_{i i}(x(\tau)) \\
& =\frac{1}{t} \ln \left|\prod_{i=1}^{d} \Lambda_{i}\left(x_{0}, t\right)\right|=\sum_{i=1}^{d} \lambda_{i}\left(x_{0}, t\right) \tag{8.20}
\end{align*}
$$

along the trajectory. If the flow is not singular (for example, the trajectory does not run head-on into the Coulomb $1 / r$ singularity), the matrix of variations elements are bounded everywhere, $\left|A_{i j}\right|<M$, and so is the trace $\sum_{i} A_{i i}$. The time integral in (8.20) grows at most linearly with $t$, hence $\left\langle\partial_{i} v_{i}\right\rangle_{t}$ is bounded for all times, and numerical estimates of the $t \rightarrow \infty$ limit of $\left\langle\partial_{i} v_{i}\right\rangle$ are not marred by any blowups.

Even if we were to insist on extracting $\left\langle\partial_{i} v_{i}\right\rangle$ from (4.29) by first multiplying stability matrices along the flow, and then taking the logarithm, we can avoid exponential blowups in $\mathbf{J}^{t}$ by using the multiplicative structure (4.31), $\operatorname{det} \mathbf{J}^{\boldsymbol{t}^{\prime}+t}\left(x_{0}\right)=\operatorname{det} \mathbf{J}^{t^{\prime}}\left(x^{\prime}\right) \operatorname{det} \mathbf{J}^{t}\left(x_{0}\right)$ to restart with $\mathbf{J}^{0}\left(x^{\prime}\right)=\mathbf{1}$ whenever the eigenvalues of $\mathbf{J}^{t}\left(x_{0}\right)$ start getting out of hand. In numerical evaluations of Lyapunov exponents, $\lambda_{i}=\lim _{t \rightarrow \infty} \lambda_{i}\left(x_{0}, t\right)$, the sum rule (8.20) can serve as a helpful check on the accuracy of the computation.

The divergence $\partial_{i} v_{i}$ is an important characterization of the flow - it describes the behavior of a phase space volume in the infinitesimal neighborhood of the trajectory. If $\partial_{i} v_{i}<0$, the flow is locally contracting, and the trajectory might be falling into an attractor. If $\partial_{i} v_{i}=0$, the flow preserves phase space volume and $\operatorname{det} \mathbf{J}^{t}=\mathbf{1}$. A flow with this property is called incompressible. An important class of such flows are the Hamiltonian flows considered in sect. 5.2.

But before we can get to that, Henri Roux, the perfect student always on alert, pipes up. He does not like our definition of the Jacobian matrix in terms of the time-ordered exponential (4.30). Depending on the signs of stability eigenvalues, the left hand side of (8.19) can be either positive or negative. But the right hand side is an exponential of a real number, and that can only be positive. What gives? As we shall see much later on in this text, in discussion of topological indices arising in semiclassical quantization, this is not at all a dumb question.

### 8.6.1 There goes the neighborhood

In what follows, our task will be to determine the size of a neighborhood of $x(t)$, and that is why we care about the stability eigenvalues, and especially the unstable (expanding) ones. Nearby points aligned along the stable (contracting) directions remain in the neighborhood of the trajectory $x(t)=f^{t}\left(x_{0}\right)$; the ones to keep an eye on are the points which leave the neighborhood along the unstable directions. The sub-volume $\left|\mathcal{M}_{i}\right|=\prod_{i}^{e} \Delta x_{i}$ of the set of points which get no further away from $f^{t}\left(x_{0}\right)$ than $L$, the typical size of the system, is fixed by the condition that $\Delta x_{i} \Lambda_{i}=O(L)$ in each expanding direction $i$. Hence the neighborhood size scales as $\propto 1 /\left|\Lambda_{p}\right|$ where $\Lambda_{p}$ is the product of expanding eigenvalues (8.3) only; contracting ones play a secondary role. So secondary that even
infinitely many of them (for example, the infinity of contracting eigendirections of the spatiotemporal dynamics of sect. 2.4.1) will not matter.

So the physically important information is carried by the expanding sub-volume, not the total volume computed so easily in (8.20). That is also the reason why the dissipative and the Hamiltonian chaotic flows are much more alike than one would have naively expected for "compressible" vs. "incompressible" flows. In hyperbolic systems what matters are the expanding directions. Whether the contracting eigenvalues are inverses of the expanding ones or not is of secondary importance. As long as the number of unstable directions is finite, the same theory applies both to the finite-dimensional ODEs and infinite-dimensional PDEs.

## Résumé

Periodic orbits play a central role in any invariant characterization of the dynamics, because (a) their existence and inter-relations are a topological, coordinate-independent property of the dynamics, and (b) their stability eigenvalues form an infinite set of metric invariants: The stability eigenvalues of a periodic orbit remain invariant under any smooth nonlinear change of coordinates $f \rightarrow h \circ f \circ h^{-1}$.

We shall show in chapter 11 that extending their local stability eigendirections into stable and unstable manifolds yields important global information about the topological organization of phase space.

The physically important information is carried by the unstable manifold, and the expanding sub-volume characterized by the product of expanding eigenvalues of $\mathbf{J}_{p}$. In hyperbolic systems what matters are the expanding directions. As long as the number of unstable directions is finite, the theory can be applied to flows of arbitrarily high dimension.

## Exercises

Exercise 8.1 Fundamental domain fixed points. Use the formula (6.11) for billiard Jacobian matrix to compute the periods $T_{p}$ and the expanding eigenvalues $\Lambda_{p}$ of the fundamental domain $\overline{0}$ (the 2-cycle of the complete 3disk space) and $\overline{1}$ (the 3-cycle of the complete 3-disk space) fixed points:

|  | $T_{p}$ | $\Lambda_{p}$ |
| :---: | :---: | :---: |
| $\overline{0}:$ | $R-2$ | $R-1+R \sqrt{1-2 / R}$ |
| $\overline{1}:$ | $R-\sqrt{3}$ | $-\frac{2 R}{\sqrt{3}}+1-\frac{2 R}{\sqrt{3}} \sqrt{1-\sqrt{3} / R}$ |

We have set the disk radius to $a=1$.

Exercise 8.2 Fundamental domain 2-cycle. Verify that for the $\overline{10}$-cycle the cycle length and the trace of the Jacobian matrix are given by

$$
\begin{align*}
L_{10} & =2 \sqrt{R^{2}-\sqrt{3} R+1}-2 \\
\operatorname{tr} \mathbf{J}_{10} & =2 L_{10}+2+\frac{1}{2} \frac{L_{10}\left(L_{10}+2\right)^{2}}{\sqrt{3} R / 2-1} \tag{8.22}
\end{align*}
$$

The $\overline{10}$-cycle is drawn in figure 11.6. The unstable eigenvalue $\Lambda_{10}$ follows from (4.18).

## Chapter 9

## Transporting densities

Paulina: I'll draw the curtain:<br>My lord's almost so far transported that<br>He'll think anon it lives.<br>W. Shakespeare: The Winter's Tale

(P. Cvitanović, R. Artuso, L. Rondoni, and E.A. Spiegel)

In chapters 2, 3, 5 and 6 we learned how to track an individual trajectory, and saw that such a trajectory can be very complicated. In chapter 4 we studied a small neighborhood of a trajectory and learned that such neighborhood can grow exponentially with time, making the concept of tracking an individual trajectory for long times a purely mathematical idealization.

While the trajectory of an individual representative point may be highly convoluted, the density of these points might evolve in a manner that is relatively smooth. The evolution of the density of representative points is for this reason (and other that will emerge in due course) of great interest. So are the behaviors of other properties carried by the evolving swarm of representative points.

We shall now show that the global evolution of the density of representative points is conveniently formulated in terms of evolution operators.

### 9.1 Measures

Do I then measure, O my God, and know not what I measure?
St. Augustine, The confessions of Saint Augustine
A fundamental concept in the description of dynamics of a chaotic system is that of measure, which we denote by $d \mu(x)=\rho(x) d x$. An intuitive way to define and construct a physically meaningful measure is by a process of coarse-graining. Consider a sequence $1,2, \ldots, n, \ldots$ of increasingly


Figure 9.1: (a) First level of partitioning: A coarse partition of $\mathcal{M}$ into regions $\mathcal{M}_{0}$, $\mathcal{M}_{1}$, and $\mathcal{M}_{2}$. (b) $n=2$ level of partitioning: A refinement of the above partition, with each region $\mathcal{M}_{i}$ subdivided into $\mathcal{M}_{i 0}, \mathcal{M}_{i 1}$, and $\mathcal{M}_{i 2}$.
refined partitions of phase space, figure 9.1, into regions $\mathcal{M}_{i}$ defined by the characteristic function

$$
\chi_{i}(x)= \begin{cases}1 & \text { if } x \in \mathcal{M}_{i},  \tag{9.1}\\ 0 & \text { otherwise } .\end{cases}
$$

A coarse-grained measure is obtained by assigning the "mass", or the fraction of trajectories contained in the $i$ th region $\mathcal{M}_{i} \subset \mathcal{M}$ at the $n$th level of partitioning of the phase space:

$$
\begin{equation*}
\Delta \mu_{i}=\int_{\mathcal{M}} d \mu(x) \chi_{i}(x)=\int_{\mathcal{M}_{i}} d \mu(x)=\int_{\mathcal{M}_{i}} d x \rho(x) \tag{9.2}
\end{equation*}
$$

The function $\rho(x)=\rho(x, t)$ denotes the density of representative points in phase space at time $t$. This density can be (and in chaotic dynamics, often is) an arbitrarily ugly function, and it may display remarkable singularities; for instance, there may exist directions along which the measure is singular with respect to the Lebesgue measure. As our intent is to sprinkle phase space with a finite number of initial points (repeat an experiment a finite number of times), we shall assume that the measure can be normalized

$$
\begin{equation*}
\sum_{i}^{(n)} \Delta \mu_{i}=1 \tag{9.3}
\end{equation*}
$$

where the sum is over subregions $i$ at the $n$th level of partitioning. The infinitesimal measure $d x \rho(x)$ can be thought of as an infinitely refined partition limit of $\Delta \mu_{i}=\left|\mathcal{M}_{i}\right| \rho\left(x_{i}\right), x_{i} \in \mathcal{M}_{i}$, with normalization

$$
\begin{equation*}
\int_{\mathcal{M}} d x \rho(x)=1 \tag{9.4}
\end{equation*}
$$

So far, any arbitrary sequence of partitions will do. What are intelligent ways of partitioning phase space? We postpone the answer to chapter 11, after we have developed some intuition about how the dynamics transports densities.

### 9.2 Perron-Frobenius operator

Given a density, the question arises as to what it might evolve into with time. Consider a swarm of representative points making up the measure contained in a region $\mathcal{M}_{i}$ at time $t=0$. As the flow evolves, this region is carried into $f^{t}\left(\mathcal{M}_{i}\right)$, as in figure $2.1(\mathrm{~b})$. No trajectory is created or destroyed, so the conservation of representative points requires that

$$
\int_{f^{t}\left(\mathcal{M}_{i}\right)} d x \rho(x, t)=\int_{\mathcal{M}_{i}} d x_{0} \rho\left(x_{0}, 0\right) .
$$

If the flow is invertible and the transformation $x_{0}=f^{-t}(x)$ is single-valued, we can transform the integration variable in the expression on the left to

$$
\int_{\mathcal{M}_{i}} d x_{0} \rho\left(f^{t}\left(x_{0}\right), t\right)\left|\operatorname{det} \mathbf{J}^{t}\left(x_{0}\right)\right| .
$$

We conclude that the density changes with time as the inverse of the Jacobian (8.19)

$$
\begin{equation*}
\rho(x, t)=\frac{\rho\left(x_{0}, 0\right)}{\left|\operatorname{det} \mathbf{J}^{t}\left(x_{0}\right)\right|}, \quad x=f^{t}\left(x_{0}\right) \tag{9.5}
\end{equation*}
$$

which makes sense: the density varies inversely to the infinitesimal volume occupied by the trajectories of the flow.

The manner in which a flow transports densities may be recast into the language of operators, by writing

$$
\begin{equation*}
\rho(x, t)=\mathcal{L}^{t} \rho(x)=\int_{\mathcal{M}} d x_{0} \delta\left(x-f^{t}\left(x_{0}\right)\right) \rho\left(x_{0}, 0\right) . \tag{9.6}
\end{equation*}
$$

Let us check this formula. Integrating Dirac delta functions is easy: $\int_{\mathcal{M}} d x \delta(x)=$ 1 if $0 \in \mathcal{M}$, zero otherwise. The integral over a one-dimensional Dirac delta function picks up the Jacobian of its argument evaluated at all of its zeros:

$$
\begin{equation*}
\int d x \delta(h(x))=\sum_{\{x: h(x)=0\}} \frac{1}{\left|h^{\prime}(x)\right|}, \tag{9.7}
\end{equation*}
$$

and in $d$ dimensions the denominator is replaced by

$$
\begin{equation*}
\int d x \delta(h(x))=\sum_{\{x: h(x)=0\}} \frac{1}{\left|\operatorname{det} \frac{\partial h(x)}{\partial x}\right|} . \tag{9.8}
\end{equation*}
$$

Figure 9.2: A piecewise-linear skew "Ulam tent" $\operatorname{map}(9.11)\left(\Lambda_{0}=4 / 3, \Lambda_{1}=-4\right)$.


Now you can check that (9.6) is just a rewrite of (9.5):

$$
\begin{align*}
\mathcal{L}^{t} \rho(x) & =\sum_{x_{0}=f^{-t}(x)} \frac{\rho\left(x_{0}\right)}{\left|f^{t^{t}}\left(x_{0}\right)\right|} \\
& =\sum_{x_{0}=f^{-t}(x)} \frac{\rho\left(x_{0}\right)}{\left|\operatorname{det} \mathbf{J}^{t}\left(x_{0}\right)\right|} \quad \text { (1-dimensional) } \tag{9.9}
\end{align*}
$$

For a deterministic, invertible flow $x$ has only one preimage $x_{0}$; allowing for multiple preimages also takes account of noninvertible mappings such as the "stretch \& fold" maps of the interval, to be discussed briefly in the next example, and in more detail in sect. 11.3.1.

We shall refer to the kernel of (9.6) as the Perron-Frobenius operator:
(2) 9.3
page 137
[-7
[空 remark 15.4

If you do not like the word "kernel" you might prefer to think of $\mathcal{L}^{t}(x, y)$ as a matrix with indices $x, y$. The Perron-Frobenius operator assembles the density $\rho(x, t)$ at time $t$ by going back in time to the density $\rho\left(x_{0}, 0\right)$ at time $t=0$.

$$
\begin{equation*}
\mathcal{L}^{t}(x, y)=\delta\left(x-f^{t}(y)\right) . \tag{9.10}
\end{equation*}
$$

> in depth:
> appendix D, p. 659

Example 9.1 Perron-Frobenius operator for a piecewise-linear map: Consider Assume the expanding 1-d map $f(x)$ of figure 9.2, a piecewise-linear 2-branch map with slopes $\Lambda_{0}>1$ and $\Lambda_{1}<-1$ :

$$
f(x)= \begin{cases}f_{0}(x)=\Lambda_{0} x, & x \in \mathcal{M}_{0}=\left[0,1 / \Lambda_{0}\right)  \tag{9.11}\\ f_{1}(x)=\frac{\Lambda_{0}}{\Lambda_{0}-1}(1-x), & x \in \mathcal{M}_{1}=\left(1 / \Lambda_{0}, 1\right] .\end{cases}
$$

Both $f\left(\mathcal{M}_{0}\right)$ and $f\left(\mathcal{M}_{1}\right)$ map onto the entire unit interval $\mathcal{M}=[0,1]$. Assume a piecewise constant density

$$
\rho(x)=\left\{\begin{array}{ll}
\rho_{0} & \text { if } x \in \mathcal{M}_{0}  \tag{9.12}\\
\rho_{1} & \text { if } x \in \mathcal{M}_{1}
\end{array} .\right.
$$

As can be easily checked using (9.9), the Perron-Frobenius operator acts on this piece- wise constant function as a [ $2 \times 2$ ] "transfer" matrix with matrix elements

$$
\binom{\rho_{0}}{\rho_{1}} \rightarrow \mathcal{L} \rho=\left(\begin{array}{cc}
\frac{1}{\left|\Lambda_{0}\right|} & \frac{1}{\left|\Lambda_{1}\right|}  \tag{9.13}\\
\frac{1}{\left|\Lambda_{0}\right|} & \frac{1}{\left|\Lambda_{1}\right|}
\end{array}\right)\binom{\rho_{0}}{\rho_{1}}
$$

stretching both $\rho_{0}$ and $\rho_{1}$ over the whole unit interval $\Lambda$. In this example the density is constant after one iteration, so $\mathcal{L}$ has only a unit eigenvalue $e^{s_{0}}=1 /\left|\Lambda_{0}\right|+1 /\left|\Lambda_{1}\right|=$ 1 , with constant density eigenvector $\rho_{0}=\rho_{1}$. The quantities $1 /\left|\Lambda_{0}\right|, 1 /\left|\Lambda_{1}\right|$ are, respectively, the fractions of phase space taken up by the $\left|\mathcal{M}_{0}\right|,\left|\mathcal{M}_{1}\right|$ intervals. This simple explicit matrix representation of the Perron-Frobenius operator is a consequence of the piecewise linearity of $f$, and the restriction of the densities $\rho$ to the space of piecewise constant functions. The example gives a flavor of the enterprise upon which we are about to embark in this book, but the full story is much subtler: in general, there will exist no such finite-dimensional representation for the Perron-Frobenius operator. (Continued in example 10.1.)

To a student with a practical bent the example suggests a strategy for constructing evolution operators for smooth maps, as limits of partitions of phase space into regions $\mathcal{M}_{i}$, with a piecewise-linear approximations $f_{i}$ to the dynamics in each region, but that would be too naive; much of the physically interesting spectrum would be missed. As we shall see, the choice of function space for $\rho$ is crucial, and the physically motivated choice is a space of smooth functions, rather than the space of piecewise constant functions.

### 9.3 Invariant measures

A stationary or invariant density is a density left unchanged by the flow

$$
\begin{equation*}
\rho(x, t)=\rho(x, 0)=\rho(x) . \tag{9.14}
\end{equation*}
$$

Conversely, if such a density exists, the transformation $f^{t}(x)$ is said to be measure-preserving. As we are given deterministic dynamics and our goal is the computation of asymptotic averages of observables, our task is to identify interesting invariant measures for a given $f^{t}(x)$. Invariant measures remain unaffected by dynamics, so they are fixed points (in the infinite-dimensional function space of $\rho$ densities) of the Perron-Frobenius operator (9.10), with the unit eigenvalue:

$$
\begin{equation*}
\mathcal{L}^{t} \rho(x)=\int_{\mathcal{M}} d y \delta\left(x-f^{t}(y)\right) \rho(y)=\rho(x) . \tag{9.15}
\end{equation*}
$$

In general, depending on the choice of $f^{t}(x)$ and the function space for $\rho(x)$, there may be no, one, or many solutions of the eigenfunction condition (9.15). For instance, a singular measure $d \mu(x)=\delta\left(x-x_{q}\right) d x$ concentrated on an equilibrium point $x_{q}=f^{t}\left(x_{q}\right)$, or any linear combination of such
measures, each concentrated on a different equilibrium point, is stationary. There are thus infinitely many stationary measures that can be constructed. Almost all of them are unnatural in the sense that the slightest perturbation will destroy them.

From a physical point of view, there is no way to prepare initial densities which are singular, so it makes sense to concentrate on measures which are limits of transformations experienced by an initial smooth distribution $\rho(x)$ under the action of $f$, rather than as a limit computed from a single trajectory,

$$
\begin{equation*}
\rho_{0}(x)=\lim _{t \rightarrow \infty} \int_{\mathcal{M}} d y \delta\left(x-f^{t}(y)\right) \rho(y, 0), \quad \int_{\mathcal{M}} d y \rho(y, 0)=1 . \tag{9.16}
\end{equation*}
$$

Intuitively, the "natural" measure (or measures) should be the least sensitive to facts of life, such as noise (no matter how weak).

### 9.3.1 Natural measure

Huang: Chen-Ning, do you think ergodic theory gives us useful insight into the foundation of statistical mechanics?
Yang: I don't think so.
Kerson Huang, C.N. Yang interview
The natural or equilibrium measure can be defined as the limit

$$
\bar{\rho}_{x_{0}}(y)= \begin{cases}\lim _{t \rightarrow \infty} \frac{1}{t} \int_{0}^{t} d \tau \delta\left(y-f^{\tau}\left(x_{0}\right)\right) & \text { flows }  \tag{9.17}\\ \lim _{n \rightarrow \infty} \frac{1}{n} \sum_{k=0}^{n-1} \delta\left(y-f^{k}\left(x_{0}\right)\right) & \text { maps }\end{cases}
$$

where $x_{0}$ is a generic inital point. Staring at an average over infinitely many Dirac deltas is not a prospect we cherish. Generated by the action of $f$, the natural measure satisfies the stationarity condition (9.15) and is thus invariant by construction. From a computational point of view, the natural measure is the visitation frequency defined by coarse-graining, integrating (9.17) over the $\mathcal{M}_{i}$ region

$$
\begin{equation*}
\Delta \bar{\mu}_{i}=\lim _{t \rightarrow \infty} \frac{t_{i}}{t}, \tag{9.18}
\end{equation*}
$$

where $t_{i}$ is the accumulated time that a trajectory of total duration $t$ spends in the $\mathcal{M}_{i}$ region, with the initial point $x_{0}$ picked from some smooth density $\rho(x)$.

Let $a=a(x)$ be any observable. In the mathematical literature $a(x)$ is a function belonging to some function space, for instance the space of
integrable functions $L^{1}$, that associates to each point in phase space a number or a set of numbers. In physical applications the observable $a(x)$ is necessarily a smooth function. The observable reports on some property of the dynamical system. Several examples will be given in sect. 10.1.

The space average of the observable $a$ with respect to a measure $\rho$ is given by the $d$-dimensional integral over the phase space $\mathcal{M}$ :

$$
\begin{align*}
\langle a\rangle_{\rho}= & \frac{1}{\left|\rho_{\mathcal{M}}\right|} \int_{\mathcal{M}} d x \rho(x) a(x) \\
& \left|\rho_{\mathcal{M}}\right|=\int_{\mathcal{M}} d x \rho(x)=\text { mass in } \mathcal{M} . \tag{9.19}
\end{align*}
$$

For now we assume that the phase space $\mathcal{M}$ has a finite dimension and a finite volume. By definition, $\langle a\rangle_{\rho}$ is a function(al) of $\rho$.

Inserting the right-hand-side of (9.17) into (9.19), we see that the natural measure corresponds to a time average of the observable $a$ along a trajectory of the initial point $x_{0}$,

$$
\begin{equation*}
\overline{a_{x_{0}}}=\lim _{t \rightarrow \infty} \frac{1}{t} \int_{0}^{t} d \tau a\left(f^{\tau}\left(x_{0}\right)\right) . \tag{9.20}
\end{equation*}
$$

Analysis of the above asymptotic time limit is the central problem of ergodic theory. The Birkhoff ergodic theorem asserts that if a natural measure $\rho$ exists, the limit $\overline{a\left(x_{0}\right)}$ for the time average (9.20) exists for all initial $x_{0}$. As we shall not rely on this result in what follows we forgo a proof here. Furthermore, if the dynamical system is ergodic, the time average over almost any trajectory tends to the space average

$$
\begin{equation*}
\lim _{t \rightarrow \infty} \frac{1}{t} \int_{0}^{t} d \tau a\left(f^{\tau}\left(x_{0}\right)\right)=\langle a\rangle \tag{9.21}
\end{equation*}
$$

for "almost all" initial $x_{0}$. By "almost all" we mean that the time average is independent of the initial point apart from a set of $\rho$-measure zero.

For future reference, we note a further property that is stronger than ergodicity: if the space average of a product of any two variables decorrelates with time,

$$
\begin{equation*}
\lim _{t \rightarrow \infty}\left\langle a(x) b\left(f^{t}(x)\right)\right\rangle=\langle a\rangle\langle b\rangle, \tag{9.22}
\end{equation*}
$$

the dynamical system is said to be mixing.

Example 9.2 The Hénon attractor natural measure: A numerical calculation of the natural measure (9.18) for the Hénon attractor (3.12) is given by the histogram in figure 9.3. The phase space is partitioned into many equal-size areas $\mathcal{M}_{i}$, and the coarse grained measure (9.18) is computed by a long-time iteration of the Hénon map, and represented by the height of the column over area $\mathcal{M}_{i}$. What we see is a typical invariant measure - a complicated, singular function concentrated on a fractal set.

Figure 9.3: Natural measure (9.18) for the Hénon map (3.12) strange attractor at parameter values $(a, b)=(1.4,0.3)$. See figure 3.4 for a sketch of the attractor without the natural measure binning. (Courtesy of J.-P. Eckmann)


If an invariant measure is quite singular (for instance a Dirac $\delta$ concentrated on a fixed point or a cycle), its existence is most likely of limited physical import. No smooth inital density will converge to this measure if the dynamics is unstable. In practice the average (9.17) is problematic and often hard to control, as generic dynamical systems are neither uniformly hyperbolic nor structurally stable: it is not known whether even the simplest model of a strange attractor, the Hénon attractor, is a strange attractor or merely a long stable cycle.

### 9.3.2 Determinism vs. stochasticity

While dynamics can lead to very singular $\rho$ 's, in any physical setting we cannot do better than to measure it averaged over some region $\mathcal{M}_{i}$; the coarse-graining is not an approximation but a physical necessity. One is free to think of a measure as a probability density, as long as one keeps in mind the distinction between deterministic and stochastic flows. In deterministic evolution the evolution kernels are not probabilistic; the density of trajectories is transported deterministically. What this distinction means will became apparent later: for deterministic flows our trace and determinant formulas will be exact, while for quantum and stochastic flows they will only be the leading saddlepoint (stationary phase, steepest descent) approximations.

Clearly, while deceptively easy to define, measures spell trouble. The good news is that if you hang on, you will never need to compute them, at least not in this book. How so? The evolution operators to which we next turn, and the trace and determinant formulas to which they will lead us, will assign the correct weights to desired averages without recourse to any explicit computation of the coarse-grained measure $\Delta \rho_{i}$.

### 9.4 Density evolution for infinitesimal times

Consider the evolution of a smooth density $\rho(x)=\rho(x, 0)$ under an infinitesimal step $\delta \tau$, by expanding the action of $\mathcal{L}^{\delta \tau}$ to linear order in $\delta \tau$ :

$$
\begin{aligned}
\mathcal{L}^{\delta \tau} \rho(y) & =\int_{\mathcal{M}} d x \delta\left(y-f^{\delta \tau}(x)\right) \rho(x) \\
& =\int_{\mathcal{M}} d x \delta(y-x-\delta \tau v(x)) \rho(x)
\end{aligned}
$$

$$
\begin{align*}
& =\frac{\rho(y-\delta \tau v(y))}{\left|\operatorname{det}\left(1+\delta \tau \frac{\partial v(y)}{\partial x}\right)\right|}=\frac{\rho(y)-\delta \tau \sum_{i=1}^{d} v_{i}(y) \partial_{i} \rho(y)}{1+\delta \tau \sum_{i=1}^{d} \partial_{i} v_{i}(y)} \\
\rho(x, \delta \tau) & =\rho(x, 0)-\delta \tau \sum_{i=1}^{d} \frac{\partial}{\partial x_{i}}\left(v_{i}(x) \rho(x, 0)\right) . \tag{9.23}
\end{align*}
$$

Here we have used the infinitesimal form of the flow (2.5), the Dirac delta Jacobian (9.9), and the $\ln$ det $=\operatorname{tr} \ln$ relation. Moving $\rho(y, 0)$ to the left hand side and dividing by $\delta \tau$, we discover that the rate of the deformation of $\rho$ under the infinitesimal action of the Perron-Frobenius operator is nothing but the continuity equation for the density:

$$
\begin{equation*}
\partial_{t} \rho+\partial \cdot(\rho v)=0 . \tag{9.24}
\end{equation*}
$$

The family of Perron-Frobenius operators operators $\left\{\mathcal{L}^{t}\right\}_{t \in \mathbb{R}_{+}}$forms a semigroup parametrized by time
(a) $\mathcal{L}^{0}=I$
(b) $\mathcal{L}^{t} \mathcal{L}^{t^{\prime}}=\mathcal{L}^{t+t^{\prime}} \quad t, t^{\prime} \geq 0 \quad$ (semigroup property).

The generator of the semigroup (that is, the generator of a time translation by an infinitesimal step $\delta \tau$ ) is determined by (9.23):

$$
\begin{equation*}
\mathcal{A} \rho(x)=+\lim _{\delta \tau \rightarrow 0^{+}} \frac{1}{\delta \tau}\left(\mathcal{L}^{\delta \tau}-I\right) \rho(x)=-\partial_{i}\left(v_{i}(x) \rho(x)\right) \tag{9.25}
\end{equation*}
$$

If the flow is finite-dimensional and invertible, $\mathcal{A}$ is a generator of a fullfledged group. The left hand side is the definition of time derivative, so the evolution equation for $\rho(x)$ is

$$
\begin{equation*}
\left(\frac{\partial}{\partial t}-\mathcal{A}\right) \rho(x)=0 \tag{9.26}
\end{equation*}
$$

The finite time Perron-Frobenius operator (9.10) can be formally expressed by exponentiating the time evolution generator $\mathcal{A}$ as

$$
\begin{equation*}
\mathcal{L}^{t}=e^{t \mathcal{A}} \tag{9.27}
\end{equation*}
$$

The generator $\mathcal{A}$ is reminiscent of the generator of translations. Indeed, for a constant velocity field dynamical evolution is nothing but a translation by (time $\times$ velocity):

$$
\begin{equation*}
e^{-t v \frac{\partial}{\partial x}} a(x)=a(x-t v) . \tag{9.28}
\end{equation*}
$$

As we will not need to implement a computational formula for general $e^{t \mathcal{A}}$ in what follows, we relegate making sense of such operators to appendix D.2.

### 9.4.1 Resolvent of $\mathcal{L}$

Here we limit ourselves to a brief remark about the notion of the "spectrum" of a linear operator.

The Perron-Frobenius operator $\mathcal{L}$ acts multiplicatively in time, so it is reasonable to suppose that there exist constants $M>0, \beta \geq 0$ such that $\left\|\mathcal{L}^{t}\right\| \leq M e^{t \beta}$ for all $t \geq 0$. What does that mean? The operator norm is defined in the same spirit in which one defines matrix norms (see appendix K.2): We are assuming that no value of $\mathcal{L}^{t} \rho(x)$ grows faster than exponentially for any choice of function $\rho(x)$, so that the fastest possible growth can be bounded by $e^{t \beta}$, a reasonable expectation in the light of the
[ig sect. K. 2 simplest example studied so far, the exact escape rate (10.20). If that is so, multiplying $\mathcal{L}^{t}$ by $e^{-t \beta}$ we construct a new operator $e^{-t \beta} \mathcal{L}^{t}=e^{t(\mathcal{A}-\beta)}$ which decays exponentially for large $t,\left\|e^{t(\mathcal{A}-\beta)}\right\| \leq M$. We say that $e^{-t \beta} \mathcal{L}^{t}$ is an element of a bounded semigroup with generator $\mathcal{A}-\beta I$. Given this bound, it follows by the Laplace transform

$$
\begin{equation*}
\int_{0}^{\infty} d t e^{-s t} \mathcal{L}^{t}=\frac{1}{s-\mathcal{A}}, \quad \operatorname{Re} s>\beta \tag{9.29}
\end{equation*}
$$

that the resolvent operator $(s-\mathcal{A})^{-1}$ is bounded ("resolvent" $=$ able to cause separation into constituents)

$$
\left\|\frac{1}{s-\mathcal{A}}\right\| \leq \int_{0}^{\infty} d t e^{-s t} M e^{t \beta}=\frac{M}{s-\beta} .
$$

If one is interested in the spectrum of $\mathcal{L}$, as we will be, the resolvent operator is a natural object to study. The main lesson of this brief aside is that for continuous time flows, the Laplace transform is the tool that brings down the generator in (9.27) into the resolvent form (9.29) and enables us to study its spectrum.

> in depth:
> appendix D.2, p. 661

### 9.5 Liouville operator

- 

A case of special interest is the Hamiltonian or symplectic flow defined by the Hamilton's equations of motion (5.1). A reader versed in quantum mechanics will have observed by now that with replacement $\mathcal{A} \rightarrow$ $-\frac{i}{\hbar} \hat{H}$, where $\hat{H}$ is the quantum Hamiltonian operator, (9.26) looks rather like the time dependent Schrödinger equation, so this is probably the right moment to figure out what all this means in the case of Hamiltonian flows.

The Hamilton's evolution equations (5.1) for any time-independent quantity $Q=Q(q, p)$ are given by

$$
\begin{equation*}
\frac{d Q}{d t}=\frac{\partial Q}{\partial q_{i}} \frac{d q_{i}}{d t}+\frac{\partial Q}{\partial p_{i}} \frac{d p_{i}}{d t}=\frac{\partial H}{\partial p_{i}} \frac{\partial Q}{\partial q_{i}}-\frac{\partial Q}{\partial p_{i}} \frac{\partial H}{\partial q_{i}} . \tag{9.30}
\end{equation*}
$$

As equations with this structure arise frequently for symplectic flows, it is convenient to introduce a notation for them, the Poisson bracket

$$
\begin{equation*}
\{A, B\}=\frac{\partial A}{\partial p_{i}} \frac{\partial B}{\partial q_{i}}-\frac{\partial A}{\partial q_{i}} \frac{\partial B}{\partial p_{i}} . \tag{9.31}
\end{equation*}
$$

In terms of Poisson brackets the time evolution equation (9.30) takes the compact form

$$
\begin{equation*}
\frac{d Q}{d t}=\{H, Q\} . \tag{9.32}
\end{equation*}
$$

The full phase space flow velocity is $\dot{x}=(\dot{q}, \dot{p})$, where the dot signifies time derivative for fixed initial point.

The discussion of sect. 9.4 applies to any deterministic flow. If the density itself is a material invariant, combining

$$
\partial_{t} I+v \cdot \partial I=0 .
$$

and (9.24) we conclude that $\partial_{i} v_{i}=0$ and $\mathbf{J}^{t}\left(x_{0}\right)=1$. An example of such incompressible flow is the Hamiltonian flow of sect. 5.2. For incompressible flows the continuity equation (9.24) becomes a statement of conservation of the phase space volume (see sect. 5.2), or the Liouville theorem

$$
\begin{equation*}
\partial_{t} \rho+v_{i} \partial_{i} \rho=0 . \tag{9.33}
\end{equation*}
$$

Hamilton's equations (5.1) imply that the flow is incompressible, $\partial_{i} v_{i}=$ 0 , so for Hamiltonian flows the equation for $\rho$ reduces to the continuity equation for the phase-space density:

$$
\begin{equation*}
\partial_{t} \rho+\partial_{i}\left(\rho v_{i}\right)=0 . \tag{9.34}
\end{equation*}
$$

Consider the evolution of the phase space density $\rho$ of an ensemble of noninteracting particles; the particles are conserved, so

$$
\frac{d}{d t} \rho(q, p, t)=\left(\frac{\partial}{\partial t}+\dot{q}_{i} \frac{\partial}{\partial q_{i}}+\dot{p}_{i} \frac{\partial}{\partial p_{i}}\right) \rho(q, p, t)=0 .
$$

Inserting Hamilton's equations (5.1) we obtain the Liouville equation, a special case of (9.26):

$$
\begin{equation*}
\frac{\partial}{\partial t} \rho(q, p, t)=-\mathcal{A} \rho(q, p, t)=\{H, \rho(q, p, t)\}, \tag{9.35}
\end{equation*}
$$

where $\{$,$\} is the Poisson bracket (9.31). The generator of the flow (9.25)$ is now the generator of infinitesimal symplectic transformations,

$$
\begin{equation*}
\mathcal{A}=\dot{q}_{i} \frac{\partial}{\partial q_{i}}+\dot{p_{i}} \frac{\partial}{\partial p_{i}}=\frac{\partial H}{\partial p_{i}} \frac{\partial}{\partial q_{i}}-\frac{\partial H}{\partial q_{i}} \frac{\partial}{\partial p_{i}} . \tag{9.36}
\end{equation*}
$$

For separable Hamiltonians of form $H=p^{2} / 2 m+V(q)$, the equations of motion are

$$
\begin{equation*}
\dot{q}_{i}=\frac{p_{i}}{m}, \quad \dot{p}_{i}=-\frac{\partial V(q)}{\partial q_{i}} . \tag{9.3}
\end{equation*}
$$

and the action of the generator

$$
\begin{equation*}
\mathcal{A}=-\frac{p_{i}}{m} \frac{\partial}{\partial q_{i}}+\partial_{i} V(q) \frac{\partial}{\partial p_{i}} . \tag{9.38}
\end{equation*}
$$

can be interpreted as a translation (9.28) in configuration space, followed by acceleration by force $\partial V(q)$ in the momentum space.

This special case of the time evolution generator (9.25) for the case of symplectic flows is called the Liouville operator. You might have encountered it in statistical mechanics, while discussing what ergodicity means for $10^{23}$ hard balls. Here its action will be very tangible; we shall apply the evolution operator to systems as small as 1 or 2 hard balls and to our surprise learn that that suffices to alredy get a bit of a grip on foundations of the classical nonequilibrium statistical mechanics.

## Commentary

Remark 9.1 Ergodic theory. An overview of ergodic theory is outside the scope of this book: the interested reader may find it useful to consult [9.1]. The existence of time average (9.20) is the basic result of ergodic theory, known as the Birkhoff theorem, see for example refs. [9.1, 1.14], or the statement of theorem 7.3.1 in ref. [9.6]. The natural measure (9.18) (more carefully defined than in the above sketch) is often referred to as the SRB or Sinai-Ruelle-Bowen measure [1.17, 1.16, 1.19].

Remark 9.2 Bounded semigroup. For a discussion of the bounded semigroups of page 132 see, for example, Marsden and Hughes [9.2].

Remark 9.3 The sign convention of the Poisson bracket. The Poisson bracket is antisymmetric in its arguments and there is a freedom to define it with either sign convention. When such freedom exists, it is certain that both conventions are in use and this is no exception. In some texts [1.4, 9.3] you will see the right hand side of (9.31) defined as $\{B, A\}$ so that (9.32) is $\frac{d Q}{d t}=\{Q, H\}$. Other equally reputable texts [25.2] employ the convention used here. Landau and Lifshitz [9.4] denote a Poisson bracket by $[A, B]$, notation that we reserve here for the quantummechanical commutator. As long as one is consistent, there should be no problem.

## Résumé

In physically realistic settings the initial state of a system can be specified only to a finite precision. If the dynamics is chaotic, it is not possible to calculate accurately the long time trajectory of a given initial point. Depending on the desired precision, and given a deterministic law of evolution, the state of the system can then be tracked for a finite time.

The study of long-time dynamics thus requires trading in the evolution of a single phase space point for the evolution of a measure, or the density of representative points in phase space, acted upon by an evolution operator. Essentially this means trading in nonlinear dynamical equations on finite dimensional spaces $x=\left(x_{1}, x_{2} \cdots x_{d}\right)$ for linear equations on infinite dimensional vector spaces of density functions $\rho(x)$. The most physical of stationary measures is the natural measure, a measure robust under perturbations by weak noise.

Reformulated this way, classical dynamics takes on a distinctly quantummechanical flavor. If the Lyapunov time (1.1), the time after which the notion of an individual deterministic trajectory loses meaning, is much shorter than the observation time, the "sharp" observables are those dual to time, the eigenvalues of evolution operators. This is very much the same situation as in quantum mechanics; as atomic time scales are so short, what is measured is the energy, the quantum-mechanical observable dual to the time. For long times the dynamics is described in terms of stationary measures, that is, fixed points of certain evolution operators. Both in classical and quantum mechanics one has a choice of implementing dynamical evolution on densities ("Schrödinger picture", sect. 9.4) or on observables ("Heisenberg picture", sect. 10.2 and chapter 14).

In what follows we shall find the second formulation more convenient, but the alternative is worth keeping in mind when posing and solving invariant density problems. However, as classical evolution operators are not unitary, their eigenstates can be quite singular and difficult to work with. In what follows we shall learn how to avoid dealing with these eigenstates
altogether. As a matter of fact, what follows will be a labor of radical deconstruction; after having argued so strenuously here that only smooth measures are "natural", we shall merrily proceed to erect the whole edifice of our theory on periodic orbits, that is, objects that are $\delta$-functions in phase space. The trick is that each comes with an interval, its neighborhood - cycle points only serve to pin these intervals, just as the millimeter marks on a measuring rod partition continuum into intervals.

## References

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[9.2] J.E. Marsden and T.J.R. Hughes, Mathematical Foundations of Elasticity (Prentice-Hall, Englewood Cliffs, New Jersey 1983)
[9.3] H. Goldstein, Classical Mechanics (Addison-Wesley, Reading, 1980).
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[9.5] P. Cvitanović, C.P. Dettmann, R. Mainieri and G. Vattay, Trace formulas for stochastic evolution operators: Weak noise perturbation theory, J. Stat. Phys. 93, 981 (1998); chao-dyn/9807034.
[9.6] A. Lasota and M.C. Mackey, Chaos, Fractals and Noise (Springer, New York 1994).
[9.7] We thank Laurette Tuckerman for providing us with the pseudospectral code.
[9.8] We have used the NAG library integration routine D02BHF. Our code and numerical results are available on ChaosBook.org/extras.

## Exercises

Exercise 9.1 Integrating over Dirac delta functions. Let us verify a few of the properties of the delta function and check (9.9), as well as the formulas (9.7) and (9.8) to be used later.
(a) If $f: \mathbb{R}^{d} \rightarrow \mathbb{R}^{d}$, show that

$$
\int_{\mathbb{R}^{d}} d x \delta(f(x))=\sum_{x \in f^{-1}(0)} \frac{1}{\left|\operatorname{det} \partial_{x} f\right|}
$$

(b) The delta function can be approximated by a sequence of Gaussians

$$
\int d x \delta(x) f(x)=\lim _{\sigma \rightarrow 0} \int d x \frac{e^{-\frac{x^{2}}{2 \sigma}}}{\sqrt{2 \pi \sigma}} f(x)
$$

Use this approximation to see whether the formal expression

$$
\int_{\mathbb{R}} d x \delta\left(x^{2}\right)
$$

makes sense.

Exercise 9.2 Derivatives of Dirac delta functions. Consider $\delta^{(k)}(x)=$ $\frac{\partial^{k}}{\partial x^{k}} \delta(x)$.
(a) Using integration by parts, determine the value of

$$
\int_{\mathbb{R}} d x \delta^{\prime}(y)
$$

where $y=f(x)-x$.
(b)

$$
\begin{equation*}
\int d x \delta^{(2)}(y)=\sum_{\{x: y(x)=0\}} \frac{1}{\left|y^{\prime}\right|}\left\{3 \frac{\left(y^{\prime \prime}\right)^{2}}{\left(y^{\prime}\right)^{4}}-\frac{y^{\prime \prime \prime}}{\left(y^{\prime}\right)^{3}}\right\} . \tag{9.39}
\end{equation*}
$$

(c)

$$
\begin{equation*}
\int d x b(x) \delta^{(2)}(y)=\sum_{\{x: y(x)=0\}} \frac{1}{\left|y^{\prime}\right|}\left\{\frac{b^{\prime \prime}}{\left(y^{\prime}\right)^{2}}-\frac{b^{\prime} y^{\prime \prime}}{\left(y^{\prime}\right)^{3}}+b\left(3 \frac{\left(y^{\prime \prime}\right)^{2}}{\left(y^{\prime}\right)^{4}}-\frac{y^{\prime \prime \prime}}{\left(y^{\prime}\right)^{3}}\right)\right\} . \tag{9.40}
\end{equation*}
$$

These formulas are useful for computing effects of weak noise on deterministic dynamics [9.5].

Exercise 9.3 $\quad \mathcal{L}^{t}$ generates a semigroup. Check that the Perron-Frobenius operator has the semigroup property,

$$
\begin{equation*}
\int_{M} d z \mathcal{L}^{t_{2}}(y, z) \mathcal{L}^{t_{1}}(z, x)=\mathcal{L}^{t_{2}+t_{1}}(y, x), \quad t_{1}, t_{2} \geq 0 \tag{9.41}
\end{equation*}
$$

As the flows in which we tend to be interested are invertible, the $\mathcal{L}$ 's that we will use often do form a group, with $t_{1}, t_{2} \in \mathbb{R}$.

## Exercise 9.4 Escape rate of the tent map.

(a) Calculate by numerical experimentation the log of the fraction of trajectories remaining trapped in the interval $[0,1]$ for the tent map

$$
f(x)=a(1-2|x-0.5|)
$$

for several values of $a$.
(b) Determine analytically the a dependence of the escape rate $\gamma(a)$.
(c) Compare your results for (a) and (b).

Exercise 9.5 Invariant measure. We will compute the invariant measure for two different piecewise linear maps.

(a) Verify the matrix $\mathcal{L}$ representation (10.19).
(b) The maximum value of the first map is 1. Compute an invariant measure for this map.
(c) Compute the leading eigenvalue of $\mathcal{L}$ for this map.
(d) For this map there is an infinite number of invariant measures, but only one of them will be found when one carries out a numerical simulation. Determine that measure, and explain why your choice is the natural measure for this map.
(e) In the second map the maximum occurs at $\alpha=(3-\sqrt{5}) / 2$ and the slopes are $\pm(\sqrt{5}+1) / 2$. Find the natural measure for this map. Show that it is piecewise linear and that the ratio of its two values is $(\sqrt{5}+1) / 2$.
(medium difficulty)

Exercise 9.6 Escape rate for a flow conserving map. Adjust $\Lambda_{0}, \Lambda_{1}$ in (10.17) so that the gap between the intervals $\mathcal{M}_{0}, \mathcal{M}_{1}$ vanishes. Show that the escape rate equals zero in this situation.

Exercise 9.7 Eigenvalues of the Perron-Frobenius operator for the skew Ulam tent map. Show that for the skew Ulam tent map

the eigenvalues are available analytically, compute the first few.
Exercise 9.8 "Kissing disks"* (continuation of exercises 6.1 and 6.2). Close off the escape by setting $R=2$, and look in real time at the density of the Poincaré section iterates for a trajectory with a randomly chosen initial condition. Does it look uniform? Should it be uniform? (Hint - phase space volumes are preserved for Hamiltonian flows by the Liouville theorem). Do you notice the trajectories that loiter near special regions of phase space for long times? These exemplify "intermittency", a bit of unpleasantness to which we shall return in chapter 21.

Exercise 9.9 Invariant measure for the Gauss map. Consider the Gauss map (we shall need this map in chapter 24):

$$
f(x)= \begin{cases}\frac{1}{x}-\left[\frac{1}{x}\right] & x \neq 0 \\ 0 & x=0\end{cases}
$$

where [ ] denotes the integer part.
(a) Verify that the density

$$
\rho(x)=\frac{1}{\log 2} \frac{1}{1+x}
$$

is an invariant measure for the map.
(b) Is it the natural measure?

Exercise $9.10 \mathcal{A}$ as a generator of translations. Verify that for a constant velocity field the evolution generator $\mathcal{A}$ in (9.28) is the generator of translations,

$$
e^{t v \frac{\partial}{\partial x}} a(x)=a(x+t v) .
$$

Exercise 9.11 Incompressible flows. Show that (9.9) implies that $\rho_{0}(x)=1$ is an eigenfunction of a volume-preserving flow with eigenvalue $s_{0}=0$. In particular, this implies that the natural measure of hyperbolic and mixing Hamiltonian flows is uniform. Compare this results with the numerical experiment of exercise 9.8.

## Chapter 10

## Averaging

For it, the mystic evolution;
Not the right only justified

- what we call evil also justified.

Walt Whitman,
Leaves of Grass: Song of the Universal

We start by discussing the necessity of studying the averages of observables in chaotic dynamics, and then cast the formulas for averages in a multiplicative form that motivates the introduction of evolution operators and further formal developments to come. The main result is that any dynamical average measurable in a chaotic system can be extracted from the spectrum of an appropriately constructed evolution operator. In order to keep our toes closer to the ground, in sect. 10.3 we try out the formalism on the first quantitative diagnosis that a system's got chaos, Lyapunov exponents.

### 10.1 Dynamical averaging

In chaotic dynamics detailed prediction is impossible, as any finitely specified initial condition, no matter how precise, will fill out the entire accessible phase space. Hence for chaotic dynamics one cannot follow individual trajectories for a long time; what is attainable is a description of the geometry of the set of possible outcomes, and evaluation of long time averages. Examples of such averages are transport coefficients for chaotic dynamical flows, such as escape rate, mean drift and diffusion rate; power spectra; and a host of mathematical constructs such as generalized dimensions, entropies and Lyapunov exponents. Here we outline how such averages are evaluated within the evolution operator framework. The key idea is to replace the expectation values of observables by the expectation values of generating functionals. This associates an evolution operator with a given observable, and relates the expectation value of the observable to the leading eigenvalue of the evolution operator.

### 10.1.1 Time averages

Let $a=a(x)$ be any observable, a function that associates to each point in phase space a number, a vector, or a tensor. The observable reports on a property of the dynamical system. It is a device, such as a thermometer or laser Doppler velocitometer. The device itself does not change during the measurement. The velocity field $a_{i}(x)=v_{i}(x)$ is an example of a vector observable; the length of this vector, or perhaps a temperature measured in an experiment at instant $\tau$ are examples of scalar observables. We define the integrated observable $A^{t}$ as the time integral of the observable $a$ evaluated along the trajectory of the initial point $x_{0}$,

$$
\begin{equation*}
A^{t}\left(x_{0}\right)=\int_{0}^{t} d \tau a\left(f^{\tau}\left(x_{0}\right)\right) . \tag{10.1}
\end{equation*}
$$

If the dynamics is given by an iterated mapping and the time is discrete, $t \rightarrow n$, the integrated observable is given by

$$
\begin{equation*}
A^{n}\left(x_{0}\right)=\sum_{k=0}^{n-1} a\left(f^{k}\left(x_{0}\right)\right) \tag{10.2}
\end{equation*}
$$

(we suppress possible vectorial indices for the time being). For example, if the observable is the velocity, $a_{i}(x)=v_{i}(x)$, its time integral $A_{i}^{t}\left(x_{0}\right)$ is the trajectory $A_{i}^{t}\left(x_{0}\right)=x_{i}(t)$. Another familiar example, for Hamiltonian flows, is the action associated with a trajectory $x(t)=[q(t), p(t)]$ passing through a phase space point $x_{0}=[q(0), p(0)]$ (this function will be the key to the semiclassical quantization of chapter 30):

$$
\begin{equation*}
A^{t}\left(x_{0}\right)=\int_{0}^{t} d \tau \dot{\mathbf{q}}(\tau) \cdot \mathbf{p}(\tau) \tag{10.3}
\end{equation*}
$$

The time average of the observable along a trajectory is defined by

$$
\begin{equation*}
\overline{a\left(x_{0}\right)}=\lim _{t \rightarrow \infty} \frac{1}{t} A^{t}\left(x_{0}\right) . \tag{10.4}
\end{equation*}
$$

If $a$ does not behave too wildly as a function of time - for example, if $a_{i}(x)$ is the Chicago temperature, bounded between $-80^{\circ} \mathrm{F}$ and $+130^{\circ} \mathrm{F}$ for all times - $A^{t}\left(x_{0}\right)$ is expected to grow not faster than $t$, and the limit (10.4) exists. For an example of a time average - the Lyapunov exponent - see sect. 10.3.

The time average depends on the trajectory, but not on the initial point on that trajectory: if we start at a later phase space point $f^{T}\left(x_{0}\right)$ we get a couple of extra finite contributions that vanish in the $t \rightarrow \infty$ limit:

$$
\overline{a\left(f^{T}\left(x_{0}\right)\right)}=\lim _{t \rightarrow \infty} \frac{1}{t} \int_{T}^{t+T} d \tau a\left(f^{\tau}\left(x_{0}\right)\right)
$$



Figure 10.1: (a) A typical chaotic trajectory explores the phase space with the long time visitation frequency building up the natural measure $\rho_{0}(x)$. (b) time average evaluated along an atypical trajectory such as a periodic orbit fails to explore the entire accessible phase space. (A. Johansen)

$$
\begin{aligned}
& =\overline{a\left(x_{0}\right)}-\lim _{t \rightarrow \infty} \frac{1}{t}\left(\int_{0}^{T} d \tau a\left(f^{\tau}\left(x_{0}\right)\right)-\int_{t}^{t+T} d \tau a\left(f^{\tau}\left(x_{0}\right)\right)\right) \\
& =\overline{a\left(x_{0}\right)} .
\end{aligned}
$$

The integrated observable $A^{t}\left(x_{0}\right)$ and the time average $\overline{a\left(x_{0}\right)}$ take a particularly simple form when evaluated on a periodic orbit. Define

$$
\begin{array}{ll}
\text { flows: } \quad A_{p} & =a_{p} T_{p}=\int_{0}^{T_{p}} a\left(f^{\tau}\left(x_{0}\right)\right) d \tau, \quad x_{0} \in p \\
\text { maps: } & =a_{p} n_{p}=\sum_{i=0}^{n_{p}-1} a\left(f^{i}\left(x_{0}\right)\right) \tag{10.5}
\end{array}
$$

where $p$ is a prime cycle, $T_{p}$ is its period, and $n_{p}$ is its discrete time period in the case of iterated map dynamics. $A_{p}$ is a loop integral of the observable along a single parcourse of a prime cycle $p$, so it is an intrinsic property of the cycle, independent of the starting point $x_{0} \in p$. (If the observable $a$ is not a scalar but a vector or matrix we might have to be more careful in defining an average which is independent of the starting point on the cycle). If the trajectory retraces itself $r$ times, we just obtain $A_{p}$ repeated $r$ times. Evaluation of the asymptotic time average (10.4) requires therefore only a single traversal of the cycle:

$$
\begin{equation*}
a_{p}=\frac{1}{T_{p}} A_{p} \tag{10.6}
\end{equation*}
$$

However, $\overline{a\left(x_{0}\right)}$ is in general a wild function of $x_{0}$; for a hyperbolic system ergodic with respect to a smooth measure, it takes the same value $\langle a\rangle$ for almost all initial $x_{0}$, but a different value (10.6) on any periodic orbit, that is, on a dense set of points (figure 10.1(b)). For example, for an open system such as the Sinai gas of sect. 23.1 (an infinite 2-dimensional periodic array of scattering disks) the phase space is dense with initial points that correspond to periodic runaway trajectories. The mean distance squared
traversed by any such trajectory grows as $x(t)^{2} \sim t^{2}$, and its contribution to the diffusion rate $D \approx x(t)^{2} / t$, (10.4) evaluated with $a(x)=x(t)^{2}$, diverges. Seemingly there is a paradox; even though intuition says the typical motion should be diffusive, we have an infinity of ballistic trajectories.

For chaotic dynamical systems, this paradox is resolved by robust averaging, that is, averaging also over the initial $x$, and worrying about the measure of the "pathological" trajectories.

### 10.1.2 Space averages

The space average of a quantity $a$ that may depend on the point $x$ of phase space $\mathcal{M}$ and on the time $t$ is given by the $d$-dimensional integral over the $d$ coordinates of the dynamical system:

$$
\begin{align*}
\langle a\rangle(t) & =\frac{1}{|\mathcal{M}|} \int_{\mathcal{M}} d x a(x(t)) \\
|\mathcal{M}| & =\int_{\mathcal{M}} d x=\text { volume of } \mathcal{M} \tag{10.7}
\end{align*}
$$

The space $\mathcal{M}$ is assumed to have finite dimension and volume (open systems like the 3 -disk game of pinball are discussed in sect. 10.1.3).

What is it we really do in experiments? We cannot measure the time average (10.4), as there is no way to prepare a single initial condition with infinite precision. The best we can do is to prepare some initial density $\rho(x)$ perhaps concentrated on some small (but always finite) neighborhood $\rho(x)=\rho(x, 0)$, so one should abandon the uniform space average (10.7), and consider instead

$$
\begin{equation*}
\langle a\rangle_{\rho}(t)=\frac{1}{|\mathcal{M}|} \int_{\mathcal{M}} d x \rho(x) a(x(t)) . \tag{10.8}
\end{equation*}
$$

We do not bother to lug the initial $\rho(x)$ around, as for the ergodic and mixing systems that we shall consider here any smooth initial density will tend to the asymptotic natural measure $t \rightarrow \infty$ limit $\rho(x, t) \rightarrow \rho_{0}(x)$, so we can just as well take the initial $\rho(x)=$ const. . The worst we can do is to start out with $\rho(x)=$ const., as in (10.7); so let us take this case and define the expectation value $\langle a\rangle$ of an observable $a$ to be the asymptotic time and space average over the phase space $\mathcal{M}$

$$
\begin{equation*}
\langle a\rangle=\lim _{t \rightarrow \infty} \frac{1}{|\mathcal{M}|} \int_{\mathcal{M}} d x \frac{1}{t} \int_{0}^{t} d \tau a\left(f^{\tau}(x)\right) . \tag{10.9}
\end{equation*}
$$

We use the same $\langle\cdots\rangle$ notation as for the space average (10.7), and distinguish the two by the presence of the time variable in the argument: if the
quantity $\langle a\rangle(t)$ being averaged depends on time, then it is a space average, if it does not, it is the expectation value $\langle a\rangle$.

The expectation value is a space average of time averages, with every $x \in \mathcal{M}$ used as a starting point of a time average. The advantage of averaging over space is that it smears over the starting points which were problematic for the time average (like the periodic points). While easy to define, the expectation value $\langle a\rangle$ turns out not to be particularly tractable in practice. Here comes a simple idea that is the basis of all that follows: Such averages are more conveniently studied by investigating instead of $\langle a\rangle$ the space averages of form

$$
\begin{equation*}
\left\langle e^{\beta \cdot A^{t}}\right\rangle=\frac{1}{|\mathcal{M}|} \int_{\mathcal{M}} d x e^{\beta \cdot A^{t}(x)} . \tag{10.10}
\end{equation*}
$$

In the present context $\beta$ is an auxiliary variable of no particular physical significance. In most applications $\beta$ is a scalar, but if the observable is a $d$-dimensional vector $a_{i}(x) \in \mathbb{R}^{d}$, then $\beta$ is a conjugate vector $\beta \in \mathbb{R}^{d}$; if the observable is a $d \times d$ tensor, $\beta$ is also a rank- 2 tensor, and so on. Here we will mostly limit the considerations to scalar values of $\beta$.

If the limit $\overline{a\left(x_{0}\right)}$ for the time average (10.4) exists for "almost all" initial $x_{0}$ and the system is ergodic and mixing (in the sense of sect. 1.3.1), we expect the time average along almost all trajectories to tend to the same value $\bar{a}$, and the integrated observable $A^{t}$ to tend to $t \bar{a}$. The space average (10.10) is an integral over exponentials, and such integral also grows exponentially with time. So as $t \rightarrow \infty$ we would expect the space average of $\left\langle\exp \left(\beta \cdot A^{t}\right)\right\rangle$ itself to grow exponentially with time

$$
\left\langle e^{\beta \cdot A^{t}}\right\rangle \propto e^{t s(\beta)},
$$

and its rate of growth to go to a limit

$$
\begin{equation*}
s(\beta)=\lim _{t \rightarrow \infty} \frac{1}{t} \ln \left\langle e^{\beta \cdot A^{t}}\right\rangle . \tag{10.11}
\end{equation*}
$$

Now we understand one reason for why it is smarter to compute $\left\langle\exp \left(\beta \cdot A^{t}\right)\right\rangle$ rather than $\langle a\rangle$ : the expectation value of the observable (10.9) and the moments of the integrated observable (10.1) can be computed by evaluating the derivatives of $s(\beta)$

$$
\begin{align*}
\left.\frac{\partial s}{\partial \beta}\right|_{\beta=0} & =\lim _{t \rightarrow \infty} \frac{1}{t}\left\langle A^{t}\right\rangle=\langle a\rangle \\
\left.\frac{\partial^{2} s}{\partial \beta^{2}}\right|_{\beta=0} & =\lim _{t \rightarrow \infty} \frac{1}{t}\left(\left\langle A^{t} A^{t}\right\rangle-\left\langle A^{t}\right\rangle\left\langle A^{t}\right\rangle\right)  \tag{10.12}\\
& =\lim _{t \rightarrow \infty} \frac{1}{t}\left\langle\left(A^{t}-t\langle a\rangle\right)^{2}\right\rangle
\end{align*}
$$

and so forth. We have written out the formulas for a scalar observable; the vector case is worked out in the exercise 10.2. If we can compute the function $s(\beta)$, we have the desired expectation value without having to estimate any infinite time limits from finite time data.

Suppose we could evaluate $s(\beta)$ and its derivatives. What are such formulas good for? A typical application is to the problem of describing a particle scattering elastically off a 2-dimensional triangular array of disks. If the disks are sufficiently large to block any infinite length free flights, the particle will diffuse chaotically, and the transport coefficient of interest is the diffusion constant given by $\left\langle x(t)^{2}\right\rangle \approx 4 D t$. In contrast to $D$ estimated numerically from trajectories $x(t)$ for finite but large $t$, the above formulas yield the asymptotic $D$ without any extrapolations to the $t \rightarrow \infty$ limit. For example, for $a_{i}=v_{i}$ and zero mean drift $\left\langle v_{i}\right\rangle=0$, the diffusion constant is given by the curvature of $s(\beta)$ at $\beta=0$,

$$
\begin{equation*}
D=\lim _{t \rightarrow \infty} \frac{1}{2 d t}\left\langle x(t)^{2}\right\rangle=\left.\frac{1}{2 d} \sum_{i=1}^{d} \frac{\partial^{2} s}{\partial \beta_{i}^{2}}\right|_{\beta=0}, \tag{10.13}
\end{equation*}
$$

so if we can evaluate derivatives of $s(\beta)$, we can compute transport coefficients that characterize deterministic diffusion. As we shall see in chapter 23, periodic orbit theory yields an explicit closed form expression for D.

fast track: sect. 10.2, p. 148

### 10.1.3 Averaging in open systems

- ${ }^{3}$
If the $\mathcal{M}$ is a compact region or set of regions to which the dynamics is confined for all times, (10.9) is a sensible definition of the expectation value. However, if the trajectories can exit $\mathcal{M}$ without ever returning,

$$
\int_{\mathcal{M}} d y \delta\left(y-f^{t}\left(x_{0}\right)\right)=0 \quad \text { for } t>t_{\text {exit }}, \quad x_{0} \in \mathcal{M}
$$

we might be in trouble. In particular, for a repeller the trajectory $f^{t}\left(x_{0}\right)$ will eventually leave the region $\mathcal{M}$, unless the initial point $x_{0}$ is on the repeller, so the identity

$$
\int_{\mathcal{M}} d y \delta\left(y-f^{t}\left(x_{0}\right)\right)=1, \quad t>0, \quad \text { iff } x_{0} \in \text { non-wandering } \operatorname{set}(10.14)
$$

might apply only to a fractal subset of initial points a set of zero Lebesgue measure. Clearly, for open systems we need to modify the definition of the

Figure 10.2: A piecewise-linear repeller (10.17): All trajectories that land in the gap between the $f_{0}$ and $f_{1}$ branches escape ( $\Lambda_{0}=4$, $\left.\Lambda_{1}=-2\right)$.

expectation value to restrict it to the dynamics on the non-wandering set, the set of trajectories which are confined for all times.

Note by $\mathcal{M}$ a phase space region that encloses all interesting initial points, say the 3 -disk Poincaré section constructed from the disk boundaries and all possible incidence angles, and denote by $|\mathcal{M}|$ the volume of $\mathcal{M}$. The volume of the phase space containing all trajectories which start out within the phase space region $\mathcal{M}$ and recur within that region at the time $t$

$$
\begin{equation*}
|\mathcal{M}(t)|=\int_{\mathcal{M}} d x d y \delta\left(y-f^{t}(x)\right) \sim|\mathcal{M}| e^{-\gamma t} \tag{10.15}
\end{equation*}
$$

is expected to decrease exponentially, with the escape rate $\gamma$. The integral over $x$ takes care of all possible initial points; the integral over $y$ checks whether their trajectories are still within $\mathcal{M}$ by the time $t$. For example, any trajectory that falls off the pinball table in figure 1.1 is gone for good.

The non-wandering set can be very difficult object to describe; but for any finite time we can construct a normalized measure from the finite-time covering volume (10.15), by redefining the space average (10.10) as

$$
\begin{equation*}
\left\langle e^{\beta \cdot A^{t}}\right\rangle=\int_{\mathcal{M}} d x \frac{1}{|\mathcal{M}(t)|} e^{\beta \cdot A^{t}(x)} \sim \frac{1}{|\mathcal{M}|} \int_{\mathcal{M}} d x e^{\beta \cdot A^{t}(x)+\gamma t} . \tag{10.16}
\end{equation*}
$$

in order to compensate for the exponential decrease of the number of surviving trajectories in an open system with the exponentially growing factor $e^{\gamma t}$. What does this mean? Once we have computed $\gamma$ we can replenish the density lost to escaping trajectories, by pumping in $e^{\gamma t}$ in such a way that the overall measure is correctly normalized at all times, $\langle 1\rangle=1$.

## Example 10.1 A piecewise-linear example: (continuation of exercise 9.1)

 What is gained by reformulating the dynamics in terms of "operators"? We start by considering a simple example in which the operator is a $[2 \times 2]$ matrix. Assume the expanding 1-d map $f(x)$ of figure 10.2, a piecewise-linear 2-branch repeller with slopes $\Lambda_{0}>1$ and $\Lambda_{1}<-1$ :$$
f(x)=\left\{\begin{array}{ll}
f_{0}=\Lambda_{0} x & \text { if } \quad x \in \mathcal{M}_{0}=\left[0,1 / \Lambda_{0}\right]  \tag{10.17}\\
f_{1}=\Lambda_{1}(x-1) & \text { if } x \in \mathcal{M}_{1}=\left[1+1 / \Lambda_{1}, 1\right]
\end{array} .\right.
$$

Both $f\left(\mathcal{M}_{0}\right)$ and $f\left(\mathcal{M}_{1}\right)$ map onto the entire unit interval $\mathcal{M}=[0,1]$. Assume a piecewise constant density

$$
\rho(x)=\left\{\begin{array}{ll}
\rho_{0} & \text { if } x \in \mathcal{M}_{0}  \tag{10.18}\\
\rho_{1} & \text { if } x \in \mathcal{M}_{1}
\end{array} .\right.
$$

There is no need to define $\rho(x)$ in the gap between $\mathcal{M}_{0}$ and $\mathcal{M}_{1}$, as any point that lands in the gap escapes.

The physical motivation for studying this kind of mapping is the pinball game: $f$ is the simplest model for the pinball escape, figure 1.7, with $f_{0}$ and $f_{1}$ modelling its two strips of survivors.

As can be easily checked using (9.9), the Perron-Frobenius operator acts on this piecewise constant function as a [2×2] "transfer" matrix with matrix elements

$$
\binom{\rho_{0}}{\rho_{1}} \rightarrow \mathcal{L} \rho=\left(\begin{array}{cc}
\frac{1}{\left|\Lambda_{0}\right|} & \frac{1}{\left|\Lambda_{1}\right|}  \tag{10.19}\\
\frac{1}{\left|\Lambda_{0}\right|} & \frac{1}{\left|\Lambda_{1}\right|}
\end{array}\right)\binom{\rho_{0}}{\rho_{1}},
$$

stretching both $\rho_{0}$ and $\rho_{1}$ over the whole unit interval $\Lambda$, and decreasing the density at every iteration. In this example the density is constant after one iteration, so $\mathcal{L}$ has only one non-zero eigenvalue $e^{s_{0}}=1 /\left|\Lambda_{0}\right|+1 /\left|\Lambda_{1}\right|$, with constant density eigenvector $\rho_{0}=\rho_{1}$. The quantities $1 /\left|\Lambda_{0}\right|, 1 /\left|\Lambda_{1}\right|$ are, respectively, the sizes of the $\left|\mathcal{M}_{0}\right|,\left|\mathcal{M}_{1}\right|$ intervals, so the exact escape rate (1.3) - the log of the fraction of survivors at each iteration for this linear repeller - is given by the sole eigenvalue of $\mathcal{L}$ :

$$
\begin{equation*}
\gamma=-s_{0}=-\ln \left(1 /\left|\Lambda_{0}\right|+1 /\left|\Lambda_{1}\right|\right) . \tag{10.20}
\end{equation*}
$$

Voila! Here is the rationale for introducing operators - in one time step we have solved the problem of evaluating escape rates at infinite time. This simple explicit matrix representation of the Perron-Frobenius operator is a consequence of the piecewise linearity of $f$, and the restriction of the densities $\rho$ to the space of piecewise constant functions. The example gives a flavor of the enterprise upon which we are about to embark in this book, but the full story is much subtler: in general, there will exist no such finite-dimensional representation for the Perron-Frobenius operator.

We now turn to the problem of evaluating $\left\langle e^{\beta \cdot A^{t}}\right\rangle$.

### 10.2 Evolution operators

The above simple shift of focus, from studying $\langle a\rangle$ to studying $\left\langle\exp \left(\beta \cdot A^{t}\right)\right\rangle$ is the key to all that follows. Make the dependence on the flow explicit by rewriting this quantity as

$$
\begin{equation*}
\left\langle e^{\beta \cdot A^{t}}\right\rangle=\frac{1}{|\mathcal{M}|} \int_{\mathcal{M}} d x \int_{\mathcal{M}} d y \delta\left(y-f^{t}(x)\right) e^{\beta \cdot A^{t}(x)} \tag{10.21}
\end{equation*}
$$

Here $\delta\left(y-f^{t}(x)\right)$ is the Dirac delta function: for a deterministic flow an initial point $x$ maps into a unique point $y$ at time $t$. Formally, all we have done above is to insert the identity

$$
\begin{equation*}
1=\int_{\mathcal{M}} d y \delta\left(y-f^{t}(x)\right) \tag{10.22}
\end{equation*}
$$

Figure 10.3: Space averaging pieces together the time average computed along the $t \rightarrow \infty$ trajectory of figure 10.1 by a space average over infinitely many short $t$ trajectory segments starting at all initial points at once. (A. Johansen)

into (10.10) to make explicit the fact that we are averaging only over the trajectories that remain in $\mathcal{M}$ for all times. However, having made this substitution we have replaced the study of individual trajectories $f^{t}(x)$ by the study of the evolution of density of the totality of initial conditions. Instead of trying to extract a temporal average from an arbitrarily long trajectory which explores the phase space ergodically, we can now probe the entire phase space with short (and controllable) finite time pieces of trajectories originating from every point in $\mathcal{M}$.

As a matter of fact (and that is why we went to the trouble of defining the generator (9.25) of infinitesimal transformations of densities) infinitesimally short time evolution can suffice to determine the spectrum and eigenvalues of $\mathcal{L}^{t}$.

We shall refer to the kernel of $\mathcal{L}^{t}=e^{t \mathcal{A}}$ in the phase-space representation (10.21) as the evolution operator

$$
\begin{equation*}
\mathcal{L}^{t}(y, x)=\delta\left(y-f^{t}(x)\right) e^{\beta \cdot A^{t}(x)} . \tag{10.23}
\end{equation*}
$$

The simplest example is the Perron-Frobenius operator introduced in section 5.2. Another example - designed to deliver the Lyapunov exponent - will be the evolution operator (10.35). The evolution operator acts on scalar functions $\phi(x)$ as

$$
\begin{equation*}
\mathcal{L}^{t} \phi(y)=\int_{\mathcal{M}} d x \delta\left(y-f^{t}(x)\right) e^{\beta \cdot A^{t}(x)} \phi(x) . \tag{10.24}
\end{equation*}
$$

In terms of the evolution operator, the expectation value of the generating function (10.21) is given by

$$
\left\langle e^{\beta \cdot A^{t}}\right\rangle=\left\langle\mathcal{L}^{t} \iota\right\rangle,
$$

where the initial density $\iota(x)$ is the constant function that always returns 1.

The evolution operator is different for different observables, as its definition depends on the choice of the integrated observable $A^{t}$ in the exponential. Its job is deliver to us the expectation value of $a$, but before showing that it accomplishes that, we need to verify the semigroup property of evolution operators.

By its definition, the integral over the observable $a$ is additive along the trajectory

$$
\begin{array}{rlrl}
\mathrm{x}(0) & \mathrm{x}^{\mathrm{x}\left(\mathrm{t}_{1}+\mathrm{t}_{2}\right)} & \left.={ }^{\mathrm{x}\left(\mathrm{t}_{1}\right)}>^{\mathrm{x}\left(\mathrm{t}_{1}+\mathrm{t}_{1}\right)}+\mathrm{t}_{2}\right) \\
A^{t_{1}+t_{2}}\left(x_{0}\right) & =\int_{0}^{t_{1}} d \tau a(x(\tau)) & +\int_{t_{1}}^{t_{1}+t_{2}} d \tau a(x(\tau)) \\
& =A^{t_{1}}\left(x_{0}\right) & & +A^{t_{2}}\left(f^{t_{1}}\left(x_{0}\right)\right) .
\end{array}
$$

If $A^{t}(x)$ is additive along the trajectory, the evolution operator generates a semigroup

$$
\begin{equation*}
\mathcal{L}^{t_{1}+t_{2}}(y, x)=\int_{\mathcal{M}} d z \mathcal{L}^{t_{2}}(y, z) \mathcal{L}^{t_{1}}(z, x) \tag{10.25}
\end{equation*}
$$

as is easily checked by substitution

$$
\mathcal{L}^{t_{2}} \mathcal{L}^{t_{1}} a(x)=\int_{\mathcal{M}} d y \delta\left(x-f^{t_{2}}(y)\right) e^{\beta \cdot A^{t_{2}}(y)}\left(\mathcal{L}^{t_{1}} a\right)(y)=\mathcal{L}^{t_{1}+t_{2}} a(x) .
$$

This semigroup property is the main reason why (10.21) is preferable to (10.9) as a starting point for evaluation of dynamical averages: it recasts averaging in form of operators multiplicative along the flow.

### 10.3 Lyapunov exponents

(J. Mathiesen and P. Cvitanović)

Let us apply the newly acquired tools to the fundamental diagnostics in this subject: Is a given system "chaotic"? And if so, how chaotic? If all points in a neighborhood of a trajectory converge toward the same trajectory, the attractor is a fixed point or a limit cycle. However, if the attractor is strange, two trajectories

$$
\begin{equation*}
x(t)=f^{t}\left(x_{0}\right) \quad \text { and } \quad x(t)+\delta x(t)=f^{t}\left(x_{0}+\delta x(0)\right) \tag{10.26}
\end{equation*}
$$

that start out very close to each other separate exponentially with time, and in a finite time their separation attains the size of the accessible phase space. This sensitivity to initial conditions can be quantified as

$$
\begin{equation*}
|\delta x(t)| \approx e^{\lambda t}|\delta x(0)| \tag{10.27}
\end{equation*}
$$

where $\lambda$, the mean rate of separation of trajectories of the system, is called the Lyapunov exponent.

### 10.3.1 Lyapunov exponent as a time average

We can start out with a small $\delta x$ and try to estimate $\lambda$ from (10.27), but now that we have quantified the notion of linear stability in chapter 4 and defined the dynamical time averages in sect. 10.1.1, we can do better. The problem with measuring the growth rate of the distance between two points is that as the points separate, the measurement is less and less a local measurement. In study of experimental time series this might be the only option, but if we have the equations of motion, a better way is to measure the growth rate of vectors transverse to a given orbit.

The mean growth rate of the distance $|\delta x(t)| /|\delta x(0)|$ between neighboring trajectories (10.27) is given by the Lyapunov exponent

$$
\begin{equation*}
\lambda=\lim _{t \rightarrow \infty} \frac{1}{t} \ln |\delta x(t)| /|\delta x(0)| \tag{10.28}
\end{equation*}
$$

(For notational brevity we shall often suppress the dependence of quantities such as $\lambda=\lambda\left(x_{0}\right), \delta x(t)=\delta x\left(x_{0}, t\right)$ on the initial point $x_{0}$ and the time $\left.t\right)$. For infinitesimal $\delta x$ we know the $\delta x_{i}(t) / \delta x_{j}(0)$ ratio exactly, as this is by definition the Jacobian matrix (4.30)

$$
\lim _{\delta x \rightarrow 0} \frac{\delta x_{i}(t)}{\delta x_{j}(0)}=\frac{\partial x_{i}(t)}{\partial x_{j}(0)}=\mathbf{J}_{i j}^{t}\left(x_{0}\right),
$$

so the leading Lyapunov exponent can be computed from the linear approximation (4.22)

$$
\begin{equation*}
\lambda=\lim _{t \rightarrow \infty} \frac{1}{t} \ln \frac{\left|\mathbf{J}^{t}\left(x_{0}\right) \delta x(0)\right|}{|\delta x(0)|}=\lim _{t \rightarrow \infty} \frac{1}{2 t} \ln \left|\hat{n}^{T}\left(\mathbf{J}^{t}\right)^{T} \mathbf{J}^{t} \hat{n}\right| . \tag{10.29}
\end{equation*}
$$

In this formula the scale of the initial separation drops out, only its orientation given by the unit vector $\hat{n}=\delta x /|\delta x|$ matters. The eigenvalues of $\mathbf{J}$ are either real or come in complex conjugate pairs. As $\mathbf{J}$ is in general not symmetric and not diagonalizable, it is more convenient to work with the symmetric and diagonalizable matrix $\mathbf{M}=\left(\mathbf{J}^{t}\right)^{T} \mathbf{J}^{t}$, with real positive eigenvalues $\left\{\left|\Lambda_{1}\right|^{2} \geq \ldots \geq\left|\Lambda_{d}\right|^{2}\right\}$, and a complete orthonormal set of eigenvectors of $\left\{u_{1}, \ldots, u_{d}\right\}$. Expanding the initial orientation $\hat{n}=\sum\left(\hat{n} \cdot u_{i}\right) u_{i}$ in the $\mathbf{M} u_{i}=\left|\Lambda_{i}\right| u_{i}$ eigenbasis, we have

$$
\begin{equation*}
\hat{n}^{T} \mathbf{M} \hat{n}=\sum_{i=1}^{d}\left(\hat{n} \cdot u_{i}\right)^{2}\left|\Lambda_{i}\right|^{2}=\left(\hat{n} \cdot u_{1}\right)^{2} e^{2 \lambda_{1} t}\left(1+O\left(e^{-2\left(\lambda_{1}-\lambda_{2}\right) t}\right)\right), \tag{10.30}
\end{equation*}
$$

where $t \lambda_{i}=\ln \left|\Lambda_{i}\left(x_{0}, t\right)\right|$, and we assume that $\lambda_{1}>\lambda_{2} \geq \lambda_{3} \cdots$. For long times the largest Lyapunov exponent dominates exponentially (10.29), provided the orientation $\hat{n}$ of the initial separation was not chosen perpendicular to the dominant expanding eigendirection $u_{1}$. The Lyapunov exponent

Figure 10.4: A numerical estimate of the leading Lyapunov exponent for the Rössler system (2.14) from the dominant expanding eigenvalue formula (10.29). The leading Lyapunov exponent $\lambda \approx 0.09$ is positive, so numerics supports the hypothesis that the Rössler attractor is strange. (J. Mathiesen)

is the time average

$$
\begin{align*}
\overline{\lambda\left(x_{0}\right)} & =\lim _{t \rightarrow \infty} \frac{1}{t}\left\{\ln \left|\hat{n} \cdot u_{1}\right|+\ln \left|\Lambda_{1}\left(x_{0}, t\right)\right|+O\left(e^{-2\left(\lambda_{1}-\lambda_{2}\right) t}\right)\right\} \\
& =\lim _{t \rightarrow \infty} \frac{1}{t} \ln \left|\Lambda_{1}\left(x_{0}, t\right)\right|, \tag{10.31}
\end{align*}
$$

where $\Lambda_{1}\left(x_{0}, t\right)$ is the leading eigenvalue of $\mathbf{J}^{t}\left(x_{0}\right)$. By choosing the initial displacement such that $\hat{n}$ is normal to the first ( $i-1$ ) eigendirections we can define not only the leading, but all Lyapunov exponents as well:

$$
\begin{equation*}
\overline{\lambda_{i}\left(x_{0}\right)}=\lim _{t \rightarrow \infty} \frac{1}{t} \ln \left|\Lambda_{i}\left(x_{0}, t\right)\right|, \quad i=1,2, \cdots, d \tag{10.32}
\end{equation*}
$$

The leading Lyapunov exponent now follows from the Jacobian matrix by numerical integration of (4.32). The equations can be integrated accurately for a finite time, hence the infinite time limit of (10.29) can be only estimated from plots of $\frac{1}{2} \ln \left|\hat{n}^{T} \mathbf{M} \hat{n}\right|$ as function of time, such as the figure 10.4 for the Rössler system (2.14). As the local expansion and contraction rates vary along the flow, the temporal dependence exhibits small and large humps. The sudden fall to a low level is caused by a close passage to a folding point of the attractor, an illustration of why numerical evaluation of the Lyapunov exponents, and proving the very existence of a strange attractor is a very difficult problem. The approximately monotone part of the curve can be used (at your own peril) to estimate the leading Lyapunov exponent by a straight line fit.

As we can already see, we are courting difficulties if we try to calculate the Lyapunov exponent by using the definition (10.31) directly. First of all, the phase space is dense with atypical trajectories; for example, if $x_{0}$ happened to lie on a periodic orbit $p, \bar{\lambda}$ would be simply $\ln \left|\Lambda_{p}\right| / T_{p}$, a local property of cycle $p$, not a global property of the dynamical system. Furthermore, even if $x_{0}$ happens to be a "generic" phase space point, it is still not obvious that $\ln \left|\Lambda\left(x_{0}, t\right)\right| / t$ should be converging to anything in particular. In a Hamiltonian system with coexisting elliptic islands and chaotic regions, a chaotic trajectory gets every so often captured in the neighborhood of an elliptic island and can stay there for arbitrarily long time; as there the orbit is nearly stable, during such episode $\ln \left|\Lambda\left(x_{0}, t\right)\right| / t$ can dip arbitrarily close to $0^{+}$. For phase space volume non-preserving flows the trajectory can traverse locally contracting regions, and $\ln \left|\Lambda\left(x_{0}, t\right)\right| / t$ can occasionally go negative; even worse, one never knows whether the asymptotic attractor is periodic or "strange", so any finite estimate of $\bar{\lambda}$ might be dead wrong.

### 10.3.2 Evolution operator evaluation of Lyapunov exponents

A cure to these problems was offered in sect. 10.2. We shall now replace time averaging along a single trajectory by action of a multiplicative evolution operator on the entire phase space, and extract the Lyapunov exponent from its leading eigenvalue. If the chaotic motion fills the whole phase space, we are indeed computing the asymptotic Lyapunov exponent. If the chaotic motion is transient, leading eventually to some long attractive cycle, our Lyapunov exponent, computed on nonwandering set, will characterize the chaotic transient; this is actually what any experiment would measure, as even very small amount of external noise will suffice to destabilize a long stable cycle with a minute immediate basin of attraction.

Due to the chain rule (4.35) for the derivative of an iterated map, the stability of a $1-d$ mapping is multiplicative along the flow, so the integral (10.1) of the observable $a(x)=\ln \left|f^{\prime}(x)\right|$, the local trajectory divergence rate, evaluated along the trajectory of $x_{0}$ is additive:

$$
\begin{equation*}
A^{n}\left(x_{0}\right)=\ln \left|f^{n \prime}\left(x_{0}\right)\right|=\sum_{k=0}^{n-1} \ln \left|f^{\prime}\left(x_{k}\right)\right| . \tag{10.33}
\end{equation*}
$$

The Lyapunov exponent is then the expectation value (10.9) given by a spatial integral (10.8) weighted by the natural measure

$$
\begin{equation*}
\lambda=\langle\ln | f^{\prime}(x)| \rangle=\int_{\mathcal{M}} d x \rho_{0}(x) \ln \left|f^{\prime}(x)\right| . \tag{10.34}
\end{equation*}
$$

The associated (discrete time) evolution operator (10.23) is

$$
\begin{equation*}
\mathcal{L}(y, x)=\delta(y-f(x)) e^{\beta \ln \left|f^{\prime}(x)\right|} \tag{10.35}
\end{equation*}
$$

Here we have restricted our considerations to $1-d$ maps, as for higherdimensional flows only the Jacobian matrices are multiplicative, not the individual eigenvalues. Construction of the evolution operator for evaluation of the Lyapunov spectra in the general case requires more cleverness than warranted at this stage in the narrative: an extension of the evolution equations to a flow in the tangent space.


All that remains is to determine the value of the Lyapunov exponent

$$
\begin{equation*}
\lambda=\langle\ln | f^{\prime}(x)| \rangle=\left.\frac{\partial s(\beta)}{\partial \beta}\right|_{\beta=1}=s^{\prime}(1) \tag{10.36}
\end{equation*}
$$

from (10.12), the derivative of the leading eigenvalue $s_{0}(\beta)$ of the evolution operator (10.35). The only question is: how?

### 10.4 Why not just run it on a computer?

(R. Artuso and P. Cvitanović)

All of the insight gained in this chapter and the preceding one was nothing but an elegant way of thinking of the evolution operator, $\mathcal{L}$, as a matrix (this point of view will be further elaborated in chapter 16). There are many textbook methods of approximating an operator $\mathcal{L}$ by sequences of finite matrix approximations $\mathcal{L}$, but in what follows the great achievement will be that we shall avoid constructing any matrix approximation to $\mathcal{L}$ altogether. Why a new method? Why not just run it on a computer, as many do with such relish in diagonalizing quantum Hamiltonians?

The simplest possible way of introducing a phase space discretization, figure 10.5 , is to partition the phase space $\mathcal{M}$ with a non-overlapping collection of sets $\mathcal{M}_{i}, i=1, \ldots, N$, and to consider densities (9.2) that are locally constant on each $\mathcal{M}_{i}$ :

$$
\rho(x)=\sum_{i=1}^{N} \rho_{i} \frac{\chi_{i}(x)}{\left|\mathcal{M}_{i}\right|}
$$

where $\chi_{i}(x)$ is the characteristic function (9.1) of the set $\mathcal{M}_{i}$. The density $\rho_{i}$ at a given instant is related to the densities at the previous step in time by the action of the Perron-Frobenius operator, as in (9.6):

$$
\begin{aligned}
\rho_{j}^{\prime} & =\int_{\mathcal{M}} d y \chi_{j}(y) \rho^{\prime}(y)=\int_{\mathcal{M}} d x d y \chi_{j}(y) \delta(y-f(x)) \rho(x) \\
& =\sum_{i=1}^{N} \rho_{i} \frac{\left|\mathcal{M}_{i} \cap f^{-1}\left(\mathcal{M}_{j}\right)\right|}{\left|\mathcal{M}_{i}\right|} .
\end{aligned}
$$

In this way

$$
\begin{equation*}
\mathbf{L}_{i j}=\frac{\left|\mathcal{M}_{i} \cap f^{-1}\left(\mathcal{M}_{j}\right)\right|}{\left|\mathcal{M}_{i}\right|}, \quad \rho^{\prime}=\rho \mathbf{L} \tag{10.37}
\end{equation*}
$$

is a matrix approximation to the Perron-Frobenius operator, and its leading left eigenvector is a piecewise constant approximation to the invariant measure. It is an old idea of Ulam that such an approximation for the Perron-Frobenius operator is a meaningful one.

The problem with such phase space discretization approaches is that they are blind, the grid knows not what parts of the phase space are more or less important. This observation motivates the next step in developing the theory of long-time dynamics of chaotic systems: in chapter 11 we shall exploit the intrinsic topology of the flow to give us both an invariant

Figure 10.5: Phase space discretization approach to computing averages.

partition of the phase space and a measure of the partition volumes, in the spirit of figure 1.9.

Furthermore, a piecewise constant $\rho$ belongs to an unphysical function space, and with such approximations one is plagued by numerical artifacts such as spurious eigenvalues. In chapter 16 we shall employ a more refined approach to extracting spectra, by expanding the initial and final densities $\rho, \rho^{\prime}$ in some basis $\varphi_{0}, \varphi_{1}, \varphi_{2}, \cdots$ (orthogonal polynomials, let us say), and replacing $\mathcal{L}(y, x)$ by its $\varphi_{\alpha}$ basis representation $\mathbf{L}_{\alpha \beta}=\left\langle\varphi_{\alpha}\right| \mathcal{L}\left|\varphi_{\beta}\right\rangle$. The art is then the subtle art of finding a "good" basis for which finite truncations of $\mathbf{L}_{\alpha \beta}$ give accurate estimates of the eigenvalues of $\mathcal{L}$.

Regardless of how sophisticated the choice of basis might be, the basic problem cannot be avoided - as illustrated by the natural measure for the Hénon map (3.12) sketched in figure 9.3 , eigenfunctions of $\mathcal{L}$ are complicated, singular functions concentrated on fractal sets, and in general cannot be represented by a nice basis set of smooth functions. We shall resort to matrix representations of $\mathcal{L}$ and the $\varphi_{\alpha}$ basis approach only insofar this helps us prove that the spectrum that we compute is indeed the correct one, and that finite periodic orbit truncations do converge.


## Commentary

Remark 10.1 "Pressure". The quantity $\left\langle\exp \left(\beta \cdot A^{t}\right)\right\rangle$ is called a "partition function" by Ruelle [15.1]. Mathematicians decorate it with considerably more Greek and Gothic letters than is the case in this treatise. Ruelle [15.2] and Bowen [10.1] had given name "pressure" $P(a)$ to $s(\beta)$ (where $a$ is the observable introduced here in sect. 10.1.1), defined by the "large system" limit (10.11). As we shall apply the theory also to computation of the physical gas pressure exerted on the walls of a container by a bouncing particle, we prefer to $s(\beta)$ as simply the leading
eigenvalue of the evolution operator introduced in sect. 9.4. The "convexity" properties such as $P(a) \leq P(|a|)$ will be pretty obvious consequence of the definition (10.11). In the case that $\mathcal{L}$ is the Perron-Frobenius operator (9.10), the eigenvalues $\left\{s_{0}(\beta), s_{1}(\beta), \cdots\right\}$ are called the Ruelle-Pollicott resonances [10.2, 10.3, 10.4], with the leading one, $s(\beta)=s_{0}(\beta)$ being the one of main physical interest. In order to aid the reader in digesting the mathematics literature, we shall try to point out the notational correspondences whenever appropriate. The rigorous formalism is replete with lims, sups, infs, $\Omega$-sets which are not really essential to understanding the physical applications of the theory, and are avoided in this presentation.

Remark 10.2 Microcanonical ensemble. In statistical mechanics the space average (10.7) performed over the Hamiltonian system constant energy surface invariant measure $\rho(x) d x=\operatorname{dqdp} \delta(H(q, p)-E)$ of volume $|\mathcal{M}|=\int_{\mathcal{M}} d q d p \delta(H(q, p)-E)$

$$
\begin{equation*}
\langle a(t)\rangle=\frac{1}{|\mathcal{M}|} \int_{\mathcal{M}} d q d p \delta(H(q, p)-E) a(q, p, t) \tag{10.38}
\end{equation*}
$$

is called the microcanonical ensemble average.

Remark 10.3 Lyapunov exponents. The Multiplicative Ergodic Theorem of Oseledec [10.5] states that the limit (10.32) exists for almost all points $x_{0}$ and all tangent vectors $\hat{n}$. There are at most $d$ distinct values of $\lambda$ as we let $\hat{n}$ range over the tangent space. These are the Lyapunov exponents [10.6] $\lambda_{i}\left(x_{0}\right)$.

There is much literature on numerical computation of the Lyapunov exponents, see for example refs. [4.1, 10.12, 10.13, 10.14].

## Résumé

The expectation value $\langle a\rangle$ of an observable $a(x)$ measured $A^{t}(x)=\int_{0}^{t} d \tau a(x(\tau))$ and averaged along the flow $x \rightarrow f^{t}(x)$ is given by the derivative $\partial s / \partial \beta$ of the leading eigenvalue $e^{t s(\beta)}$ of the corresponding evolution operator $\mathcal{L}^{t}$.

Using the Perron-Frobenius operator (9.10) whose leading eigenfunction, the natural measure, once computed, yields expectation value (9.19) of any observable $a(x)$ a separate evolution operator $\mathbf{L}$ has to be constructed for each and every observable. However, by the time the scaffolding is removed both $\mathcal{L}$ 's and their eigenfunctions will be gone, and only the formulas for expectation value of observables will remain.

The next question is: how do we evaluate the eigenvalues of $\mathcal{L}$ ? We saw in example 10.1, in the case of piecewise-linear dynamical systems, that these operators reduce to finite matrices, but for generic smooth flows, they are infinite-dimensional linear operators, and finding smart ways of computing their eigenvalues requires some thought. In chapter 11 we take the first step, and replace the ad hoc partitioning (10.37) by the intrinsic, topologically invariant partitioning. In chapter 13 we apply this information
to our first application of the evolution operator formalism, evaluation of the topological entropy, the growth rate of the number of topologically distinct orbits. This small victory will then be refashioned in chapters 14 and 15 into a systematic method for computing eigenvalues of evolution operators in terms of periodic orbits.

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## Exercises

## Exercise 10.1 How unstable is the Hénon attractor?

(a) Evaluate numerically the Lyapunov exponent by iterating the Hénon map

$$
\left[\begin{array}{l}
x^{\prime} \\
y^{\prime}
\end{array}\right]=\left[\begin{array}{l}
1-a x^{2}+y \\
b x
\end{array}\right]
$$

for $a=1.4, b=0.3$.
(b) Now check how robust is the Lyapunov exponent for the Hénon attractor? Evaluate numerically the Lyapunov exponent by iterating the Hénon map for $a=1.39945219, b=0.3$. How much do you trust now your result for the part (a) of this exercise?

Exercise 10.2 Expectation value of a vector observable and its moments. Check and extend the expectation value formulas (10.12) by evaluating the derivatives of $s(\beta)$ up to 4 -th order for the space average $\left\langle\exp \left(\beta \cdot A^{t}\right)\right\rangle$ with $a_{i}$ a vector quantity:
(a)

$$
\begin{equation*}
\left.\frac{\partial s}{\partial \beta_{i}}\right|_{\beta=0}=\lim _{t \rightarrow \infty} \frac{1}{t}\left\langle A_{i}^{t}\right\rangle=\left\langle a_{i}\right\rangle, \tag{10.39}
\end{equation*}
$$

(b)

$$
\begin{align*}
\left.\frac{\partial^{2} s}{\partial \beta_{i} \partial \beta_{j}}\right|_{\beta=0} & =\lim _{t \rightarrow \infty} \frac{1}{t}\left(\left\langle A_{i}^{t} A_{j}^{t}\right\rangle-\left\langle A_{i}^{t}\right\rangle\left\langle A_{j}^{t}\right\rangle\right) \\
& =\lim _{t \rightarrow \infty} \frac{1}{t}\left\langle\left(A_{i}^{t}-t\left\langle a_{i}\right\rangle\right)\left(A_{j}^{t}-t\left\langle a_{j}\right\rangle\right)\right\rangle \tag{10.40}
\end{align*}
$$

Note that the formalism is cmart: it automatically yields the variance from the mean, rather than simply the 2nd moment $\left\langle a^{2}\right\rangle$.
(c) compute the third derivative of $s(\beta)$.
(d) compute the fourth derivative assuming that the mean in (10.39) vanishes, $\left\langle a_{i}\right\rangle=0$. The 4-th order moment formula

$$
\begin{equation*}
K(t)=\frac{\left\langle x^{4}(t)\right\rangle}{\left\langle x^{2}(t)\right\rangle^{2}}-3 \tag{10.41}
\end{equation*}
$$

that you have derived is known as kurtosis: it measures a deviation from what the 4-th order moment would be were the distribution a pure Gaussian (see (23.22) for a concrete example). If the observable is a vector, the kurtosis is given by

$$
K(t)=\frac{\sum_{i j}\left[\left\langle A_{i} A_{i} A_{j} A_{j}\right\rangle+2\left(\left\langle A_{i} A_{j}\right\rangle\left\langle A_{j} A_{i}\right\rangle-\left\langle A_{i} A_{i}\right\rangle\left\langle A_{j} A_{j}\right\rangle\right)\right]}{\left(\sum_{i}\left\langle A_{i} A_{i}\right\rangle\right)^{2}}(10.42)
$$

Exercise 10.3 Pinball escape rate from numerical simulation*. Estimate the escape rate for $R: a=6$ 3-disk pinball by shooting 100,000 randomly initiated pinballs into the 3-disk system and plotting the logarithm of the number of trapped orbits as function of time. For comparison, a numerical simulation of ref. [6.2] yields $\gamma=.410 \ldots$.

## Chapter 11

## Qualitative dynamics, for pedestrians

The classification of the constituents of a chaos, nothing less is here essayed.
Herman Melville, Moby Dick, chapter 32

In this chapter we begin to learn how to use qualitative properties of a flow in order to partition the phase space in a topologically invariant way, and name topologically distinct orbits. This will enable us - in chapter 13 - to count the distinct orbits, and in the process touch upon all the main themes of this book, going the whole distance from diagnosing chaotic dynamics to computing zeta functions.

We start by a simple physical example, symbolic dynamics of a 3-disk game of pinball, and then show that also for smooth flows the qualitative dynamics of stretching and folding flows enables us to partition the phase space and assign symbolic dynamics itineraries to trajectories. Here we illustrate the method on a $1-d$ approximation to Rössler flow. In chapter 13 we turn this topological dynamics into a multiplicative operation on the phase space partitions by means of transition matrices/Markov graphs, the simplest examples of evolution operators. Deceptively simple, this subject can get very difficult very quickly, so in this chapter we do the first pass, at a pedestrian level, postponing the discussion of higher-dimensional, cyclist level issues to chapter 12.

Even though by inclination you might only care about the serious stuff, like Rydberg atoms or mesoscopic devices, and resent wasting time on things formal, this chapter and chapter 13 are good for you. Read them.

### 11.1 Qualitative dynamics

(R. Mainieri and P. Cvitanović)

Figure 11.1: A trajectory with itinerary 021012.


What can a flow do to the phase space points? This is a very difficult question to answer because we have assumed very little about the evolution function $f^{t}$; continuity, and differentiability a sufficient number of times. Trying to make sense of this question is one of the basic concerns in the study of dynamical systems. One of the first answers was inspired by the motion of the planets: they appear to repeat their motion through the firmament. Motivated by this observation, the first attempts to describe dynamical systems were to think of them as periodic.

However, periodicity is almost never quite exact. What one tends to observe is recurrence. A recurrence of a point $x_{0}$ of a dynamical system is a return of that point to a neighborhood of where it started. How close the point $x_{0}$ must return is up to us: we can choose a volume of any size and shape, and call it the neighborhood $\mathcal{M}_{0}$, as long as it encloses $x_{0}$. For chaotic dynamical systems, the evolution might bring the point back to the starting neighborhood infinitely often. That is, the set

$$
\begin{equation*}
\left\{y \in \mathcal{M}_{0}: \quad y=f^{t}\left(x_{0}\right), \quad t>t_{0}\right\} \tag{11.1}
\end{equation*}
$$

will in general have an infinity of recurrent episodes.
To observe a recurrence we must look at neighborhoods of points. This suggests another way of describing how points move in phase space, which turns out to be the important first step on the way to a theory of dynamical systems: qualitative, topological dynamics, or, as it is usually called, symbolic dynamics. As the subject can get quite technical, a summary of the basic notions and definitions of symbolic dynamics is relegated to sect. 11.6; check that section whenever you run into obscure symbolic dynamics jargon.

We start by cutting up the phase space up into regions $\mathcal{M}_{A}, \mathcal{M}_{B}, \ldots, \mathcal{M}_{Z}$. This can be done in many ways, not all equally clever. Any such division of the phase space into topologically distinct regions is a partition, and we associate with each region (sometimes referred to as a state) a symbol $s$ from an $N$-letter alphabet or state set $\mathcal{A}=\{A, B, C, \cdots, Z\}$. As the dynamics moves the point through the phase space, different regions will be visited. The visitation sequence - forthwith referred to as the itinerary can be represented by the letters of the alphabet $\mathcal{A}$. If, as in the example sketched in figure 11.1, the phase space is divided into three regions $\mathcal{M}_{0}$,

Figure 11.2: Two pinballs that start out very close to each other exhibit the same qualitative dynamics _2313_ for the first three bounces, but due to the exponentially growing separation of trajectories with time, follow different itineraries thereafter: one escapes after _2313_, the other one escapes after _23132321_

$\mathcal{M}_{1}$, and $\mathcal{M}_{2}$, the "letters" are the integers $\{0,1,2\}$, and the itinerary for the trajectory sketched in the figure is $0 \mapsto 2 \mapsto 1 \mapsto 0 \mapsto 1 \mapsto 2 \mapsto \cdots$.

If there is no way to reach partition $\mathcal{M}_{i}$ from partition $\mathcal{M}_{j}$, and conversely, partition $\mathcal{M}_{j}$ from partition $\mathcal{M}_{i}$, the phase space consists of at least two disconnected pieces, and we can analyze it piece by piece. An interesting partition should be dynamically connected, that is, one should be able to go from any region $\mathcal{M}_{i}$ to any other region $\mathcal{M}_{j}$ in a finite number of steps. A dynamical system with such partition is said to be metrically indecomposable.

In general one also encounters transient regions - regions to which the dynamics does not return to once they are exited. Hence we have to distinguish between (for us uninteresting) wandering trajectories that never return to the initial neighborhood, and the non-wandering set (2.2) of the recurrent trajectories.

The allowed transitions between the regions of a partition are encoded in the $[N \times N]$-dimensional transition matrix whose elements take values

$$
T_{i j}= \begin{cases}1 & \text { if a transition } \mathcal{M}_{j} \rightarrow \mathcal{M}_{i} \text { is possible }  \tag{11.2}\\ 0 & \text { otherwise }\end{cases}
$$

The transition matrix encodes the topological dynamics as an invariant law of motion, with the allowed transitions at any instant independent of the trajectory history, requiring no memory.

Example 11.1 Complete $N$-ary dynamics: All transition matrix entries equal unity (one can reach any region from any other region in one step):

$$
T_{c}=\left(\begin{array}{cccc}
1 & 1 & \ldots & 1  \tag{11.3}\\
1 & 1 & \ldots & 1 \\
\vdots & \vdots & \ddots & \vdots \\
1 & 1 & \ldots & 1
\end{array}\right) .
$$

Further examples of transition matrices, such as the 3-disk transition matrix (11.5) and the 1-step memory sparse matrix (11.15), are peppered throughout the text.
(a)

(b)


Figure 11.3: The 3-disk game of pinball Poincaré section, trajectories emanating from the disk 1 with $x_{0}=($ arclength, parallel momentum $)=\left(s_{0}, p_{0}\right)$, disk radius : center separation ratio $a: R=1: 2.5$. (a) Strips of initial points $\mathcal{M}_{12}, \mathcal{M}_{13}$ which reach disks 2,3 in one bounce, respectively. (b) Strips of initial points $\mathcal{M}_{121}, \mathcal{M}_{131} \mathcal{M}_{132}$ and $\mathcal{M}_{123}$ which reach disks $1,2,3$ in two bounces, respectively. (Y. Lan)

However, knowing that a point from $\mathcal{M}_{i}$ reaches $\mathcal{M}_{j}$ in one step is not quite good enough. We would be happier if we knew that any point in $\mathcal{M}_{i}$ reaches $\mathcal{M}_{j}$; otherwise we have to subpartition $\mathcal{M}_{i}$ into the points which land in $\mathcal{M}_{j}$, and those which do not, and often we will find ourselves partitioning ad infinitum.

Such considerations motivate the notion of a Markov partition, a partition for which no memory of preceding steps is required to fix the transitions allowed in the next step. Dynamically, finite Markov partitions can be generated by expanding $d$-dimensional iterated mappings $f: \mathcal{M} \rightarrow \mathcal{M}$, if $\mathcal{M}$ can be divided into $N$ regions $\left\{\mathcal{M}_{0}, \mathcal{M}_{1}, \ldots, \mathcal{M}_{N-1}\right\}$ such that in one step points from an initial region $\mathcal{M}_{i}$ either fully cover a region $\mathcal{M}_{j}$, or miss it altogether,

$$
\begin{equation*}
\text { either } \quad \mathcal{M}_{j} \cap f\left(\mathcal{M}_{i}\right)=\emptyset \quad \text { or } \quad \mathcal{M}_{j} \subset f\left(\mathcal{M}_{i}\right) . \tag{11.4}
\end{equation*}
$$

Let us illustrate what this means by our favorite example, the game of pinball.

Example 11.2 3-disk symbolic dynamics: Consider the motion of a free point particle in a plane with 3 elastically reflecting convex disks. After a collision with a disk a particle either continues to another disk or escapes, and any trajectory can be labeled by the disk sequence. For example, if we label the three disks by 1,2 and 3 , the two trajectories in figure 11.2 have itineraries_2313_, _23132321_ respectively. The 3-disk prime cycles given in figures 1.5 and 11.6 are further examples of such itineraries.

At each bounce a cone of initially nearby trajectories defocuses (see figure 1.7), and in order to attain a desired longer and longer itinerary of bounces the initial point $x_{0}=\left(s_{0}, p_{0}\right)$ has to be specified with a larger and larger precision, and lie within initial phase space strips drawn in figure 11.3. Similarly, it is intuitively clear that as we go backward in time (in this case, simply reverse the velocity vector), we also need increasingly precise specification of $x_{0}=\left(s_{0}, p_{0}\right)$ in order to follow a given past itinerary. Another way to look at the survivors after two bounces is to plot $\mathcal{M}_{s_{1} . s_{2}}$, the intersection of $\mathcal{M}_{. s_{2}}$ with the strips $\mathcal{M}_{s_{1}}$. obtained by time reversal (the velocity changes $\operatorname{sign} \sin \phi \rightarrow-\sin \phi$ ). $\mathcal{M}_{s_{1} . s_{2}}$, figure 11.4, is a "rectangle" of nearby trajectories which have arrived from the disk $s_{1}$ and are heading for the disk $s_{2}$.


Figure 11.4: The Poincaré section of the phase space for the binary labeled pinball. For definitiveness, this set is generated by starting from disk 1, preceded by disk 2. Indicated are the fixed points $\overline{0}, \overline{1}$ and the 2 -cycle periodic points $\overline{01}, \overline{10}$, together with strips which survive $1,2, \ldots$ bounces. Iteration corresponds to the decimal point shift; for example, all points in the rectangle [01.01] map into the rectangle [010.1] in one iteration. See also figure 11.6 (b).

We see that a finite length trajectory is not uniquely specified by its finite itinerary, but an isolated unstable cycle is: its itinerary is an infinitely repeating block of symbols. More generally, for hyperbolic flows the intersection of the future and past itineraries, the bi-infinite itinerary $S^{-} . S^{+}=\cdots s_{-2} s_{-1} s_{0} \cdot s_{1} s_{2} s_{3} \cdots$ specifies a unique trajectory. This is intuitively clear for our 3 -disk game of pinball, and is stated more formally in the definition (11.4) of a Markov partition. The definition requires that the dynamics be expanding forward in time in order to ensure that the cone of trajectories with a given itinerary becomes sharper and sharper as the number of specified symbols is increased.

Example 11.3 Pruning rules for a 3-disk alphabet: As the disks are convex, there can be no two consecutive reflections off the same disk, hence the covering symbolic dynamics consists of all sequences which include no symbol repetitions _11_, _22_, _33_. This is a finite set of finite length pruning rules, hence, the dynamics is a subshift of finite type (see (11.24) for definition), with the transition matrix (11.2) given by

$$
T=\left(\begin{array}{lll}
0 & 1 & 1  \tag{11.5}\\
1 & 0 & 1 \\
1 & 1 & 0
\end{array}\right)
$$

For convex disks the separation between nearby trajectories increases at every reflection, implying that the stability matrix has an expanding eigenvalue. By the Liouville phasespace volume conservation (5.23), the other transverse eigenvalue is contracting. This example demonstrates that finite Markov partitions can be constructed for hyperbolic dynamical systems which are expanding in some directions, contracting in others. Further examples are the 1-dimensional expanding mapping sketched in figure 11.9, and more examples are worked out in sect. 23.2.

Determining whether the symbolic dynamics is complete (as is the case for sufficiently separated disks), pruned (for example, for touching or overlapping disks), or only a first coarse graining of the topology (as, for example, for smooth potentials with islands of stability) requires case-by-case investigation, a discussion we postpone to sect. 11.4 and chapter 12. For the

Figure 11.5: Binary labeling of the 3-disk pinball trajectories; a bounce in which the trajectory returns to the preceding disk is labeled 0 , and a bounce which results in continuation to the third disk is labeled 1.

time being we assume that the disks are sufficiently separated that there is no additional pruning beyond the prohibition of self-bounces.


### 11.2 A brief detour; recoding, symmetries, tilings

- 

Though a useful tool, Markov partitioning is not without drawbacks. One glaring shortcoming is that Markov partitions are not unique: any of many different partitions might do the job. The 3-disk system offers a simple illustration of different Markov partitioning strategies for the same dynamical system.

The $\mathcal{A}=\{1,2,3\}$ symbolic dynamics for 3 -disk system is neither unique, nor necessarily the smartest one - before proceeding it pays to exploit the symmetries of the pinball in order to obtain a more efficient description. In chapter 22 we shall be handsomely rewarded for our labors.

As the three disks are equidistantly spaced, our game of pinball has a sixfold symmetry. For instance, the cycles $\overline{12}, \overline{23}$, and $\overline{13}$ are related to each other by rotation by $\pm 2 \pi / 3$ or, equivalently, by a relabeling of the disks. Further examples of such symmetries are shown in figure 1.5. The disk labels are arbitrary; what is important is how a trajectory evolves as it hits subsequent disks, not what label the starting disk had. We exploit this symmetry by recoding, in this case replacing the absolute disk labels by relative symbols, indicating the type of the collision. For the 3 -disk game of pinball there are two topologically distinct kinds of collisions, figure 11.5:

$$
s_{i}=\left\{\begin{array}{lll}
0 & : & \text { pinball returns to the disk it came from }  \tag{11.6}\\
1 & : & \text { pinball continues to the third disk }
\end{array}\right.
$$

This binary symbolic dynamics has two immediate advantages over the ternary one; the prohibition of self-bounces is automatic, and the coding utilizes the symmetry of the 3 -disk pinball game in elegant manner. If the disks are sufficiently far apart there are no further restrictions on symbols, the symbolic dynamics is complete, and all binary sequences are admissible itineraries. As this type of symbolic dynamics pops up frequently, we list
page 184

| $n_{p}$ | $p$ | $n_{p}$ | $p$ | $n_{p}$ | $p$ | $n_{p}$ | $p$ | $n_{p}$ | $p$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| , | 0 | 7 | 0001001 | 8 | 00001111 | 9 | 000001101 | 9 | 001001111 |
|  | 1 |  | 0000111 |  | 00010111 |  | 000010011 |  | 001010111 |
| 2 | 01 |  | 0001011 |  | 00011011 |  | 000010101 |  | 001011011 |
| 3 | 001 |  | 0001101 |  | 00011101 |  | 000011001 |  | 001011101 |
|  | 011 |  | 0010011 |  | 00100111 |  | 000100011 |  | 001100111 |
| 4 | 0001 |  | 0010101 |  | 00101011 |  | 000100101 |  | 001101011 |
|  | 0011 |  | 0001111 |  | 00101101 |  | 000101001 |  | 001101101 |
|  | 0111 |  | 0010111 |  | 00110101 |  | 000001111 |  | 001110101 |
| 5 | 00001 |  | 0011011 |  | 00011111 |  | 000010111 |  | 010101011 |
|  | 00011 |  | 0011101 |  | 00101111 |  | 000011011 |  | 000111111 |
|  | 00101 |  | 0101011 |  | 00110111 |  | 000011101 |  | 001011111 |
|  | 00111 |  | 0011111 |  | 00111011 |  | 000100111 |  | 001101111 |
|  | 01011 |  | 0101111 |  | 00111101 |  | 000101011 |  | 001110111 |
|  | 01111 |  | 0110111 |  | 01010111 |  | 000101101 |  | 001111011 |
| 6 | 000001 |  | 0111111 |  | 01011011 |  | 000110011 |  | 001111101 |
|  | 000011 | 8 | 00000001 |  | 00111111 |  | 000110101 |  | 010101111 |
|  | 000101 |  | 00000011 |  | 01011111 |  | 000111001 |  | 010110111 |
|  | 000111 |  | 00000101 |  | 01101111 |  | 001001011 |  | 010111011 |
|  | 001011 |  | 00001001 |  | 01111111 |  | 001001101 |  | 001111111 |
|  | 001101 |  | 00000111 | 9 | 000000001 |  | 001010011 |  | 010111111 |
|  | 001111 |  | 00001011 |  | 000000011 |  | 001010101 |  | 011011111 |
|  | 010111 |  | 00001101 |  | 000000101 |  | 000011111 |  | 011101111 |
|  | 011111 |  | 00010011 |  | 000001001 |  | 000101111 |  | 011111111 |
| 7 | 0000001 |  | 00010101 |  | 000010001 |  | 000110111 |  |  |
|  | 0000011 |  | 00011001 |  | 000000111 |  | 000111011 |  |  |
|  | 0000101 |  | 00100101 |  | 000001011 |  | 000111101 |  |  |

Table 11.1: Prime cycles for the binary symbolic dynamics up to length 9.

Example 11.4 Recoding ternary symbolic dynamics in binary: Given a ternary sequence and labels of 2 preceding disks, rule (11.6) fixes the subsequent binary symbols. Here we list an arbitrary ternanry itinerary, and the corresponding binary sequence:

```
ternary : 4
    binary : . . 1 1 0
```

The first 2 disks initialize the trajectory and its direction; $3 \mapsto 1 \mapsto 2 \mapsto \cdots$. Due to the 3-disk symmetry the six 3-disk sequences initialized by 12, 13, 21, 23, 31, 32 respectively have the same weights, the same size partitions, and are coded by a single binary sequence. For periodic orbits, the equivalent ternary cycles reduce to binary cycles of $1 / 3,1 / 2$ or the same length. How this works is best understood by inspection of table 11.2, figure 11.6 and figure 22.3.

The 3-disk game of pinball is tiled by six copies of the fundamental domain, a one-sixth slice of the full 3-disk system, with the symmetry axes acting as reflecting mirrors, see figure 11.6 (b). Every global 3-disk trajectory has a corresponding fundamental domain mirror trajectory obtained by replacing every crossing of a symmetry axis by a reflection. Depending on the symmetry of the global trajectory, a repeating binary symbols block corresponds either to the full periodic orbit or to an irreducible segment (examples are shown in figure 11.6 and table 11.2). An irreducible segment corresponds to a periodic orbit in the fundamental domain. Table 11.2 lists some of the shortest binary periodic orbits, together with the corresponding full 3-disk symbol sequences and orbit symmetries. For a number of reasons that will be elucidated in chapter 22 , life is much simpler in


Figure 11.6: The 3-disk game of pinball with the disk radius : center separation ratio $\mathrm{a}: \mathrm{R}=1: 2.5$. (a) The three disks, with $\overline{12}, \overline{123}$ and $\overline{121232313}$ cycles indicated. (b) The fundamental domain, that is, the small $1 / 6$ th wedge indicated in (a), consisting of a section of a disk, two segments of symmetry axes acting as straight mirror walls, and an escape gap. The above cycles restricted to the fundamental domain are now the two fixed points $\overline{0}, \overline{1}$, and the $\overline{100}$ cycle.
the fundamental domain than in the full system, so whenever possible our computations will be carried out in the fundamental domain.

Inspecting the figure 11.3 we see that the relative ordering of regions with differing finite itineraries is a qualitative, topological property of the flow, so it makes sense to define a simple "canonical" representative partition which in a simple manner exhibits spatial ordering common to an entire class of topologically similar nonlinear flows.


### 11.3 Stretch and fold

Symbolic dynamics for $N$-disk game of pinball is so straightforward that one may altogether fail to see the connection between the topology of hyperbolic flows and their symbolic dynamics. This is brought out more clearly by the 1-dimensional visualization of "stretch \& fold" flows to which we turn now.

Suppose concentrations of certain chemical reactants worry you, or the variations in the Chicago temperature, humidity, pressure and winds affect your mood. All such properties vary within some fixed range, and so do their rates of change. Even if we are studying an open system such as the 3 -disk pinball game, we tend to be interested in a finite region around the disks and ignore the escapees. So a typical dynamical system that we care about is bounded. If the price for keeping going is high - for example, we try to stir up some tar, and observe it come to a dead stop the moment we cease our labors - the dynamics tends to settle into a simple limiting state. However, as the resistance to change decreases - the tar is heated up and

| $\tilde{p}$ | $p$ | $\mathbf{g}_{\tilde{p}}$ | $\tilde{p}$ | $p$ | $\mathbf{g}_{\tilde{p}}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 0 | 12 | $\sigma_{12}$ | 000001 | 121212131313 | $\sigma_{23}$ |
| 1 | 123 | $C_{3}$ | 000011 | 121212313131232323 | $C_{3}^{2}$ |
| 01 | 1213 | $\sigma_{23}$ | 000101 | 121213 | $e$ |
| 001 | 121232313 | $C_{3}$ | 000111 | 121213212123 | $\sigma_{12}$ |
| 011 | 121323 | $\sigma_{13}$ | 001011 | 121232131323 | $\sigma_{23}$ |
| 0001 | 12121313 | $\sigma_{23}$ | 001101 | 121231323213 | $\sigma_{13}$ |
| 0011 | 121231312323 | $C_{3}^{2}$ | 001111 | 121231232312313123 | $C_{3}$ |
| 0111 | 12132123 | $\sigma_{12}$ | 010111 | 121312313231232123 | $C_{3}^{2}$ |
| 00001 | 121212323231313 | $C_{3}$ | 011111 | 121321323123 | $\sigma_{13}$ |
| 00011 | 1212132323 | $\sigma_{13}$ | 0000001 | 121212123232323131313 | $C_{3}$ |
| 00101 | 1212321213 | $\sigma_{12}$ | 0000011 | 12121213232323 | $\sigma_{13}$ |
| 00111 | 12123 | $e$ | 0000101 | 12121232121213 | $\sigma_{12}$ |
| 01011 | 121312321231323 | $C_{3}$ | 0000111 | 1212123 |  |
| 01111 | 1213213123 | $\sigma_{23}$ | . . 0 | . . | $\ldots$ |

Table 11.2: $C_{3 v}$ correspondence between the binary labeled fundamental domain prime cycles $\tilde{p}$ and the full 3-disk ternary labeled cycles $p$, together with the $C_{3 v}$ transformation that maps the end point of the $\tilde{p}$ cycle into the irreducible segment of the $p$ cycle, see sect. 22.2.2. Breaks in the ternary sequences mark repeats of the irreducible segment. The degeneracy of $p$ cycle is $m_{p}=6 n_{\tilde{p}} / n_{p}$. The shortest pair of the fundamental domain cycles related by time reversal (but no spatial symmetry) are the 6 -cycles $\overline{001011}$ and $\overline{001101}$.
we are more vigorous in our stirring - the dynamics becomes unstable.
If a flow is locally unstable but globally bounded, any open ball of initial points will be stretched out and then folded back.

At this juncture we show how this works on the simplest example: unimodal mappings of the interval. The erudite reader should skim through this chapter and then take a more demanding path, via the Smale horseshoes of chapter 12. Unimodal maps are easier, but physically less motivated. The Smale horseshoes are the high road, more complicated, but the right tool to generalize what we learned from the 3-disk dynamics, and begin analysis of general dynamical systems. It is up to you - unimodal maps suffice to get quickly to the heart of this treatise.

### 11.3.1 Temporal ordering: itineraries

In this section we learn how to name (and, in chapter 13 , how to count) periodic orbits for the simplest, and nevertheless very instructive case, for 1-dimensional maps of an interval.

Suppose that the compression of the folded interval in figure 11.7 is so fierce that we can neglect the thickness of the attractor. For example, the Rössler flow (2.14) is volume contracting, and an interval transverse to the attractor is stretched, folded and pressed back into a nearly 1-dimensional interval, typically compressed transversally by a factor of $\approx 10^{13}$ in one Poincaré section return. In such cases it makes sense to approximate the return map of a "stretch \& fold" flow by a 1-dimensional map.

The simplest mapping of this type is unimodal; interval is stretched


Figure 11.7: (a) A recurrent flow that stretches and folds. (b) The "stretch \& fold" return map on the Poincaré section.

Figure 11.8: The Poincaré return map of the Kuramoto-Sivashinsky system (2.21) figure 3.3, from the unstable manifold of the $\overline{1}$ fixed point to the (neighborhood of) the unstable manifold. Also indicated are the periodic points $\overline{0}$ and $\overline{01}$.



Figure 11.9: (a) The complete tent map together with intervals that follow the indicated itinerary for $n$ steps. (b) A unimodal repeller with the remaining intervals after 1, 2 and 3 iterations. Intervals marked $s_{1} s_{2} \cdots s_{n}$ are unions of all points that do not escape in $n$ iterations, and follow the itinerary $S^{+}=s_{1} s_{2} \cdots s_{n}$. Note that the spatial ordering does not respect the binary ordering; for example $x_{00}<x_{01}<x_{11}<$ $x_{10}$. Also indicated: the fixed points $x_{0}, x_{1}$, the 2-cycle $\overline{01}$, and the 3-cycle $\overline{011}$.
and folded only once, with at most two points mapping into a point in the refolded interval. A unimodal map $f(x)$ is a 1 -dimensional function $\mathbb{R} \rightarrow \mathbb{R}$ defined on an interval $\mathcal{M} \in \mathbb{R}$ with a monotonically increasing (or decreasing) branch, a critical point (or interval) $x_{c}$ for which $f\left(x_{c}\right)$ attains the maximum (minimum) value, followed by a monotonically decreasing (increasing) branch. Uni-modal means that the map is a one-humped map with one critical point within interval $\mathcal{M}$. A multi-modal map has several critical points within interval $\mathcal{M}$.

Example 11.5 Complete tent map, logistic map: The simplest examples of unimodal maps are the complete tent map, figure 11.9 (a),

$$
\begin{equation*}
f(\gamma)=1-2|\gamma-1 / 2|, \tag{11.8}
\end{equation*}
$$

and the quadratic map (sometimes also called the logistic map)

$$
\begin{equation*}
x_{t+1}=1-a x_{t}^{2} \tag{11.9}
\end{equation*}
$$

with the one critical point at $x_{c}=0$. Furthe example is the repelling unimodal map of figure 11.9 (b).

Such dynamical systems are irreversible (the inverse of $f$ is double-valued), but, as we shall show in sect. 12.2, they may nevertheless serve as effective descriptions of invertible 2-dimensional hyperbolic flows.

For the unimodal maps of figure 11.9 a Markov partition of the unit interval $\mathcal{M}$ is given by the two intervals $\left\{\mathcal{M}_{0}, \mathcal{M}_{1}\right\}$. We refer to (11.8) as the "complete" tent map because its symbolic dynamics is complete binary: as both $f\left(\mathcal{M}_{0}\right)$ and $f\left(\mathcal{M}_{1}\right)$ fully cover $\mathcal{M}_{0}$ and $\mathcal{M}_{1}$, the corresponding transition matrix is a [2×2] matrix with all entries equal to 1, as in (11.3). As binary symbolic dynamics pops up frequently in applications, we list the shortest binary prime cycles in table 11.1.

The critical value denotes either the maximum or the minimum value of $f(x)$ on the defining interval; we assume here that it is a maximum, $f\left(x_{c}\right) \geq f(x)$ for all $x \in \mathcal{M}$. The critical value $f\left(x_{c}\right)$ belongs neither to the left nor to the right partition $\mathcal{M}_{i}$, and is denoted by its own symbol $s=C$. As we shall see, its preimages serve as partition boundary points.

The trajectory $x_{1}, x_{2}, x_{3}, \ldots$ of the initial point $x_{0}$ is given by the iteration $x_{n+1}=f\left(x_{n}\right)$. Iterating $f$ and checking whether the point lands to the left or to the right of $x_{c}$ generates a temporally ordered topological itinerary (11.17) for a given trajectory,

$$
s_{n}=\left\{\begin{array}{ll}
1 & \text { if } x_{n}>x_{c}  \tag{11.10}\\
0 & \text { if } x_{n}<x_{c}
\end{array} .\right.
$$

We shall refer to $S^{+}\left(x_{0}\right)=. s_{1} s_{2} s_{3} \cdots$ as the future itinerary. Our next task is to answer the reverse problem: given an itinerary, what is the corresponding spatial ordering of points that belong to a given trajectory?

### 11.3.2 Spatial ordering, 1- $d$ maps

Suppose you have succeeded in constructing a covering symbolic dynamics, such as for a well-separated 3 -disk system. Now start moving the disks toward each other. At some critical separation a disk will start blocking families of trajectories traversing the other two disks. The order in which trajectories disappear is determined by their relative ordering in space; the ones closest to the intervening disk will be pruned first. Determining inadmissible itineraries requires that we relate the spatial ordering of trajectories to their time ordered itineraries.

The easiest point of departure is to start out by working out this relation for the symbolic dynamics of 1-dimensional mappings. As it appears impossible to present this material without getting bogged down in a sea of 0 's, 1's and subscripted subscripts, we announce the main result before embarking upon its derivation:

The admissibility criterion stated in sect. 11.4 eliminates all itineraries that cannot occur for a given unimodal map.

The tent map (11.8) consists of two straight segments joined at $x=1 / 2$. The symbol $s_{n}$ defined in (11.10) equals 0 if the function increases, and 1 if the function decreases. The piecewise linearity of the map makes it possible to analytically determine an initial point given its itinerary, a property that we now use to define a topological coordinatization common to all unimodal maps.

Here we have to face the fundamental problem of pedagogy: combinatorics cannot be taught. The best one can do is to state the answer, and then hope that you will figure it out by yourself. Once you figure it out, feel free to complain that the way the rule is stated here is incomprehensible.

Figure 11.10: Alternating binary tree relates the itinerary labeling of the unimodal map figure 11.9 intervals to their spatial ordering. Dotted line stands for 0 , full line for 1 ; the binary sub-tree whose root is a full line (symbol 1) reverses the orientation, due to the orientation reversing fold in figures 11.7 and 11.9.


The tent map point $\gamma\left(S^{+}\right)$with future itinerary $S^{+}$is given by converting the sequence of $s_{n}$ 's into a binary number by the following algorithm:

$$
\left.\begin{array}{rl}
w_{n+1} & =\left\{\begin{array}{ll}
w_{n} & \text { if } s_{n}=0 \\
1-w_{n} & \text { if } s_{n}=1
\end{array}, \quad w_{1}=s_{1}\right.
\end{array}\right] \begin{aligned}
& \gamma\left(S^{+}\right)=0 . w_{1} w_{2} w_{3} \ldots=\sum_{n=1}^{\infty} w_{n} / 2^{n}
\end{aligned}
$$

This follows by inspection from the binary tree of figure 11.10.

> Example 11.6 $\quad \gamma$ whose itinerary is $S^{+}=0110000 \cdots$ is given by the binary number $\gamma=.010000 \cdots$. Conversely, the itinerary of $\gamma=.01$ is $s_{1}=0, f(\gamma)=.1 \rightarrow s_{2}=1$, $f^{2}(\gamma)=f(.1)=1 \rightarrow s_{3}=1$, etc..

We shall refer to $\gamma\left(S^{+}\right)$as the (future) topological coordinate. $w_{t}$ 's are the digits in the binary expansion of the starting point $\gamma$ for the complete tent map (11.8). In the left half-interval the map $f(x)$ acts by multiplication by 2 , while in the right half-interval the map acts as a flip as well as multiplication by 2 , reversing the ordering, and generating in the process the sequence of $s_{n}$ 's from the binary digits $w_{n}$.

The mapping $\quad x_{0} \rightarrow S^{+}\left(x_{0}\right) \rightarrow \gamma_{0}=\gamma\left(S^{+}\right) \quad$ is a topological conjugacy which maps the trajectory of an initial point $x_{0}$ under iteration of a given unimodal map to that initial point $\gamma$ for which the trajectory of the "canonical" unimodal map (11.8) has the same itinerary. The virtue of this conjugacy is that it preserves the ordering for any unimodal map in the sense that if $\bar{x}>x$, then $\bar{\gamma}>\gamma$.

### 11.4 Kneading theory

## (K.T. Hansen and P. Cvitanović)

The main motivation for being mindful of spatial ordering of temporal itineraries is that this spatial ordering provides us with criteria that separate inadmissible orbits from those realizable by the dynamics. For 1dimensional mappings the kneading theory provides such criterion of admissibility.

Figure 11.11: The "dike" map obtained by slicing of a top portion of the tent map figure 11.9 (a). Any orbit that visits the primary pruning interval $(\kappa, 1]$ is inadmissible. The admissible orbits form the Cantor set obtained by removing from the unit interval the primary pruning interval and all its iterates. Any admissible orbit has the same topological coordinate and itinerary as the corresponding tent map figure 11.9 (a) orbit.


If the parameter in the quadratic map (11.9) is $a>2$, then the iterates of the critical point $x_{c}$ diverge for $n \rightarrow \infty$. As long as $a \geq 2$, any sequence $S^{+}$composed of letters $s_{i}=\{0,1\}$ is admissible, and any value of $0 \leq \gamma<1$ corresponds to an admissible orbit in the non-wandering set of the map. The corresponding repeller is a complete binary labeled Cantor set, the $n \rightarrow \infty$ limit of the $n$th level covering intervals sketched in figure 11.9.

For $a<2$ only a subset of the points in the interval $\gamma \in[0,1]$ corresponds to admissible orbits. The forbidden symbolic values are determined by observing that the largest $x_{n}$ value in an orbit $x_{1} \rightarrow x_{2} \rightarrow x_{3} \rightarrow \ldots$ has to be smaller than or equal to the image of the critical point, the critical value $f\left(x_{c}\right)$. Let $K=S^{+}\left(x_{c}\right)$ be the itinerary of the critical point $x_{c}$, denoted the kneading sequence of the map. The corresponding topological coordinate is called the kneading value

$$
\begin{equation*}
\kappa=\gamma(K)=\gamma\left(S^{+}\left(x_{c}\right)\right) . \tag{11.12}
\end{equation*}
$$

A map with the same kneading sequence $K$ as $f(x)$, such as the dike map figure 11.11, is obtained by slicing off all $\gamma\left(S^{+}\left(x_{0}\right)\right)>\kappa$,

$$
f(\gamma)=\left\{\begin{array}{ll}
f_{0}(\gamma)=2 \gamma & \gamma \in I_{0}=[0, \kappa / 2)  \tag{11.13}\\
f_{c}(\gamma)=\kappa & \gamma \in I_{c}=[\kappa / 2,1-\kappa / 2] \\
f_{1}(\gamma)=2(1-\gamma) & \gamma \in I_{1}=[1-\kappa / 2,1]
\end{array} .\right.
$$

The dike map is the complete tent map figure 11.9 (a) with the top sliced off. It is convenient for coding the symbolic dynamics, as those $\gamma$ values that survive the pruning are the same as for the complete tent map figure 11.9 (a), and are easily converted into admissible itineraries by (??).

If $\gamma\left(S^{+}\right)>\gamma(K)$, the point $x$ whose itinerary is $S^{+}$would exceed the critical value, $x>f\left(x_{c}\right)$, and hence cannot be an admissible orbit. Let

$$
\begin{equation*}
\hat{\gamma}\left(S^{+}\right)=\sup _{m} \gamma\left(\sigma^{m}\left(S^{+}\right)\right) \tag{11.14}
\end{equation*}
$$

be the maximal value, the highest topological coordinate reached by the orbit $x_{1} \rightarrow x_{2} \rightarrow x_{3} \rightarrow \ldots$. We shall call the interval ( $\left.\kappa, 1\right]$ the primary pruned interval. The orbit $S^{+}$is inadmissible if $\gamma$ of any shifted sequence of $S^{+}$falls into this interval.

Criterion of admissibility: Let $\kappa$ be the kneading value of the critical point, and $\hat{\gamma}\left(S^{+}\right)$be the maximal value of the orbit $S^{+}$. Then the orbit $S^{+}$ is admissible if and only if $\hat{\gamma}\left(S^{+}\right) \leq \kappa$.

While a unimodal map may depend on many arbitrarily chosen parameters, its dynamics determines the unique kneading value $\kappa$. We shall call $\kappa$ the topological parameter of the map. Unlike the parameters of the original dynamical system, the topological parameter has no reason to be either smooth or continuous. The jumps in $\kappa$ as a function of the map parameter such as $a$ in (11.9) correspond to inadmissible values of the topological parameter. Each jump in $\kappa$ corresponds to a stability window associated with a stable cycle of a smooth unimodal map. For the quadratic map (11.9) $\kappa$ increases monotonically with the parameter $a$, but for a general unimodal map such monotonicity need not hold.

For further details of unimodal dynamics, the reader is referred to appendix E.1. As we shall see in sect. 12.4, for higher dimensional maps and flows there is no single parameter that orders dynamics monotonically; as a matter of fact, there is an infinity of parameters that need adjustment for a given symbolic dynamics. This difficult subject is beyond our current ambition horizon.

### 11.5 Markov graphs

### 11.5.1 Finite memory

In the complete $N$-ary symbolic dynamics case (see example (11.3)) the choice of the next symbol requires no memory of the previous ones. However, any further refinement of the partition requires finite memory.

For example, for the binary labeled repeller with complete binary symbolic dynamics, we might chose to partition the phase space into four regions $\left\{\mathcal{M}_{00}, \mathcal{M}_{01}, \mathcal{M}_{10}, \mathcal{M}_{11}\right\}$, a 1-step refinement of the initial partition $\left\{\mathcal{M}_{0}, \mathcal{M}_{1}\right\}$. Such partitions are drawn in figure 11.4, as well as figure 1.8. Topologically $f$ acts as a left shift (12.7), and its action on the rectangle [.01] is to move the decimal point to the right, to [0.1], forget the past, [.1], and land in either of the two rectangles $\{[.10],[.11]\}$. Filling in the matrix elements for the other three initial states we obtain the 1-step memory transition matrix acting on the 4 -state vector

[^2]\[

\phi^{\prime}=T \phi=\left($$
\begin{array}{cccc}
T_{00,00} & 0 & T_{00,10} & 0  \tag{11.15}\\
T_{01,00} & 0 & T_{01,10} & 0 \\
0 & T_{10,01} & 0 & T_{10,11} \\
0 & T_{11,01} & 0 & T_{11,11}
\end{array}
$$\right)\left($$
\begin{array}{c}
\phi_{00} \\
\phi_{01} \\
\phi_{10} \\
\phi_{11}
\end{array}
$$\right)
\]



Figure 11.12: (a) The self-similarity of the complete binary symbolic dynamics represented by a binary tree (b) identification of nodes $B=A, C=A$ leads to the finite 1-node, 2-links Markov graph. All admissible itineraries are generated as walks on this finite Markov graph.

(a)


Figure 11.13: (a) The 2-step memory Markov graph, links version obtained by identifying nodes $A=D=E=F=G$ in figure 11.12(a). Links of this graph correspond to the matrix entries in the transition matrix (11.15). (b) the 2-step memory Markov graph, node version.

By the same token, for $M$-step memory the only nonvanishing matrix elements are of the form $T_{s_{1} s_{2} \ldots s_{M+1}, s_{0} s_{1} \ldots s_{M}}, s_{M+1} \in\{0,1\}$. This is a sparse matrix, as the only non vanishing entries in the $m=s_{0} s_{1} \ldots s_{M}$ column of $T_{d m}$ are in the rows $d=s_{1} \ldots s_{M} 0$ and $d=s_{1} \ldots s_{M} 1$. If we increase the number of steps remembered, the transition matrix grows big quickly, as the $N$-ary dynamics with $M$-step memory requires an $\left[N^{M+1} \times N^{M+1}\right]$ matrix. Since the matrix is very sparse, it pays to find a compact representation for $T$. Such representation is afforded by Markov graphs, which are not only compact, but also give us an intuitive picture of the topological dynamics.

Construction of a good Markov graph is, like combinatorics, unexplainable. The only way to learn is by some diagrammatic gymnastics, so we work our way through a sequence of exercises in lieu of plethora of baffling definitions.

To start with, what do finite graphs have to do with infinitely long trajectories? To understand the main idea, let us construct a graph that enumerates all possible itineraries for the case of complete binary symbolic dynamics.

Mark a dot "." on a piece of paper. Draw two short lines out of the


Figure 11.14: (a) The self-similarity of the _00_ pruned binary tree: trees originating from nodes $C$ and $E$ are the same as the entire tree. (b) Identification of nodes $A=C=E$ leads to the finite 2-node, 3-links Markov graph; as 0 is always followed by 1 , the walks on this graph generate only the admissible itineraries.
dot, end each with a dot. The full line will signify that the first symbol in an itinerary is " 1 ", and the dotted line will signifying " 0 ". Repeat the procedure for each of the two new dots, and then for the four dots, and so on. The result is the binary tree of figure 11.12(a). Starting at the top node, the tree enumerates exhaustively all distinct finite itineraries

$$
\{0,1\},\{00,01,10,11\},\{000,001,010, \cdots\}, \cdots
$$

The $M=4$ nodes in figure 11.12(a) correspond to the 16 distinct binary strings of length 4, and so on. By habit we have drawn the tree as the alternating binary tree of figure 11.10, but that has no significance as far as enumeration of itineraries is concerned - an ordinary binary tree would serve just as well.

The trouble with an infinite tree is that it does not fit on a piece of paper. On the other hand, we are not doing much - at each node we are turning either left or right. Hence all nodes are equivalent, and can be identified. To say it in other words, the tree is self-similar; the trees originating in nodes $B$ and $C$ are themselves copies of the entire tree. The result of identifying $B=A, C=A$ is a single node, 2-link Markov graph of figure $11.12(\mathrm{~b})$ : any itinerary generated by the binary tree figure 11.12(a), no matter how long, corresponds to a walk on this graph.

This is the most compact encoding of the complete binary symbolic dynamics. Any number of more complicated Markov graphs can do the job as well, and might be sometimes preferable. For example, identifying the trees originating in $D, E, F$ and $G$ with the entire tree leads to the 2step memory Markov graph of figure 11.13a. The corresponding transition matrix is given by (11.15).

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### 11.6 Symbolic dynamics, basic notions

In this section we collect the basic notions and definitions of symbolic dynamics. The reader might prefer to skim through this material on first reading, return to it later as the need arises.

Shifts. We associate with every initial point $x_{0} \in \mathcal{M}$ the future itinerary, a sequence of symbols $S^{+}\left(x_{0}\right)=s_{1} s_{2} s_{3} \cdots$ which indicates the order in which the regions are visited. If the trajectory $x_{1}, x_{2}, x_{3}, \ldots$ of the initial point $x_{0}$ is generated by

$$
\begin{equation*}
x_{n+1}=f\left(x_{n}\right), \tag{11.16}
\end{equation*}
$$

then the itinerary is given by the symbol sequence

$$
\begin{equation*}
s_{n}=s \quad \text { if } \quad x_{n} \in \mathcal{M}_{s} . \tag{11.17}
\end{equation*}
$$

Similarly, the past itinerary $S^{-}\left(x_{0}\right)=\cdots s_{-2} s_{-1} s_{0}$ describes the history of $x_{0}$, the order in which the regions were visited before arriving to the point $x_{0}$. To each point $x_{0}$ in the dynamical space we thus associate a bi-infinite itinerary

$$
\begin{equation*}
S\left(x_{0}\right)=\left(s_{k}\right)_{k \in \mathbb{Z}}=S^{-} \cdot S^{+}=\cdots s_{-2} s_{-1} s_{0} \cdot s_{1} s_{2} s_{3} \cdots . \tag{11.18}
\end{equation*}
$$

The itinerary will be finite for a scattering trajectory, entering and then escaping $\mathcal{M}$ after a finite time, infinite for a trapped trajectory, and infinitely repeating for a periodic trajectory.

The set of all bi-infinite itineraries that can be formed from the letters of the alphabet $\mathcal{A}$ is called the full shift

$$
\begin{equation*}
\mathcal{A}^{\mathbb{Z}}=\left\{\left(s_{k}\right)_{k \in \mathbb{Z}}: s_{k} \in \mathcal{A} \text { for all } \mathrm{k} \in \mathbb{Z}\right\} . \tag{11.19}
\end{equation*}
$$

The jargon is not thrilling, but this is how professional dynamicists talk to each other. We will stick to plain English to the extent possible.

We refer to this set of all conceivable itineraries as the covering symbolic dynamics. The name shift is descriptive of the way the dynamics acts on these sequences. As is clear from the definition (11.17), a forward iteration $x \rightarrow x^{\prime}=f(x)$ shifts the entire itinerary to the left through the "decimal point". This operation, denoted by the shift operator $\sigma$,

$$
\begin{equation*}
\sigma\left(\cdots s_{-2} s_{-1} s_{0} \cdot s_{1} s_{2} s_{3} \cdots\right)=\cdots s_{-2} s_{-1} s_{0} s_{1} \cdot s_{2} s_{3} \cdots \tag{11.20}
\end{equation*}
$$

demoting the current partition label $s_{1}$ from the future $S^{+}$to the "has been" itinerary $S^{-}$. The inverse shift $\sigma^{-1}$ shifts the entire itinerary one step to the right.

A finite sequence $b=s_{k} s_{k+1} \cdots s_{k+n_{b}-1}$ of symbols from $\mathcal{A}$ is called a block of length $n_{b}$. A phase space trajectory is periodic if it returns to its initial point after a finite time; in the shift space the trajectory is periodic if its itinerary is an infinitely repeating block $p^{\infty}$. We shall refer to the set of periodic points that belong to a given periodic orbit as a cycle

$$
\begin{equation*}
p=\overline{s_{1} s_{2} \cdots s_{n_{p}}}=\left\{x_{s_{1} s_{2} \cdots s_{n_{p}}}, x_{s_{2} \cdots s_{n_{p}} s_{1}}, \cdots, x_{s_{n_{p}} s_{1} \cdots s_{n_{p}-1}}\right\} . \tag{11.21}
\end{equation*}
$$

By its definition, a cycle is invariant under cyclic permutations of the symbols in the repeating block. A bar over a finite block of symbols denotes a periodic itinerary with infinitely repeating basic block; we shall omit the bar whenever it is clear from the context that the trajectory is periodic. Each cycle point is labeled by the first $n_{p}$ steps of its future itinerary. For example, the 2 nd cycle point is labeled by

$$
x_{s_{2} \cdots s_{n_{p}} s_{1}}=x_{\overline{s_{2} \cdots s_{n_{p}} s_{1}} \cdot \overline{s_{2} \cdots s_{n_{p}} s_{1}}} .
$$

A prime cycle $p$ of length $n_{p}$ is a single traversal of the orbit; its label is a block of $n_{p}$ symbols that cannot be written as a repeat of a shorter block (in literature such cycle is sometimes called prime; we shall refer to it as "prime" throughout this text).

Partitions. A partition is called generating if every infinite symbol sequence corresponds to a distinct point in the phase space. Finite Markov partition (11.4) is an example. Constructing a generating partition for a given system is a difficult problem. In examples to follow we shall concentrate on cases which allow finite partitions, but in practice almost any generating partition of interest is infinite.

A mapping $f: \mathcal{M} \rightarrow \mathcal{M}$ together with a partition $\mathcal{A}$ induces topological dynamics $(\Sigma, \sigma)$, where the subshift

$$
\begin{equation*}
\Sigma=\left\{\left(s_{k}\right)_{k \in \mathbb{Z}}\right\}, \tag{11.22}
\end{equation*}
$$

is the set of all admissible infinite itineraries, and $\sigma: \Sigma \rightarrow \Sigma$ is the shift operator (11.20). The designation "subshift" comes form the fact that $\Sigma \subset \mathcal{A}^{\mathbb{Z}}$ is the subset of the full shift (11.19). One of our principal tasks in developing symbolic dynamics of dynamical systems that occur in nature will be to determine $\Sigma$, the set of all bi-infinite itineraries $S$ that are actually realized by the given dynamical system.

A partition too coarse, coarser than, for example, a Markov partition, would assign the same symbol sequence to distinct dynamical trajectories. To avoid that, we often find it convenient to work with partitions finer than strictly necessary. Ideally the dynamics in the refined partition assigns a unique infinite itinerary $\cdots s_{-2} s_{-1} s_{0} \cdot s_{1} s_{2} s_{3} \cdots$ to each distinct trajectory, but there might exist full shift symbol sequences (11.19) which are not realized as trajectories; such sequences are called inadmissible, and we say
that the symbolic dynamics is pruned. The word is suggested by "pruning" of branches corresponding to forbidden sequences for symbolic dynamics organized hierarchically into a tree structure, as explained in sect. 11.5.

Pruning. If the dynamics is pruned, the alphabet must be supplemented by a grammar, a set of pruning rules. After the inadmissible sequences have been pruned, it is often convenient to parse the symbolic strings into words of variable length - this is called coding. Suppose that the grammar can be stated as a finite number of pruning rules, each forbidding a block of finite length,

$$
\begin{equation*}
\mathcal{G}=\left\{b_{1}, b_{2}, \cdots b_{k}\right\}, \tag{11.23}
\end{equation*}
$$

where a pruning block $b$ is a sequence of symbols $b=s_{1} s_{2} \cdots s_{n_{b}}, s \in \mathcal{A}$, of finite length $n_{b}$. In this case we can always construct a finite Markov partition (11.4) by replacing finite length words of the original partition by letters of a new alphabet. In particular, if the longest forbidden block is of length $M+1$, we say that the symbolic dynamics is a shift of finite type with $M$-step memory. In that case we can recode the symbolic dynamics in terms of a new alphabet, with each new letter given by an admissible block of at most length $M$. In the new alphabet the grammar rules are implemented by setting $T_{i j}=0$ in (11.3) for forbidden transitions.

A topological dynamical system $(\Sigma, \sigma)$ for which all admissible itineraries are generated by a finite transition matrix

$$
\begin{equation*}
\Sigma=\left\{\left(s_{k}\right)_{k \in \mathbb{Z}}: T_{s_{k} s_{k+1}}=1 \quad \text { for all } k\right\} \tag{11.24}
\end{equation*}
$$

is called a subshift of finite type. Such systems are particularly easy to handle; the topology can be converted into symbolic dynamics by representing the transition matrix by a finite directed Markov graph, a convenient visualization of topological dynamics.

Markov graphs. A Markov graph describes compactly the ways in which the phase-space regions map into each other, accounts for finite memory effects in dynamics, and generates the totality of admissible trajectories as the set of all possible walks along its links.

A Markov graph consists of a set of nodes (or vertices, or states), one for each state in the alphabet $\mathcal{A}=\{A, B, C, \cdots, Z\}$, connected by a set of directed links (edges, arcs). Node $i$ is connected by a directed link to node $j$ whenever the transition matrix element (11.2) takes value $T_{i j}=1$. There might be a set of links connecting two nodes, or links that originate and terminate on the same node. Two graphs are isomorphic if one can be obtained from the other by relabeling links and nodes; for us they are one and the same graph. As we are interested in recurrent dynamics, we restrict our attention to irreducible or strongly connected graphs, that is, graphs for which there is a path from any node to any other node.

The simplest example is given in figure 11.15.
(a) $T=\left(\begin{array}{ll}1 & 1 \\ 1 & 0\end{array}\right)(b$


Figure 11.15: (a) The transition matrix for a simple subshift on two-state partition $\mathcal{A}=\{0,1\}$, with grammar $\mathcal{G}$ given by a single pruning block $b=11$ (consecutive repeat of symbol 1 is inadmissible): the state $\mathcal{M}_{0}$ maps both onto $\mathcal{M}_{0}$ and $\mathcal{M}_{1}$, but the state $\mathcal{M}_{1}$ maps only onto $\mathcal{M}_{0}$. (b) The corresponding finite 2-node, 3-links Markov graph, with nodes coding the symbols. All admissible itineraries are generated as walks on this finite Markov graph.

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## Commentary

Remark 11.1 Symbolic dynamics, history and good taste. For a brief history of symbolic dynamics, from J. Hadamard in 1898 onward, see Notes to chapter 1 of Kitchens monograph [11.1], a very clear and enjoyable mathematical introduction to topics discussed here. Finite Markov graphs or finite automata are discussed in refs. $[11.2,11.3,11.4,11.5]$. They belong to the category of regular languages. A good hands-on introduction to symbolic dynamics is given in ref. [11.6].

The binary labeling of the once-folding map periodic points was introduced by Myrberg [11.7] for one-dimensional maps, and its utility to two-dimensional maps has been emphasized in refs. [3.5, 3.9]. For one-dimensional maps it is now customary to use the $R$ - $L$ notation of Metropolis, Stein and Stein [11.8, 11.9], indicating that the point $x_{n}$ lies either to the left or to the right of the critical point in figure 11.9. The symbolic dynamics of such mappings has been extensively studied by means of the Smale horseshoes, see for example ref. [11.10]. Using letters rather than numerals in symbol dynamics alphabets probably reflects good taste. We prefer numerals for their computational convenience, as they speed up the implementation of conversions into the topological coordinates $(\delta, \gamma)$ introduced in sect. 12.3.1.

Remark 11.2 Inflating Markov graphs. In the above examples the symbolic dynamics has been encoded by labeling links in the Markov graph. Alternatively one can encode the dynamics by labeling the nodes, as in figure 11.13, where the 4 nodes refer to 4 Markov partition regions $\left\{\mathcal{M}_{00}, \mathcal{M}_{01}, \mathcal{M}_{10}, \mathcal{M}_{11}\right\}$, and the 8 links to the 8 non-zero entries in the 2-step memory transition matrix (11.15).

## Résumé

In chapters 14 and 15 we will establish that spectra of evolution operators can be extracted from periodic orbit sums:

$$
\sum(\text { spectral eigenvalues })=\sum(\text { periodic orbits })
$$

In order to implement this theory we need to know what periodic orbits can exist, and the symbolic dynamics developed above and in chapter 12 is an invaluable tool toward this end.


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## Exercises

Exercise 11.1 Binary symbolic dynamics. Verify that the shortest prime binary cycles of the unimodal repeller of figure 11.9 are $\overline{0}, \overline{1}, \overline{01}, \overline{001}, \overline{011}, \cdots$. Compare with table 11.1. Try to sketch them in the graph of the unimodal function $f(x)$; compare ordering of the periodic points with figure 11.10. The point is that while overlayed on each other the longer cycles look like a hopeless jumble, the cycle points are clearly and logically ordered by the alternating binary tree.

Exercise 11.2 3-disk fundamental domain symbolic dynamics. Try to sketch $\overline{0}, \overline{1}, \overline{01}, \overline{001}, \overline{011}, \cdots$. in the fundamental domain, figure 11.6 , and interpret the symbols $\{0,1\}$ by relating them to topologically distinct types of collisions. Compare with table 11.2. Then try to sketch the location of periodic points in the Poincaré section of the billiard flow. The point of this exercise is that while in the configuration space longer cycles look like a hopeless jumble, in the Poincaré section they are clearly and logically ordered. The Poincaré section is always to be preferred to projections of a flow onto the configuration space coordinates, or any other subset of phase space coordinates which does not respect the topological organization of the flow.

Exercise 11.3 Generating prime cycles. Write a program that generates all binary prime cycles up to given finite length.

Exercise 11.4 A contracting baker's map. Consider a contracting (or "dissipative") baker's defined in exercise 4.4.

The symbolic dynamics encoding of trajectories is realized via symbols $0(y \leq 1 / 2)$ and $1(y>1 / 2)$. Consider the observable $a(x, y)=x$. Verify that for any periodic orbit $p\left(\epsilon_{1} \ldots \epsilon_{n_{p}}\right), \epsilon_{i} \in\{0,1\}$

$$
A_{p}=\frac{3}{4} \sum_{j=1}^{n_{p}} \delta_{j, 1} .
$$

## Exercise 11.5 Reduction of 3-disk symbolic dynamics to binary.

(a) Verify that the 3-disk cycles
$\{\overline{12}, \overline{13}, \overline{23}\},\{\overline{123}, \overline{132}\},\{\overline{1213}+2$ perms. $\}$, $\{\overline{121232313}+5$ perms. $\},\{\overline{121323}+2$ perms. $\}, \cdots$, correspond to the fundamental domain cycles $\overline{0}, \overline{1}, \overline{01}, \overline{001}, \overline{011}, \ldots$ respectively.
(b) Check the reduction for short cycles in table 11.2 by drawing them both in the full 3 -disk system and in the fundamental domain, as in figure 11.6.
(c) Optional: Can you see how the group elements listed in table 11.2 relate irreducible segments to the fundamental domain periodic orbits?

Exercise 11.6 Unimodal map symbolic dynamics. Show that the tent map point $\gamma\left(S^{+}\right)$with future itinerary $S^{+}$is given by converting the sequence of $s_{n}$ 's into a binary number by the algorithm (11.11). This follows by inspection from the binary tree of figure 11.10.

Exercise 11.7 "Golden mean" pruned map. Consider a symmetrical tent map on the unit interval such that its highest point belongs to a 3-cycle:

(a) Find the absolute value $\Lambda$ for the slope (the two different slopes $\pm \Lambda$ just differ by a sign) where the maximum at $1 / 2$ is part of a period three orbit, as in the figure.
(b) Show that no orbit of this map can visit the region $x>(1+\sqrt{5}) / 4$ more than once. Verify that once an orbit exceeds $x>(\sqrt{5}-1) / 4$, it does not reenter the region $x<(\sqrt{5}-1) / 4$.
(c) If an orbit is in the interval $(\sqrt{5}-1) / 4<x<1 / 2$, where will it be on the next iteration?
(d) If the symbolic dynamics is such that for $x<1 / 2$ we use the symbol 0 and for $x>1 / 2$ we use the symbol 1 , show that no periodic orbit will have the substring _00_ in it.
(e) On the second thought, is there a periodic orbit that violates the above _00_ pruning rule?

For continuation, see exercise 13.6 and exercise 13.8. See also exercise 13.7 and exercise 13.9.

Exercise 11.8 Binary 3 -step transition matrix. Construct [ $8 \times 8$ ] binary 3 -step transition matrix analogous to the 2 -step transition matrix (11.15). Convince yourself that the number of terms of contributing to $\operatorname{tr} T^{n}$ is independent of the memory length, and that this $\left[2^{m} \times 2^{m}\right]$ trace is well defined in the infinite memory limit $m \rightarrow \infty$.

## Chapter 12

## Qualitative dynamics, for cyclists

I.1. Introduction to conjugacy problems for diffeomorphisms. This is a survey article on the area of global analysis defined by differentiable dynamical systems or equivalently the action (differentiable) of a Lie group $G$ on a manifold $M$. Here $\operatorname{Diff}(M)$ is the group of all diffeomorphisms of $M$ and a diffeomorphism is a differentiable map with a differentiable inverse. (...) Our problem is to study the global structure, that is, all of the orbits of $M$.
Stephen Smale, Differentiable Dynamical Systems
In sects. 9.1 and 11.1 we introduced the concept of partitioning the phase space, in any way you please. In chapter 8 we established that stability eigenvalues of periodic orbits are invariants of a given flow. The invariance of stabilities of a periodic orbit is a local property of the flow.

Both in the Rössler flow of example 3.3, and in the Kuramoto-Sivashinsky system of example 3.4 we have learned that the attractor is very thin, but otherwise the return maps that we found were disquieting - neither figure 3.2 nor figure 3.3 appeared to be one-to-one maps. This apparent loss of invertibility is an artifact of projection of higher-dimensional return maps onto lower-dimensional subspaces. As the choice of lower-dimensional subspace is arbitrary, the resulting snapshots of return maps look rather arbitrary, too. Other projections might look even less suggestive.

Such observations beg a question: Does there exist a "natural", intrinsically optimal coordinate system in which we should plot of a return map?

As we shall now argue (see also sect. 17.1), the answer is yes: The intrinsic coordinates are given by the stable/unstable manifolds, and a return map should be plotted as a map from the unstable manifold back onto the immediate neighborhood of the unstable manifold.

In this chapter we show that every equilibrium point and every peri-
odic orbit carries with it stable and unstable manifolds which provide a topologically invariant global foliation of the phase space. This qualitative dynamics of stretching and mixing enables us to partition the phase space and assign symbolic dynamics itineraries to trajectories.

Given an itinerary, the topology of stretching and folding fixes the relative spatial ordering of trajectories, and separates the admissible and inadmissible itineraries. The level is distinctly cyclist, in distincition to the pedestrian tempo of the preceeding chapter. Skip this chapter unless you really need to get into nitty-gritty details of symbolic dynamics.


### 12.1 Going global: Stable/unstable manifolds

A neighborhood of a trajectory deforms as it is transported by the flow. In the linear approximation, the matrix of variations $A$ describes this shearing of an infinitesimal neighborhood in an infinitesimal time step. The shearing after finite time is described by the Jacobian matrix $\mathbf{J}^{t}$. Its eigenvalues and eigendirections describe deformation of an initial infinitesimal sphere of neighboring trajectories into an ellipsoid time $t$ later. Nearby trajectories separate exponentially along the unstable directions, approach each other along the stable directions, and maintain their distance along the marginal directions.

The fixed or periodic point $x^{*}$ stability matrix $\mathbf{J}_{p}\left(x^{*}\right)$ eigenvectors (8.9) form a rectilinear coordinate frame in which the flow into, out of, or encircling the fixed point is linear in the sense of sect. 4.2. These eigendirections are numerically continued into global curvilinear invariant manifolds as follows.

The global continuations of the local stable, unstable eigendirections are called the stable, respectively unstable manifolds. They consist of all points which march into the fixed point forward, respectively backward in time

$$
\begin{align*}
W^{s} & =\left\{x \in \mathcal{M}: f^{t}(x)-x^{*} \rightarrow 0 \text { as } t \rightarrow \infty\right\} \\
W^{u} & =\left\{x \in \mathcal{M}: f^{-t}(x)-x^{*} \rightarrow 0 \text { as } t \rightarrow \infty\right\} . \tag{12.1}
\end{align*}
$$

The stable/unstable manifolds of a flow are rather hard to visualize, so as long as we are not worried about a global property such as the number of times they wind around a periodic trajectory before completing a parcourse, we might just as well look at their Poincaré section return maps. Stable, unstable manifolds for maps are defined by

$$
\begin{align*}
W^{s} & =\left\{x \in \mathcal{P}: f^{n}(x)-x^{*} \rightarrow 0 \text { as } n \rightarrow \infty\right\} \\
W^{u} & =\left\{x \in \mathcal{P}: f^{-n}(x)-x^{*} \rightarrow 0 \text { as } n \rightarrow \infty\right\} . \tag{12.2}
\end{align*}
$$

For $n \rightarrow \infty$ any finite segment of $W^{s}$, respectively $W^{u}$ converges to the linearized map eigenvector $\mathbf{e}^{(e)}$, respectively $\mathbf{e}^{(c)}$. In this sense each eigenvector defines a (curvilinear) axis of the stable, respectively unstable manifold.

Conversely, we can use an arbitrarily small segment of a fixed point eigenvector to construct a finite segment of the associated manifold. Precise construction depends on the type of the eigenvalue(s).

Expanding real and positive eigendirection. Consider $i$ th expanding eigenvalue, eigenvector pair ( $\Lambda_{i}, \mathbf{e}_{i}$ ) computed from $\mathbf{J}$ evaluated at a cycle point,

$$
\begin{equation*}
\mathbf{J}(x) \mathbf{e}_{i}(x)=\Lambda_{i} \mathbf{e}_{i}(x), \quad x \in p, \quad \Lambda_{i}>1 \tag{12.3}
\end{equation*}
$$

Take an infinitesimal eigenvector $\epsilon \mathbf{e}_{i}(x), \epsilon \ll 1$, and its image $\mathbf{J}_{p}(x) \epsilon \mathbf{e}_{i}(x)=$ $\Lambda_{i} \epsilon \mathbf{e}_{i}(x)$. Sprinkle the interval $\left|\Lambda_{i}-1\right| \epsilon$ with a large number of points $x_{m}$, equidistantly spaced on logarithmic scale $\ln \left|\Lambda_{i}-1\right|+\ln \epsilon$. The successive images of these points $f\left(x_{j}\right), f^{2}\left(x_{j}\right), \cdots, f^{m}\left(x_{j}\right)$ trace out the curvilinear unstable manifold in direction $\mathbf{e}_{i}$. Repeat for $-\epsilon \mathbf{e}_{i}(x)$.

Contractiong real, positive eigendirection. Reverse the action of the map backwards in time. This turns a contracting direction into an expanding one, tracing out the curvilinear stable manifold in continuation of $\epsilon \mathbf{e}_{j}$.

Expanding/contracting real negative eigendirection. As above, but every even iterate $f^{2}\left(x_{j}\right), f^{4}\left(x_{j}\right), f^{6}\left(x_{j}\right)$ continues in the direction $\mathbf{e}_{i}$, every odd one in the direction $-\mathbf{e}_{i}$.

Complex eigenvalue pair. Construct an orthonormal pair of eigenvectors spanning the plane $\left\{\epsilon \mathbf{e}_{j}, \epsilon \mathbf{e}_{j+1}\right\}$. Iteration of the annulus between an infinitesimal circle and its image by $\mathbf{J}$ spans the spiralling/circle unstable manifold of the complex eigenvalue pair $\left\{\Lambda_{i}, \Lambda_{i+1}=\Lambda_{i}^{*}\right\}$.

### 12.2 Horseshoes

If a flow is locally unstable but globally bounded, any open ball of initial points will be stretched out and then folded back. An example is a 3dimensional invertible flow sketched in figure 11.7 which returns an area of a Poincaré section of the flow stretched and folded into a "horseshoe", such that the initial area is intersected at most twice. Run backwards, the flow generates the backward horseshoe which intersects the forward horseshoe at most 4 times, and so forth. Such flows exist, and are easily constructed - an example is the Rössler system, discussed in example 3.3.

Now we shall construct an example of a locally unstable but globally bounded mapping which returns an initial area stretched and folded into
a "horseshoe", such that the initial area is intersected at most twice. We shall refer to such mappings with at most $2^{n}$ transverse self-intersections at the $n$th iteration as the once-folding maps.

As an example is afforded by the 2-dimensional Hénon map

$$
\begin{align*}
x_{n+1} & =1-a x_{n}^{2}+b y_{n} \\
y_{n+1} & =x_{n} . \tag{12.4}
\end{align*}
$$

The Hénon map models qualitatively the Poincaré section return map of figure 11.7. For $b=0$ the Hénon map reduces to the parabola (11.9), and, as shown in sects. 3.3 and 31.1, for $b \neq 0$ it is kind of a fattened parabola; by construction, it takes a rectangular initial area and returns it bent as a horseshoe.

For definitiveness, fix the parameter values to $a=6, b=0.9$. The map is quadratic, so it has 2 fixed points $x_{0}=f\left(x_{0}\right), x_{1}=f\left(x_{1}\right)$ indicated in figure 12.1 (a). For the parameter values at hand, they are both unstable. If you start with a small ball of initial points centered around $x_{1}$, and iterate the map, the ball will be stretched and squashed along the line $W_{1}^{u}$.
Similarly, a small ball of initial points centered around the other fixed point $x_{0}$ iterated backward in time,

$$
\begin{align*}
x_{n-1} & =y_{n} \\
y_{n-1} & =-\frac{1}{b}\left(1-a y_{n}^{2}-x_{n}\right) \tag{12.5}
\end{align*}
$$

traces out the line $W_{0}^{s}$. $W_{0}^{s}$ is the stable manifold of $x_{0}$ fixed point, and $W_{1}^{u}$ is the unstable manifold of $x_{1}$ fixed point, defined in sect. 12.1. Their intersections enclose the crosshatched region $\mathcal{M}$. Any point outside $W_{1}^{u}$ border of $\mathcal{M}$. escapes to infinity forward in time, while any point outside $W_{0}^{s}$ border escapes to infinity backwards in time. In this way the unstable - stable manifolds define topologically, invariant and optimal $\mathcal{M}$. initial region; all orbits that stay confined for all times are confined to $\mathcal{M}$. .

Iterated one step forward, the region $\mathcal{M}$. is stretched and folded into a smale horseshoe drawn in figure 12.1 (b). The horseshoe fattened parabolla shape is the consequence og the quadratic form $x^{2}$ in (12.4). Parameter $a$ controls the amount of stretching, while the parameter $b$ controls the amount of compression of the folded horseshoe. The case $a=6, b=0.9$ considered here corresponds to strong stretching and weak compression. Label the two forward intersections $f(\mathcal{M}.) \cap \mathcal{M}$. by $\mathcal{M}_{s}$., with $s \in\{0,1\}$, figure 12.1 (b). The horseshoe consists of the two strips $\mathcal{M}_{0 .}, \mathcal{M}_{1}$, and the bent segment that lies entirely outside the $W_{1}^{u}$ line. As all points in this segment escape to infinity under forward iteration, this region can safely be cut out and thrown away.

Iterated one step backwards, the region $\mathcal{M}$. is again stretched and folded into a horseshoe, figure 12.1 (c). As stability and instability are interchanged under time reversal, this horseshoe is transverse to the forward


Figure 12.1: The Hénon map for $a=6, b=.9$. (a) The fixed points $\overline{0}, \overline{1}$, and the segments of the $W_{0}^{s}$ stable manifold, $W_{1}^{u}$ unstable manifold that enclose the initial (crosshatched) region $\mathcal{M}$. (b) The forward horseshoe $f(\mathcal{M}$.). (c) The backward horseshoe $f^{-1}(\mathcal{M}$.$) . Iteration yields a complete Smale horseshoe, with every forward$ fold intersecting every backward fold.
one. Again the points in the horseshoe bend wonder off to infinity as $n \rightarrow-\infty$, and we are left with the two (backward) strips $\mathcal{M}_{.0}, \mathcal{M}_{.1}$. Iterating two steps forward we obtain the four strips $\mathcal{M}_{11}, \mathcal{M}_{01}, \mathcal{M}_{00}, \mathcal{M}_{10}$, and iterating backwards we obtain the four strips $\mathcal{M}_{.00}, \mathcal{M}_{.01}, \mathcal{M}_{.11}, \mathcal{M}_{.10}$ transverse to the forward ones just as for 3-disk pinball game figure 11.3. Iterating three steps forward we get an 8 strips, and so on ad infinitum.

What is the significance of the subscript .011 which labels the $\mathcal{M} .011$ backward strip? The two strips $\mathcal{M}_{.0}, \mathcal{M}_{.1}$ partition the phase space into two regions labeled by the two-letter alphabet $\mathcal{A}=\{0,1\} . \quad S^{+}=.011$ is the future itinerary for all $x \in \mathcal{M} .011$. Likewise, for the forward strips all $x \in \mathcal{M}_{s_{-m} \cdots s_{-1} s_{0}}$. have the past itinerary $S^{-}=s_{-m} \cdots s_{-1} s_{0}$. Which partition we use to present pictorially the regions that do not escape in $m$ iterations is a matter of taste, as the backward strips are the preimages of the forward ones

$$
\mathcal{M}_{0 .}=f\left(\mathcal{M}_{.0}\right), \quad \mathcal{M}_{1 .}=f\left(\mathcal{M}_{.1}\right)
$$

$\Omega$, the non-wandering set (2.2) of $\mathcal{M}$., is the union of all points whose forward and backward trajectories remain trapped for all time. given by the intersections of all images and preimages of $\mathcal{M}$ :

$$
\begin{equation*}
\Omega=\left\{x: x \in \lim _{m, n \rightarrow \infty} f^{m}(\mathcal{M} .) \bigcap f^{-n}(\mathcal{M} .)\right\} \tag{12.6}
\end{equation*}
$$

Two important properties of the Smale horseshoe are that it has a complete binary symbolic dynamics and that it is structurally stable.

For a complete Smale horseshoe every forward fold $f^{n}(\mathcal{M})$ intersects transversally every backward fold $f^{-m}(\mathcal{M})$, so a unique bi-infinite binary sequence can be associated to every element of the non-wandering set. A point $x \in \Omega$ is labeled by the intersection of its past and future itineraries $S(x)=\cdots s_{-2} s_{-1} s_{0} \cdot s_{1} s_{2} \cdots$, where $s_{n}=s \quad$ if $\quad f^{n}(x) \in \mathcal{M}_{. s}, s \in\{0,1\}$ and $n \in \mathbb{Z}$. For sufficiently separated disks, the 3 -disk game of pinball figure 11.3, is another example of a complete Smale horseshoe; in this case the "folding" region of the horseshoe is cut out of the picture by allowing the pinballs that fly between the disks to fall off the table and escape.

The system is said to be structurally stable if all intersections of forward and backward iterates of $\mathcal{M}$ remain transverse for sufficiently small perturbations $f \rightarrow f+\delta$ of the flow, for example, for slight displacements of the disks, or sufficiently small variations of the Hénon map parameters $a, b$ while structural stability is exceedingly desirable, it is also exceedingly rare. About this, more later.

### 12.3 Spatial ordering

Consider a system for which you have succeeded in constructing a covering symbolic dynamics, such as a well-separated 3 -disk system. Now start moving the disks toward each other. At some critical separation a disk will start blocking families of trajectories traversing the other two disks. The order in which trajectories disappear is determined by their relative ordering in space; the ones closest to the intervening disk will be pruned first. Determining inadmissible itineraries requires that we relate the spatial ordering of trajectories to their time ordered itineraries.

So far we have rules that, given a phase space partition, generate a temporally ordered itinerary for a given trajectory. Our next task is the reverse: given a set of itineraries, what is the spatial ordering of corresponding points along the trajectories? In answering this question we will be aided by Smale's visualization of the relation between the topology of a flow and its symbolic dynamics by means of "horseshoes".

### 12.3.1 Symbol square

For a better visualization of 2-dimensional non-wandering sets, fatten the intersection regions until they completely cover a unit square, as in figure 12.2 . We shall refer to such a "map" of the topology of a given "stretch \& fold" dynamical system as the symbol square. The symbol square is a topologically accurate representation of the non-wandering set and serves as a street map for labeling its pieces. Finite memory of $m$ steps and finite foresight of $n$ steps partitions the symbol square into rectangles


FIG. 4. Iterative construction of the symbol plane.

Figure 12.2: Kneading Danish Pastry: symbol square representation of an orientation reversing once-folding map obtained by fattening the Smale horseshoe intersections of figure 12.1 into a unit square. In the symbol square the dynamics maps rectangles into rectangles by a decimal point shift.
$\left[s_{-m+1} \cdots s_{0} \cdot s_{1} s_{2} \cdots s_{n}\right]$. In the binary dynamics symbol square the size of such rectangle is $2^{-m} \times 2^{-n}$; it corresponds to a region of the dynamical phase space which contains all points that share common $n$ future and $m$ past symbols. This region maps in a nontrivial way in the phase space, but in the symbol square its dynamics is exceedingly simple; all of its points are mapped by the decimal point shift (11.20)

$$
\begin{equation*}
\sigma\left(\cdots s_{-2} s_{-1} s_{0} \cdot s_{1} s_{2} s_{3} \cdots\right)=\cdots s_{-2} s_{-1} s_{0} s_{1} \cdot s_{2} s_{3} \cdots \tag{12.7}
\end{equation*}
$$

For example, the square [01.01] gets mapped into the rectangle $\sigma[01.01]=$ [010.1].

As the horseshoe mapping is a simple repetitive operation, we expect a simple relation between the symbolic dynamics labeling of the horseshoe strips, and their relative placement. The symbol square points $\gamma\left(S^{+}\right)$with future itinerary $S^{+}$are constructed by converting the sequence of $s_{n}$ 's into a binary number by the algorithm (??). This follows by inspection from figure 12.2. In order to understand this relation between the topology of horseshoes and their symbolic dynamics, it might be helpful to backtrace to sect. 11.3.2 and work through and understand first the symbolic dynamics of one-dimensional unimodal mappings.

Under backward iteration the roles of 0 and 1 symbols are interchanged; $\mathcal{M}_{0}^{-1}$ has the same orientation as $\mathcal{M}$, while $\mathcal{M}_{1}^{-1}$ has the opposite orientation. We assign to an orientation preserving once-folding map the past
topological coordinate $\delta=\delta\left(S^{-}\right)$by the algorithm:

$$
\begin{align*}
& w_{n-1}=\left\{\begin{array}{ll}
w_{n} & \text { if } s_{n}=0 \\
1-w_{n} & \text { if } s_{n}=1
\end{array}, \quad w_{0}=s_{0}\right. \\
& \delta\left(S^{-}\right)=0 . w_{0} w_{-1} w_{-2} \ldots=\sum_{n=1}^{\infty} w_{1-n} / 2^{n} . \tag{12.8}
\end{align*}
$$

Such formulas are best derived by quiet contemplation of the action of a folding map, in the same way we derived the future topological coordinate (??).

The coordinate pair $(\delta, \gamma)$ maps a point $(x, y)$ in the phase space Cantor set of figure 12.1 into a point in the symbol square of figure 12.2, preserving the topological ordering; $(\delta, \gamma)$ serves as a topologically faithful representation of the non-wandering set of any once-folding map, and aids us in partitioning the set and ordering the partitions for any flow of this type.

### 12.4 Pruning

The complexity of this figure will be striking, and I shall not even try to draw it.
H. Poincaré, on his discovery of homoclinic tangles, Les méthodes nouvelles de la méchanique céleste

In general, not all possible itineraries are realized as physical trajectories. Trying to get from "here" to "there" we might find that a short path is excluded by some obstacle, such as a disk that blocks the path, or a potential ridge. To count correctly, we need to prune the inadmissible trajectories, that is, specify the grammar of the admissible itineraries.

While the complete Smale horseshoe dynamics discussed so far is rather straightforward, we had to get through it in order to be able to approach a situation that resembles more the real life: adjust the parameters of a oncefolding map so that the intersection of the backward and forward folds is still transverse, but no longer complete, as in figure 13.2 (a). The utility of the symbol square lies in the fact that the surviving, admissible itineraries still maintain the same relative spatial ordering as for the complete case.

In the example of figure 13.2 (a) the rectangles [10.1], [11.1] have been pruned, and consequently any trajectory containing blocks $b_{1}=101, b_{2}=$ 111 is pruned. We refer to the border of this primary pruned region as the pruning front; another example of a pruning front is drawn in figure 13.2 (d). We call it a "front" as it can be visualized as a border between admissible and inadmissible; any trajectory whose periodic point would fall to the right of the front in figure 13.2 is inadmissible, that is, pruned. The pruning front is a complete description of the symbolic dynamics of once-folding maps. For now we need this only as a concrete illustration of how pruning rules arise.

In the example at hand there are total of two forbidden blocks 101, 111, so the symbol dynamics is a subshift of finite type (11.24). For now we concentrate on this kind of pruning because it is particularly clean and simple. Unfortunately, for a generic dynamical system a subshift of finite type is the exception rather than the rule. Only some repelling sets (like our game of pinball) and a few purely mathematical constructs (called Anosov flows) are structurally stable - for most systems of interest an infinitesimal perturbation of the flow destroys and/or creates an infinity of trajectories, and specification of the grammar requires determination of pruning blocks of arbitrary length. The repercussions are dramatic and counterintuitive; for example, due to the lack of structural stability the transport coefficients such as the deterministic diffusion constant of sect. 23.2 are emphatically not smooth functions of the system parameters. This generic lack of structural stability is what makes nonlinear dynamics so hard.

The conceptually simpler finite subshift Smale horseshoes suffice to motivate most of the key concepts that we shall need for time being.

### 12.4.1 Converting pruning blocks into Markov graphs

The complete binary symbolic dynamics is too simple to be illuminating, so we turn next to the simplest example of pruned symbolic dynamics, the finite subshift obtained by prohibition of repeats of one of the symbols, let us say _00_. This situation arises, for example, in studies of the circle maps, where this kind of symbolic dynamics describes "golden mean" rotations (we shall return to this example in chapter 24). Now the admissible itineraries are enumerated by the pruned binary tree of figure 11.14 (a), or the corresponding Markov graph figure 11.14 (b). We recognize this as the Markov graph example of figure 11.15.

So we can already see the main ingradients of a general algorithm: (1) Markov graph encodes self-similarities of the tree of all itineraries, and (2) if we have a pruning block of length $M$, we need to descend $M$ levels before we can start identifying the self-similar sub-trees.

Suppose now that, by hook or crook, you have been so lucky fishing for pruning rules that you now know the grammar (11.23) in terms of a finite set of pruning blocks $\mathcal{G}=\left\{b_{1}, b_{2}, \cdots b_{k}\right\}$, of lengths $n_{b_{m}} \leq M$. Our task is to generate all admissible itineraries. What to do?

## A Markov graph algorithm.

1. Starting with the root of the tree, delineate all branches that correspond to all pruning blocks; implement the pruning by removing the last node in each pruning block.
2. Label all nodes internal to pruning blocks by the itinerary connecting the root point to the internal node. Why? So far we have pruned forbidden branches by looking $n_{b}$ steps into future for all pruning
blocks. into future for pruning block $b=$ [.10010]. However, the blocks with a right combination of past and future [1.0110], [10.110], [101.10] and [1011.0] are also pruned. In other words, any node whose near past coincides with the begining of a pruning block is potentially dangerous - a branch further down the tree might get pruned.
3. Add to each internal node all remaining branches allowed by the alphabet, and label them. Why? Each one of them is the beginning point of an infinite tree, a tree that should be similar to another one originating closer to the root of the whole tree.
4. Pick one of the free external nodes closest to the root of the entire tree, forget the most distant symbol in its past. Does the truncated itinerary correspond to an internal node? If yes, identify the two nodes. If not, forget the next symbol in the past, repeat. If no such truncated past corresponds to any internal node, identify with the root of the tree.

This is a little bit abstract, so let's say the free external node in question is [1010.]. Three time steps back the past is [010.]. That is not dangerous, as no pruning block in this example starts with 0 . Now forget the third step in the past: [10.] is dangerous, as that is the start of the pruning block [10.110]. Hence the free external node [1010.] should be identified with the internal node [10.].
5. Repeat until all free nodes have been tied back into the internal nodes.
6. Clean up: check whether every node can be reached from every other node. Remove the transient nodes, that is, the nodes to which dynamics never returns.
7. The result is a Markov diagram. There is no guarantee that this is the smartest, most compact Markov diagram possible for given pruning (if you have a better algorithm, teach us), but walks around it do generate all admissible itineraries, and nothing else.

## Heavy pruning.

We complete this training by examples by implementing the pruning of figure 13.2 (d). The pruning blocks are
[100.10], [10.1], [010.01], [011.01], [11.1], [101.10].

Blocks 01101, 10110 contain the forbidden block 101, so they are redundant as pruning rules. Draw the pruning tree as a section of a binary tree with 0 and 1 branches and label each internal node by the sequence of 0 's and 1's connecting it to the root of the tree (figure 13.3 (a). These nodes are the potentially dangerous nodes - beginnings of blocks that might end up pruned. Add the side branches to those nodes (figure 13.3 (b). As we continue down such branches we have to check whether the pruning
imposes constraints on the sequences so generated: we do this by knocking off the leading bits and checking whether the shortened strings coincide with any of the internal pruning tree nodes: $00 \rightarrow 0 ; 110 \rightarrow 10 ; 011 \rightarrow 11$; $0101 \rightarrow 101$ (pruned); $1000 \rightarrow 00 \rightarrow 00 \rightarrow 0 ; 10011 \rightarrow 0011 \rightarrow 011 \rightarrow 11$; $01000 \rightarrow 0$.

As in the previous two examples, the trees originating in identified nodes are identical, so the tree is "self-similar". Now connect the side branches to the corresponding nodes, figure 13.3 (d). Nodes "." and 1 are transient nodes; no sequence returns to them, and as you are interested here only in infinitely recurrent sequences, delete them. The result is the finite Markov graph of figure 13.3 (d); the admissible bi-infinite symbol sequences are generated as all possible walks along this graph.

## Commentary

Remark 12.1 Smale horseshoe. S. Smale understood clearly that the crucial ingredient in the description of a chaotic flow is the topology of its non-wandering set, and he provided us with the simplest visualization of such sets as intersections of Smale horseshoes. In retrospect, much of the material covered here can already be found in Smale's fundamental paper [1.15], but a physicist who has run into a chaotic time series in his laboratory might not know that he is investigating the action (differentiable) of a Lie group $G$ on a manifold $M$, and that the Lefschetz trace formula is the way to go. If you find yourself mystified by Smale's article abstract about "the action (differentiable) of a Lie group $G$ on a manifold $M$ ", quoted on page 189, rereading chapter 9 might help; for example, the Liouville operators form a Lie group (of symplectic, or canonical transformations) acting on the manifold $(p, q)$.

Remark 12.2 Kneading theory. The admissible itineraries are studied in refs. [12.14, 11.8, 11.10, 11.11], as well as many others. We follow here the Milnor-Thurston exposition [12.15]. They study the topological zeta function for piecewise monotone maps of the interval, and show that for the finite subshift case it can be expressed in terms of a finite dimensional kneading determinant. As the kneading determinant is essentially the topological zeta function that we introduce in (13.4), we shall not discuss it here. Baladi and Ruelle have reworked this theory in a series of papers $[12.17,12.18,12.19]$ and in ref. [12.20] replaced it by a power series manipulation. The kneading theory is covered here in P. Dahlqvist's appendix E.1.

Remark 12.3 Pruning fronts. The notion of a pruning front was introduced in ref. [12.21], and developed by K.T. Hansen for a number of dynamical systems in his Ph.D. thesis [1.3] and a series of papers [12.27]-[12.31]. Detailed studies of pruning fronts are carried out in refs. [12.22, 12.23, 12.52]; ref. [31.5] is the most detailed study carried out so far. The rigorous theory of pruning fronts has been developed by Y. Ishii [12.24, 12.25] for the Lozi map, and A. de Carvalho [12.26] in a very general setting.

Remark 12.4 The unbearable growth of Markov graphs. A construction of finite Markov partitions is described in refs. [12.56, 12.57], as well as in the innumerably many other references.

If two regions in a Markov partition are not disjoint but share a boundary, the boundary trajectories require special treatment in order to avoid overcounting, see sect. 22.3.1. If the image of a trial partition region cuts across only a part of another trial region and thus violates the Markov partition condition (11.4), a further refinement of the partition is needed to distinguish distinct trajectories figure 13.2 is an example of such refinements.

The finite Markov graph construction sketched above is not necessarily the minimal one; for example, the Markov graph of figure 13.3 does not generate only the "fundamental" cycles (see chapter 18), but shadowed cycles as well, such as $t_{00011}$ in (13.17). For methods of reduction to a minimal graph, consult refs. [11.5, 12.51, 12.53]. Furthermore, when one implements the time reversed dynamics by the same algorithm, one usually gets a graph of very different topology even though both graphs generate the same admissible sequences, and have the same determinant. The algorithm described here makes some sense for 1- $d$ dynamics, but is unnatural for $2-d$ maps whose dynamics it treats as one-dimensional. In practice, generic pruning grows longer and longer, and more plentiful pruning rules. For generic flows the refinements might never stop, and almost always we might have to deal with infinite Markov partitions, such as those that will be discussed in sect. 13.6. Not only do the Markov graphs get more and more unwieldy, they have the unpleasant property that every time we add a new rule, the graph has to be constructed from scratch, and it might look very different form the previous one, even though it leads to a minute modification of the topological entropy. The most determined effort to construct such graphs may be the one of ref. [12.22]. Still, this seems to be the best technology available, unless the reader alerts us to something superior.

## Résumé

Given a partition $\mathcal{A}$ of the phase space $\mathcal{M}$, a dynamical system $(\mathcal{M}, f)$ induces topological dynamics $(\Sigma, \sigma)$ on the space $\Sigma$ of all admissible biinfinite itineraries. The itinerary describes the time evolution of an orbit, while (for 2-d hyperbolic maps) the symbol square describes the spatial ordering of points along the orbit. The rule that everything to one side of the pruning front is forbidden might (in hindsight) seem obvious, but if you have ever tried to work out symbolic dynamics of some "generic" dynamical system, you should be struck by its simplicity: instead of pruning a Cantor set embedded within some larger Cantor set, the pruning front cleanly cuts out a compact region in the symbol square and that is all - there are no additional pruning rules.

The symbol square is a useful tool in transforming topological pruning into pruning rules for inadmissible sequences; those are implemented by constructing transition matrices and/or Markov graphs. These matrices are the simplest examples of evolution operators prerequisite to developing a theory of averaging over chaotic flows.

Importance of symbolic dynamics is often grossly unappreciated; as we shall see in chapters 16 and 18 , coupled with uniform hyperbolicity, the existence of a finite grammar is the crucial prerequisite for construction of zeta functions with nice analyticity properties.

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## Exercises

## Exercise 12.1 A Smale horseshoe. The Hénon map

$$
\left[\begin{array}{l}
x^{\prime}  \tag{12.10}\\
y^{\prime}
\end{array}\right]=\left[\begin{array}{l}
1-a x^{2}+y \\
b x
\end{array}\right]
$$

maps the $(x, y)$ plane into itself - it was constructed by Hénon [3.1] in order to mimic the Poincaré section of once-folding map induced by a flow like the one sketched in figure 11.7. For definitivness fix the parameters to $a=6, b=-1$.
a) Draw a rectangle in the $(x, y)$ plane such that its $n$th iterate by the Hénon map intersects the rectangle $2^{n}$ times.
b) Construct the inverse of the (12.10).
c) Iterate the rectangle back in the time; how many intersections are there between the $n$ forward and $m$ backward iterates of the rectangle?
d) Use the above information about the intersections to guess the $(x, y)$ coordinates for the two fixed points, a 2-cycle point, and points on the two distinct 3-cycles from table 11.1. The exact cycle points are computed in exercise 17.11.

Exercise 12.2 Kneading Danish pastry. Write down the $(x, y) \rightarrow(x, y)$ mapping that implements the baker's map of figure 12.2, together with the inverse mapping. Sketch a few rectangles in symbol square and their forward and backward images. (Hint: the mapping is very much like the tent map (11.8)).

Exercise 12.3 Kneading Danish without flipping. The baker's map of figure 12.2 includes a flip - a map of this type is called an orientation reversing oncefolding map. Write down the $(x, y) \rightarrow(x, y)$ mapping that implements an orientation preserving baker's map (no flip; Jacobian determinant $=1$ ). Sketch and label the first few foldings of the symbol square.

Exercise 12.4 Fix this manuscript. Check whether the layers of the baker's map of figure 12.2 are indeed ordered as the branches of the alternating binary tree of figure ??. (They might not be - we have not rechecked them). Draw the correct binary trees that order both the future and past itineraries.

For once-folding maps there are four topologically distinct ways of laying out the stretched and folded image of the starting region,
(a) orientation preserving: stretch, fold upward, as in figure 12.3
(b) orientation preserving: stretch, fold downward, as in figure 13.2
(c) orientation reversing: stretch, fold upward, flip, as in figure 12.4
(d) orientation reversing: stretch, fold downward, flip, as in figure 12.2,


Figure 12.3: A complete Smale horseshoe iterated forwards and backwards, orientation preserving case: function $f$ maps the dashed border square $\mathcal{M}$ into the vertical horseshoe, while the inverse map $f^{-1}$ maps it into the horizontal horseshoe. a) One iteration, b) two iterations, c) three iterations. The non-wandering set is contained within the intersection of the forward and backward iterates (crosshatched). (from K.T. Hansen [1.3])


Figure 12.4: An orientation reversing Smale horseshoe map. Function $f=\{$ stretch,fold,flip $\}$ maps the dashed border square $\mathcal{M}$ into the vertical horseshoe, while the inverse map $f^{-1}$ maps it into the horizontal horseshoe. a) one iteration, b) two iterations, c) the non-wandering set cover by 16 rectangles, each labeled by the 2 past and the 2 future steps. (from K.T. Hansen [1.3])
with the corresponding four distinct binary-labeled symbol squares. For $n$-fold "stretch \& fold" flows the labeling would be nary. The intersection $\mathcal{M}_{0}$ for the orientation preserving Smale horseshoe, figure 12.3a, is oriented the same way as $\mathcal{M}$, while $\mathcal{M}_{1}$ is oriented opposite to $\mathcal{M}$. Brief contemplation of figure 12.2 indicates that the forward iteration strips are ordered relative to each other as the branches of the alternating binary tree in figure ??.

Check the labeling for all four cases.
Exercise 12.5 Orientation reversing once-folding map. By adding a reflection around the vertical axis to the horseshoe map $g$ we get the orientation reversing map $\tilde{g}$ shown in figure 12.4. $\tilde{Q}_{0}$ and $\tilde{Q}_{1}$ are oriented as $Q_{0}$ and $Q_{1}$, so the definition of the future topological coordinate $\gamma$ is identical to the $\gamma$ for the orientation preserving horseshoe. The inverse intersections $\tilde{Q}_{0}^{-1}$ and $\tilde{Q}_{1}^{-1}$ are oriented so that $\tilde{Q}_{0}^{-1}$ is opposite to $Q$, while $\tilde{Q}_{1}^{-1}$ has the same orientation as $Q$. Check that the past topological coordinate $\delta$ is given by

$$
\begin{align*}
w_{n-1} & =\left\{\begin{array}{ll}
1-w_{n} & \text { if } s_{n}=0 \\
w_{n} & \text { if } s_{n}=1
\end{array}, \quad w_{0}=s_{0}\right. \\
\delta(x) & =0 . w_{0} w_{-1} w_{-2} \ldots=\sum_{n=1}^{\infty} w_{1-n} / 2^{n} \tag{12.11}
\end{align*}
$$

Exercise 12.6 Infinite symbolic dynamics. Let $\sigma$ be a function that returns zero or one for every infinite binary string: $\sigma:\{0,1\}^{\mathbb{N}} \rightarrow\{0,1\}$. Its value is represented by $\sigma\left(\epsilon_{1}, \epsilon_{2}, \ldots\right)$ where the $\epsilon_{i}$ are either 0 or 1 . We will now define an operator $\mathcal{T}$ that acts on observables on the space of binary strings. A function $a$ is an observable if it has bounded variation, that is, if

$$
\|a\|=\sup _{\left\{\epsilon_{i}\right\}}\left|a\left(\epsilon_{1}, \epsilon_{2}, \ldots\right)\right|<\infty .
$$

For these functions

$$
\mathcal{T} a\left(\epsilon_{1}, \epsilon_{2}, \ldots\right)=a\left(0, \epsilon_{1}, \epsilon_{2}, \ldots\right) \sigma\left(0, \epsilon_{1}, \epsilon_{2}, \ldots\right)+a\left(1, \epsilon_{1}, \epsilon_{2}, \ldots\right) \sigma\left(1, \epsilon_{1}, \epsilon_{2}, \ldots\right) .
$$

(a) (easy) Consider a finite version $T_{n}$ of the operator $\mathcal{T}$ :

$$
\begin{aligned}
& T_{n} a\left(\epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{1, n}\right)= \\
& \quad a\left(0, \epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n-1}\right) \sigma\left(0, \epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n-1}\right)+ \\
& \quad a\left(1, \epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n-1}\right) \sigma\left(1, \epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n-1}\right) .
\end{aligned}
$$

Show that $T_{n}$ is a $2^{n} \times 2^{n}$ matrix. Show that its trace is bounded by a number independent of $n$.
(b) (medium) With the operator norm induced by the function norm, show that $\mathcal{T}$ is a bounded operator.
(c) (hard) Show that $\mathcal{T}$ is not trace class. (Hint: check if $\mathcal{T}$ is compact "trace class" is defined in appendix K.)

Exercise 12.7 Time reversability.** Hamiltonian flows are time reversible. Does that mean that their Markov graphs are symmetric in all node $\rightarrow$ node links, their transition matrices are adjacency matrices, symmetric and diagonalizable, and that they have only real eigenvalues?

Exercise 12.8 Alphabet $\{0,1\}$, prune only the fixed point $\overline{0}$. This is equivalent to the infinite alphabet $\{1,2,3,4, \ldots\}$ unrestricted symbolic dynamics. The prime cycles are labeled by all non-repeating sequences of integers, ordered lexically: $t_{n}, n>0 ; t_{m n}, t_{m m n}, \ldots, n>m>0 ; t_{m n r}, r>n>m>0, \ldots$ (see sect. 21.3). Now the number of fundamental cycles is infinite as well:

$$
\begin{align*}
1 / \zeta= & 1-\sum_{n>0} t_{n}-\sum_{n>m>0}\left(t_{m n}-t_{n} t_{m}\right) \\
& -\sum_{n>m>0}\left(t_{m m n}-t_{m} t_{m n}\right)-\sum_{n>m>0}\left(t_{m n n}-t_{m n} t_{n}\right)  \tag{12.12}\\
& -\sum_{r>n>m>0}\left(t_{m n r}+t_{m r n}-t_{m n} t_{r}-t_{m r} t_{n}-t_{m} t_{n r}+t_{m} t_{n} t_{r}\right\rangle 1
\end{align*}
$$

As shown in sect. 21.3, this grammar plays an important role in description of fixed points of marginal stability.

Exercise 12.9 3-disk pruning (Not easy) Show that for 3-disk game of pinball the pruning of orbits starts at $R: a=2.04821419 \ldots$..

(Kai T. Hansen)

Exercise 12.10 Alphabet $\{\mathbf{0 , 1}\}$, prune _1000_, _00100_, _01100_. This example is motivated by the pruning front description of the symbolic dynamics for the Hénon-type maps.
step 1. _1000_ prunes all cycles with a _000_ subsequence with the exception of the fixed point $\overline{0}$; hence we factor out $\left(1-t_{0}\right)$ explicitly, and prune _000_ from the rest. This means that $x_{0}$ is an isolated fixed point - no cycle stays in its vicinity for more than 2 iterations. In the notation of sect. 12.4.1, the alphabet is $\{1,2,3 ; \overline{0}\}$, and the remaining pruning rules have to be rewritten in terms of symbols $2=10,3=100$ :
step 2. alphabet $\{1,2,3 ; \overline{0}\}$, prune _33_, _213_,_313_. This means that the 3 -cycle $\overline{3}=\overline{100}$ is pruned and no long cycles stay close enough to it for a single _100_ repeat. As in example 1?!, prohibition of _33_ is implemented by dropping the symbol " 3 " and extending the alphabet by the allowed blocks 13, 23:
step 3. alphabet $\{1,2, \underline{13}, \underline{23} ; \overline{0}\}$, prune _213_, $-\underline{23} \underline{13}-, \underline{13} \underline{13}-$, where $\underline{13}=13$, $\underline{23}=23$ are now used as single letters. Pruning of the repetitions _13 $\underline{13}$ - (the 4-cycle $\overline{13}=\overline{1100}$ is pruned) yields the
result: alphabet $\{1,2, \underline{23}, \underline{113} ; \overline{0}\}$, unrestricted 4-ary dynamics. The other remaining possible blocks _213_, _-2313_ are forbidden by the rules of step 3. The cycle expansion is given by

$$
\begin{equation*}
1 / \zeta=\left(1-t_{0}\right)\left(1-t_{1}-t_{2}-t_{23}-t_{113}\right) \tag{12.14}
\end{equation*}
$$

for unrestricted 4-letter alphabet $\{1,2, \underline{23}, \underline{113}\}$.

## Chapter 13

## Counting, for pedestrians

That which is crooked cannot be made straight: and that which is wanting cannot be numbered. Ecclestiastes 1.15

We are now in position to develop our first prototype application of the periodic orbit theory: cycle counting. This is the simplest illustration of raison d'etre of the periodic orbit theory; we shall develop a duality transformation that relates local information - in this case the next admissible symbol in a symbol sequence - to global averages, in this case the mean rate of growth of the number of admissible itineraries with increasing itinerary length. We shall turn the topological dynamics of the preceding chapter into a multiplicative operation by means of transition matrices/Markov graphs, and show that the powers of a transition matrix count distinct itineraries. The asymptotic growth rate of the number of admissible itineraries is therefore given by the leading eigenvalue of the transition matrix; the leading eigenvalue is given by the leading zero of the characteristic determinant of the transition matrix, which is in this context called the topological zeta function. For flows with finite Markov graphs this determinant is a finite polynomial which can be read off the Markov graph.

The method goes well beyond the problem at hand, and forms the core of the entire treatise, making tangible a rather abstract notion of "spectral determinants" yet to come.

### 13.1 Counting itineraries

In the 3 -disk system the number of admissible trajectories doubles with every iterate: there are $K_{n}=3 \cdot 2^{n}$ distinct itineraries of length $n$. If disks are too close and some part of trajectories is pruned, this is only an upper bound and explicit formulas might be hard to come by, but we still might be able to establish a lower exponential bound of form $K_{n} \geq C e^{n \hat{h}}$. Hence, it is natural to characterize the growth of the number of trajectories as a
function of the itinerary length by the topological entropy:

$$
\begin{equation*}
h=\lim _{n \rightarrow \infty} \frac{1}{n} \ln K_{n} . \tag{13.1}
\end{equation*}
$$

We shall now relate this quantity to the leading eigenvalue of the transition matrix.

The transition matrix element $T_{i j} \in\{0,1\}$ in (11.2) indicates whether the transition from the starting partition $j$ into partition $i$ in one step is allowed or not, and the $(i, j)$ element of the transition matrix iterated $n$ times

$$
\left(T^{n}\right)_{i j}=\sum_{k_{1}, k_{2}, \ldots, k_{n-1}} T_{i k_{1}} T_{k_{1} k_{2}} \ldots T_{k_{n-1} j}
$$

receives a contribution 1 from every admissible sequence of transitions, so $\left(T^{n}\right)_{i j}$ is the number of admissible $n$ symbol itineraries starting with $j$ and ending with $i$.

## Example 13.1 3-disk itinerary counting

$$
\text { The }\left(T^{2}\right)_{13}=1 \text { element of } T^{2} \text { for the 3-disk transition matrix (11.5) }
$$

$$
\left(\begin{array}{lll}
c & 1 & 1  \tag{13.2}\\
1 & 0 & 1 \\
1 & 1 & 0
\end{array}\right)^{2}=\left(\begin{array}{lll}
2 & 1 & 1 \\
1 & 2 & 1 \\
1 & 1 & 2
\end{array}\right) .
$$

corresponds to $3 \rightarrow 2 \rightarrow 1$, the only 2-step path from 3 to 1 , while $\left(T^{2}\right)_{33}=2$ counts the two itineraries 313 and 323.

The total number of admissible itineraries of $n$ symbols is

$$
K_{n}=\sum_{i j}\left(T^{n}\right)_{i j}=(1,1, \ldots, 1) T^{n}\left(\begin{array}{c}
1  \tag{13.3}\\
1 \\
\vdots \\
1
\end{array}\right) .
$$

We can also count the number of prime cycles and pruned periodic points, but in order not to break up the flow of the main argument, we relegate these pretty results to sects. 13.5.2 and 13.5.3. Recommended reading if you ever have to compute lots of cycles.
$T$ is a matrix with non-negative integer entries. A matrix $M$ is said to be Perron-Frobenius if some power $k$ of $M$ has strictly positive entries, $\left(M^{k}\right)_{r s}>0$. In the case of the transition matrix $T$ this means that every partition eventually reaches all of the partitions, that is, the partition is dynamically transitive or indecomposable, as assumed in (2.2). The notion of transitivity is crucial in ergodic theory: a mapping is transitive if it has
a dense orbit, and the notion is obviously inherited by the shift once we introduce a symbolic dynamics. If that is not the case, phase space decomposes into disconnected pieces, each of which can be analyzed separately by a separate indecomposable Markov graph. Hence it suffices to restrict our considerations to the transition matrices of the Perron-Frobenius type.

A finite matrix $T$ has eigenvalues $T \varphi_{\alpha}=\lambda_{\alpha} \varphi_{\alpha}$ and (right) eigenvectors $\left\{\varphi_{0}, \varphi_{1}, \cdots, \varphi_{N-1}\right\}$. Expressing the initial vector in (13.3) in this basis

$$
T^{n}\left(\begin{array}{c}
1 \\
1 \\
\vdots \\
1
\end{array}\right)=T^{n} \sum_{\alpha=0}^{N-1} b_{\alpha} \varphi_{\alpha}=\sum_{\alpha=0}^{N-1} b_{\alpha} \lambda_{\alpha}^{n} \varphi_{\alpha}
$$

and contracting with $(1,1, \ldots, 1)$ we obtain

$$
K_{n}=\sum_{\alpha=0}^{N-1} c_{\alpha} \lambda_{\alpha}^{n}
$$

The constants $c_{\alpha}$ depend on the choice of initial and final partitions: In this example we are sandwiching $T^{n}$ between the vector $(1,1, \ldots, 1)$ and its transpose, but any other pair of vectors would do, as long as they are not orthogonal to the leading eigenvector $\varphi_{0}$. Perron theorem states that a Perron-Frobenius matrix has a nondegenerate positive real eigenvalue $\lambda_{0}>1$ (with a positive eigenvector) which exceeds the moduli of all other eigenvalues. Therefore as $n$ increases, the sum is dominated by the leading eigenvalue of the transition matrix, $\lambda_{0}>\left|\operatorname{Re} \lambda_{\alpha}\right|, \alpha=1,2, \cdots, N-1$, and the topological entropy (13.1) is given by

$$
\begin{align*}
h & =\lim _{n \rightarrow \infty} \frac{1}{n} \ln c_{0} \lambda_{0}^{n}\left[1+\frac{c_{1}}{c_{0}}\left(\frac{\lambda_{1}}{\lambda_{0}}\right)^{n}+\cdots\right] \\
& =\ln \lambda_{0}+\lim _{n \rightarrow \infty}\left[\frac{\ln c_{0}}{n}+\frac{1}{n} \frac{c_{1}}{c_{0}}\left(\frac{\lambda_{1}}{\lambda_{0}}\right)^{n}+\cdots\right] \\
& =\ln \lambda_{0} . \tag{13.4}
\end{align*}
$$

What have we learned? The transition matrix $T$ is a one-step local operator, advancing the trajectory from a partition to the next admissible partition. Its eigenvalues describe the rate of growth of the total number of trajectories at the asymptotic times. Instead of painstakingly counting $K_{1}, K_{2}, K_{3}, \ldots$ and estimating (13.1) from a slope of a log-linear plot, we have the exact topological entropy if we can compute the leading eigenvalue of the transition matrix $T$. This is reminiscent of the way the free energy is computed from transfer matrix for one-dimensional lattice models with finite range interactions: this analogy with statistical mechanics will be developed further in chapter 20.

### 13.2 Topological trace formula

There are two standard ways of getting at eigenvalues of a matrix - by evaluating the trace $\operatorname{tr} T^{n}=\sum \lambda_{\alpha}^{n}$, or by evaluating the determinant $\operatorname{det}(1-z T)$. We start by evaluating the trace of transition matrices.

Consider an $M$-step memory transition matrix, like the 1-step memory example (11.15). The trace of the transition matrix counts the number of partitions that map into themselves. In the binary case the trace picks up only two contributions on the diagonal, $T_{0 \cdots 0,0 \cdots 0}+T_{1 \cdots 1,1 \cdots 1}$, no matter how much memory we assume (check (11.15) and exercise 11.8). We can even take $M \rightarrow \infty$, in which case the contributing partitions are shrunk to the fixed points, $\operatorname{tr} T=T_{\overline{0}, \overline{0}}+T_{\overline{1}, \overline{1}}$.

More generally, each closed walk through $n$ concatenated entries of $T$ contributes to $\operatorname{tr} T^{n}$ a product of the matrix entries along the walk. Each step in such walk shifts the symbolic string by one symbol; the trace ensures that the walk closes into a periodic string $c$. Define $t_{c}$ to be the local trace, the product of matrix elements along a cycle $c$, each term being multiplied by a book keeping variable $z \cdot z^{n} \operatorname{tr} T^{n}$ is then the sum of $t_{c}$ for all cycles of length $n$. For example, for $[8 \times 8]$ transition matrix $T_{s_{1} s_{2} s_{3}, s_{0} s_{1} s_{2}}$ version of (11.15), or any refined partition $\left[2^{n} \times 2^{n}\right]$ transition matrix, $n$ arbitrarily large, the periodic point $\overline{100}$ contributes $t_{100}=z^{3} T_{\overline{100}, \overline{010}} T_{\overline{010}, \overline{001}} T_{\overline{001}, \overline{100}}$ to $z^{3} \operatorname{tr} T^{3}$. This product is manifestly cyclically symmetric, $t_{100}=t_{010}=t_{001}$, and so a prime cycle $p$ of length $n_{p}$ contributes $n_{p}$ times, once for each periodic point along its orbit. For the binary labeled non-wandering set the first few traces are given by (consult tables 11.1 and 13.1)

$$
\begin{align*}
z \operatorname{tr} T & =t_{0}+t_{1} \\
z^{2} \operatorname{tr} T^{2} & =t_{0}^{2}+t_{1}^{2}+2 t_{10} \\
z^{3} \operatorname{tr} T^{3} & =t_{0}^{3}+t_{1}^{3}+3 t_{100}+3 t_{101} \\
z^{4} \operatorname{tr} T^{4} & =t_{0}^{4}+t_{1}^{4}+2 t_{10}^{2}+4 t_{1000}+4 t_{1001}+4 t_{1011} . \tag{13.5}
\end{align*}
$$

For complete binary symbolic dynamics $t_{p}=z^{n_{p}}$ for every binary prime cycle $p$; if there is pruning $t_{p}=z^{n_{p}}$ if $p$ is admissible cycle and $t_{p}=0$ otherwise. Hence $\operatorname{tr} T^{n}$ counts the number of admissible periodic points of period $n$. In general, the $n$th order trace (13.5) picks up contributions from all repeats of prime cycles, with each cycle contributing $n_{p}$ periodic points, so the total number of periodic points of period $n$ is given by

$$
\begin{equation*}
N_{n}=\operatorname{tr} T^{n}=\sum_{n_{p} \mid n} n_{p} t_{p}^{n / n_{p}}=\sum_{p} n_{p} \sum_{r=1}^{\infty} \delta_{n, n_{p} r} r r_{p}^{r} . \tag{13.6}
\end{equation*}
$$

Here $m \mid n$ means that $m$ is a divisor of $n$, and we have taken $z=1$ so $t_{p}=1$ if the cycle is admissible, and $t_{p}=0$ otherwise. In order to get rid of the

| $n$ | $N_{n}$ | \# of prime cycles of length $n_{p}$ |  |  |  |  |  |  |  |  |  |
| ---: | ---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 10 |
| 1 | 2 | 2 |  |  |  |  |  |  |  |  |  |
| 2 | 4 | 2 | 1 |  |  |  |  |  |  |  |  |
| 3 | 8 | 2 |  | 2 |  |  |  |  |  |  |  |
| 4 | 16 | 2 | 1 |  | 3 |  |  |  |  |  |  |
| 5 | 32 | 2 |  |  |  | 6 |  |  |  |  |  |
| 6 | 64 | 2 | 1 | 2 |  |  | 9 |  |  |  |  |
| 7 | 128 | 2 |  |  |  |  |  | 18 |  |  |  |
| 8 | 256 | 2 | 1 |  | 3 |  |  |  | 30 |  |  |
| 9 | 512 | 2 |  | 2 |  |  |  |  |  | 56 |  |
| 10 | 1024 | 2 | 1 |  |  | 6 |  |  |  |  | 99 |

Table 13.1: The total numbers of periodic points $N_{n}$ of period $n$ for binary symbolic dynamics. The numbers of prime cycles contributing illustrates the preponderance of long prime cycles of length $n$ over the repeats of shorter cycles of lengths $n_{p}, n=r n_{p}$. Further listings of binary prime cycles are given in tables 11.1 and 13.2. (L. Rondoni)
awkward divisibility constraint $n=n_{p} r$ in the above sum, we introduce the generating function for numbers of periodic points

$$
\begin{equation*}
\sum_{n=1}^{\infty} z^{n} N_{n}=\operatorname{tr} \frac{z T}{1-z T} \tag{13.7}
\end{equation*}
$$

Substituting (13.6) into the left hand side, and replacing the right hand side by the eigenvalue sum $\operatorname{tr} T^{n}=\sum \lambda_{\alpha}^{n}$, we obtain our first example of a trace formula, the topological trace formula

$$
\begin{equation*}
\sum_{\alpha=0} \frac{z \lambda_{\alpha}}{1-z \lambda_{\alpha}}=\sum_{p} \frac{n_{p} t_{p}}{1-t_{p}} . \tag{13.8}
\end{equation*}
$$

A trace formula relates the spectrum of eigenvalues of an operator - in this case the transition matrix - to the spectrum of periodic orbits of the dynamical system. The $z^{n}$ sum in (13.7) is a discrete version of the Laplace transform, (see chapter 14), and the resolvent on the left hand side is the antecedent of the more sophisticated trace formulas (14.9), (14.19) and (30.3). We shall now use this result to compute the spectral determinant of the transition matrix.

### 13.3 Determinant of a graph

Our next task is to determine the zeros of the spectral determinant of an $[M \times M]$ transition matrix

$$
\begin{equation*}
\operatorname{det}(1-z T)=\prod_{\alpha=0}^{M-1}\left(1-z \lambda_{\alpha}\right) \tag{13.9}
\end{equation*}
$$

We could now proceed to diagonalize $T$ on a computer, and get this over with. Nevertheless, it pays to dissect $\operatorname{det}(1-z T)$ with some care; understanding this computation in detail will be the key to understanding the cycle expansion computations of chapter 18 for arbitrary dynamical averages. For $T$ a finite matrix (13.9) is just the characteristic equation for $T$. However, we shall be able to compute this object even when the dimension of $T$ and other such operators goes to $\infty$, and for that reason we prefer to refer to (13.9) as the "spectral determinant".

There are various definitions of the determinant of a matrix; they mostly reduce to the statement that the determinant is a certain sum over all possible permutation cycles composed of the traces $\operatorname{tr} T^{k}$, in the spirit of the determinant-trace relation of chapter 1 :

$$
\begin{align*}
\operatorname{det}(1-z T) & =\exp (\operatorname{tr} \ln (1-z T))=\exp \left(-\sum_{n=1} \frac{z^{n}}{n} \operatorname{tr} T^{n}\right) \\
& =1-z \operatorname{tr} T-\frac{z^{2}}{2}\left((\operatorname{tr} T)^{2}-\operatorname{tr}\left(T^{2}\right)\right)-\ldots \tag{13.10}
\end{align*}
$$

This is sometimes called a cumulant expansion. Formally, the right hand is an infinite sum over powers of $z^{n}$. If $T$ is an $[M \times M]$ finite matrix, then the characteristic polynomial is at most of order $M$. Coefficients of $z^{n}, n>M$ vanish exactly.

We now proceed to relate the determinant in (13.10) to the corresponding Markov graph of chapter 11: to this end we start by the usual algebra textbook expression

$$
\begin{equation*}
\operatorname{det}(1-z T)=\sum_{\{\pi\}}(-1)^{P_{\pi}}(1-z T)_{1, \pi_{1}} \cdot(1-z T)_{2, \pi_{2}} \cdots(1-z T)_{M, \pi_{M}}( \tag{13.11}
\end{equation*}
$$

where once again we suppose $T$ is an $[M \times M]$ finite matrix, $\{\pi\}$ denotes the set of permutations of $M$ symbols, $\pi_{k}$ is what $k$ is permuted into by the permutation $k$, and $P_{\pi}$ is the parity of the considered permutation. The right hand side of (13.11) yields a polynomial of order $M$ in $z$ : a contribution of order $n$ in $z$ picks up $M-n$ unit factors along the diagonal, the remaining matrix elements yielding

$$
\begin{equation*}
(-z)^{n}(-1)^{P \tilde{\pi}} T_{\eta_{1}, \tilde{\pi}_{\eta_{1}}} \cdots T_{\eta_{n}, \tilde{\pi}_{\eta_{n}}} \tag{13.12}
\end{equation*}
$$

where $\tilde{\pi}$ is the permutation of the subset of $n$ distinct symbols $\eta_{1} \ldots \eta_{n}$ indexing $T$ matrix elements. As in (13.5), we refer to any combination $t_{i}=T_{\eta_{1} \eta_{2}} T_{\eta_{2} \eta_{3}} \cdots T_{\eta_{k} \eta_{1}}, c=\eta_{1}, \eta_{2}, \cdots, \eta_{k}$ fixed, as a local trace associated with a closed loop c on the Markov graph. Each term of form (13.12) may be factored in terms of local traces $t_{c_{1}} t_{c_{2}} \cdots t_{c_{k}}$, that is loops on the Markov graph. These loops are non-intersecting, as each node may only be reached by one link, and they are indeed loops, as if a node is reached by a link, it
has to be the starting point of another single link, as each $\eta_{j}$ must appear exactly once as a row and column index. So the general structure is clear, a little more thinking is only required to get the sign of a generic contribution. We consider only the case of loops of length 1 and 2 , and leave to the reader the task of generalizing the result by induction. Consider first a term in which only loops of unit length appear on (13.12) that is, only the diagonal elements of $T$ are picked up. We have $k=n$ loops and an even permutation $\tilde{\pi}$ so the sign is given by $(-1)^{k}, k$ being the number of loops. Now take the case in which we have $i$ single loops and $j$ loops of length 2 (we must thus have $n=2 j+i)$. The parity of the permutation gives $(-1)^{j}$ and the first factor in (13.12) gives $(-1)^{n}=(-1)^{2 j+i}$. So once again these terms combine into $(-1)^{k}$, where $k=i+j$ is the number of loops. We may summarize our findings as follows:

The characteristic polynomial of a transition matrix/Markov graph is given by the sum of all possible partitions $\pi$ of the graph into products of non-intersecting loops, with each loop trace $t_{p}$ carrying a minus sign:

$$
\begin{equation*}
\operatorname{det}(1-z T)=\sum_{k=0}^{f} \sum_{\pi}^{\prime}(-1)^{k} t_{p_{1}} \cdots t_{p_{k}} \tag{13.13}
\end{equation*}
$$

Any self-intersecting loop is shadowed by a product of two loops that share the intersection point. As both the long loop $t_{a b}$ and its shadow $t_{a} t_{b}$ in the case at hand carry the same weight $z^{n_{a}+n_{b}}$, the cancellation is exact, and the loop expansion (13.13) is finite, with $f$ the maximal number of non-intersecting loops.

We refer to the set of all non-self-intersecting loops $\left\{t_{p_{1}}, t_{p_{2}}, \cdots t_{p_{f}}\right\}$ as the the fundamental cycles. This is not a very good definition, as the Markov graphs are not unique - the most we know is that for a given finitegrammar language, there exist Markov graph(s) with the minimal number of loops. Regardless of how cleverly a Markov graph is constructed, it is always true that for any finite Markov graph the number of fundamental cycles $f$ is finite. If you know a better way to define the "fundamental cycles", let us know.


### 13.3.1 Topological polynomials: learning by examples

The above definition of the determinant in terms of traces is most easily grasped by a working through a few examples. The complete binary dynamics Markov graph of figure 11.12(b) is a little bit too simple, but anyway, let us start humbly; there are only two non-intersecting loops, yielding

$$
\begin{equation*}
\operatorname{det}(1-z T)=1-t_{0}-t_{1}=1-2 z \tag{13.14}
\end{equation*}
$$

Figure 13.1: The golden mean pruning rule Markov graph, see also figure 11.14.


Figure 13.2: (a) An incomplete Smale horseshoe: the inner forward fold does not intersect the two rightmost backward folds. (b) The primary pruned region in the symbol square and the corresponding forbidden binary blocks. (c) An incomplete Smale horseshoe which illustrates (d) the monotonicity of the pruning front: the thick line which delineates the left border of the primary pruned region is monotone on each half of the symbol square. The backward folding in figures (a) and (c) is only schematic - in invertible mappings there are further missing intersections, all obtained by the forward and backward iterations of the primary pruned region.

The leading (and only) zero of this characteristic polynomial yields the topological entropy $e^{h}=2$. As we know that there are $K_{n}=2^{n}$ binary strings of length $N$, we are not surprised. Similarly, for complete symbolic dynamics of $N$ symbols the Markov graph has one node and $N$ links, yielding

$$
\begin{equation*}
\operatorname{det}(1-z T)=1-N z, \tag{13.15}
\end{equation*}
$$

whence the topological entropy $h=\ln N$.
A more interesting example is the "golden mean" pruning of figure 13.1. There is only one grammar rule, that a repeat of symbol 0 is forbidden.

The non-intersecting loops are of length 1 and 2 , so the topological polynomial is given by

$$
\begin{equation*}
\operatorname{det}(1-z T)=1-t_{1}-t_{01}=1-z-z^{2} \tag{13.16}
\end{equation*}
$$

The leading root of this polynomial is the golden mean, so the entropy (13.4) is the logarithm of the golden mean, $h=\ln \frac{1+\sqrt{5}}{2}$.

Finally, the non-self-intersecting loops of the Markov graph of figure 13.3(d) are indicated in figure 13.3(e). The determinant can be written down by inspection, as the sum of all possible partitions of the graph into products of non-intersecting loops, with each loop carrying a minus sign:

$$
\operatorname{det}(1-T)=1-t_{0}-t_{0011}-t_{0001}-t_{00011}+t_{0} t_{0011}+t_{0011} t_{0001}(13.17)
$$

With $t_{p}=z^{n_{p}}$, where $n_{p}$ is the length of the $p$-cycle, the smallest root of


Figure 13.3: Conversion of the pruning front of figure 13.2d into a finite Markov graph. (a) Starting with the start node ".", delineate all pruning blocks on the binary tree. A solid line stands for " 1 " and a dashed line for " 0 ". Ends of forbidden strings are marked with $\times$. Label all internal nodes by reading the bits connecting ".", the base of the tree, to the node. (b) Indicate all admissible starting blocks by arrows. (c) Drop recursively the leading bits in the admissible blocks; if the truncated string corresponds to an internal node in (a), connect them. (d) Delete the transient, noncirculating nodes; all admissible sequences are generated as walks on this finite Markov graph. (e) Identify all distinct loops and construct the determinant (13.17).

$$
\begin{equation*}
0=1-z-2 z^{4}+z^{8} \tag{13.18}
\end{equation*}
$$

yields the topological entropy $h=-\ln z, z=0.658779 \ldots, h=0.417367 \ldots$, significantly smaller than the entropy of the covering symbolic dynamics, the complete binary shift $h=\ln 2=0.693 \ldots$

### 13.4 Topological zeta function

What happens if there is no finite-memory transition matrix, if the Markov graph is infinite? If we are never sure that looking further into future will reveal no further forbidden blocks? There is still a way to define the determinant, and the idea is central to the whole treatise: the determinant page 78 is then defined by its cumulant expansion (13.10)

$$
\begin{equation*}
\operatorname{det}(1-z T)=1-\sum_{n=1}^{\infty} \hat{c}_{n} z^{n} \tag{13.19}
\end{equation*}
$$

For finite dimensional matrices the expansion is a finite polynomial, and (13.19) is an identity; however, for infinite dimensional operators the cumulant expansion coefficients $\hat{c}_{n}$ define the determinant.

Let us now evaluate the determinant in terms of traces for an arbitrary transition matrix. In order to obtain an expression for the spectral determinant (13.9) in terms of cycles, substitute (13.6) into (13.19) and sum over the repeats of prime cycles

$$
\begin{equation*}
\operatorname{det}(1-z T)=\exp \left(-\sum_{p} \sum_{r=1}^{\infty} \frac{t_{p}^{r}}{r}\right)=\prod_{p}\left(1-t_{p}\right) . \tag{13.20}
\end{equation*}
$$

where for the topological entropy the weight assigned to a prime cycle $p$ of length $n_{p}$ is $t_{p}=z^{n_{p}}$ if the cycle is admissible, or $t_{p}=0$ if it is pruned. This determinant is called the topological or the Artin-Mazur zeta function, conventionally denoted by

$$
\begin{equation*}
1 / \zeta_{\text {top }}=\prod_{p}\left(1-z^{n_{p}}\right)=1-\sum_{n=1} \hat{c}_{n} z^{n} . \tag{13.21}
\end{equation*}
$$

Counting cycles amounts to giving each admissible prime cycle $p$ weight $t_{p}=z^{n_{p}}$ and expanding the Euler product (13.21) as a power series in $z$. As the precise expression for coefficients $\hat{c}_{n}$ in terms of local traces $t_{p}$ is more general than the current application to counting, we shall postpone its derivation to chapter 18 .

The topological entropy $h$ can now be determined from the leading zero $z=e^{-h}$ of the topological zeta function. For a finite $[N \times N]$ transition matrix, the number of terms in the characteristic equation (13.13) is finite, and we refer to this expansion as the topological polynomial of order $\leq N$. The power of defining a determinant by the cumulant expansion is that it works even when the partition is infinite, $N \rightarrow \infty$; an example is given in sect. 13.6, and many more later on.


### 13.4.1 Topological zeta function for flows

We now apply the method we used in deriving (14.19) to the problem of deriving the topological zeta functions for flows. By analogy to (14.17), the time-weighted density of prime cycles of period $t$ is

$$
\begin{equation*}
\Gamma(t)=\sum_{p} \sum_{r=1} T_{p} \delta\left(t-r T_{p}\right) . \tag{13.22}
\end{equation*}
$$

A Laplace transform smooths the sum over Dirac delta spikes and yields the topological trace formula

$$
\begin{equation*}
\sum_{p} \sum_{r=1} T_{p} \int_{0_{+}}^{\infty} d t e^{-s t} \delta\left(t-r T_{p}\right)=\sum_{p} T_{p} \sum_{r=1}^{\infty} e^{-s T_{p} r} \tag{13.23}
\end{equation*}
$$

and the topological zeta function for flows:

$$
\begin{align*}
1 / \zeta_{\text {top }}(s) & =\prod_{p}\left(1-e^{-s T_{p}}\right) \\
\sum_{p} T_{p} \sum_{r=1}^{\infty} e^{-s T_{p} r} & =-\frac{\partial}{\partial s} \ln 1 / \zeta_{\text {top }}(s) . \tag{13.24}
\end{align*}
$$

This is the continuous time version of the discrete time topological zeta function (13.21) for maps; its leading zero $s=-h$ yields the topological entropy for a flow.

### 13.5 Counting cycles

In what follows we shall occasionally need to compute all cycles up to topological length $n$, so it is handy to know their exact number.

### 13.5.1 Counting periodic points

$N_{n}$, the number of periodic points of period $n$ can be computed from (13.19) and (13.7) as a logarithmic derivative of the topological zeta function

$$
\begin{align*}
\sum_{n=1} N_{n} z^{n} & =\operatorname{tr}\left(-z \frac{d}{d z} \ln (1-z T)\right)=-z \frac{d}{d z} \ln \operatorname{det}(1-z T) \\
& =\frac{-z \frac{d}{d z} 1 / \zeta_{\text {top }}}{1 / \zeta_{\text {top }}} . \tag{13.25}
\end{align*}
$$

We see that the trace formula (13.8) diverges at $z \rightarrow e^{-h}$, as the denominator has a simple zero there.

As a check of formula (13.19) in the finite grammar context, consider the complete $N$-ary dynamics (11.3) for which the number of periodic points of period $n$ is simply $\operatorname{tr} T_{c}^{n}=N^{n}$. Substituting

$$
\sum_{n=1}^{\infty} \frac{z^{n}}{n} \operatorname{tr} T_{c}^{n}=\sum_{n=1}^{\infty} \frac{(z N)^{n}}{n}=\ln (1-z N)
$$

into (13.19) we verify (13.15). The logarithmic derivative formula (13.25) in this case does not buy us much either, we recover

$$
\sum_{n=1} N_{n} z^{n}=\frac{N z}{1-N z}
$$

However, consider instead the nontrivial pruning of figure 13.3(e). Substituting (13.18) we obtain

$$
\begin{equation*}
\sum_{n=1} N_{n} z^{n}=\frac{z+8 z^{4}-8 z^{8}}{1-z-2 z^{4}+z^{8}} . \tag{13.26}
\end{equation*}
$$

Now the topological zeta function is not merely a tool for extracting the asymptotic growth of $N_{n}$; it actually yields the exact and not entirely trivial recursion relation for the numbers of periodic points: $N_{1}=N_{2}=N_{3}=1$, $N_{n}=2 n+1$ for $n=4,5,6,7,8$, and $N_{n}=N_{n-1}+2 N_{n-4}-N_{n-8}$ for $n>8$.

### 13.5.2 Counting prime cycles

Having calculated the number of periodic points, our next objective is to evaluate the number of prime cycles $M_{n}$ for a dynamical system whose symbolic dynamics is built from $N$ symbols. The problem of finding $M_{n}$ is classical in combinatorics (counting necklaces made out of $n$ beads out of $N$ different kinds) and is easily solved. There are $N^{n}$ possible distinct strings

| n | $M_{n}(N)$ | $M_{n}(2)$ | $M_{n}(3)$ | $M_{n}(4)$ |
| ---: | :---: | ---: | ---: | ---: |
| 1 | N | 2 | 3 | 4 |
| 2 | $N(N-1) / 2$ | 1 | 3 | 6 |
| 3 | $N\left(N^{2}-1\right) / 3$ | 2 | 8 | 20 |
| 4 | $N^{2}\left(N^{2}-1\right) / 4$ | 3 | 18 | 60 |
| 5 | $\left(N^{5}-N\right) / 5$ | 6 | 48 | 204 |
| 6 | $\left(N^{6}-N^{3}-N^{2}+N\right) / 6$ | 9 | 116 | 670 |
| 7 | $\left(N^{7}-N\right) / 7$ | 18 | 312 | 2340 |
| 8 | $N^{4}\left(N^{4}-1\right) / 8$ | 30 | 810 | 8160 |
| 9 | $N^{3}\left(N^{6}-1\right) / 9$ | 56 | 2184 | 29120 |
| 10 | $\left(N^{10}-N^{5}-N^{2}+N\right) / 10$ | 99 | 5880 | 104754 |

Table 13.2: Number of prime cycles for various alphabets and grammars up to length 10. The first column gives the cycle length, the second the formula (13.28) for the number of prime cycles for complete $N$-symbol dynamics, columns three through five give the numbers for $N=2,3$ and 4 .
of length $n$ composed of $N$ letters. These $N^{n}$ strings include all $M_{d}$ prime $d-$ cycles whose period $d$ equals or divides $n$. A prime cycle is a non-repeating symbol string: for example, $p=\overline{011}=\overline{101}=\overline{110}=\ldots 011011 \ldots$ is prime, but $\overline{0101}=010101 \ldots=\overline{01}$ is not. A prime $d$-cycle contributes $d$ strings to the sum of all possible strings, one for each cyclic permutation. The total number of possible periodic symbol sequences of length $n$ is therefore related to the number of prime cycles by

$$
\begin{equation*}
N_{n}=\sum_{d \mid n} d M_{d} \tag{13.27}
\end{equation*}
$$

where $N_{n}$ equals $\operatorname{tr} T^{n}$. The number of prime cycles can be computed recursively

$$
M_{n}=\frac{1}{n}\left(N_{n}-\sum_{d \mid n}^{d<n} d M_{d}\right)
$$

or by the Möbius inversion formula

$$
\begin{equation*}
M_{n}=n^{-1} \sum_{d \mid n} \mu\left(\frac{n}{d}\right) N_{d} . \tag{13.28}
\end{equation*}
$$

where the Möbius function $\mu(1)=1, \mu(n)=0$ if $n$ has a squared factor, and $\mu\left(p_{1} p_{2} \ldots p_{k}\right)=(-1)^{k}$ if all prime factors are different.

We list the number of prime cycles up to length 10 for 2 -, 3- and 4letter complete symbolic dynamics in table 13.2. The number of prime cycles follows by Möbius inversion (13.28).

### 13.5.3 Counting $N$-disk periodic points

A simple example of pruning is the exclusion of "self-bounces" in the $N$-disk game of pinball. The number of points that are mapped back onto themselves after $n$ iterations is given by $N_{n}=\operatorname{tr} T^{n}$. The pruning of self-bounces eliminates the diagonal entries, $T_{N-d i s k}=T_{c}-\mathbf{1}$, so the number of the $N$-disk periodic points is

$$
\begin{equation*}
N_{n}=\operatorname{tr} T_{N-\text { disk }}^{n}=(N-1)^{n}+(-1)^{n}(N-1) \tag{13.29}
\end{equation*}
$$

(here $T_{c}$ is the complete symbolic dynamics transition matrix (11.3)). For the $N$-disk pruned case (13.29) Möbius inversion (13.28) yields

$$
\begin{align*}
M_{n}^{N-d i s k} & =\frac{1}{n} \sum_{d \mid n} \mu\left(\frac{n}{d}\right)(N-1)^{d}+\frac{N-1}{n} \sum_{d \mid n} \mu\left(\frac{n}{d}\right)(-1)^{d} \\
& =M_{n}^{(N-1)} \quad \text { for } \quad n>2 . \tag{13.30}
\end{align*}
$$

There are no fixed points, $M_{1}^{N-d i s k}=0$. The number of periodic points of period 2 is $N^{2}-N$, hence there are $M_{2}^{N-d i s k}=N(N-1) / 2$ prime cycles of length 2 ; for lengths $n>2$, the number of prime cycles is the same as for the complete ( $N-1$ )-ary dynamics of table 13.2.

### 13.5.4 Pruning individual cycles

Consider the 3 -disk game of pinball. The prohibition of repeating a symbol affects counting only for the fixed points and the 2-cycles. Everything else is the same as counting for a complete binary dynamics (eq (13.30)). To obtain the topological zeta function, just divide out the binary 1 - and 2 -cycles $\left(1-z t_{0}\right)\left(1-z t_{1}\right)\left(1-z^{2} t_{01}\right)$ and multiply with the correct 3 -disk 2 -cycles $\left(1-z^{2} t_{12}\right)\left(1-z^{2} t_{13}\right)\left(1-z^{2} t_{23}\right)$ :

$$
\begin{align*}
1 / \zeta_{3-\text { disk }} & =(1-2 z) \frac{\left(1-z^{2}\right)^{3}}{(1-z)^{2}\left(1-z^{2}\right)} \\
& =(1-2 z)(1+z)^{2}=1-3 z^{2}-2 z^{3} . \tag{13.31}
\end{align*}
$$

The factorization reflects the underlying 3 -disk symmetry; we shall rederive it in (22.25). As we shall see in chapter 22, symmetries lead to factorizations of topological polynomials and topological zeta functions.

The example of exercise 13.17 with the alphabet $\left\{a, c b^{k} ; \bar{b}\right\}$ is more interesting. In the cycle counting case, the dynamics in terms of $a \rightarrow z$, $c b^{k} \rightarrow \frac{z}{1-z}$ is a complete binary dynamics with the explicit fixed point factor $\left(1-t_{b}\right)=(1-z)$ :

$$
1 / \zeta_{\text {top }}=(1-z)\left(1-z-\frac{z}{1-z}\right)=1-3 z+z^{2}
$$

| n | $M_{n}$ | $N_{n}$ | $S_{n}$ | $m_{p} \cdot \hat{p}$ |
| :---: | :---: | :--- | ---: | :--- |
| 1 | 0 | 0 | 0 |  |
| 2 | 3 | $6=3 \cdot 2$ | 1 | $3 \cdot 12$ |
| 3 | 2 | $6=2 \cdot 3$ | 1 | $2 \cdot 123$ |
| 4 | 3 | $18=3 \cdot 2+3 \cdot 4$ | 1 | $3 \cdot 1213$ |
| 5 | 6 | $30=6 \cdot 5$ | 1 | $6 \cdot 12123$ |
| 6 | 9 | $66=3 \cdot 2+2 \cdot 3+9 \cdot 6$ | 2 | $6 \cdot 121213+3 \cdot 121323$ |
| 7 | 18 | $126=18 \cdot 7$ | 3 | $6 \cdot 1212123+6 \cdot 1212313+6 \cdot 1213123$ |
| 8 | 30 | $258=3 \cdot 2+3 \cdot 4+30 \cdot 8$ | 6 | $6 \cdot 12121213+3 \cdot 12121313+6 \cdot 12121323$ |
|  |  |  |  | $+6 \cdot 12123123+6 \cdot 12123213+3 \cdot 12132123$ |
| 9 | 56 | $510=2 \cdot 3+56 \cdot 9$ | 10 | $6 \cdot 121212123+6 \cdot(121212313+121212323)$ |
|  |  |  |  | $+6 \cdot(121213123+121213213)+6 \cdot 121231323$ |
|  |  |  |  | $+6 \cdot(121231213+121232123)+2 \cdot 121232313$ |
|  |  |  |  |  |
| 10 | 99 | 1022 |  |  |

Table 13.3: List of the 3-disk prime cycles up to length 10 . Here $n$ is the cycle length, $M_{n}$ the number of prime cycles, $N_{n}$ the number of periodic points and $S_{n}$ the number of distinct prime cycles under the $C_{3 v}$ symmetry (see chapter 22 for further details). Column 3 also indicates the splitting of $N_{n}$ into contributions from orbits of lengths that divide $n$. The prefactors in the fifth column indicate the degeneracy $m_{p}$ of the cycle; for example, 3.12 stands for the three prime cycles $\overline{12}, \overline{13}$ and $\overline{23}$ related by $2 \pi / 3$ rotations. Among symmetry related cycles, a representative $\hat{p}$ which is lexically lowest was chosen. The cycles of length 9 grouped by parenthesis are related by time reversal symmetry, but not by any other $C_{3 v}$ transformation.

| n | $M_{n}$ | $N_{n}$ | $S_{n}$ | $m_{p} \cdot \hat{p}$ |
| :---: | ---: | :--- | ---: | :--- |
| 1 | 0 | 0 | 0 |  |
| 2 | 6 | $12=6 \cdot 2$ | 2 | $4 \cdot 12+2 \cdot 13$ |
| 3 | 8 | $24=8 \cdot 3$ | 1 | $8 \cdot 123$ |
| 4 | 18 | $84=6 \cdot 2+18 \cdot 4$ | 4 | $8 \cdot 1213+4 \cdot 1214+2 \cdot 1234+4 \cdot 1243$ |
| 5 | 48 | $240=48 \cdot 5$ | 6 | $8 \cdot(12123+12124)+8 \cdot 12313$ |
|  |  |  |  | $+8 \cdot(12134+12143)+8 \cdot 12413$ |
| 6 | 116 | $732=6 \cdot 2+8 \cdot 3+116 \cdot 6$ | 17 | $8 \cdot 121213+8 \cdot 121214+8 \cdot 121234$ |
|  |  |  |  | $+8 \cdot 121243+8 \cdot 121313+8 \cdot 121314$ |
|  |  |  |  | $+4 \cdot 121323+8 \cdot(121324+121423)$ |
|  |  |  |  | $+8 \cdot 121343+8 \cdot 121424+4 \cdot 121434$ |
|  |  |  | 39 | $+4 \cdot 123124+8 \cdot 123134+4 \cdot 123143$ |
|  |  |  | 108 |  |
| 7 | 312 | 2184 |  |  |
| 8 | 810 | 6564 |  |  |

Table 13.4: List of the 4-disk prime cycles up to length 8. The meaning of the symbols is the same as in table 13.3. Orbits related by time reversal symmetry (but no other symmetry) already appear at cycle length 5. List of the cycles of length 7 and 8 has been omitted.



Figure 13.4: (a) The logarithm of the difference between the leading zero of the finite polynomial approximations to topological zeta function and our best estimate, as a function of the length for the quadratic map $A=3.8$. (b) The 90 zeroes of the characteristic polynomial for the quadratic map $A=3.8$ approximated by symbolic strings up to length 90. (from ref. [1.3])

### 13.6 Topological zeta function for an infinite partition

(K.T. Hansen and P. Cvitanović)

Now consider an example of a dynamical system which (as far as we know - there is no proof) has an infinite partition, or an infinity of longer and longer pruning rules. Take the 1-d quadratic map

$$
f(x)=A x(1-x)
$$

with $A=3.8$. It is easy to check numerically that the itinerary or the "kneading sequence" of the critical point $x=1 / 2$ is

$$
K=1011011110110111101011110111110 \ldots
$$

where the symbolic dynamics is defined by the partition of figure 11.9. How this kneading sequence is converted into a series of pruning rules is a dark art, relegated to appendix E. 1 For the moment it suffices to state the result, to give you a feeling for what a "typical" infinite partition topological zeta function looks like. Approximating the dynamics by a Markov graph corresponding to a repeller of the period 29 attractive cycle close to the $A=3.8$ strange attractor (or, much easier, following the algorithm of appendix E.1) yields a Markov graph with 29 nodes and the characteristic polynomial

$$
\begin{align*}
1 / \zeta_{\text {top }}^{(29)}= & 1-z^{1}-z^{2}+z^{3}-z^{4}-z^{5}+z^{6}-z^{7}+z^{8}-z^{9}-z^{10} \\
& +z^{11}-z^{12}-z^{13}+z^{14}-z^{15}+z^{16}-z^{17}-z^{18}+z^{19}+z^{20} \\
& -z^{21}+z^{22}-z^{23}+z^{24}+z^{25}-z^{26}+z^{27}-z^{28} . \quad(13.32) \tag{13.32}
\end{align*}
$$

The smallest real root of this approximate topological zeta function is

$$
\begin{equation*}
z=0.62616120 \ldots \tag{13.33}
\end{equation*}
$$

Constructing finite Markov graphs of increasing length corresponding to $A \rightarrow 3.8$ we find polynomials with better and better estimates for the topological entropy. For the closest stable period 90 orbit we obtain our best estimate of the topological entropy of the repeller:

$$
\begin{equation*}
h=-\ln 0.62616130424685 \ldots=0.46814726655867 \ldots \tag{13.34}
\end{equation*}
$$

Figure 13.4 illustrates the convergence of the truncation approximations to the topological zeta function as a plot of the logarithm of the difference between the zero of a polynomial and our best estimate (13.34), plotted as a function of the length of the stable periodic orbit. The error of the estimate (13.33) is expected to be of order $z^{29} \approx e^{-14}$ because going from length 28 to a longer truncation yields typically combinations of loops with 29 and more nodes giving terms $\pm z^{29}$ and of higher order in the polynomial. Hence the convergence is exponential, with exponent of $-0.47=-h$, the topological entropy itself.

In figure 13.4(b) we plot the zeroes of the polynomial approximation to the topological zeta function obtained by accounting for all forbidden strings of length 90 or less. The leading zero giving the topological entropy is the point closest to the origin. Most of the other zeroes are close to the unit circle; we conclude that for infinite Markov partitions the topological zeta function has a unit circle as the radius of convergence. The convergence is controlled by the ratio of the leading to the next-to-leading eigenvalues, which is in this case indeed $\lambda_{1} / \lambda_{0}=1 / e^{h}=e^{-h}$.

### 13.7 Shadowing

The topological zeta function is a pretty function, but the infinite product (13.20) should make you pause. For finite transfer matrices the left hand side is a determinant of a finite matrix, therefore a finite polynomial; but the right hand side is an infinite product over the infinitely many prime periodic orbits of all periods?

The way in which this infinite product rearranges itself into a finite polynomial is instructive, and crucial for all that follows. You can already take a peek at the full cycle expansion (18.5) of chapter 18; all cycles beyond the fundamental $t_{0}$ and $t_{1}$ appear in the shadowing combinations such as

$$
t_{s_{1} s_{2} \cdots s_{n}}-t_{s_{1} s_{2} \cdots s_{m}} t_{s_{m+1} \cdots s_{n}}
$$

For subshifts of finite type such shadowing combinations cancel exactly, if we are counting cycles as we do here, or if the dynamics is piecewise
linear, as in exercise 15.2. As we have already argued in sect. 1.5.4 and appendix J.1.2, for nice hyperbolic flows whose symbolic dynamics is a subshift of finite type, the shadowing combinations almost cancel, and the spectral determinant is dominated by the fundamental cycles from (13.13), with longer cycles contributing only small "curvature" corrections.

These exact or nearly exact cancellations depend on the flow being smooth and the symbolic dynamics being a subshift of finite type. If the dynamics requires infinite Markov partition with pruning rules for longer and longer blocks, most of the shadowing combinations still cancel, but the few corresponding to the forbidden blocks do not, leading to a finite radius of convergence for the spectral determinant as in figure 13.4(b).

One striking aspect of the pruned cycle expansion (13.32) compared to the trace formulas such as (13.7) is that coefficients are not growing exponentially - indeed they all remain of order 1 , so instead having a radius of convergence $e^{-h}$, in the example at hand the topological zeta function has the unit circle as the radius of convergence. In other words, exponentiating the spectral problem from a trace formula to a spectral determinant as in (13.19) increases the analyticity domain: the pole in the trace (13.8) at $z=e^{-h}$ is promoted to a smooth zero of the spectral determinant with a larger radius of convergence.

A detailed discussion of the radius of convergence is given in appendix E.1.
The very sensitive dependence of spectral determinants on whether the symbolic dynamics is or is not a subshift of finite type is the bad news that we should announce already now. If the system is generic and not structurally stable (see sect. 12.2), a smooth parameter variation is in no sense a smooth variation of topological dynamics - infinities of periodic orbits are created or destroyed, Markov graphs go from being finite to infinite and back. That will imply that the global averages that we intend to compute are generically nowhere differentiable functions of the system parameters, and averaging over families of dynamical systems can be a highly nontrivial enterprise; a simple illustration is the parameter dependence of the diffusion constant computed in a remark in chapter 23.

You might well ask: What is wrong with computing an entropy from (13.1)? Does all this theory buy us anything? If we count $K_{n}$ level by level, we ignore the self-similarity of the pruned tree - examine for example figure 11.14, or the cycle expansion of (13.26) - and the finite estimates of $h_{n}=\ln K_{n} / n$ converge nonuniformly to $h$, and on top of that with a slow rate of convergence, $\left|h-h_{n}\right| \approx O(1 / n)$ as in (13.4). The determinant (13.9) is much smarter, as by construction it encodes the self-similarity of the dynamics, and yields the asymptotic value of $h$ with no need for any finite $n$ extrapolations.

So, the main lesson of learning how to count well, a lesson that will be affirmed over and over, is that while the trace formulas are a conceptually essential step in deriving and understanding periodic orbit theory, the spectral determinant is the right object to use in actual computations. Instead
of resumming all of the exponentially many periodic points required by trace formulas at each level of truncation, spectral determinants incorporate only the small incremental corrections to what is already known - and that makes them more convergent and economical to use.

## Commentary

Remark 13.1 "Entropy". The ease with which the topological entropy can be motivated obscures the fact that our definition does not lead to an invariant characterization of the dynamics, as the choice of symbolic dynamics is largely arbitrary: the same caveat applies to other entropies to be discussed in chapter 20, and to get proper invariants one needs to evaluate a supremum over all possible partitions. The key mathematical point that eliminates the need of such search is the existence of generators, that is, partitions that under dynamics are able to probe the whole phase space on arbitrarily small scales: more precisely a generator is a finite partition $\Omega=\omega_{1} \ldots \omega_{N}$, with the following property: take $\mathcal{M}$ the subalgebra of the phase space generated by $\Omega$, and consider the partition built upon all possible intersections of sets $\phi^{k}\left(\beta_{i}\right)$, where $\phi$ is dynamical evolution, $\beta_{i}$ is an element of $\mathcal{M}$ and $k$ takes all possible integer values (positive as well as negative), then the closure of such a partition coincides with the algebra of all measurable sets. For a thorough (and readable) discussion of generators and how they allow a computation of the Kolmogorov entropy, see ref. [13.1] and chapter 20.

Remark 13.2 Perron-Frobenius matrices. For a proof of Perron theorem on the leading eigenvalue see ref. [1.14]. Ref. [13.2], sect. A4.1 contains a clear discussion of the spectrum of the transition matrix.

Remark 13.3 Determinant of a graph. Many textbooks offer derivations of the loop expansions of characteristic polynomials for transition matrices and their Markov graphs, see for example refs. [13.3, 13.4, 13.5].

Remark 13.4 $T$ is not trace class. Note to the erudite reader: the transition matrix $T$ (in the infinite partition limit (13.19)) is not trace class in the sense of appendix K. Still the trace is well defined in the $n \rightarrow \infty$ limit.

Remark 13.5 Artin-Mazur zeta functions. Motivated by A. Weil's zeta function for the Frobenius map [13.6], Artin and Mazur [15.13] introduced the zeta function (13.21) that counts periodic points for diffeomorphisms (see also ref. [13.7] for their evaluation for maps of the interval). Smale [13.8] conjectured rationality of the zeta functions for Axiom A diffeomorphisms, later proved by Guckenheimer [13.9] and Manning [13.10]. See remark 15.4 on page 259 for more zeta function history.

Remark 13.6 Ordering periodic orbit expansions. In sect. 18.4 we will introduce an alternative way of hierarchically organizing cumulant expansions, in which the order is dictated by stability rather than cycle length: such a procedure may be better suited to perform computations when the symbolic dynamics is not well understood.

## Résumé

What have we accomplished? We have related the number of topologically distinct paths from "this region" to "that region" in a chaotic system to the leading eigenvalue of the transition matrix $T$. The eigenspectrum of $T$ is given by a certain sum over traces $\operatorname{tr} T^{n}$, and in this way the periodic orbit theory has entered the arena, already at the level of the topological dynamics, the crudest description of dynamics.

The main result of this chapter is the cycle expansion (13.21) of the topological zeta function (that is, the spectral determinant of the transition matrix):

$$
1 / \zeta_{\text {top }}(z)=1-\sum_{k=1} \hat{c}_{k} z^{k} .
$$

For subshifts of finite type, the transition matrix is finite, and the topological zeta function is a finite polynomial evaluated by the loop expansion (13.13) of $\operatorname{det}(1-z T)$. For infinite grammars the topological zeta function is defined by its cycle expansion. The topological entropy $h$ is given by the smallest zero $z=e^{-h}$. This expression for the entropy is exact; in contrast to the definition (13.1), no $n \rightarrow \infty$ extrapolations of $\ln K_{n} / n$ are required.

Historically, these topological zeta functions were the inspiration for applying the transfer matrix methods of statistical mechanics to the problem of computation of dynamical averages for chaotic flows. The key result were the dynamical zeta functions that derived in chapter 14, the weighted generalizations of the topological zeta function.

Contrary to claims one sometimes encounters in the literature, "exponential proliferation of trajectories" is not the problem; what limits the convergence of cycle expansions is the proliferation of the grammar rules, or the "algorithmic complexity", as illustrated by sect. 13.6, and figure 13.4 in particular.

## References

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## Exercises

## Exercise 13.1 A transition matrix for 3-disk pinball.

a) Draw the Markov graph corresponding to the 3-disk ternary symbolic dynamics, and write down the corresponding transition matrix corresponding to the graph. Show that iteration of the transition matrix results in two coupled linear difference equations, - one for the diagonal and one for the off diagonal elements. (Hint: relate $\operatorname{tr} T^{n}$ to $\operatorname{tr} T^{n-1}+\ldots$.)
b) Solve the above difference equation and obtain the number of periodic orbits of length $n$. Compare with table 13.3.
c) Find the eigenvalues of the transition matrix $\mathbf{T}$ for the 3-disk system with ternary symbolic dynamics and calculate the topological entropy. Compare this to the topological entropy obtained from the binary symbolic dynamics $\{0,1\}$.

Exercise 13.2 Sum of $A_{i j}$ is like a trace. Let $A$ be a matrix with eigenvalues $\lambda_{k}$. Show that

$$
\Gamma_{n}=\sum_{i, j}\left[A^{n}\right]_{i j}=\sum_{k} c_{k} \lambda_{k}^{n} .
$$

(a) Use this to show that $\ln \left|\operatorname{tr} A^{n}\right|$ and $\ln \left|\Gamma_{n}\right|$ have the same asymptotic behavior as $n \rightarrow \infty$, that is, their ratio converges to one.
(b) Do eigenvalues $\lambda_{k}$ need to be distinct, $\lambda_{k} \neq \lambda_{l}$ for $k \neq l$ ?

Exercise 13.3 Loop expansions. Prove by induction the sign rule in the determinant expansion (13.13):

$$
\operatorname{det}(1-z \mathbf{T})=\sum_{k \geq 0} \sum_{p_{1}+\cdots+p_{k}}(-1)^{k} t_{p_{1}} t_{p_{2}} \cdots t_{p_{k}} .
$$

Exercise 13.4 Transition matrix and cycle counting. Suppose you are given the Markov graph


This diagram can be encoded by a matrix $T$, where the entry $T_{i j}$ means that there is a link connecting node $i$ to node $j$. The value of the entry is the weight of the link.
a) Walks on the graph are given the weight that is the product of the weights of all links crossed by the walk. Convince yourself that the transition matrix for this graph is:

$$
T=\left[\begin{array}{ll}
a & b \\
c & 0
\end{array}\right]
$$

b) Enumerate all the walks of length three on the Markov graph. Now compute $T^{3}$ and look at the entries. Is there any relation between the terms in $T^{3}$ and all the walks?
c) Show that $T_{i j}^{n}$ is the number of walks from point $i$ to point $j$ in $n$ steps. (Hint: one might use the method of induction.)
d) Try to estimate the number $N(n)$ of walks of length $n$ for this simple Markov graph.
e) The topological entropy $h$ measures the rate of exponential growth of the total number of walks $N(n)$ as a function of $n$. What is the topological entropy for this Markov graph?

Exercise 13.5 3-disk prime cycle counting. A prime cycle $p$ of length $n_{p}$ is a single traversal of the orbit; its label is a non-repeating symbol string of $n_{p}$ symbols. For example, $\overline{12}$ is prime, but $\overline{2121}$ is not, since it is $\overline{21}=\overline{12}$ repeated.

Verify that a 3-disk pinball has 3, 2, 3, 6, 9, $\cdots$ prime cycles of length 2, 3, 4, 5, $6, \cdots$.

Exercise 13.6 "Golden mean" pruned map. Continuation of exercise 11.7: Show that the total number of periodic orbits of length $n$ for the "golden mean" tent map is

$$
\frac{(1+\sqrt{5})^{n}+(1-\sqrt{5})^{n}}{2^{n}}
$$

For continuation, see exercise 13.8. See also exercise 13.9.
Exercise 13.7 Alphabet $\{\mathbf{0}, \mathbf{1}\}$, prune _00_ . The Markov diagram figure 11.14(b) implements this pruning rule. The pruning rule implies that " 0 " must always be bracketed by " 1 "s; in terms of a new symbol $2=10$, the dynamics becomes unrestricted symbolic dynamics with with binary alphabet $\{1,2\}$. The cycle expansion (13.13) becomes

$$
\begin{align*}
1 / \zeta & =\left(1-t_{1}\right)\left(1-t_{2}\right)\left(1-t_{12}\right)\left(1-t_{112}\right) \ldots \\
& =1-t_{1}-t_{2}-\left(t_{12}-t_{1} t_{2}\right)-\left(t_{112}-t_{12} t_{1}\right)-\left(t_{122}-t_{12} t_{2}\right) \ldots \tag{13.35}
\end{align*}
$$

In the original binary alphabet this corresponds to:

$$
\begin{align*}
1 / \zeta= & 1-t_{1}-t_{10}-\left(t_{110}-t_{1} t_{10}\right) \\
& -\left(t_{1110}-t_{110} t_{1}\right)-\left(t_{11010}-t_{110} t_{10}\right) \ldots \tag{13.36}
\end{align*}
$$

This symbolic dynamics describes, for example, circle maps with the golden mean winding number, see chapter 24. For unimodal maps this symbolic dynamics is realized by the tent map of exercise 13.6.


Figure 13.5: (a) A unimodal map for which the critical point maps into the right hand fixed point in three iterations, and (b) the corresponding Markov graph (Kai T. Hansen).

## Exercise 13.8 Spectrum of the "golden mean" pruned map. (medium

 - Exercise 13.6 continued)(a) Determine an expression for $\operatorname{tr} \mathcal{L}^{n}$, the trace of powers of the PerronFrobenius operator (9.10) for the tent map of exercise 13.6.
(b) Show that the spectral determinant for the Perron-Frobenius operator is

$$
\begin{equation*}
\operatorname{det}(1-z \mathcal{L})=\prod_{k \text { even }}\left(1-\frac{z}{\Lambda^{k+1}}-\frac{z^{2}}{\Lambda^{2 k+2}}\right) \prod_{k \text { odd }}\left(1+\frac{z}{\Lambda^{k+1}}+\frac{z^{2}}{\Lambda^{2 k+2}}\right) . \tag{13.37}
\end{equation*}
$$

Exercise 13.9 A unimodal map example. Consider a unimodal map of figure 13.5(a) for which the critical point maps into the right hand fixed point in three iterations, $S^{+}=100 \overline{1}$. Show that the admissible itineraries are generated by the Markov graph figure 13.5(b).
(Kai T. Hansen)

Exercise 13.10 Glitches in shadowing.** Note that the combination $t_{00011}$ minus the "shadow" $t_{0} t_{0011}$ in (13.17) cancels exactly, and does not contribute to the topological polynomial (13.18). Are you able to construct a smaller Markov graph than figure 13.3(e)?

Exercise 13.11 Whence Möbius function? To understand where the Möbius function comes from consider the function

$$
\begin{equation*}
f(n)=\sum_{d \mid n} g(d) \tag{13.38}
\end{equation*}
$$

where $d \mid n$ stands for sum over all divisors $d$ of $n$. Invert recursively this infinite tower of equations and derive the Möbius inversion formula

$$
\begin{equation*}
g(n)=\sum_{d \mid n} \mu(n / d) f(d) \tag{13.39}
\end{equation*}
$$

Exercise 13.12 Counting prime binary cycles. In order to get comfortable with Möbius inversion reproduce the results of the second column of table 13.2.

Write a program that determines the number of prime cycles of length $n$. You might want to have this program later on to be sure that you have missed no 3-pinball prime cycles.

Exercise 13.13 Counting subsets of cycles. The techniques developed above can be generalized to counting subsets of cycles. Consider the simplest example of a dynamical system with a complete binary tree, a repeller map (11.8) with two straight branches, which we label 0 and 1. Every cycle weight for such map factorizes, with a factor $t_{0}$ for each 0 , and factor $t_{1}$ for each 1 in its symbol string. Prove that the transition matrix traces (13.5) collapse to $\operatorname{tr}\left(T^{k}\right)=\left(t_{0}+t_{1}\right)^{k}$, and $1 / \zeta$ is simply

$$
\begin{equation*}
\prod_{p}\left(1-t_{p}\right)=1-t_{0}-t_{1} \tag{13.40}
\end{equation*}
$$

Substituting (13.40) into the identity

$$
\prod_{p}\left(1+t_{p}\right)=\prod_{p} \frac{1-t_{p}^{2}}{1-t_{p}}
$$

we obtain

$$
\begin{align*}
\prod_{p}\left(1+t_{p}\right) & =\frac{1-t_{0}^{2}-t_{1}^{2}}{1-t_{0}-t_{1}}=1+t_{0}+t_{1}+\frac{2 t_{0} t_{1}}{1-t_{0}-t_{1}} \\
& =1+t_{0}+t_{1}+\sum_{n=2}^{\infty} \sum_{k=1}^{n-1} 2\binom{n-2}{k-1} t_{0}^{k} t_{1}^{n-k} \tag{13.41}
\end{align*}
$$

Hence for $n \geq 2$ the number of terms in the cumulant expansion with $k 0$ 's and $n-k$ 1 's in their symbol sequences is $2\binom{n-2}{k-1}$.

In order to count the number of prime cycles in each such subset we denote with $M_{n, k} \quad(n=1,2, \ldots ; k=\{0,1\}$ for $n=1 ; k=1, \ldots, n-1$ for $n \geq 2)$ the number of prime $n$-cycles whose labels contain $k$ zeros. Show that

$$
\begin{aligned}
M_{1,0} & =M_{1,1}=1 \\
n M_{n, k} & =\sum_{m \left\lvert\, \frac{n}{k}\right.} \mu(m)\binom{n / m}{k / m}, \quad n \geq 2, k=1, \ldots, n-1
\end{aligned}
$$

where the sum is over all $m$ which divide both $n$ and $k$.

Exercise $13.14 \quad$ Logarithmic periodicity of $\ln N_{n}{ }^{*}$. Plot $\ln N_{n}-n h$ for a system with a nontrivial finite Markov graph. Do you see any periodicity? If yes, why?

Exercise 13.15 4-disk pinball topological polynomial. Show that the 4disk pinball topological polynomial (the pruning affects only the fixed points and the 2-cycles) is given by

$$
\begin{align*}
1 / \zeta_{4-\text { disk }} & =(1-3 z) \frac{\left(1-z^{2}\right)^{6}}{(1-z)^{3}\left(1-z^{2}\right)^{3}} \\
& =(1-3 z)(1+z)^{3}=1-6 z^{2}-8 z^{3}-3 z^{4} \tag{13.42}
\end{align*}
$$

Exercise $13.16 \quad N$-disk pinball topological polynominal. Show that for an $N$-disk pinball, the topological polynominal is given by

$$
\begin{align*}
1 / \zeta_{N-d i s k} & =(1-(N-1) z) \frac{\left(1-z^{2}\right)^{N(N-1) / 2}}{(1-z)^{N-1}\left(1-z^{2}\right)^{(N-1)(N-2) / 2}} \\
& =(1-(N-1) z)(1+z)^{N-1} \tag{13.43}
\end{align*}
$$

The topological polynomial has a root $z^{-1}=N-1$, as we already know it should from (13.29) or (13.15). We shall see in sect. 22.4 that the other roots reflect the symmetry factorizations of zeta functions.

Exercise 13.17 Alphabet $\{a, b, c\}$, prune _ $a b_{\mathbf{Z}}$. The pruning rule implies that any string of " $b$ "s must be preceeded by a " $c$ "; so one possible alphabet is $\left\{a, c b{ }^{k} ; \bar{b}\right\}$, $k=0,1,2 \ldots$ As the rule does not prune the fixed point $\bar{b}$, it is explicitly included in the list. The cycle expansion (13.13) becomes

$$
\begin{aligned}
1 / \zeta & =\left(1-t_{a}\right)\left(1-t_{b}\right)\left(1-t_{c}\right)\left(1-t_{c b}\right)\left(1-t_{a c}\right)\left(1-t_{c b b}\right) \ldots \\
& =1-t_{a}-t_{b}-t_{c}+t_{a} t_{b}-\left(t_{c b}-t_{c} t_{b}\right)-\left(t_{a c}-t_{a} t_{c}\right)-\left(t_{c b b}-t_{c b} t_{b}\right) \ldots
\end{aligned}
$$

The effect of the _ab_ pruning is essentially to unbalance the 2 cycle curvature $t_{a b}-t_{a} t_{b}$; the remainder of the cycle expansion retains the curvature form.

## Exercise 13.18 Alphabet $\{0,1\}$, prune $n$ repeats. of " 0 " $\quad 000 \ldots 00$. .

This is equivalent to the $n$ symbol alphabet $\{1,2, \ldots, n\}$ unrestricted symbolic dynamics, with symbols corresponding to the possible 10... 00 block lengths: $2=10$, $3=100, \ldots, n=100 \ldots$. . . The cycle expansion (13.13) becomes

$$
\begin{equation*}
1 / \zeta=1-t_{1}-t_{2} \ldots-t_{n}-\left(t_{12}-t_{1} t_{2}\right) \ldots-\left(t_{1 n}-t_{1} t_{n}\right) \ldots \tag{13.44}
\end{equation*}
$$

Exercise 13.19 Alphabet $\{\mathbf{0 , 1}\}$, prune _1000_, _00100_, _01100_.
Show that the topological zeta function is given by

$$
\begin{equation*}
1 / \zeta=\left(1-t_{0}\right)\left(1-t_{1}-t_{2}-t_{23}-t_{113}\right) \tag{13.45}
\end{equation*}
$$

with the unrestricted 4-letter alphabet $\{1,2, \underline{23}, \underline{113}\}$. Here 2, 3, refer to 10, 100 respectively, as in exercise 13.18.

Exercise 13.20 Alphabet $\{\mathbf{0}, \mathbf{1}\}$, prune _1000_, _00100_, _01100_, _10011_. The first three pruning rules were incorporated in the preceeding exercise.
(a) Show that the last pruning rule _10011_ leads (in a way similar to exercise 13.19) to the alphabet $\left\{\underline{21^{k}}, \underline{23}, \underline{21^{k} 113} ; \overline{1}, \overline{0}\right\}$, and the cycle expansion

$$
\begin{equation*}
1 / \zeta=\left(1-t_{0}\right)\left(1-t_{1}-t_{2}-t_{23}+t_{1} t_{23}-t_{2113}\right) \tag{13.46}
\end{equation*}
$$

Note that this says that 1, 23, 2, 2113 are the fundamental cycles; not all cycles up to length 7 are needed, only 2113.
(b) Show that the topological polynomial is

$$
\begin{equation*}
1 / \zeta_{\text {top }}=(1-z)\left(1-z-z^{2}-z^{5}+z^{6}-z^{7}\right) \tag{13.47}
\end{equation*}
$$

and check that it yields the exact value of the entropy $h=0.522737642 \ldots$

## Chapter 14

## Trace formulas

The trace formula is not a formula, it is an idea. Martin Gutzwiller

Dynamics is posed in terms of local equations, but the ergodic averages require global information. How can we use a local description of a flow to learn something about the global behavior? We have given a quick sketch of this program in sects. 1.5 and 1.6 ; now we redo the same material in greater depth. In chapter 10 we have related global averages to the eigenvalues of appropriate evolution operators. We shall show that the traces of evolution operators can be evaluated as integrals over Dirac delta functions, and in this way relate the spectra of evolution operators to periodic orbits. If there is one idea that one should learn about chaotic dynamics, it happens in this chapter, and it is this: there is a fundamental local $\leftrightarrow$ global duality which says that
the spectrum of eigenvalues is dual to the spectrum of periodic orbits

For linear dynamics on the circle, this is called Fourier analysis; for dynamics on well-tiled manifolds, Selberg traces and zetas; and for generic nonlinear dynamical systems the duality is embodied in the trace formulas that we will now derive. These objects are to dynamics what partition functions are to statistical mechanics.

### 14.1 Trace of an evolution operator

Our extraction of the spectrum of $\mathcal{L}$ commences with the evaluation of the trace. To compute an expectation value using (10.21) we have to integrate over all the values of the kernel $\mathcal{L}^{t}(x, y)$. If $\mathcal{L}^{t}$ were a matrix we would be computing a weighted sum of its eigenvalues which is dominated by the leading eigenvalue as $t \rightarrow \infty$. As the trace of $\mathcal{L}^{t}$ is also dominated by the leading eigenvalue as $t \rightarrow \infty$, we might just as well look at the trace

$$
\begin{equation*}
\operatorname{tr} \mathcal{L}^{t}=\int d x \mathcal{L}^{t}(x, x)=\int d x \delta\left(x-f^{t}(x)\right) e^{\beta \cdot A^{t}(x)} . \tag{14.1}
\end{equation*}
$$

Assume that $\mathcal{L}$ has a spectrum of discrete eigenvalues $s_{0}, s_{1}, s_{2}, \cdots$ ordered so that $\operatorname{Re} s_{\alpha} \geq \operatorname{Re} s_{\alpha+1}$. We ignore for the time being the question of what function space the eigenfunctions belong to, as we shall compute the eigenvalue spectrum without constructing any explicit eigenfunctions.

On one hand, by definition, the trace is the sum over eigenvalues (for the time being we choose not to worry about convergence of such sums):

$$
\begin{equation*}
\operatorname{tr} \mathcal{L}^{t}=\sum_{\alpha=0}^{\infty} e^{s_{\alpha} t} . \tag{14.2}
\end{equation*}
$$

On the other hand, we have learned in sect. 9.2 how to evaluate the delta-function integral (14.1).

As the case of discrete time mappings is somewhat simpler, we first evaluate this integral for maps, and then for flows. The final formula (14.19) covers both cases.

### 14.1.1 Hyperbolicity assumption

According to (9.8) the trace (14.1) picks up a contribution whenever $x-$ $f^{n}(x)=0$, that is, whenever $x$ belongs to a periodic orbit. For reasons which we will explain in sect. 14.3, it is wisest to start by focusing on discrete time systems. The contribution of an isolated prime cycle $p$ of period $n_{p}$ for a map $f$ can be evaluated by restricting the integration to an infinitesimal open neighborhood $\mathcal{M}_{p}$ around the cycle points

$$
\begin{align*}
\operatorname{tr}_{p} \mathcal{L}^{n_{p}} & =\int_{\mathcal{M}_{p}} d x \delta\left(x-f^{n_{p}}(x)\right) \\
& =\frac{n_{p}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}\right)\right|}=n_{p} \prod_{i=1}^{d} \frac{1}{\left|1-\Lambda_{p, i}\right|} . \tag{14.3}
\end{align*}
$$

We set the observable $e^{\beta A_{p}}=1$ for the time being. The periodic orbit Jacobian matrix $\mathbf{J}_{p}$ is also known as the monodromy matrix (from Greek mono$=$ alone, single, and dromo $=$ run, racecourse), and its eigenvalues $\Lambda_{p, 1}$, $\Lambda_{p, 2}, \ldots, \Lambda_{p, d}$ are called the Floquet multipliers. We sort the eigenvalues $\Lambda_{p, 1}, \Lambda_{p, 2}, \ldots, \Lambda_{p, d}$ of the $p$-cycle $[d \times d]$ Jacobian matrix $\mathbf{J}_{p}$ into expanding, marginal and contracting sets $\{e, m, c\}$, as in (8.2). As the integral (14.3) can be carried out only if $\mathbf{J}_{p}$ has no eigenvalue of unit magnitude, we assume that no eigenvalue is marginal (we shall show in sect. 14.3 that the longitudinal $\Lambda_{p, d+1}=1$ eigenvalue for flows can be eliminated by restricting the
consideration to the transverse Jacobian matrix $\mathbf{J}_{p}$ ), and factorize the trace (14.3) into a product over the expanding and the contracting eigenvalues

$$
\begin{equation*}
\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}\right)\right|^{-1}=\frac{1}{\left|\Lambda_{p}\right|} \prod_{e} \frac{1}{1-1 / \Lambda_{p, e}} \prod_{c} \frac{1}{1-\Lambda_{p, c}} \tag{14.4}
\end{equation*}
$$

where $\Lambda_{p}=\prod_{e} \Lambda_{p, e}$ is the product of expanding eigenvalues. Both $\Lambda_{p, c}$ and $1 / \Lambda_{p, e}$ are smaller than 1 in absolute value, and as they are either real or come in complex conjugate pairs we are allowed to drop the absolute value brackets $|\cdots|$ in the above products.

The hyperbolicity assumption requires that the stabilities of all cycles included in the trace sums are exponentially bounded away from unity:

$$
\begin{array}{ll}
\left|\Lambda_{p, e}\right|>e^{\lambda_{e} T_{p}} & \text { any } p, \text { any expanding }\left|\Lambda_{p, e}\right|>1 \\
\left|\Lambda_{p, c}\right|<e^{-\lambda_{c} T_{p}} & \text { any } p, \text { any contracting }\left|\Lambda_{p, c}\right|<1 \tag{14.5}
\end{array}
$$

where $\lambda_{e}, \lambda_{c}>0$ form strictly positive bounds on the expanding and contracting cycle Lyapunov exponents. If a dynamical system satisfies the hyperbolicity assumption (for example, the well separated 3-disk system clearly does), the $\mathcal{L}^{t}$ spectrum is relatively easy to control. If the expansion/contraction is slower than exponential, let us say $\left|\Lambda_{p, i}\right| \sim T_{p}{ }^{2}$, the system may exhibit "phase transitions", and the analysis is much harder we shall turn to such problems in chapter 21.

It follows from (14.4) that for long times, $t=r T_{p} \rightarrow \infty$, only the product of expanding eigenvalues matters, $\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right| \rightarrow\left|\Lambda_{p}\right|^{r}$. We shall use this fact to motivate the construction of dynamical zeta functions in sect. 15.3. However, for evaluation of the full spectrum (14.2) the exact cycle weight (14.3) has to be kept.

### 14.2 A trace formula for maps

If the evolution is given by a discrete time mapping, and all periodic points have stability eigenvalues $\left|\Lambda_{p, i}\right| \neq 1$ strictly bounded away from unity, the trace $\mathcal{L}^{n}$ is given by the sum over all periodic points $i$ of period $n$ :

$$
\begin{equation*}
\operatorname{tr} \mathcal{L}^{n}=\int d x \mathcal{L}^{n}(x, x)=\sum_{x_{i} \in \operatorname{Fix} f^{n}} \frac{e^{\beta \cdot A_{i}}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}^{n}\left(x_{i}\right)\right)\right|} \tag{14.6}
\end{equation*}
$$

Here Fix $f^{n}=\left\{x: f^{n}(x)=x\right\}$ is the set of all periodic points of period $n$, and $A_{i}$ is the observable (10.5) evaluated over $n$ discrete time steps along the cycle to which the periodic point $x_{i}$ belongs. The weight follows from the properties of the Dirac delta function (9.8) by taking the determinant of $\partial_{i}\left(x_{j}-f^{n}(x)_{j}\right)$. We showed in (??) that if a trajectory retraces itself
$r$ times, its Jacobian matrix is $\mathbf{J}_{p}^{r}$, where $\mathbf{J}_{p}$ is the $[d \times d]$ Jacobian matrix (4.5) evaluated along a single traversal of the prime cycle $p$. As we saw in (10.5), the integrated observable $A^{n}$ is additive along the cycle: If a prime cycle $p$ trajectory retraces itself $r$ times, $n=r n_{p}$, we obtain $A_{p}$ repeated $r$ times, $A_{i}=A^{n}\left(x_{i}\right)=r A_{p}, x_{i} \in p$.

A prime cycle is a single traversal of the orbit, and its label is a nonrepeating symbol string. There is only one prime cycle for each cyclic permutation class. For example, the four cycle points $\overline{0011}=\overline{1001}=\overline{1100}$ $=\overline{0110}$ belong to the same prime cycle $p=0011$ of length 4 . As both the stability of a cycle and the weight $A_{p}$ are the same everywhere along the orbit, each prime cycle of length $n_{p}$ contributes $n_{p}$ terms to the sum, one for each cycle point. Hence (14.6) can be rewritten as a sum over all prime cycles and their repeats

$$
\begin{equation*}
\operatorname{tr} \mathcal{L}^{n}=\sum_{p} n_{p} \sum_{r=1}^{\infty} \frac{e^{r \beta \cdot A_{p}}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|} \delta_{n, n_{p} r}, \tag{14.7}
\end{equation*}
$$

with the Kronecker delta $\delta_{n, n_{p} r}$ projecting out the periodic contributions of total period $n$. This constraint is awkward, and will be more awkward still for the continuous time flows, where it will yield a series of Dirac delta spikes (14.17). Such sums are familiar from the density-of-states sums of statistical mechanics, where they are handled with in the same way as we shall do here: we smooth this distribution by taking a Laplace transform to remove the $\delta_{n, n_{p} r}$ constraint.

We define the trace formula for maps to be the Laplace transform of $\operatorname{tr} \mathcal{L}^{n}$ which, for discrete time mappings, is simply the generating function for the trace sums

$$
\begin{equation*}
\sum_{n=1}^{\infty} z^{n} \operatorname{tr} \mathcal{L}^{n}=\operatorname{tr} \frac{z \mathcal{L}}{1-z \mathcal{L}}=\sum_{p} n_{p} \sum_{r=1}^{\infty} \frac{z^{n_{p} r} e^{r \beta \cdot A_{p}}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|} \tag{14.8}
\end{equation*}
$$

Expressing the trace as in (14.2), in terms of the sum of the eigenvalues of $\mathcal{L}$, we obtain the trace formula for maps:

$$
\begin{equation*}
\sum_{\alpha=0}^{\infty} \frac{z e^{s_{\alpha}}}{1-z e^{s_{\alpha}}}=\sum_{p} n_{p} \sum_{r=1}^{\infty} \frac{z^{n_{p} r} e^{r \beta \cdot A_{p}}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|} . \tag{14.9}
\end{equation*}
$$

This is our second example of the duality between the spectrum of eigenvalues and the spectrum of periodic orbits, announced in the introduction to this chapter. (The first example was the topological entropy trace formula...).


## Example 14.1 A trace formula for a piecewise linear transfer operator

(10) For a piecewise-linear map (10.17), the trace formula can be explicitly evaluated. By the piecewise linearity and the chain rule $\Lambda_{p}=\Lambda_{0}^{n_{0}} \Lambda_{1}^{n_{1}}$, where the cycle $p$ contains $n_{0}$ symbols 0 and $n_{1}$ symbols 1 , the trace (14.6) reduces to

$$
\begin{equation*}
\operatorname{tr} \mathcal{L}^{n}=\sum_{m=0}^{n}\binom{n}{m} \frac{1}{\left|1-\Lambda_{0}^{m} \Lambda_{1}^{n-m}\right|}=\sum_{k=0}^{\infty}\left(\frac{1}{\left|\Lambda_{0}\right| \Lambda_{0}^{k}}+\frac{1}{\left|\Lambda_{1}\right| \Lambda_{1}^{k}}\right)^{n} \tag{14.10}
\end{equation*}
$$

The eigenvalues are simply

$$
\begin{equation*}
e^{s_{k}}=\frac{1}{\left|\Lambda_{0}\right| \Lambda_{0}^{k}}+\frac{1}{\left|\Lambda_{1}\right| \Lambda_{1}^{k}} \tag{14.11}
\end{equation*}
$$

For $k=0$ this agrees with the eigenvalue (10.20) of the Perron-Frobenius operator explicit transfer matrix $(10.19)$ representation. But what about the infinity of other eigenvalues?

The alert reader should experience anxiety at this point. Is it not true that we have already written down explicitly the transfer operator in (10.19), and that it is clear by inspection that it has only one non-zero eigenvalue $e^{s_{0}}=1 /\left|\Lambda_{0}\right|+1 /\left|\Lambda_{1}\right|$ ? The example at hand is one of the simplest illustrations of the necessity of defining the space on which the operator acts in order to define the spectrum. The transfer operator (10.19) is the correct operator on the space of functions piecewise constant on the two defining intervals $\left\{\mathcal{M}_{0}, \mathcal{M}_{1}\right\}$; on this space the operator indeed has only one non-zero eigenvalue $e^{s_{0}}$. As we shall see in sect. 16.1, the full spectrum (14.11) corresponds to the action of the transfer operator on the space of real analytic functions.

The Perron-Frobenius operator trace formula for the piecewise-linear map (10.17) follows from (14.8)

$$
\begin{equation*}
\operatorname{tr} \frac{z \mathcal{L}}{1-z \mathcal{L}}=\frac{z\left(\frac{1}{\left|\Lambda_{0}-1\right|}+\frac{1}{\left|\Lambda_{1}-1\right|}\right)}{1-z\left(\frac{1}{\left|\Lambda_{0}-1\right|}+\frac{1}{\left|\Lambda_{1}-1\right|}\right)} \tag{14.12}
\end{equation*}
$$

explicitly verifying the trace formula (14.9).

### 14.3 A trace formula for flows

Amazing! I did not understand a single word.
Fritz Haake
(R. Artuso and P. Cvitanović)

As any pair of nearby points on a cycle returns to itself exactly at each cycle period, the eigenvalue of the Jacobian matrix corresponding to the eigenvector along the flow necessarily equals unity for all periodic orbits. This violates the hyperbolicity assumption (14.5), so for flows the trace integral $\operatorname{tr} \mathcal{L}^{t}$ requires a separate treatment for the longitudinal direction. To evaluate the contribution of an isolated prime cycle $p$ of period $T_{p}$, restrict the integration to an infinitesimally thin tube $\mathcal{M}_{p}$ enveloping the cycle (see
figure 1.10), and choose a local coordinate system with a longitudinal coordinate $d x_{\|}$along the direction of the flow, and $d$ transverse coordinates $x_{\perp}$

$$
\begin{equation*}
\operatorname{tr}_{p} \mathcal{L}^{t}=\int_{\mathcal{M}_{p}} d x_{\perp} d x_{\|} \delta\left(x_{\perp}-f_{\perp}^{t}(x)\right) \delta\left(x_{\|}-f_{\|}^{t}(x)\right) . \tag{14.13}
\end{equation*}
$$

Here we again set the observable $\exp \left(\beta \cdot A^{t}\right)=1$ for the time being. Let $v(x)$ be the magnitude of the velocity at the point $x$ along the flow. $v(x)$ is strictly positive, as otherwise the orbit would stagnate for infinite time at $v(x)=0$ points. Therefore we can parametrize the longitudinal coordinate $x_{\|}$by the flight time

$$
x_{\|}(\tau)=\left.\int_{0}^{\tau} d \sigma v(\sigma)\right|_{\bmod L_{p}}
$$

where $v(\sigma)=v\left(x_{\|}(\sigma)\right)$, and $L_{p}$ is the length of the circuit on which the periodic orbit lies (for the time being the mod operation in the above definition is redundant, as $\tau \in\left[0, T_{p}\right]$ ). With this parametrization

$$
\left(f_{\|}^{t}(x)-x_{\|}\right)=\left.\int_{\tau}^{t+\tau} d \sigma v(\sigma)\right|_{\bmod L_{p}}
$$

the integral along the longitudinal coordinate can be written as

$$
\begin{equation*}
\int_{0}^{L_{p}} d x_{\|} \delta\left(x_{\|}-f_{\|}^{t}(x)\right)=\int_{0}^{T_{p}} d \tau v(\tau) \delta\left(\left.\int_{\tau}^{t+\tau} d \sigma v(\sigma)\right|_{\bmod L_{p}}\right) .( \tag{14.14}
\end{equation*}
$$

The zeroes of the argument of the delta function do not depend on $\tau$, as $v$ is positive, so we may rewrite (14.14) as

$$
\int_{0}^{L_{p}} d x_{\|} \delta\left(x_{\|}-f_{\|}^{t}(x)\right)=\sum_{r=1}^{\infty} \delta\left(t-r T_{p}\right) \int_{0}^{T_{p}} d \tau v(\tau) \frac{1}{v(\tau+t)},
$$

having used (9.7). The $r$ sum starts from $\mathrm{r}=1$, as we are considering strictly positive times. Now we use another elementary property of delta functions,

$$
h(x) \delta\left(x-x_{0}\right)=h\left(x_{0}\right) \delta\left(x-x_{0}\right) .
$$

The velocities cancel; we then get

$$
\begin{equation*}
\oint_{p} d x_{\|} \delta\left(x_{\|}-f_{\|}^{t}(x)\right)=T_{p} \sum_{r=1}^{\infty} \delta\left(t-r T_{p}\right) . \tag{14.15}
\end{equation*}
$$

The prime period arises also for repeated orbits, because the space integration (14.14) sweeps the periodic orbit in phase space: This observation will also be important for the derivation of the semiclassical trace formula in chapter 30 .

For the remaining transverse integration variables the Jacobian matrix is defined in a reduced Poincaré surface of section $\mathcal{P}$ of constant $x_{\|}$. Linearization of the periodic flow transverse to the orbit yields

$$
\begin{equation*}
\int_{\mathcal{P}} d x_{\perp} \delta\left(x_{\perp}-f_{\perp}^{r T_{p}}(x)\right)=\frac{1}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|} \tag{14.16}
\end{equation*}
$$

where $\mathbf{J}_{p}$ is the $p$-cycle $[d \times d]$ transverse Jacobian matrix, and as in (14.5) we have to assume hyperbolicity, that is, that the magnitudes of all transverse eigenvalues are bounded away from unity.

Substituting (14.15), (14.16) into (14.13), we obtain an expression for $\operatorname{tr} \mathcal{L}^{t}$ as a sum over all prime cycles $p$ and their repetitions

$$
\begin{equation*}
\operatorname{tr} \mathcal{L}^{t}=\sum_{p} T_{p} \sum_{r=1}^{\infty} \frac{e^{r \beta \cdot A_{p}}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|} \delta\left(t-r T_{p}\right) . \tag{14.17}
\end{equation*}
$$

A trace formula follows by taking a Laplace transform. This is a delicate step, since the transfer operator becomes the identity in the $t \rightarrow 0^{+}$limit. In order to make sense of the trace we regularize the Laplace transform by a lower cutoff $\epsilon$ smaller than the period of any periodic orbit, and write

$$
\begin{align*}
\int_{\epsilon}^{\infty} d t e^{-s t} \operatorname{tr} \mathcal{L}^{t} & =\operatorname{tr} \frac{e^{-(s-\mathcal{A}) \epsilon}}{s-\mathcal{A}}=\sum_{\alpha=0}^{\infty} \frac{e^{-\left(s-s_{\alpha}\right) \epsilon}}{s-s_{\alpha}} \\
& =\sum_{p} T_{p} \sum_{r=1}^{\infty} \frac{e^{r\left(\beta \cdot A_{p}-s T_{p}\right)}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|} \tag{14.18}
\end{align*}
$$

where $\mathcal{A}$ is the generator of the semigroup of dynamical evolution (See sect. 9.4). The classical trace formula for flows is the $\epsilon \rightarrow 0$ limit of the above expression:

$$
\begin{equation*}
\sum_{\alpha=0}^{\infty} \frac{1}{s-s_{\alpha}}=\sum_{p} T_{p} \sum_{r=1}^{\infty} \frac{e^{r\left(\beta \cdot A_{p}-s T_{p}\right)}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|} . \tag{14.19}
\end{equation*}
$$

This formula is still another example of the duality between the (local) cycles and (global) eigenvalues. If $T_{p}$ takes only integer values, we can replace $e^{-s} \rightarrow z$, so the trace formula for maps (14.9) is a special case of the trace formula for flows. The relation between continuous and discrete
time cases can be summarized as follows:

$$
\begin{align*}
T_{p} & \leftrightarrow n_{p} \\
e^{-s} & \leftrightarrow z \\
e^{t \mathcal{A}} & \leftrightarrow \mathcal{L}^{n} . \tag{14.20}
\end{align*}
$$

This is as it should be - any flow can be equally well, represented by Poincari section return maps, and in section? we shall make the relation between averages computed in continuous and discrete time precise. We could now proceed to estimate the location of the leading singularity of $\operatorname{tr}(s-\mathcal{A})^{-1}$ by extrapolating finite cycle length truncations of (14.19) by methods such as Padé approximants. However, it pays to first perform a simple resummation which converts this divergent trace into a zero of a spectral determinant. We shall do this in sect. 15.2, but a brief discussion of how all this relates to the heuristic derivation offered in the introduction.

### 14.4 An asymptotic trace formula



In order to illuminate the manipulations of sect. 14.2 and relate them to something we already know, we now rederive the heuristic sum of sect. 1.5.1 from the exact trace formula (14.9). The Laplace transforms (14.9), (14.19) are designed to capture the time $\rightarrow \infty$ asymptotic behavior of the trace sums. By the hyperbolicity assumption (14.5), for $t=T_{p} r$ large the cycle weight approaches

$$
\begin{equation*}
\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right| \rightarrow\left|\Lambda_{p}\right|^{r}, \tag{14.21}
\end{equation*}
$$

where $\Lambda_{p}$ is the product of the expanding eigenvalues of $\mathbf{J}_{p}$. Denote the corresponding approximation to the $n$th trace (14.6) by

$$
\begin{equation*}
\Gamma_{n}=\sum_{i}^{(n)} \frac{1}{\left|\Lambda_{i}\right|}, \tag{14.22}
\end{equation*}
$$

and denote the approximate trace formula obtained by replacing the cycle weights $\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|$ by $\left|\Lambda_{p}\right|^{r}$ in (14.9) by $\Gamma(z)$. Equivalently, think of this as a replacement of the evolution operator (10.23) by a transfer operator (as in example 14.1). For concreteness consider a dynamical system whose symbolic dynamics is complete binary, such as the 3 -disk system in figure 1.4. In this case distinct periodic points that contribute to the $n$th periodic points sum (14.7) are labeled by all admissible itineraries composed of sequences of letters $s_{i} \in\{0,1\}$ :

$$
\Gamma(z)=\sum_{n=1}^{\infty} z^{n} \Gamma_{n}=\sum_{n=1}^{\infty} z^{n} \sum_{x_{i} \in \operatorname{Fix} f^{n}} \frac{e^{\beta \cdot A^{n}\left(x_{i}\right)}}{\left|\Lambda_{i}\right|}
$$

$$
\begin{align*}
= & z\left\{\frac{e^{\beta \cdot A_{0}}}{\left|\Lambda_{0}\right|}+\frac{e^{\beta \cdot A_{1}}}{\left|\Lambda_{1}\right|}\right\} \\
& +z^{2}\left\{\frac{e^{2 \beta \cdot A_{0}}}{\left|\Lambda_{0}\right|^{2}}+\frac{e^{\beta \cdot A_{01}}}{\left|\Lambda_{01}\right|}+\frac{e^{\beta \cdot A_{10}}}{\left|\Lambda_{10}\right|}+\frac{e^{2 \beta \cdot A_{1}}}{\left|\Lambda_{1}\right|^{2}}\right\} \\
& +z^{3}\left\{\frac{e^{3 \beta \cdot A_{0}}}{\left|\Lambda_{0}\right|^{3}}+\frac{e^{\beta \cdot A_{001}}}{\left|\Lambda_{001}\right|}+\frac{e^{\beta \cdot A_{010}}}{\left|\Lambda_{010}\right|}+\frac{e^{\beta \cdot A_{100}}}{\left|\Lambda_{100}\right|}+\ldots\right\} \tag{14.23}
\end{align*}
$$

Both the cycle averages $A_{i}$ and the stabilities $\Lambda_{i}$ are the same for all points $x_{i} \in p$ in a cycle $p$. Summing over repeats of all prime cycles yields

$$
\begin{equation*}
\Gamma(z)=\sum_{p} \frac{n_{p} t_{p}}{1-t_{p}}, \quad t_{p}=z^{n_{p}} e^{\beta \cdot A_{p}} /\left|\Lambda_{p}\right| . \tag{14.24}
\end{equation*}
$$

This is precisely our initial heuristic estimate (1.8). Note that we could not perform such a sum over $r$ in the exact trace formula (14.9) as $\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right| \neq$ $\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}\right)\right|^{r}$; the correct way to resum the exact trace formulas is to first expand the factors $1 /\left|1-\Lambda_{p, i}\right|$, as we shall do in (15.9).

If the weights $e^{\beta A^{n}(x)}$ are multiplicative along the flow, and the flow is hyperbolic, for given $\beta$ the magnitude of each $\left|e^{\beta A^{n}\left(x_{i}\right)} / \Lambda_{i}\right|$ term is bounded by some constant $M^{n}$. The total number of cycles grows as $2^{n}$ (or as $e^{h n}$, where $h$ is the topological entropy), and the sum is convergent for $z$ sufficiently small, $|z|<1 / 2 M$. For large $n$ the $n$th level sum (14.6) tends to the leading eigenvalue $e^{n s_{0}}$ of $\mathcal{L}^{n}$. Summing this asymptotic estimate level by level yields

$$
\begin{equation*}
\Gamma(z) \approx \sum_{n=1}^{\infty}\left(z e^{s_{0}}\right)^{n}=\frac{z e^{s_{0}}}{1-z e^{s_{0}}} \tag{14.25}
\end{equation*}
$$

we see that we should be able to determine $s_{0}$ by determining the smallest value of $z=e^{-s_{0}}$ for which the cycle expansion (14.24) diverges.

If one is interested only in the leading eigenvalue of $\mathcal{L}$, it suffices to consider the approximate trace $\Gamma(z)$. We will use this fact in section xxxx to motivate the introduction of dynamical zeta functions (15.14), and in sect. 15.5 we shall give the exact relation between the exact and the approximate trace formulas.

## Commentary

Remark 14.1 Continuous time cycle weights Continuous time flow traces weighted by cycle periods were introduced by Bowen [14.1] who treated them as Poincaré section suspensions weighted by the "time ceiling" function (3.4). They were used by Parry and Pollicott [14.2]. The derivation presented here [14.3] parallels closely to the derivation of the Gutzwiller semiclassical trace formula of chapters 28 and 30.

Remark 14.2 Flat and sharp traces. In the above formal derivation of trace formulas we cared very little whether our sums were well posed. In Fredholm theory traces like (14.1) require compact operators with continuous function kernels. This is not the case for our Dirac delta evolution operators: nevertheless, there is a large class of dynamical systems for which our results may be shown to be perfectly legal. In the mathematical literature expressions like (14.6) are called flat traces (see chapter 16 and the review [14.4]). Other names for traces appear as well: for instance, in the context of $1-d$ mappings, sharp traces refer to generalizations of (14.6) where periodic point contributions are weighted by the Lefschetz sign $\pm 1$, reflecting whether the periodic point sits on a branch of $n$th iterate of the map which crosses the diagonal starting from below or starting from above [15.12]. Such traces are connected to the theory of kneading invariants (see ref. [14.4] and references therein). Traces weighted by $\pm 1$ sign of the derivative of the fixed point have been used to study the period doubling repeller, leading to high precision estimates of the Feigenbaum constant $\delta$, refs. [14.5, 18.6, 14.6].

## Résumé

The description of a chaotic dynamical system in terms of cycles can be visualized as a tessellation of the dynamical system, figure 1.9, with a smooth flow approximated by its periodic orbit skeleton. Each region $\mathcal{M}_{i}$ is centered on a periodic point $x_{i}$ of the topological length $n$, and the size of the region is determined by the linearization of the flow around the periodic point. The integral over this topologically partitioned phase space yields the classical trace formula

$$
\sum_{\alpha=0}^{\infty} \frac{1}{s-s_{\alpha}}=\sum_{p} T_{p} \sum_{r=1}^{\infty} \frac{e^{r\left(\beta \cdot A_{p}-s T_{p}\right)}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|}
$$

Now that we have a trace formula, we might ask for what is it good? As it stands, it is little more than a scary divergent formula which relates the unspeakable infinity of global eigenvalues to the unthinkable infinity of local unstable cycles. However, it is a good stepping stone on the way to the construction of spectral determinants (to which we turn next), and a first hint that when the going is good, sometimes the theory turn out to be convergent beyond our wildest dreams (chapter 16). In order to implement such formulas, we will have to determine "all" prime cycles. This task we postpone to chapters 12 and 17 .

## References

[14.1] R. Bowen, Equilibrium states and the ergodic theory of Anosov diffeomorphisms, Springer Lecture Notes in Math. 470 (1975).
[14.2] W. Parry and M. Pollicott, Zeta Functions and the periodic Structure of Hyperbolic Dynamics, Astérisque 187-188 (Société Mathématique de France, Paris 1990).
[14.3] P. Cvitanović and B. Eckhardt, J. Phys. A 24, L237 (1991).
[14.4] V. Baladi and D. Ruelle, Ergodic Theory Dynamical Systems 14, 621 (1994).
[14.5] R. Artuso, E. Aurell and P. Cvitanović, Nonlinearity 3, 325 (1990); ibidem 361 (1990)
[14.6] M. Pollicott, J. Stat. Phys. 62, 257 (1991).

## Exercises

Exercise 14.1 $t \rightarrow 0_{+}$regularization of eigenvalue sums**. In taking the Laplace transform (14.19) we have ignored the $t \rightarrow 0_{+}$divergence, as we do not know how to regularize the delta function kernel in this limit. In the quantum (or heat kernel) case this limit gives rise to the Weyl or Thomas-Fermi mean eigenvalue spacing (see sect. 30.1.1). Regularize the divergent sum in (14.19) following (for example) the prescription of appendix K. 6 and assign to such volume term some interesting role in the theory of classical resonance spectra. E-mail the solution to the authors.

Exercise 14.2 General weights. (easy) Let $f^{t}$ be a flow and $\mathcal{L}^{t}$ the operator

$$
\mathcal{L}^{t} g(x)=\int d y \delta\left(x-f^{t}(y)\right) w(t, y) g(y)
$$

where $w$ is a weight function. In this problem we will try and determine some of the properties $w$ must satisfy.
(a) Compute $\mathcal{L}^{s} \mathcal{L}^{t} g(x)$ to show that

$$
w\left(s, f^{t}(x)\right) w(t, x)=w(t+s, x)
$$

(b) Restrict $t$ and $s$ to be integers and show that the most general form of $w$ is

$$
w(n, x)=g(x) g(f(x)) g\left(f^{2}(x)\right) \cdots g\left(f^{n-1}(x)\right)
$$

for some $g$ that can be multiplied. Could $g$ be a function from $\mathbb{R}^{n_{1}} \mapsto$ $\mathbb{R}^{n_{2}} ?\left(n_{i} \in \mathbb{N}\right.$. $)$

## Chapter 15

## Spectral determinants

"It seems very pretty," she said when she had finished it, "but it's rather hard to understand!" (You see she didn't like to confess, even to herself, that she couldn't make it out at all.) "Somehow it seems to fill my head with ideas - only I don't exactly know what they are!"
Lewis Carroll, Through the Looking Glass
The problem with the trace formulas (14.9), (14.19) and (14.24) is that they diverge at $z=e^{-s_{0}}$, respectively $s=s_{0}$, that is, precisely where one would like to use them. While this does not prevent numerical estimation of some "thermodynamic" averages for iterated mappings, in the case of the Gutzwiller trace formula of chapter 30 this leads to a perplexing observation that crude estimates of the radius of convergence seem to put the entire physical spectrum out of reach. We shall now cure this problem by thinking, at no extra computational cost; while traces and determinats are formally
[183 chapter 30
[10) chapter 16 equivalent, determinants are the tool of choice when it comes to computing spectra. The idea is illustrated by figure 1.11: Determinants tend to have larger analyticity domains because if $\operatorname{tr} \mathcal{L} /(1-z \mathcal{L})=-\frac{d}{d z} \ln \operatorname{det}(1-z \mathcal{L})$ diverges at a particular value of $z$, then $\operatorname{det}(1-z \mathcal{L})$ might have an isolated zero there, and a zero of a function is easier to determine numerically than its poles.

### 15.1 Spectral determinants for maps

The eigenvalues $z_{k}$ of a linear operator are given by the zeros of the determinant

$$
\begin{equation*}
\operatorname{det}(1-z \mathcal{L})=\prod_{k}\left(1-z / z_{k}\right) \tag{15.1}
\end{equation*}
$$

For finite matrices this is the characteristic determinant; for operators this is the Hadamard representation of the spectral determinant (sparing the
reader from pondering possible regularization factors). Consider first the case of maps, for which the evolution operator advances the densities by integer steps in time. In this case we can use the formal matrix identity

$$
\begin{equation*}
\ln \operatorname{det}(1-M)=\operatorname{tr} \ln (1-M)=-\sum_{n=1}^{\infty} \frac{1}{n} \operatorname{tr} M^{n} \tag{15.2}
\end{equation*}
$$

to relate the spectral determinant of an evolution operator for a map to its traces (14.7), that is, periodic orbits:

$$
\begin{align*}
\operatorname{det}(1-z \mathcal{L}) & =\exp \left(-\sum_{n}^{\infty} \frac{z^{n}}{n} \operatorname{tr} \mathcal{L}^{n}\right) \\
& =\exp \left(-\sum_{p} \sum_{r=1}^{\infty} \frac{1}{r} \frac{z^{n_{p} r} e^{r \beta \cdot A_{p}}}{\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right) \mid}\right) \tag{15.3}
\end{align*}
$$

Going the other way, the trace formula (14.9) can be recovered from the spectral determinant by taking a derivative

$$
\begin{equation*}
\operatorname{tr} \frac{z \mathcal{L}}{1-z \mathcal{L}}=-z \frac{d}{d z} \ln \operatorname{det}(1-z \mathcal{L}) \tag{15.4}
\end{equation*}
$$


fast track: sect. 15.2, p. 249

## Example 15.1 Spectral determinants of transfer operators:

For a piecewise-linear map (10.17) with a finite Markov partition, an explicit formula for the spectral determinant follows by substituting the trace formula (14.10) into (15.3):

$$
\begin{equation*}
\operatorname{det}(1-z \mathcal{L})=\prod_{k=0}^{\infty}\left(1-\frac{t_{0}}{\Lambda_{0}^{k}}-\frac{t_{1}}{\Lambda_{1}^{k}}\right) \tag{15.5}
\end{equation*}
$$

where $t_{s}=z /\left|\Lambda_{s}\right|$. The eigenvalues are - as they should be - (14.11), the ones that we already determined from the trace formula (14.9).

The exponential spacing of eigenvalues guarantees that the spectral determinant (15.5) is an entire function. It is this property that will generalize to piecewise smooth flows with finite Markov parititions, and single out spectral determinants rather than the trace formulas or dynamical zeta functions as the tool of choice for evaluation of spectra.

### 15.2 Spectral determinant for flows

... an analogue of the [Artin-Mazur] zeta function for diffeomorphisms seems quite remote for flows. However we will mention a wild idea in this direction. [...] define $l(\gamma)$ to be the minimal period of $\gamma[\cdots]$ then define formally (another zeta function!) $Z(s)$ to be the infinite product

$$
Z(s)=\prod_{\gamma \in \Gamma} \prod_{k=0}^{\infty}\left(1-[\exp l(\gamma)]^{-s-k}\right)
$$

Stephen Smale, Differentiable Dynamical Systems
We write the formula for the spectral determinant for flows by analogy to (15.3)

$$
\begin{equation*}
\operatorname{det}(s-\mathcal{A})=\exp \left(-\sum_{p} \sum_{r=1}^{\infty} \frac{1}{r} \frac{e^{r\left(\beta \cdot A_{p}-s T_{p}\right)}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|}\right), \tag{15.6}
\end{equation*}
$$

and then check that the trace formula (14.19) is the logarithmic derivative of the spectral determinant so defined

$$
\begin{equation*}
\operatorname{tr} \frac{1}{s-\mathcal{A}}=\frac{d}{d s} \ln \operatorname{det}(s-\mathcal{A}) . \tag{15.7}
\end{equation*}
$$

With $z$ set to $z=e^{-s}$ as in (14.20), the spectral determinant (15.6) has the same form for both maps and flows. We shall refer to (15.6) as spectral determinant, as the spectrum of the operator $\mathcal{A}$ is given by the zeros of

$$
\begin{equation*}
\operatorname{det}(s-\mathcal{A})=0 \tag{15.8}
\end{equation*}
$$

We now note that the $r$ sum in (15.6) is close in form to the expansion of a logarithm. This observation enables us to recast the spectral determinant into an infinite product over periodic orbits as follows:

Let $\mathbf{J}_{p}$ be the $p$-cycle $[d \times d]$ transverse Jacobian matrix, with eigenvalues $\Lambda_{p, 1}, \Lambda_{p, 2}, \ldots, \Lambda_{p, d}$. Expanding the expanding eigenvalue factors $1 /\left(1-1 / \Lambda_{p, e}\right)$ and the contracting eigenvalue factors $1 /\left(1-\Lambda_{p, c}\right)$ in (14.4) as geometric series, substituting back into (15.6), and resumming the logarithms, we find that the spectral determinant is formally given by the infinite product

$$
\begin{align*}
\operatorname{det}(s-\mathcal{A}) & =\prod_{k_{1}=0}^{\infty} \cdots \prod_{l_{c}=0}^{\infty} \frac{1}{\zeta_{k_{1} \cdots l_{c}}} \\
1 / \zeta_{k_{1} \cdots l_{c}} & =\prod_{p}\left(1-t_{p} \frac{\Lambda_{p, e+1}^{l_{1}} \Lambda_{p, e+2}^{l_{2}} \cdots \Lambda_{p, d}^{l_{c}}}{\Lambda_{p, 1}^{k_{1}} \Lambda_{p, 2}^{k_{2}} \cdots \Lambda_{p, e}^{k_{e}}}\right)  \tag{15.9}\\
t_{p} & =t_{p}(z, s, \beta)=\frac{1}{\left|\Lambda_{p}\right|} e^{\beta \cdot A_{p}-s T_{p}} z^{n_{p}} . \tag{15.10}
\end{align*}
$$

In such formulas $t_{p}$ is a weight associated with the $p$ cycle (letter $t$ refers to the "local trace" evaluated along the $p$ cycle trajectory), and the index $p$ runs through all distinct prime cycles. When convenient, we inserts the $z^{n_{p}}$ factor into cycle weights, as a formal parameter which keeps track of the topological cycle lengths. These factors will assists us in expanding zeta functions and determinants, eventually we shall set $z=1$. The subscripts $e, c$ indicate that there are $e$ expanding eigenvalues, and $c$ contracting eigenvalues. The observable whose average we wish to compute contributes through the $A^{t}(x)$ term in the $p$ cycle multiplicative weight $e^{\beta \cdot A_{p}}$. By its definition (10.1), for maps the weight is a product along the cycle points

$$
e^{A_{p}}=\prod_{j=0}^{n_{p}-1} e^{a\left(f^{j}\left(x_{p}\right)\right)},
$$

and for the flows the weight is an exponential of the integral (10.5) along the cycle

$$
e^{A_{p}}=\exp \left(\int_{0}^{T_{p}} a(x(\tau)) d \tau\right) .
$$

This formula is correct for scalar weighting functions; more general matrip valued weights require a time-ordering prescription as in the Jacobian matrix of sect. 4.1.

Example 15.2 Expanding 1-d map:


For expanding 1-d mappings the spectral determinant (15.9) takes the form

$$
\begin{equation*}
\operatorname{det}(1-z \mathcal{L})=\prod_{p} \prod_{k=0}^{\infty}\left(1-t_{p} / \Lambda_{p}^{k}\right), \quad t_{p}=\frac{e^{\beta A_{p}}}{\left|\Lambda_{p}\right|} z^{n_{p}} . \tag{15.11}
\end{equation*}
$$

Example 15.3 Two-degree of freedom Hamiltonian flows: For a 2-degree of freedom Hamiltonian flows the energy conservation eliminates on phase-space variable, and restriction to a Poincare section eliminates the marginal logitudinal eigenvalue $\Lambda=$ 1, so a periodic orbit of 2-degree of freedom hyperbolic Hamiltonian flow has one expanding transverse eigenvalue $\Lambda,|\Lambda|>1$, and one contracting transverse eigenvalue $1 / \Lambda$. The weight in (14.4) is expanded as follows:

$$
\begin{equation*}
\frac{1}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|}=\frac{1}{|\Lambda|^{r}\left(1-1 / \Lambda_{p}^{r}\right)^{2}}=\frac{1}{|\Lambda|^{r}} \sum_{k=0}^{\infty} \frac{k+1}{\Lambda_{p}^{k r}} \tag{15.12}
\end{equation*}
$$

The spectral determinant exponent can be resumed,

$$
-\sum_{r=1}^{\infty} \frac{1}{r} \frac{e^{\left(\beta A_{p}-s T_{p}\right) r}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|}=\sum_{k=0}^{\infty}(k+1) \log \left(1-\frac{e^{\beta A_{p}-s T_{p}}}{\left|\Lambda_{p}\right| \Lambda_{p}^{k}}\right),
$$

and the spectral determinant for a 2-dimensional hyperbolic Hamiltonian flow rewritten as an infinite product over prime cycles

$$
\begin{equation*}
\operatorname{det}(s-\mathcal{A})=\prod_{p} \prod_{k=0}^{\infty}\left(1-t_{p} / \Lambda_{p}^{k}\right)^{k+1} \tag{15.13}
\end{equation*}
$$

Now we are finally poised to deal with the problem posed at the beginning of chapter 14; how do we actually evaluate the averages introduced in sect. 10.1? The eigenvalues of the dynamical averaging evolution operator are given by the values of $s$ for which the spectral determinant (15.6) of the evolution operator (10.23) vanishes. If we can compute the leading eigenvalue $s_{0}(\beta)$ and its derivatives, we are done. Unfortunately, the infinite product formula (15.9) is no more than a shorthand notation for the periodic orbit weights contributing to the spectral determinant; more work will be needed to bring such formulas into a tractable form. This we shall accomplish in chapter 18, but this point in the narrative is a natural point to introduce a still another variant of a determinant, the dynamical zeta function.

### 15.3 Dynamical zeta functions

It follows from sect. 14.1.1 that if one is interested only in the leading eigenvalue of $\mathcal{L}^{t}$, the size of the $p$ cycle neighborhood can be approximated by $1 /\left|\Lambda_{p}\right|^{r}$, the dominant term in the $r T_{p}=t \rightarrow \infty$ limit, where $\Lambda_{p}=$ $\prod_{e} \Lambda_{p, e}$ is the product of the expanding eigenvalues of the Jacobian matrix $\mathbf{J}_{p}$. With this replacement the spectral determinant (15.6) is replaced by the dynamical zeta function

$$
\begin{equation*}
1 / \zeta=\exp \left(-\sum_{p} \sum_{r=1}^{\infty} \frac{1}{r} r_{p}^{r}\right) \tag{15.14}
\end{equation*}
$$

that we have already derived heuristically in sect. 1.5.2. Resumming the logarithms using $\sum_{r} t_{p}^{r} / r=-\ln \left(1-t_{p}\right)$ we obtain the Euler product representation of the dynamical zeta function:

$$
\begin{equation*}
1 / \zeta=\prod_{p}\left(1-t_{p}\right) \tag{15.15}
\end{equation*}
$$

For reasons of economy of the notation, we shall usually omit the explicit dependence of $1 / \zeta, t_{p}$ on $z, s, \beta$ whenever the dependence is clear from the context.

The approximate trace formula (14.24) plays the same role vis-à-vis the dynamical zeta function (15.7)

$$
\begin{equation*}
\Gamma(s)=\frac{d}{d s} \ln \zeta^{-1}=\sum_{p} \frac{T_{p} t_{p}}{1-t_{p}}, \tag{15.16}
\end{equation*}
$$

as the exact trace formula (14.19) plays vis-à-vis the spectral determinant (15.6). The heuristically derived dynamical zeta function of sect. 1.5.2 now re-emerges as the $1 / \zeta_{0 \ldots 0}(z)$ part of the exact spectral determinant; other factors in the infinite product (15.9) affect the non-leading eigenvalues of $\mathcal{L}$.

To summarize: the dynamical zeta function (15.15) associated with the flow $f^{t}(x)$ is defined as the product over all prime cycles $p . T_{p}, n_{p}$ and $\Lambda_{p}$ are the period, topological length and product of the expanding stability eigenvalues of prime cycle $p, A_{p}$ is the integrated observable $a(x)$ evaluated on a single traversal of cycle $p$ (see (10.5)), $s$ is a variable dual to the time $t$, $z$ is dual to the discrete "topological" time $n$, and $t_{p}(z, s, \beta)$ is the local trace over the cycle $p$. We have included the factor $z^{n_{p}}$ in the definition of the cycle weight in order to keep track of the number of times a cycle traverses the surface of section. The dynamical zeta function is useful because

$$
\begin{equation*}
1 / \zeta(s)=0 \tag{15.17}
\end{equation*}
$$

vanishes at $s$ equal to $s_{0}$, the leading eigenvalue of $\mathcal{L}^{t}=e^{t \mathcal{A}}$, and often the leading eigenvalue is all that is needed in applications. The above completes our derivation of the trace and determinant formulas for classical chaotic flows. In chapters that follow we shall make these formulas tangible by working out a series of simple examples.

The remainder of this chapter offers examples of zeta functions.


### 15.3.1 A contour integral formulation

The following observation is sometimes useful, in particular when the zeta functions have richer analytic structure than just zeros and poles, as in the case of intermittency (chapter 21): $\Gamma_{n}$, the trace sum (14.22), can be expressed in terms of the dynamical zeta function (15.15)

$$
\begin{equation*}
1 / \zeta(z)=\prod_{p}\left(1-\frac{z^{n_{p}}}{\left|\Lambda_{p}\right|}\right) \tag{15.18}
\end{equation*}
$$

as a contour integral

$$
\begin{equation*}
\Gamma_{n}=\frac{1}{2 \pi i} \oint_{\gamma_{r}^{-}} z^{-n}\left(\frac{d}{d z} \log \zeta^{-1}(z)\right) d z \tag{15.19}
\end{equation*}
$$

where a small contour $\gamma_{r}^{-}$encircles the origin in negative (clockwise) direction. If the contour is small enough, that is, it lies inside the unit circle

Figure 15.1: The survival probability $\Gamma_{n}$ can be split into contributions from poles (x) and zeros (o) between the small and the large circle and a contribution from the large circle.

$|z|=1$, we may write the logarithmic derivative of $\zeta^{-1}(z)$ as a convergent sum over all periodic orbits. Integrals and sums can be interchanged, the integrals can be solved term by term, and the trace formula (14.22) is recovered. For hyperbolic maps, cycle expansions or other techniques provide an analytic extension of the dynamical zeta function beyond the leading zero; we may therefore deform the orignal contour into a larger circle with radius $R$ which encircles both poles and zeros of $\zeta^{-1}(z)$, see figure 15.1. Residue calculus turns this into a sum over the zeros $z_{\alpha}$ and poles $z_{\beta}$ of the dynamical zeta function, that is

$$
\begin{equation*}
\Gamma_{n}=\sum_{\left|z_{\alpha}\right|<R}^{\text {zeros }} \frac{1}{z_{\alpha}^{n}}-\sum_{\left|z_{\beta}\right|<R}^{\text {poles }} \frac{1}{z_{\beta}^{n}}+\frac{1}{2 \pi i} \oint_{\gamma_{R}^{-}} d z z^{-n} \frac{d}{d z} \log \zeta^{-1}, \tag{15.20}
\end{equation*}
$$

where the last term gives a contribution from a large circle $\gamma_{R}^{-}$. It would be a miracle if you still remebered this, but in sect. 1.4.2 we interpreted $\Gamma_{n}$ as fraction of survivors after $n$ bounces, and defined the escape rate $\gamma$ as the rate of the find exponential decay of $\Gamma_{n}$. We now see that this exponential decay is dominated by the leading zero or pole of $\zeta^{-1}(z)$.

### 15.3.2 Dynamical zeta functions for transfer operators

Ruelle's original dynamical zeta function was a generalization of the topological zeta function (13.21) to a function that assigns different weights to different cycles:

$$
\zeta(z)=\exp \sum_{n=1}^{\infty} \frac{z^{n}}{n}\left(\sum_{x_{i} \in \operatorname{Fix} f^{n}} \operatorname{tr} \prod_{j=0}^{n-1} g\left(f^{j}\left(x_{i}\right)\right)\right) .
$$

Here the sum goes over all periodic points $x_{i}$ of period $n$, and $g(x)$ is any page 246 (matrix valued) weighting function, with weight evaluated multiplicatively along the trajectory of $x_{i}$.

By the chain rule (4.34) the stability of any $n$-cycle of a $1-d$ map is given by $\Lambda_{p}=\prod_{j=1}^{n} f^{\prime}\left(x_{i}\right)$, so the $1-d$ map cycle stability is the simplest example
of a multiplicative cycle weight $g\left(x_{i}\right)=1 /\left|f^{\prime}\left(x_{i}\right)\right|$, and indeed - via the Perron-Frobenius evolution operator (9.9) - the historical motivation for Ruelle's more abstract construction.

In particular, for a piecewise-linear map with a finite Markov partition such as the map of exercise 9.1, the dynamical zeta function is given by a finite polynomials, a straightforward generalization of the topological transition matrix determinant (11.2). As explained in sect. 13.3, for a finite [ $N \times N$ ] dimensional matrix the determinant is given by

$$
\prod_{p}\left(1-t_{p}\right)=\sum_{n=1}^{N} z^{n} c_{n}
$$

where $c_{n}$ is given by the sum over all non-self-intersecting closed paths of length $n$ together with products of all non-intersecting closed paths of total length $n$. One illustration is the piecewise linear repeller of sect. 14.1. (10.17).

Example 15.4 A piecewise linear repeller: Due to the piecewise linearity, the stability of any $n$-cycle of the piecewise linear repeller (10.17) factorizes as $\Lambda_{s_{1} s_{2} \ldots s_{n}}=$ $\Lambda_{0}^{m} \Lambda_{1}^{n-m}$, where $m$ is the total number of times the letter $s_{j}=0$ appears in the $p$ symbol sequence, so the traces in the sum (14.24) are of a particularly simple form

$$
\operatorname{tr} T^{n}=\Gamma_{n}=\left(\frac{1}{\left|\Lambda_{0}\right|}+\frac{1}{\left|\Lambda_{1}\right|}\right)^{n}
$$

The dynamical zeta function (15.14) evaluated by resumming the traces,

$$
\begin{equation*}
1 / \zeta(z)=1-z /\left|\Lambda_{0}\right|-z /\left|\Lambda_{1}\right| \tag{15.21}
\end{equation*}
$$

is indeed the determinant $\operatorname{det}(1-z T)$ of the transfer operator (10.19), almost as simple as the topological zeta function (13.25).

More generally, piecewise-linear approximations to dynamical systems yield polynomial or rational polynomial cycle expansions, provided that the symbolic dynamics is a subshift of finite type.

We see that the exponential proliferation of cycles so dreaded by quantum chaoticists is a bogus anxiety; we are dealing with exponentially many cycles of increasing length and instability, but all that really matters in this example are the stabilities of the two fixed points. Clearly the information carried by the infinity of longer cycles is highly redundant; we shall learn in chapter 18 how to exploit systematically this redundancy.

### 15.4 False zeros

Compare (15.21) with the Euler product (15.15). For simplicity take the two scales equal, $\left|\Lambda_{0}\right|=\left|\Lambda_{1}\right|=e^{\lambda}$. Our task is to determine the leading
zero $z=e^{\gamma}$ of the Euler product. It is a novice error to assume that the infinite Euler product (15.15) vanishes whenever one of its factors vanishes. If that were true, each factor $\left(1-z^{n_{p}} /\left|\Lambda_{p}\right|\right)$ would yield

$$
\begin{equation*}
0=1-e^{n_{p}\left(\gamma-\lambda_{p}\right)}, \tag{15.22}
\end{equation*}
$$

that is, the escape rate $\gamma$ would equal the stability exponent of a repulsive cycle, one eigenvalue $\gamma=\gamma_{p}$ for each prime cycle $p$. False! The exponentially growing number of cycles with growing period conspires to shift the zeros of the infinite product. The correct formula follows from (15.21)

$$
\begin{equation*}
0=1-e^{\gamma-\lambda+h}, \quad h=\ln 2 . \tag{15.23}
\end{equation*}
$$

This particular formula for the escape rate is a special case of a general relation between escape rates, Lyapunov exponents and entropies that is not yet included into this book. The physical interpretation is that the escape induced by repulsion by each unstable fixed point is diminished by the rate of backscatter from other repelling regions, that is, the entropy $h$; the positive entropy of orbits shifts the "false zeros" $z=e^{\lambda_{p}}$ of the Euler product (15.15) to the true zero $z=e^{\lambda-h}$.

### 15.5 Spectral determinants vs. dynamical zeta functions

In sect. 15.3 we derived the dynamical zeta function as an approximation to the spectral determinant. Here we relate dynamical zeta functions to the spectral determinants exactly, by showing that a dynamical zeta function can be expressed as a ratio of products of spectral determinants.

The elementary identity for $d$-dimensional matrices

$$
\begin{equation*}
1=\frac{1}{\operatorname{det}(1-\mathbf{J})} \sum_{k=0}^{d}(-1)^{k} \operatorname{tr}\left(\wedge^{k} \mathbf{J}\right), \tag{15.24}
\end{equation*}
$$

inserted into the exponential representation (15.14) of the dynamical zeta function, relates the dynamical zeta function to weighted spectral determinants.

Example 15.5 Dynamical zeta function in terms of determinants, $1-d$ maps: For 1-d maps the identity

$$
1=\frac{1}{(1-1 / \Lambda)}-\frac{1}{\Lambda} \frac{1}{(1-1 / \Lambda)}
$$

substituted into (15.14) yields an expression for the dynamical zeta function for 1-d maps as a ratio of two spectral determinants

$$
\begin{equation*}
1 / \zeta=\frac{\operatorname{det}(1-z \mathcal{L})}{\operatorname{det}\left(1-z \mathcal{L}_{(1)}\right)} \tag{15.25}
\end{equation*}
$$

where the cycle weight in $\mathcal{L}_{(1)}$ is given by replacement $t_{p} \rightarrow t_{p} / \Lambda_{p}$. As we shall see in chapter 16, this establishes that for nice hyperbolic flows $1 / \zeta$ is meromorphic, with poles given by the zeros of $\operatorname{det}\left(1-z \mathcal{L}_{(1)}\right)$. The dynamical zeta function and the spectral determinant have the same zeros - only in exceptional circumstances some zeros of $\operatorname{det}\left(1-z \mathcal{L}_{(1)}\right)$ might be cancelled by coincident zeros of $\operatorname{det}\left(1-z \mathcal{L}_{(1)}\right)$. Hence even though we have derived the dynamical zeta function in sect. 15.3 as an "approximation" to the spectral determinant, the two contain the same spectral information.

Example 15.6 Dynamical zeta function in terms of determinants, 2- $d$ Hamiltonian maps: For 2-dimensional Hamiltonian flows the above identity yields

$$
\frac{1}{|\Lambda|}=\frac{1}{|\Lambda|(1-1 / \Lambda)^{2}}\left(1-2 / \Lambda+1 / \Lambda^{2}\right),
$$

so

$$
\begin{equation*}
1 / \zeta=\frac{\operatorname{det}(1-z \mathcal{L}) \operatorname{det}\left(1-z \mathcal{L}_{(2)}\right)}{\operatorname{det}\left(1-z \mathcal{L}_{(1)}\right)} . \tag{15.26}
\end{equation*}
$$

This establishes that for nice hyperbolic flows dynamical zeta function is meromorphic in 2-d.

Example 15.7 Dynamical zeta functions for 2-d Hamiltonian flows: The relation (15.26) is not particularly useful for our purposes. Instead we insert the identity

$$
1=\frac{1}{(1-1 / \Lambda)^{2}}-\frac{2}{\Lambda} \frac{1}{(1-1 / \Lambda)^{2}}+\frac{1}{\Lambda^{2}} \frac{1}{(1-1 / \Lambda)^{2}}
$$

into the exponential representation (15.14) of $1 / \zeta_{k}$, and obtain

$$
\begin{equation*}
1 / \zeta_{k}=\frac{\operatorname{det}\left(1-z \mathcal{L}_{(k)}\right) \operatorname{det}\left(1-z \mathcal{L}_{(k+2)}\right)}{\operatorname{det}\left(1-z \mathcal{L}_{(k+1)}\right)^{2}} . \tag{15.27}
\end{equation*}
$$

Even though we have no guarantee that $\operatorname{det}\left(1-z \mathcal{L}_{(k)}\right)$ are entire, we do know that the upper bound on the leading zeros of $\operatorname{det}\left(1-z \mathcal{L}_{(k+1)}\right)$ lies strictly below the leading zeros of $\operatorname{det}\left(1-z \mathcal{L}_{(k)}\right)$, and therefore we expect that for 2-dimensional Hamiltonian flows the dynamical zeta function $1 / \zeta_{k}$ has generically a double leading pole coinciding with the leading zero of the $\operatorname{det}\left(1-z \mathcal{L}_{(k+1)}\right)$ spectral determinant. This might fail if the poles and leading eigenvalues come in wrong order, but we have not encountered such situation in our numerical investigations. This result can also be stated as follows: the theorem that establishes that the spectral determinant (15.13) is entire, implies that the poles in $1 / \zeta_{k}$ must have right multiplicities in order that they be cancelled in the $\operatorname{det}(1-z \mathcal{L})=\prod 1 / \zeta_{k}^{k+1}$ product.


Figure 15.2: (a) A game of pinball consisting of two disks of equal size in a plane, with its only periodic orbit (A. Wirzba). (b) The classical resonances $\alpha=\{k, n\}$ for a 2-disk game of pinball, equation (15.28).

### 15.6 All too many eigenvalues?

2. 

What does the 2-dimensional hyperbolic Hamiltonian flow spectral determinant (15.13) tell us? Consider one of the simplest conceivable hyperbolic flows: the game of pinball of figure 15.2 (a) consisting of two disks of equal size in a plane. There is only one periodic orbit, with the period $T$ and the expanding eigenvalue $\Lambda$ is given by elementary considerations (see exercise 8.1), and the resonances $\operatorname{det}\left(s_{\alpha}-\mathcal{A}\right)=0, \alpha=\{k, n\}$ plotted in figure 15.2 (b)

$$
\begin{equation*}
s_{\alpha}=-(k+1) \lambda+n \frac{2 \pi i}{T}, \quad n \in \mathbb{Z}, k \in \mathbb{Z}_{+}, \quad \text { multiplicity } k+1, \tag{15.28}
\end{equation*}
$$

can be read off the spectral determinant (15.13) for a single unstable cycle:

$$
\begin{equation*}
\operatorname{det}(s-\mathcal{A})=\prod_{k=0}^{\infty}\left(1-e^{-s T} /|\Lambda| \Lambda^{k}\right)^{k+1} \tag{15.29}
\end{equation*}
$$

In the above $\lambda=\ln |\Lambda| / T$ is the cycle Lyapunov exponent. For an open system, the real part of the eigenvalue $s_{\alpha}$ gives the decay rate of $\alpha$ th eigenstate, and the imaginary part gives the "node number" of the eigenstate. The negative real part of $s_{\alpha}$ indicates that the resonance is unstable, and the decay rate in this simple case (zero entropy) equals to the cycle Lyapunov exponent.

Fast decaying eigenstates with large negative $\operatorname{Re} s_{\alpha}$ are not a problem, but as there are eigenvalues arbitrarily far in the imaginary direction, this might seem like all too many eigenvalues. However, they are necessary we can check this by explicit computation of the right hand side of (14.19), the trace formula for flows:

$$
\sum_{\alpha=0}^{\infty} e^{s_{\alpha} t}=\sum_{k=0}^{\infty} \sum_{n=-\infty}^{\infty}(k+1) e^{(k+1) \lambda t+i 2 \pi n t / T}
$$

$$
\begin{align*}
& =\sum_{k=0}^{\infty}(k+1)\left(\frac{1}{|\Lambda| \Lambda^{k}}\right)^{t / T} \sum_{n=-\infty}^{\infty} e^{i 2 \pi n / T} \\
& =\sum_{k=0}^{\infty} \frac{k+1}{|\Lambda|^{r} \Lambda^{k r}} \sum_{r=-\infty}^{\infty} \delta(r-t / T) \\
& =T \sum_{r=-\infty}^{\infty} \frac{\delta(t-r T)}{|\Lambda|\left(1-1 / \Lambda^{r}\right)^{2}} \tag{15.30}
\end{align*}
$$

So the two sides of the trace formula (14.19) check. The formula is fine for $t>0$; for $t \rightarrow 0_{+}$both sides are divergent and need regularization.

The reason why such sums do not occur for maps is that for discrete time we work in the variable $z=e^{s}$, an infinite strip along $\operatorname{Im} s$ maps into an anulus in the complex $z$ plane, and the Dirac delta sum in the above is replaced by the Kronecker delta sum in (14.7). In the case at hand there is only one time scale $T$, and we could just as well replace $s$ by the variable $z=e^{-s T}$. In general, a continuous time flow has an infinity of irrationally related cycle periods, and the resonance arrays are more irregular, cf. figure 18.1.

## Commentary

Remark 15.1 Piecewise monotone maps. A partial list of cases for which the transfer operator is well defined: expanding Hölder case, weighted subshifts of finite type, expanding differentiable case, see Bowen [1.16]: expanding holomorphic case, see Ruelle [16.9]; piecewise monotone maps of the interval, see Hofbauer and Keller [15.15] and Baladi and Keller [15.18].

Remark 15.2 Smale's wild idea. Smale's wild idea quoted on page 249 was technically wrong because 1) the Selberg zeta yields the spectrum of a quantum mechanical Laplacian rather than the classical resonances, 2) the spectral determinant weights are different from what Smale conjectured, as the individual cycle weights also depend on the stability of the cycle, 3 ) the formula is not dimensionally correct, as $k$ is an integer and $s$ is dimensionally inverse time. Only for spaces of constant negative curvature do all cycles have the same Lyapunov exponent $\lambda=\ln \left|\Lambda_{p}\right| / T_{p}$. In this case, normalizing the time so that $\lambda=1$, the factors $e^{-s T_{p}} / \Lambda_{p}^{k}$ in (15.9) simplify to $s^{-(s+k) T_{p}}$, as intuited in Smale's quote on page 249 (where $l(\gamma)$ is the cycle period denoted here by $T_{p}$ ). Nevertheless, Smale's intuition was remarkably on the target.

Remark 15.3 Is this a generalization of the Fourier analysis? The Fourier analysis is a theory of the space $\leftrightarrow$ eigenfunctions duality for dynamics on a circle. The sense in which the periodic orbit theory is the generalization of the Fourier analysis to nonlinear flows is discussed in ref. [15.4], a very readable introduction to the Selberg Zeta function.

Remark 15.4 Zeta functions, antecedents. For a function to be deserving of the appellation "zeta function", one expects it to have an Euler product representation (15.15), and perhaps also satisfy a functional equation. Various kinds of zeta functions are reviewed in refs. [15.8, 15.9, 15.10]. Historical antecedents of the dynamical zeta function are the fixed-point counting functions introduced by Weil [15.11], Lefschetz [15.12] and Artin and Mazur [15.13], and the determinants of transfer operators of statistical mechanics [1.17].

In his review article Smale [1.15] already intuited, by analogy to the Selberg Zeta function, that the spectral determinant is the right generalization for continuous time flows. In dynamical systems theory dynamical zeta functions arise naturally only for piecewise linear mappings; for smooth flows the natural object for the study of classical and quantal spectra are the spectral determinants. Ruelle had derived the relation (15.3) between spectral determinants and dynamical zeta functions, but since he was motivated by the Artin-Mazur zeta function (13.21) and the statistical mechanics analogy, he did not consider the spectral determinant a more natural object than the dynamical zeta function. This has been put right in papers on "flat traces" [12.20, 16.24].

The nomenclature has not settled down yet; what we call evolution operators here is elsewhere called transfer operators [1.19], Perron-Frobenius operators [15.6] and/or Ruelle-Araki operators.

Here we refer to kernels such as (10.23) as evolution operators. We follow Ruelle in usage of the term "dynamical zeta function", but elsewhere in the literature the function (15.15) is often called the Ruelle zeta function. Ruelle [1.20] points out that the corresponding transfer operator $T$ was never considered by either Perron or Frobenius; a more appropriate designation would be the Ruelle-Araki operator. Determinants similar to or identical with our spectral determinants are sometimes called Selberg Zetas, Selberg-Smale zetas [1.4], functional determinants, Fredholm determinants, or even - to maximize confusion - dynamical zeta functions [15.14]. A Fredholm determinant is a notion that applies only to trace class operators - as we consider here a somewhat wider class of operators, we prefer to refer to their determinants losely as "spectral determinants".

## Résumé

The eigenvalues of evolution operators are given by the zeros of corresponding determinants, and one way to evaluate determinants is to expand them in terms of traces, using the matrix identity $\log$ det $=$ tr log. Traces of evolution operators can be evaluated as integrals over Dirac delta functions, and in this way the spectra of evolution operators become related to periodic orbits. The spectral problem is now recast into a problem of determining zeros of either the spectral determinant

$$
\operatorname{det}(s-\mathcal{A})=\exp \left(-\sum_{p} \sum_{r=1}^{\infty} \frac{1}{r} \frac{e^{\left(\beta \cdot A_{p}-s T_{p}\right) r}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|}\right)
$$

or the leading zeros of the dynamical zeta function

$$
1 / \zeta=\prod_{p}\left(1-t_{p}\right), \quad t_{p}=\frac{1}{\left|\Lambda_{p}\right|} e^{\beta \cdot A_{p}-s T_{p}}
$$

The spectral determinant is the tool of choice in actual calculations, as it has superior convergence properties (this will be discussed in chapter 16 and is illustrated, for example, by table 18.2). In practice both spectral determinants and dynamical zeta functions are preferable to trace formulas because they yield the eigenvalues more readily; the main difference is that while a trace diverges at an eigenvalue and requires extrapolation methods, determinants vanish at $s$ corresponding to an eigenvalue $s_{\alpha}$, and are analytic in $s$ in an open neighborhood of $s_{\alpha}$.

The critical step in the derivation of the periodic orbit formulas for spectral determinants and dynamical zeta functions is the hyperbolicity assumption (14.5), that is, the requirement that all cycle stability eigenvalues are bounded away from unity, $\left|\Lambda_{p, i}\right| \neq 1$. By dropping the prefactors in (1.4), we have given up on any possibility of recovering the precise distribution of starting $x$ (return to the past is rendered moot by the chaotic mixing and the exponential growth of errors), but in exchange we gain an effective description of the asymptotic behavior of the system. The pleasant surprise (to be demonstrated in chapter 18) is that the infinite time behavior of an unstable system turns out to be as easy to determine as its short time behavior.

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## Exercises

## Exercise 15.1 Escape rate for a 1- $d$ repeller, numerically. Consider

 the quadratic map$$
\begin{equation*}
f(x)=A x(1-x) \tag{15.31}
\end{equation*}
$$

on the unit interval. The trajectory of a point starting in the unit interval either stays in the interval forever or after some iterate leaves the interval and diverges to minus infinity. Estimate numerically the escape rate (19.8), the rate of exponential decay of the measure of points remaining in the unit interval, for either $A=9 / 2$ or $A=6$. Remember to compare your numerical estimate with the solution of the continuation of this exercise, exercise 18.2.

## Exercise 15.2 Dynamical zeta functions. (easy)

(a) Evaluate in closed form the dynamical zeta function

$$
1 / \zeta(z)=\prod_{p}\left(1-\frac{z^{n_{p}}}{\left|\Lambda_{p}\right|}\right),
$$

for the piecewise-linear map (10.17) with the left branch slope $\Lambda_{0}$, the right branch slope $\Lambda_{1}$.

(b) What if there are four different slopes $s_{00}, s_{01}, s_{10}$, and $s_{11}$ instead of just two, with the preimages of the gap adjusted so that junctions of branches $s_{00}, s_{01}$ and $s_{11}, s_{10}$ map in the gap in one iteration? What would the dynamical zeta function be?

Exercise 15.3 Dynamical zeta functions from Markov graphs. Extend sect. 13.3 to evaluation of dynamical zeta functions for piecewise linear maps with finite Markov graphs. This generalizes the results of exercise 15.2.

Exercise 15.4 Zeros of infinite products. Determination of the quantities of interest by periodic orbits involves working with infinite product formulas.
(a) Consider the infinite product

$$
F(z)=\prod_{k=0}^{\infty}\left(1+f_{k}(z)\right)
$$

where the functions $f_{k}$ are "sufficiently nice." This infinite product can be converted into an infinite sum by the use of a logarithm. Use the properties of infinite sums to develop a sensible definition of infinite products.
(b) If $z^{*}$ is a root of the function $F$, show that the infinite product diverges when evaluated at $z^{*}$.
(c) How does one compute a root of a function represented as an infinite product?
(d) Let $p$ be all prime cycles of the binary alphabet $\{0,1\}$. Apply your definition of $F(z)$ to the infinite product

$$
F(z)=\prod_{p}\left(1-\frac{z^{n_{p}}}{\Lambda^{n_{p}}}\right)
$$

(e) Are the roots of the factors in the above product the zeros of $F(z)$ ?

> (Per Rosenqvist)

Exercise 15.5 Dynamical zeta functions as ratios of spectral determinants.
(medium) Show that the zeta function

$$
1 / \zeta(z)=\exp \left(-\sum_{p} \sum_{r=1} \frac{1}{r} \frac{z^{n_{p}}}{\left|\Lambda_{p}\right|^{r}}\right)
$$

can be written as the ratio $1 / \zeta(z)=\operatorname{det}\left(1-z \mathcal{L}_{(0)}\right) / \operatorname{det}\left(1-z \mathcal{L}_{(1)}\right)$, where $\operatorname{det}\left(1-z \mathcal{L}_{(s)}\right)=\prod_{p} \prod_{k=0}^{\infty}\left(1-z^{n_{p}} /\left|\Lambda_{p}\right| \Lambda_{p}^{k+s}\right)$.

Exercise 15.6 Contour integral for survival probability. Perform explicitly the contour integral appearing in (15.19).

Exercise 15.7 Dynamical zeta function for maps. In this problem we will compare the dynamical zeta function and the spectral determinant. Compute the exact dynamical zeta function for the skew Ulam tent map (9.42)

$$
1 / \zeta(z)=\prod_{p \in P}\left(1-\frac{z^{n_{p}}}{\left|\Lambda_{p}\right|}\right) .
$$

What are its roots? Do they agree with those computed in exercise 9.7?

Exercise 15.8 Dynamical zeta functions for Hamiltonian maps. Starting from

$$
1 / \zeta(s)=\exp \left(-\sum_{p} \sum_{r=1}^{\infty} \frac{1}{r} t_{p}^{r}\right)
$$

for a 2-dimensional Hamiltonian map. Using the equality

$$
1=\frac{1}{(1-1 / \Lambda)^{2}}\left(1-2 / \Lambda+1 / \Lambda^{2}\right)
$$

show that

$$
1 / \zeta=\operatorname{det}(1-\mathcal{L}) \operatorname{det}\left(1-\mathcal{L}_{(2)}\right) / \operatorname{det}\left(1-\mathcal{L}_{(1)}\right)^{2} .
$$

In this expression $\operatorname{det}\left(1-z \mathcal{L}_{(k)}\right)$ is the expansion one gets by replacing $t_{p} \rightarrow t_{p} / \Lambda_{p}^{k}$ in the spectral determinant.

Exercise 15.9 Riemann $\zeta$ function. The Riemann $\zeta$ function is defined as the sum

$$
\zeta(s)=\sum_{n=1}^{\infty} \frac{1}{n^{s}}, \quad s \in \mathbb{C}
$$

(a) Use factorization into primes to derive the Euler product representation

$$
\zeta(s)=\prod_{p} \frac{1}{1-p^{-s}} .
$$

The dynamical zeta function exercise 15.15 is called a "zeta" function because it shares the form of the Euler product representation with the Riemann zeta function.
(b) (Not trivial:) For which complex values of $s$ is the Riemann zeta sum convergent?
(c) Are the zeros of the terms in the product, $s=-\ln p$, also the zeros of the Riemann $\zeta$ function? If not, why not?

Exercise 15.10 Finite truncations. (easy) Suppose we have a one-dimensional system with complete binary dynamics, where the stability of each orbit is given by a simple multiplicative rule:

$$
\Lambda_{p}=\Lambda_{0}^{n_{p, 0}} \Lambda_{1}^{n_{p, 1}}, \quad n_{p, 0}=\# 0 \text { in } p, n_{p, 1}=\# 1 \text { in } p
$$

so that, for example, $\Lambda_{00101}=\Lambda_{0}^{3} \Lambda_{1}^{2}$.
(a) Compute the dynamical zeta function for this system; perhaps by creating a transfer matrix analogous to (10.19), with the right weights.
(b) Compute the finite $p$ truncations of the cycle expansion, that is take the product only over the $p$ up to given length with $n_{p} \leq N$, and expand as a series in $z$

$$
\prod_{p}\left(1-\frac{z^{n_{p}}}{\left|\Lambda_{p}\right|}\right)
$$

Do they agree? If not, how does the disagreement depend on the truncation length $N$ ?

## Chapter 16

## Why does it work?

Bloch: "Space is the field of linear operators." Heisenberg: "Nonsense, space is blue and birds fly through it."
Felix Bloch, Heisenberg and the early days of quantum mechanics
(R. Artuso, H.H. Rugh and P. Cvitanović)

The trace formulas and spectral determinants work well, sometimes very well indeed. The question is: why? The heuristic manipulations of chapter 14 were naive and reckless, as we are facing infinite-dimensional vector spaces and singular integral kernels.

We now outline the key ingredients of proofs that put the trace and determinant formulas on solid mathematical footing. This requires taking a closer look at the evolution operators from a mathematical point of view, since up to now we have talked about eigenvalues without any reference to what kind of a function space the corresponding eigefunctions belong to. We shall restrict our considerations to the spectral properties of the PerronFrobenius operator for maps, as proofs for more general evolution operators follow along the same lines. What we refer to as a "the set of eigenvalues" acquires meaning only within a precisely specified functional setting: this sets the stage for a discussion of analyticity properties of spectral determinants. In sect. 16.1 we show on a very simple example that the spectrum is quite sensitive to the regularity properties of the functions considered. In sect. 16.3 we review the basic facts of the classical Fredholm theory of integral equations. The program is sketched in sect. 16.4, motivated by an explicit study of eigenspectrum of the Bernoulli shift map, and in sect. 16.5 generalized to piecewise real-analytic hyperbolic maps acting on appropriate densities. For expanding and hyperbolic finite-subshift maps analyticity leads to a very strong result; not only do the determinants have better analyticity properties than the trace formulas, but the spectral determinants are singled out as being entire functions in the complex $s$ plane.

The goal of this chapter is not to provide an exhaustive review of the
rigorous theory of the Perron-Frobenius operators and their spectral determinants, but rather to give you a feeling for how our heuristic considerations can be put on a firm basis. The mathematics underpinning the theory is both hard and profound.

If you are primarily interested in applications of periodic orbit theory, you should skip this chapter on the first reading.


### 16.1 Linear maps: exact spectra

We start gently; in example 16.1 we work out the exact eigenvalues and eigenfunctions of the Perron-Frobenius operator for the simplest example of unstable (expanding) dynamics, a linear 1- $d$ map with one unstable fixed point. Not only that, but in example 16.5 we compute the exact spectrum for the simplest example of a dynamical system with an infinity of unstable periodic orbits, the Bernoulli shift. In example 16.7 we show that these results carry over to any analytic single-branch $1-d$ repeller, and motivate by these examples the steps that lead to a proof that spectral determinants for more general analytic 1-dimensional expanding maps, and - in sect. 16.5, for 2-dimensional hyperbolic mappings - are also entire functions.

Example 16.1 The simplest of eigenspectra - a single fixed point: In order to get some feeling for the determinants defined so formally in sect. 15.2, let us work out a trivial example: a repeller with only one expanding linear branch

$$
f(x)=\Lambda x, \quad|\Lambda|>1,
$$

and only one fixed point $x^{*}=0$. The action of the Perron-Frobenius operator (9.10) is

$$
\begin{equation*}
\mathcal{L} \phi(y)=\int d x \delta(y-\Lambda x) \phi(x)=\frac{1}{|\Lambda|} \phi(y / \Lambda) . \tag{16.1}
\end{equation*}
$$

From this one immediately gets that the monomials $y^{k}$ are eigenfunctions:

$$
\begin{equation*}
\mathcal{L} y^{k}=\frac{1}{|\Lambda| \Lambda^{k}} y^{k}, \quad k=0,1,2, \ldots \tag{16.2}
\end{equation*}
$$

What are these eigenfunctions? Think of eigenfunctions of the Schrödinger equation: $k$ labels the $k$ th eigenfunction $x^{k}$ in the same spirit in which the number of nodes labels the $k$ th quantum-mechanical eigenfunction. A quantum-mechanical amplitude with more nodes has more variability, hence a higher kinetic energy. Analogously, for a Perron-Frobenius operator, a higher $k$ eigenvalue $1 /|\Lambda| \Lambda^{k}$ is getting exponentially smaller because densities that vary more rapidly decay more rapidly under the expanding action of the map.

Example 16.2 The trace formula for a single fixed point: The eigenvalues $\Lambda^{-k-1}$ fall off exponentially with $k$, so the trace of $\mathcal{L}$ is a convergent sum

$$
\operatorname{tr} \mathcal{L}=\frac{1}{|\Lambda|} \sum_{k=0}^{\infty} \Lambda^{-k}=\frac{1}{|\Lambda|\left(1-\Lambda^{-1}\right)}=\frac{1}{\left|f(0)^{\prime}-1\right|}
$$

in agreement with (14.6). A similar result follows for powers of $\mathcal{L}$, yielding the singlefixed point version of the trace formula for maps (14.9):

$$
\begin{equation*}
\sum_{k=0}^{\infty} \frac{z e^{s_{k}}}{1-z e^{s_{k}}}=\sum_{r=1}^{\infty} \frac{z^{r}}{\left|1-\Lambda^{r}\right|}, \quad e^{s_{k}}=\frac{1}{|\Lambda| \Lambda^{k}} \tag{16.3}
\end{equation*}
$$

The left hand side of (16.3) is a meromorphic function, with the leading zero at $z=|\Lambda|$.

Example 16.3 Meromorphic functions and exponential convergence: As an illustration of how exponential convergence of a truncated series is related to analytic properties of functions, consider, as the simplest possible example of a meromorphic function, the ratio

$$
h(z)=\frac{z-a}{z-b}
$$

with $a, b$ real and positive and $a<b$. Within the spectral radius $|z|<b$ the function $h$ can be represented by the power series

$$
h(z)=\sum_{k=0}^{\infty} \sigma_{k} z^{k}
$$

where $\sigma_{0}=a / b$, and the higher order coefficients are given by $\sigma_{j}=(a-b) / b^{j+1}$. Consider now the truncation of order $N$ of the power series

$$
h_{N}(z)=\sum_{k=0}^{N} \sigma_{k} z^{k}=\frac{a}{b}+\frac{z(a-b)\left(1-z^{N} / b^{N}\right)}{b^{2}(1-z / b)} .
$$

Let $\hat{z}_{N}$ be the solution of the truncated series $h_{N}\left(\hat{z}_{N}\right)=0$. To estimate the distance between $a$ and $\hat{z}_{N}$ it is sufficient to calculate $h_{N}(a)$. It is of order $(a / b)^{N+1}$, so finite order estimates converge exponentially to the asymptotic value.

This example shows that: (1) an estimate of the leading pole (the leading eigenvalue of $\mathcal{L}$ ) from a finite truncation of a trace formula converges exponentially, and (2) the non-leading eigenvalues of $\mathcal{L}$ lie outside of the radius of convergence of the trace formula and cannot be computed by means of such cycle expansion. However, as we shall now see, the whole spectrum is reachable at no extra effort, by computing it from a determinant rather than a trace.

Example 16.4 The spectral determinant for a single fixed point: The spectral determinant (15.3) follows from the trace formulas of example 16.2:

$$
\begin{equation*}
\operatorname{det}(1-z \mathcal{L})=\prod_{k=0}^{\infty}\left(1-\frac{z}{|\Lambda| \Lambda^{k}}\right)=\sum_{n=0}^{\infty}(-t)^{n} Q_{n}, \quad t=\frac{z}{|\Lambda|} \tag{16.4}
\end{equation*}
$$

where the cummulants $Q_{n}$ are given explicitly by the Euler formula

$$
\begin{equation*}
Q_{n}=\frac{1}{1-\Lambda^{-1}} \frac{\Lambda^{-1}}{1-\Lambda^{-2}} \cdots \frac{\Lambda^{-n+1}}{1-\Lambda^{-n}} \tag{16.5}
\end{equation*}
$$

(If you cannot figure out how to derive this formula, the solutions on p. 801 offer several proofs.)

The main lesson to glean from this simple example is that the cummulants $Q_{n}$ decay asymptotically faster than exponentially, as $\Lambda^{-n(n-1) / 2}$. For example, if we approximate series such as (16.4) by the first 10 terms, the error in the estimate of the leading zero is $\approx 1 / \Lambda^{50}$ !

So far all is well for a rather boring example, a dynamical system with a single repelling fixed point. What about chaos? Systems where the number of unstable cycles increases exponentially with their length? We now turn to the simplest example of a dynamical system with an infinity of unstable periodic orbits, in order to motivate the steps that will eventually lead to a proof that spectral determinants for more general nonlinear 1-dimensional expanding maps are entire functions.

Example 16.5 Bernoulli shift: Consider next the Bernoulli shift map

$$
\begin{equation*}
x \mapsto 2 x(\bmod 1), \quad x \in[0,1] . \tag{16.6}
\end{equation*}
$$

The associated Perron-Frobenius operator (9.9) assambles $\rho(y)$ from its two preimages

$$
\begin{equation*}
\mathcal{L} \rho(y)=\frac{1}{2} \rho\left(\frac{y}{2}\right)+\frac{1}{2} \rho\left(\frac{y+1}{2}\right) . \tag{16.7}
\end{equation*}
$$

For this simple example the eigenfunctions can be written down explicitly: they coincide, up to constant prefactors, with the Bernoulli polynomials $B_{n}(x)$. These polynomials are generated by the Taylor expansion of

$$
\mathcal{G}_{t}(x)=\frac{t e^{x t}}{e^{t}-1}=\sum_{k=0}^{\infty} B_{k}(x) \frac{t^{k}}{k!}, \quad B_{0}(x)=1, \quad B_{1}(x)=x-\frac{1}{2}, \ldots
$$

The Perron-Frobenius operator (16.7) acts on the generating function $\mathcal{G}$ as

$$
\mathcal{L G}_{t}(x)=\frac{1}{2}\left(\frac{t e^{x t / 2}}{e^{t}-1}+\frac{t e^{t / 2} e^{x t / 2}}{e^{t}-1}\right)=\frac{t / 2 e^{x t / 2}}{e^{t / 2}-1}=\sum_{k=1}^{\infty} B_{k}(x) \frac{(t / 2)^{k}}{k!},
$$

hence each $B_{k}(x)$ is an eigenfunction of $\mathcal{L}$ with eigenvalue $1 / 2^{k}$.
The full operator has two components corresponding to the two branches. For the $n$ times iterated operator we have a full binary shift, and for each of the $2^{n}$ branches the above calculations carry over, yielding the same trace $\left(2^{n}-1\right)^{-1}$ for every cycle on length $n$. Without further ado we substitute everything back and obtain the determinant,

$$
\begin{equation*}
\operatorname{det}(1-z \mathcal{L})=\exp \left(-\sum_{n=1} \frac{z^{n}}{n} \frac{2^{n}}{2^{n}-1}\right)=\prod_{k=0}\left(1-\frac{z}{2^{k}}\right) \tag{16.8}
\end{equation*}
$$

verifying that the Bernoulli polynomials are eigenfunctions with eigenvalues $1,1 / 2, \ldots$, $1 / 2^{n}, \ldots$.

The Bernoulli map spectrum looks reminiscent of the single fixed-point spectrum (16.2), with the difference that the leading eigenvalue here is 1 , rather than $1 /|\Lambda|$. The difference is significant: the single fixed-point map is a repeller, with escape rate (1.6) given by the $\mathcal{L}$ leading eigenvalue $\gamma=\ln |\Lambda|$, while there is no escape in the case of the Bernoulli map. As already noted in discussion of the relation (15.23), for bound systems the local expansion rate (here $\ln |\Lambda|=\ln 2$ ) is balanced by the entropy (here $\ln 2$, the $\log$ of the number of preimages $F_{s}$ ), yielding zero escape rate.

So far we have demonstrated that our periodic orbit formulas are correct for two piecewise linear maps in 1 dimension, one with a single fixed point, and one with a full binary shift chaotic dynamics. What about higher dimensions? We check our formulas on a $2-d$ hyperbolic map next.

Example 16.6 The simplest of 2- $d$ maps - a single hyperbolic fixed point: We start by considering a very simple linear hyperbolic map with a single hyperbolic fixed point,

$$
f(x)=\left(f_{1}\left(x_{1}, x_{2}\right), f_{2}\left(x_{1}, x_{2}\right)\right)=\left(\Lambda_{s} x_{1}, \Lambda_{u} x_{2}\right), \quad 0<\left|\Lambda_{s}\right|<1, \quad\left|\Lambda_{u}\right|>1 .
$$

The Perron-Frobenius operator (9.10) acts on the 2-d density functions as

$$
\begin{equation*}
\mathcal{L} \rho\left(x_{1}, x_{2}\right)=\frac{1}{\left|\Lambda_{s} \Lambda_{u}\right|} \rho\left(x_{1} / \Lambda_{s}, x_{2} / \Lambda_{u}\right) \tag{16.9}
\end{equation*}
$$

What are good eigenfunctions? Cribbing the 1-d eigenfunctions for the stable, contracting $x_{1}$ direction from example 16.1 is not a good idea, as under the iteration of $\mathcal{L}$ the high terms in a Taylor expansion of $\rho\left(x_{1}, x_{2}\right)$ in the $x_{1}$ variable would get multiplied by exponentially exploding eigenvalues $1 / \Lambda_{s}^{k}$. This makes sense, as in the contracting directions hyperbolic dynamics crunches up initial densities, rather then smoothing them. So we guess instead that eigenfunctions are of form

$$
\begin{equation*}
\varphi_{k_{1} k_{2}}\left(x_{1}, x_{2}\right)=x_{2}^{k_{2}} / x_{1}^{k_{1}+1}, \quad k_{1}, k_{2}=0,1,2, \ldots, \tag{16.10}
\end{equation*}
$$

a mixture of the Laurent series in the contraction direction, $x_{1}$ variable, and the Taylor series in the expanding direction, the $x_{2}$ variable. The action of Perron-Frobenius operator on this set of basis functions

$$
\mathcal{L} \varphi_{k_{1} k_{2}}\left(x_{1}, x_{2}\right)=\frac{1}{\left|\Lambda_{u}\right|} \frac{\Lambda_{s}^{k_{1}}}{\Lambda_{u}^{k_{2}}} \varphi_{k_{1} k_{2}}\left(x_{1}, x_{2}\right)
$$

is smoothing, with the higher $k_{1}, k_{2}$ eigenvectors decaying exponentially faster, by $\Lambda_{s}^{k_{1}} / \Lambda_{u}^{k_{2}+1}$ factor in the eigenvalue. One verifies by an explicit calculation (undoing the geometric series expansions to lead to (15.9)) that the trace of $\mathcal{L}$ indeed equals $1 /|\operatorname{det}(\mathbf{1}-\mathbf{J})|=1 /\left|\left(1-\Lambda_{u}\right)\left(1-\Lambda_{s}\right)\right|$, from which it follows that all our trace and spectral determinant formulas apply.

So far we have checked the trace and spectral determinant formulas derived heuristically in chapters 14 and 15 , but only for the case of 1 - and $2-d$ linear maps. But for infinite-dimensional vector spaces this game is fraught with dangers, and we have already been mislead by piecewise linear examples into spectral confusions: contrast the spectra of example 9.1 and example 10.1 with the spectrum computed in sect. 14.1.

We show next that the above results do carry over to a sizable class of piecewise analytic expanding maps.

### 16.2 Evolution operator in a matrix representation

The standard, and for numerical purposes sometimes very effective way to look at operators is through their matrix representations. Evolution operators are moving density functions defined over some phase space, and as in general we can implement this only numerically, the temptation is to discretize the phase space as in sect. 10.4. The problem with such phase space discretization approaches that they sometimes yield plainly wrong spectra (compare example 10.1 with the result of sect. 14.1), so we have to think through carefully what is it that we really measure.

An expanding map $f(x)$ takes an initial smooth density $\phi_{n}(x)$, defined on a subinterval, stretches it out and overlays it over a larger interval, resulting in a new, smoother density $\phi_{n+1}(x)$. Repetition of this process smoothes the initial density, so it is natural to represent densities $\phi_{n}(x)$ by their Taylor series. By expanding both

$$
\begin{aligned}
& \phi_{n}(x)=\sum_{k=0}^{\infty} \phi_{n}^{(k)}(0) \frac{y^{k}}{k!}, \quad \phi_{n+1}(y)_{k}=\sum_{\ell=0}^{\infty} \phi_{n+1}^{(\ell)}(0) \frac{y^{\ell}}{\ell!}, \quad \text { and } \\
& \phi_{n+1}^{(\ell)}(0)=\left.\int d x \delta^{(\ell)}(0-f(x)) \phi_{n}(x)\right|_{y=0}, \quad x=f^{-1}(0),
\end{aligned}
$$

and substituting the two Taylor series into (9.6)

$$
\phi_{n+1}(y)=\left(\mathcal{L} \phi_{n}\right)(y)=\int_{\mathcal{M}} d x \delta(y-f(x)) \phi_{n}(x) .
$$

The matrix elements follow by evaluating the integral

$$
\begin{equation*}
\mathbf{L}_{\ell k}=\left.\frac{\partial^{\ell}}{\partial y^{\ell}} \int d x \mathcal{L}(y, x) \frac{x^{\ell}}{\ell!}\right|_{x=0} . \tag{16.11}
\end{equation*}
$$

we obtain a matrix representation of the evolution operator

$$
\int d x \mathcal{L}(y, x) \frac{x^{k}}{k!}=\sum_{k^{\prime}} \frac{y^{k^{\prime}}}{k^{\prime}!} \mathbf{L}_{k^{\prime} k}, \quad k, k^{\prime}=0,1,2, \ldots
$$

which maps the $x^{k}$ component of the density of trajectories $\phi_{n}(x)$ into the $y^{k^{\prime}}$ component of the density $\phi_{n+1}(y)$ one time step later, with $y=f(x)$.

We already have some practice with evaluating derivatives $\delta^{(\ell)}(y)=$ $\frac{\partial^{l}}{\partial y^{\ell}} \delta(y)$ from (??). This yields a representation of the evolution operator
centered on the fixed point, evaluated recursively in terms of derivatives of the map $f$ :

$$
\begin{align*}
(\mathbf{L})_{\ell k} & =\left.\int d x \delta^{(\ell)}(x-f(x)) \frac{x^{k}}{k!}\right|_{x=0}  \tag{16.12}\\
& =\left.\frac{1}{\left|f^{\prime}\right|}\left(\frac{d}{d x} \frac{1}{f^{\prime}(x)}\right)^{\ell} \frac{x^{k}}{k!}\right|_{x=0}
\end{align*}
$$

The matrix elements vanish for $\ell<k$, so $\mathbf{L}$ is a lower triangular matrix. The diagonal and the successive off-diagonal matrix elements are easily evaluated iteratively by computer algebra

$$
\mathbf{L}_{k k}=\frac{1}{|\Lambda|(\Lambda)^{k}}, \quad \mathbf{L}_{k+1, k}=-\frac{(k+2)!f^{\prime \prime}}{2 k!|\Lambda|(\Lambda)^{k+2}}, \quad \cdots
$$

For chaotic systems the map is expanding, $|\Lambda|>1$. Hence the diagonal terms drop off exponentially, as $1 /|\Lambda|^{k+1}$, the terms below the diagonal fall off even faster, and truncating $\mathbf{L}$ to a finite matrix introduces only exponentially small errors.

The trace formula (16.3) takes now a matrix form

$$
\begin{equation*}
\operatorname{tr} \frac{z \mathcal{L}}{1-z \mathcal{L}}=\operatorname{tr} \frac{\mathbf{L}}{1-z \mathbf{L}} . \tag{16.13}
\end{equation*}
$$

In order to illustrate how this works, we work out a simple example.

Example 16.7 Perron-Frobenius operator in a matrix representation: As in example 16.1, we start with a map with a single fixed point, but this time with a nonlinear piecewise analytic map $f$ with a nonlinear inverse $F=f^{-1}$, sign of the derivative $\sigma=\sigma\left(F^{\prime}\right)=F^{\prime} /\left|F^{\prime}\right|$, and the Perron-Frobenius operator acting on densities analytic in an open domain enclosing the fixed point $x=w^{*}$,

$$
\mathcal{L} \phi(y)=\int d x \delta(y-f(x)) \phi(x)=\sigma F^{\prime}(y) \phi(F(y))
$$

Assume that $F$ is a contraction of the unit disk in the complex plane, that is,

$$
\begin{equation*}
|F(z)|<\theta<1 \quad \text { and } \quad\left|F^{\prime}(z)\right|<C<\infty \quad \text { for } \quad|z|<1 \tag{16.14}
\end{equation*}
$$

and expand $\phi$ in a polynomial basis by means of the Cauchy formula

$$
\phi(z)=\sum_{n \geq 0} z^{n} \phi_{n}=\oint \frac{d w}{2 \pi i} \frac{\phi(w)}{w-z}, \quad \phi_{n}=\oint \frac{d w}{2 \pi i} \frac{\phi(w)}{w^{n+1}}
$$

Combining this with (16.22), we see that in this basis Perron-Frobenius operator $\mathcal{L}$ is represented by the matrix

$$
\begin{equation*}
\mathcal{L} \phi(w)=\sum_{m, n} w^{m} L_{m n} \phi_{n}, \quad L_{m n}=\oint \frac{d w}{2 \pi i} \frac{\sigma F^{\prime}(w)(F(w))^{n}}{w^{m+1}} \tag{16.15}
\end{equation*}
$$

Figure 16.1: A nonlinear one-branch repeller with a single fixed point $w^{*}$.


Taking the trace and summing we get:

$$
\operatorname{tr} \mathcal{L}=\sum_{n \geq 0} L_{n n}=\oint \frac{d w}{2 \pi i} \frac{\sigma F^{\prime}(w)}{w-F(w)} .
$$

This integral has but one simple pole at the unique fixed point $w^{*}=F\left(w^{*}\right)=f\left(w^{*}\right)$. Hence

$$
\operatorname{tr} \mathcal{L}=\frac{\sigma F^{\prime}\left(w^{*}\right)}{1-F^{\prime}\left(w^{*}\right)}=\frac{1}{\left|f^{\prime}\left(w^{*}\right)-1\right|}
$$

This super-exponential decay of cummulants $Q_{k}$ ensures that for a repeller consisting of a single repelling point the spectral determinant (16.4) is entire in the complex $z$ plane.

In retrospect, the matrix representation method for solving the density evolution problems is eminently sensible - after all, that is the way one solves a close relative to classical density evolution equations, the Schrödinger equation. When available, matrix representations for $\mathcal{L}$ enable us to compute many more orders of cumulant expansions of spectral determinants and many more eigenvalues of evolution operators than the cycle expensions approach.

Now, if the spectral determinant is entire, formulas such as (15.25) imply that the dynamical zeta function is a meromorphic function. The practical import of this observation is that it guarantees that finite order estimates of zeroes of dynamical zeta functions and spectral determinants converge exponentially, or - in cases such as (16.4) - super-exponentially to the exact values, and so the cycle expansions to be discussed in chapter 18 represent a true perturbative approach to chaotic dynamics.

Before turning to specifics we summarize a few facts about classical theory of integral equations, something you might prefer to skip on first reading. The purpose of this exercise is to understand that the Fredholm theory, a theory that works so well for the Hilbert spaces of quantum mechanics does not necessarily work for deterministic dynamics - the ergodic theory is much harder.

### 16.3 Classical Fredholm theory

He who would valiant be<br>'Gainst all disaster<br>Let him in constancy<br>Follow the Master.<br>John Bunyan, Pilgrim's Progress

$$
\mathcal{L} \phi(x)=\int d y \delta(x-f(y)) \phi(y)
$$

has the same appearance as a classical Fredholm integral operator

$$
\begin{equation*}
\mathcal{K} \varphi(x)=\int_{\mathcal{M}} d y \mathcal{K}(x, y) \varphi(y) \tag{16.16}
\end{equation*}
$$

and one is tempted to resort to the classical Fredholm theory in order to establish analyticity properties of spectral determinants. This path to enlightenment is blocked by the singular nature of the kernel, which is a distribution, whereas the standard theory of integral equations usually concerns itself with regular kernels $\mathcal{K}(x, y) \in L^{2}\left(\mathcal{M}^{2}\right)$. Here we briefly recall some steps of the Fredholm theory, before working out the example of example 16.5.

The general form of Fredholm integral equations of the second kind is

$$
\begin{equation*}
\varphi(x)=\int_{\mathcal{M}} d y \mathcal{K}(x, y) \varphi(y)+\xi(x) \tag{16.17}
\end{equation*}
$$

where $\xi(x)$ is a given function in $L^{2}(\mathcal{M})$ and the kernel $\mathcal{K}(x, y) \in L^{2}\left(\mathcal{M}^{2}\right)$ (Hilbert-Schmidt condition). The natural object to study is then the linear integral operator (16.16), acting on the Hilbert space $L^{2}(\mathcal{M})$ : and the fundamental property that follows from the $L^{2}(Q)$ nature of the kernel is that such an operator is compact, that is close to a finite rank operator (see appendix K). A compact operator has the property that for every $\delta>0$ only a finite number of linearly independent eigenvectors exist corresponding to eigenvalues whose absolute value exceeds $\delta$, so we immediately realize (figure 16.4) that much work is needed to bring Perron-Frobenius operators into this picture.

We rewrite (16.17) in the form

$$
\begin{equation*}
\mathcal{T} \varphi=\xi, \quad \mathcal{T}=\mathbb{1}-\mathcal{K} . \tag{16.18}
\end{equation*}
$$

The Fredholm alternative is now stated as follows: the equation $\mathcal{T} \varphi=\xi$ as a unique solution for every $\xi \in L^{2}(\mathcal{M})$ or there exists a non-zero solution of $\mathcal{T} \varphi_{0}=0$, with an eigenvector of $\mathcal{K}$ corresponding to the eigenvalue 1 .

The theory remains the same if instead of $\mathcal{T}$ we consider the operator $\mathcal{T}_{\lambda}=\mathbb{1}-\lambda \mathcal{K}$ with $\lambda \neq 0$. As $\mathcal{K}$ is a compact operator there will be at most a denumerable set of $\lambda$ for which the second part of Fredholm alternative holds: so apart from this set the inverse operator $(\mathbb{1}-\lambda \mathcal{T})^{-1}$ exists and is a bounded operator. When $\lambda$ is sufficiently small we may look for a perturbative expression for such an inverse, as a geometric series

$$
\begin{equation*}
(\mathbb{1}-\lambda \mathcal{K})^{-1}=\mathbb{1}+\lambda \mathcal{K}+\lambda^{2} \mathcal{K}^{2}+\cdots=\mathbb{1}+\lambda \mathcal{W}, \tag{16.19}
\end{equation*}
$$

where each $\mathcal{K}^{n}$ is still a compact integral operator with kernel

$$
\mathcal{K}^{n}(x, y)=\int_{\mathcal{M}^{n-1}} d z_{1} \ldots d z_{n-1} \mathcal{K}\left(x, z_{1}\right) \cdots \mathcal{K}\left(z_{n-1}, y\right)
$$

and $\mathcal{W}$ is also compact, as it is given by the convergent sum of compact operators. The problem with (16.19) is that the series has a finite radius of convergence, while apart from a denumerable set of $\lambda$ 's the inverse operator is well defined. A fundamental result in the theory of integral equations consists in rewriting the resolving kernel $\mathcal{W}$ as a ratio of two analytic functions of $\lambda$

$$
\mathcal{W}(x, y)=\frac{\mathcal{D}(x, y ; \lambda)}{D(\lambda)}
$$

If we introduce the notation

$$
\mathcal{K}\binom{x_{1} \ldots x_{n}}{y_{1} \ldots y_{n}}=\left|\begin{array}{ccc}
\mathcal{K}\left(x_{1}, y_{1}\right) & \ldots & \mathcal{K}\left(x_{1}, y_{n}\right) \\
\ldots & \ldots & \ldots \\
\mathcal{K}\left(x_{n}, y_{1}\right) & \ldots & \mathcal{K}\left(x_{n}, y_{n}\right)
\end{array}\right|
$$

we may write the explicit expressions

$$
\begin{align*}
D(\lambda) & =1+\sum_{n=1}^{\infty}(-1)^{n} \frac{\lambda^{n}}{n!} \int_{\mathcal{M}^{n}} d z_{1} \ldots d z_{n} \mathcal{K}\binom{z_{1} \ldots z_{n}}{z_{1} \ldots z_{n}} \\
& =\exp \left(-\sum_{m=1}^{\infty} \frac{\lambda^{m}}{m} \operatorname{tr} \mathcal{K}^{m}\right)  \tag{16.20}\\
\mathcal{D}(x, y ; \lambda) & =\mathcal{K}\binom{x}{y}+\sum_{n=1}^{\infty} \frac{(-\lambda)^{n}}{n!} \int_{\mathcal{M}^{n}} d z_{1} \ldots d z_{n} \mathcal{K}\left(\begin{array}{llll}
x & z_{1} & \ldots & z_{n} \\
y & z_{1} & \ldots & z_{n}
\end{array}\right)
\end{align*}
$$

$D(\lambda)$ is known as the Fredholm determinant (see (15.24) and appendix K): it is an entire analytic function of $\lambda$, and $D(\lambda)=0$ only if $1 / \lambda$ is an eigenvalue of $\mathcal{K}$.

Worth emphasizing again: the Fredholm theory is based on the compactness of the integral operator, that is, on the functional properties (summability) of its kernel.

### 16.4 Analyticity of spectral determinants


#### Abstract

They savored the strange warm glow of being much more ignorant than ordinary people, who were only ignorant of ordinary things. Terry Pratchett


Spaces of functions integrable $L^{1}$, or square-integrable $L^{2}$ on interval $[0,1]$ are mapped into themselves by the Perron-Frobenius operator, and in both cases the constant function $\phi_{0} \equiv 1$ is an eigenfunction with eigenvalue 1. If we focus our attention on $L^{1}$ we also have a family of $L^{1}$ eigenfunctions,

$$
\begin{equation*}
\phi_{\theta}(y)=\sum_{k \neq 0} \exp (2 \pi i k y) \frac{1}{|k|^{\theta}} \tag{16.21}
\end{equation*}
$$

with complex eigenvalue $2^{-\theta}$, parametrized by complex $\theta$ with $\operatorname{Re} \theta>0$. By varying $\theta$ one realizes that such eigenvalues fill out the entire unit disk. Such essential spectrum, the case $k=0$ of figure 16.4, hides all fine details of the spectrum.

What's going on? Spaces $L^{1}$ and $L^{2}$ contain arbitrarily ugly functions, allowing any singularity as long as it is (square) integrable - and there is no way that expanding dynamics can smooth a kinky function with a nondifferentiable singularity, let's say a discontinuous step, and that is why the eigenspectrum is dense rather than discrete. Mathematicians love to wallow in this kind of muck, but there is no way for an experimental physicist to prepare a nowhere differentiable $L^{1}$ initial density. The only thing we can prepare and measure are piecewise smooth (real-analytic) density functions.

For a bounded linear operator $\mathcal{A}$ on a Banach space $\boldsymbol{\Omega}$, the spectral radius is the smallest positive number $\rho_{\text {spec }}$ such the spectrum is inside the disk of radius $\rho_{\text {spec }}$, while the essential spectral radius is the smallest positive number $\rho_{\text {ess }}$ such that outside the disk of radius $\rho_{\text {ess }}$ the spectrum consists only of isolated eigenvalues of finite multiplicity (see figure 16.4).

We may shrink the essential spectrum by letting the Perron-Frobenius operator act on a space of smoother functions, exactly as in the one-branch repeller case of sect. 16.1 . We thus consider a smaller space, $\mathbb{C}^{k+\alpha}$, the space of $k$ times differentiable functions whose $k$ 'th derivatives are Hölder continuous with an exponent $0<\alpha \leq 1$ : the expansion property guarantees that such a space is mapped into itself by the Perron-Frobenius operator. In the strip $0<\operatorname{Re} \theta<k+\alpha$ most $\phi_{\theta}$ will cease to be eigenfunctions in the space $\mathbb{C}^{k+\alpha}$; the function $\phi_{n}$ survives only for integer valued $\theta=n$. In this way we arrive at a finite set of isolated eigenvalues $1,2^{-1}, \cdots, 2^{-k}$, and an essential spectral radius $\rho_{\text {ess }}=2^{-(k+\alpha)}$.

We follow a simpler path and restrict the function space even further, namely to a space of analytic functions, that is, functions for which the

Taylor expansion is convergent at each point of the interval $[0,1]$. With this choice things turn out easy and elegant. To be more specific let $\phi$ be a holomorphic and bounded function on the disk $D=B(0, R)$ of radius $R>0$ centered at the origin. Our Perron-Frobenius operator preserves the space of such functions provided $(1+R) / 2<R$ so all we need is to choose $R>1$. If $F_{s}, s \in\{0,1\}$, denotes the $s$ inverse branch of the Bernoulli shift (16.6), the corresponding action of the Perron-Frobenius operator is given by $\mathcal{L}_{s} h(y)=\sigma F_{s}^{\prime}(y) h \circ F_{s}(y)$, using the Cauchy integral formula along the $\partial D$ boundary contour:

$$
\begin{equation*}
\mathcal{L}_{s} h(y)=\sigma \oint \frac{d w}{2 \pi i} \partial D \frac{h(w) F_{s}^{\prime}(y)}{w-F_{s}(y)} . \tag{16.22}
\end{equation*}
$$

For reasons that will be made clear later we have introduced a sign $\sigma=$ $\pm 1$ of the given real branch $\left|F^{\prime}(y)\right|=\sigma F^{\prime}(y)$. For both branches of the Bernoulli shift $s=1$, but in general one is not allowed to take absolute values as this could destroy analyticity. In the above formula one may also replace the domain $D$ by any domain containing $[0,1]$ such that the inverse branches maps the closure of $D$ into the interior of $D$. Why? simply because the kernel stays non-singular under this condition, that is, $w-F(y) \neq 0$ whenever $w \in \partial D$ and $y \in \mathrm{Cl} D$.

The problem is now reduced to the standard theory for Fredholm determinants, sect. 16.3. The integral kernel is no longer singular, traces and determinants are well-defined, and we can evaluate the trace of $\mathcal{L}_{F}$ by means of the Cauchy contour integral formula:

$$
\operatorname{tr} \mathcal{L}_{F}=\oint \frac{d w}{2 \pi i} \frac{\sigma F^{\prime}(w)}{w-F(w)}
$$

Elementary complex analysis shows that since $F$ maps the closure of $D$ into its own interior, $F$ has a unique (real-valued) fixed point $x^{*}$ with a multiplier strictly smaller than one in absolute value. Residue calculus therefore yields

$$
\operatorname{tr} \mathcal{L}_{F}=\frac{\sigma F^{\prime}\left(x^{*}\right)}{1-F^{\prime}\left(x^{*}\right)}=\frac{1}{\left|f^{\prime}\left(x^{*}\right)-1\right|},
$$

justifying our previous ad hoc calculations of traces by means of Dirac delta functions. In example 16.7 we have shown how this works for a nonlinear piecewise analytic map with a single fixed point.

We worked out a very specific example, yet our conclusions can be generalized, provided a number of restrictive requirements are met by the dynamical system under investigation:

1) the evolution operator is multiplicative along the flow,
2) the symbolic dynamics is a finite subshift,
3) all cycle eigenvalues are hyperbolic (exponentially bounded away from 1),
4) the map (or the flow) is real analytic, that is, it has a piecewise analytic continuation to a complex extension of the phase space.

These assumptions are romantic projections not lived up to by the dynamical systems that we actually desire to understand. Still, they are not devoid of physical interest; for example, nice repellers like our 3-disk game of pinball do satisfy the above requirements.

Properties 1 and 2 enable us to represent the evolution operator as a matrix in an appropriate basis space; properties 3 and 4 enable us to bound the size of the matrix elements and control the eigenvalues. To see what can go wrong consider the following examples:

Property 1 is violated for flows in 3 or more dimensions by the following weighted evolution operator

$$
\mathcal{L}^{t}(y, x)=\left|\Lambda^{t}(x)\right|^{\beta} \delta\left(y-f^{t}(x)\right)
$$

where $\Lambda^{t}(x)$ is an eigenvalue of the Jacobian matrix transverse to the flow. Semiclassical quantum mechanics suggest operators of this form with $\beta=$ $1 / 2$, (see chapter 30 ). The problem with such operators is due to the fact that when considering the Jacobian matrices $\mathbf{J}_{a b}=\mathbf{J}_{a} \mathbf{J}_{b}$ for two successive trajectory segments $a$ and $b$, the corresponding eigenvalues are in general not multiplicative, $\Lambda_{a b} \neq \Lambda_{a} \Lambda_{b}$ (unless $a, b$ are repeats of the same prime cycle $p$, so $\mathbf{J}_{a} \mathbf{J}_{b}=\mathbf{J}_{p}^{r_{a}+r_{b}}$ ). Consequently, this evolution operator is not multiplicative along the trajectory. The theorems require that the evolution be represented as a matrix in an appropriate polynomial basis, and thus cannot be applied to non-multiplicative kernels, that is, kernels that do not satisfy the semi-group property $\mathcal{L}^{t^{\prime}} \mathcal{L}^{t}=\mathcal{L}^{t^{\prime}+t}$. Cure for this problem in this particular case is given in appendix H.1.

Property 2 is violated by the $1-d$ tent map (see figure 16.2 (a))

$$
f(x)=\alpha(1-|1-2 x|), \quad 1 / 2<\alpha<1
$$

All cycle eigenvalues are hyperbolic, but in general the critical point $x_{c}=$ $1 / 2$ is not a pre-periodic point, there is no finite Markov partition and the symbolic dynamics does not have a finite grammar (see sect. 12.4 for definitions). In practice this means that while the leading eigenvalue of $\mathcal{L}$ might be computable, the rest of the spectrum is very hard to control; as the parameter $\alpha$ is varied, non-leading zeros of the spectral determinant move wildly about.

Property 3 is violated by the map (see figure 16.2 (b))

$$
f(x)=\left\{\begin{array}{lll}
x+2 x^{2} & , & x \in I_{0}=\left[0, \frac{1}{2}\right] \\
2-2 x & , & x \in I_{1}=\left[\frac{1}{2}, 1\right]
\end{array} .\right.
$$


(a)

(b)

Figure 16.2: (a) A (hyperbolic) tent map without a finite Markov partition. (b) A Markov map with a marginal fixed point.

Here the interval $[0,1]$ has a Markov partition into the two subintervals $I_{0}$ and $I_{1} ; f$ is monotone on each. However, the fixed point at $x=0$ has marginal stability $\Lambda_{0}=1$, and violates the condition 3. This type of map is called intermittent and necessitates much extra work. The problem is that the dynamics in the neighborhood of a marginal fixed point is very slow, with correlations decaying as power laws rather than exponentially. We will discuss such flows in chapter 21.

The property 4 is required as the heuristic approach of chapter 14 faces two major hurdles:

1. The trace (14.7) is not well defined since the integral kernel is singular.
2. The existence and properties of eigenvalues are by no means clear.

Actually this property is quite restrictive, but we need it in the present approach, in order that the Banach space of analytic functions in a disk is preserved by the Perron-Frobenius operator.

In attempting to generalize the results we encounter several problems. First, in higher dimensions life is not as simple. Multi-dimensional residue calculus is at our disposal but in general requires that we find poly-domains (direct product of domains in each coordinate) and this need not be the case. Second, and perhaps somewhat surprisingly, the 'counting of periodic orbits' presents a difficult problem. For example, instead of the Bernoulli shift consider the doubling map of the circle, $x \mapsto 2 x \bmod 1, x \in R / Z$. Compared to the shift on the interval $[0,1]$ the only difference is that the endpoints 0 and 1 are now glued together. But since these endpoints are fixed points of the map the number of cycles of length $n$ decreases by 1 . The determinant becomes:

$$
\begin{equation*}
\operatorname{det}(1-z \mathcal{L})=\exp \left(-\sum_{n=1} \frac{z^{n}}{n} \frac{2^{n}-1}{2^{n}-1}\right)=1-z \tag{16.23}
\end{equation*}
$$

The value $z=1$ still comes from the constant eigenfunction but the Bernoulli polynomials no longer contribute to the spectrum (they are not periodic).

Proofs of these facts, however, are difficult if one sticks to the space of analytic functions.

Third, our Cauchy formulas a priori work only when considering purely expanding maps. When stable and unstable directions co-exist we have to resort to stranger function spaces, as shown in the next section.

### 16.5 Hyperbolic maps

(H.H. Rugh)

Moving onto hyperbolic systems, one faces the following paradox: If $f$ is an area-preserving hyperbolic and real-analytic map of, for example, a 2-dimensional torus then the Perron-Frobenius operator is unitary on the space of $L^{2}$ functions, and its spectrum is confined to the unit circle. On the other hand, when we compute determinants we find eigenvalues scattered around inside the unit disk. Thinking back on our Bernoulli shift example one would like to imagine these eigenvalues as popping up from the $L^{2}$ spectrum by shrinking the function space. Shrinking the space, however, can only make the spectrum smaller so this is obviously not what happens. Instead one needs to introduce a 'mixed' function space where in the unstable direction one resort to analytic functions, as before, but in the stable direction one considers a 'dual space' of distributions on analytic functions. Such a space is neither included in nor does it include the $L^{2}$ space and we have thus resolved the paradox. But it still remains to be seen how traces and determinants are calculated.

The linear hyperbolic fixed point example example 16.6 is somewhat misleading, as we have made explicit use of a map that acts independently along the stable and unstable directions. For a more general hyperbolic map, there is no way to implement such direct product structure, and the whole argument falls apart. Her comes an idea; use the analyticity of the map to rewrite the Perron-Frobenius operator acting as follows ( $\sigma$ is the sign of the derivative in the unstable direction):

$$
\begin{equation*}
\mathcal{L} h\left(z_{1}, z_{2}\right)=\oint \oint \frac{\sigma h\left(w_{1}, w_{2}\right)}{\left(z_{1}-f_{1}\left(w_{1}, w_{2}\right)\left(f_{2}\left(w_{1}, w_{2}\right)-z_{2}\right)\right.} \frac{d w_{1}}{2 \pi i} \frac{d w_{2}}{2 \pi i} . \tag{16.24}
\end{equation*}
$$

Here the function $\phi$ should belong to a space of functions analytic respectively outside a disk and inside a disk in the first and the second coordinate, with the additional property that the function decays to zero as the first coordinate tends to infinity. The contour integrals are along the boundaries of these disks. It is but an exercise in multi-dimensional residue calculus to verify that for the above linear example this expression reduces to (16.9). Such operators form the building bricks in the calculation of traces and determinants and one is able to prove the following:

Theorem: The spectral determinant for hyperbolic analytic maps is entire.


The proof, apart from the Markov property which is the same as for the purely expanding case, relies heavily on analyticity of the map in the explicit construction of the function space. As we have also seen in the previous example the basic idea is to view the hyperbolicity as a cross product of a contracting map in the forward time and another contracting map in the backward time. In this case the Markov property introduced above has to be elaborated a bit. Instead of dividing the phase space into intervals, one divides it into rectangles. The rectangles should be viewed as a direct product of intervals (say horizontal and vertical), such that the forward map is contracting in, for example, the horizontal direction, while the inverse map is contracting in the vertical direction. For Axiom A systems (see remark 16.9) one may choose coordinate axes close to the stable/unstable manifolds of the map. With the phase space divided into $N$ rectangles $\left\{\mathcal{M}_{1}, \mathcal{M}_{2}, \ldots, \mathcal{M}_{N}\right\}, \mathcal{M}_{i}=I_{i}^{h} \times I_{i}^{v}$ one needs complex extension $D_{i}^{h} \times D_{i}^{v}$, with which the hyperbolicity condition (which at the same time guarantees the Markov property) can be formulated as follows:

Analytic hyperbolic property: Either $f\left(\mathcal{M}_{i}\right) \cap \operatorname{Int}\left(\mathcal{M}_{j}\right)=\emptyset$, or for each pair $w_{h} \in \mathrm{Cl}\left(D_{i}^{h}\right), z_{v} \in \mathrm{Cl}\left(D_{j}^{v}\right)$ there exist unique analytic functions of $w_{h}, z_{v}: w_{v}=w_{v}\left(w_{h}, z_{v}\right) \in \operatorname{Int}\left(D_{i}^{v}\right), z_{h}=z_{h}\left(w_{h}, z_{v}\right) \in \operatorname{Int}\left(D_{j}^{h}\right)$, such that $f\left(w_{h}, w_{v}\right)=\left(z_{h}, z_{v}\right)$. Furthermore, if $w_{h} \in I_{i}^{h}$ and $z_{v} \in I_{j}^{v}$, then $w_{v} \in I_{i}^{v}$ and $z_{h} \in I_{j}^{h}$ (see figure 16.3).

What this means for the iterated map is that one replaces coordinates $z_{h}, z_{v}$ at time $n$ by the contracting pair $z_{h}, w_{v}$, where $w_{v}$ is the contracting coordinate at time $n+1$ for the 'partial' inverse map.

In two dimensions the operator in (16.24) acts on functions analytic outside $D_{i}^{h}$ in the horizontal direction (and tending to zero at infinity) and inside $D_{i}^{v}$ in the vertical direction. The contour integrals are precisely along the boundaries of these domains.

A map $f$ satisfying the above condition is called analytic hyperbolic and the theorem states that the associated spectral determinant is entire,
and that the trace formula (14.7) is correct.
Examples of analytic hyperbolic maps are provided by small analytic perturbations of the cat map (where the Markov partitioning is non-trivial [11.11]), the 3-disk repeller, and the 2-d baker's map.

### 16.6 Physics of eigenvalues and eigenfunctions



We appreciate by now that any honest attempt to look at spectral properties of the Perron-Frobenius operator involves hard mathematics: but the effort is rewarded by the fact that we are finally able to control analyticity properties of dynamical zeta functions and spectral determinants, and thus substantiate the claim that these objects provide a powerful and well founded perturbation theory.

Often (see chapter 10) the physically important part of the spectrum is just the leading eigenvalue, which gives us the escape rate from a repeller, or, for a general evolution operator, formulas for expectation values of observables and their higher moments. Also the eigenfunction associated to the leading eigenvalue has a physical interpretation (see chapter 9): it is the density of the natural measure, with singular measures ruled out by the proper choice of the function space. Such a conclusion is coherent with a the validity of a generalized Perron-Frobenius theorem for evolution operators. In the finite dimensional setting such theorem is formulated as follows:

- Perron-Frobenius theorem: let $L_{n m}$ be a nonnegative matrix, such that some $n$ exists for which $\left(L^{n}\right)_{i j}>0 \forall i, j$ : then

1. the maximal modulus eigenvalue is non degenerate, real and positive
2. the corresponding eigenvector (defined up to a constant) has nonnegative coordinates

We may ask what physical information is contained in eigenvalues beyond the leading one: suppose that we have a probability conserving system (so that the dominant eigenvalue is 1 ), for which the essential spectral radius is such that $0<\rho_{\text {ess }}<\theta<1$ on some Banach space $\mathcal{B}$ and denote by $\mathbf{P}$ the projection corresponding to the part of the spectrum inside a disk of radius $\theta$. We denote by $\lambda_{1}, \lambda_{2} \ldots \lambda_{M}$ the eigenvalues outside of this disk, ordered by the size of their absolute value (so that $\lambda_{1}=1$ ). Then we have the following decomposition

$$
\begin{equation*}
\mathcal{L} \varphi=\sum_{i=1}^{M} \lambda_{i} \psi_{i} L_{i} \psi_{i}^{*} \varphi+\mathbf{P} \mathcal{L} \varphi \tag{16.25}
\end{equation*}
$$

when $L_{i}$ are (finite) matrices in Jordan normal form ( $L_{0}=0$ is a $[1 \times 1]$ matrix, as $\lambda_{0}$ is simple, due to Perron-Frobenius theorem), while $\psi_{i}$ is a row vector whose elements are a basis on the eigenspace corresponding to $\lambda_{i}$, and $\psi_{i}^{*}$ is a column vector of elements of $\mathcal{B}^{*}$ (the dual space, of linear functionals over $\mathcal{B}$ ) spanning the eigenspace of $\mathcal{L}^{*}$ corresponding to $\lambda_{i}$. For iterates of Perron-Frobenius operator (16.25) becomes

$$
\begin{equation*}
\mathcal{L}^{n} \varphi=\sum_{i=1}^{M} \lambda_{i}^{n} \psi_{i} L_{i}^{n} \psi_{i}^{*} \varphi+\mathbf{P} \mathcal{L}^{n} \varphi \tag{16.26}
\end{equation*}
$$

If we now consider expressions like

$$
\begin{equation*}
\langle\xi| \mathcal{L}^{n}|\varphi\rangle=\int_{\mathcal{M}} d y \xi(y)\left(\mathcal{L}^{n} \varphi\right)(y)=\int_{\mathcal{M}} d w\left(\xi \circ f^{n}\right)(w) \varphi(w) \tag{16.27}
\end{equation*}
$$

we have

$$
\begin{equation*}
C_{\xi, \varphi}^{(n)}=\lambda_{1}^{n} \omega_{1}(\xi, \varphi)+\sum_{i=2}^{L} \lambda_{i}^{n} \omega_{i}^{(n)}(\xi, \varphi)+\mathcal{O}\left(\theta^{n}\right) \tag{16.28}
\end{equation*}
$$

where

$$
\omega_{i}^{(n)}(\xi, \varphi)=\int_{\mathcal{M}} d y \xi(y) \psi_{i} L_{i}^{n} \psi_{i}^{*} \varphi
$$

In this way we see how eigenvalues beyond the leading one provide a twofold piece of information: they rule the convergence of expressions containing high powers of evolution operator to the leading order (the $\lambda_{1}$ contribu-
page 289 tion). Moreover if $\omega_{1}(\xi, \varphi)=0$ then (16.27) defines a correlation function: as each term in (16.28) vanishes exponentially in the $n \rightarrow \infty$ limit, the eigenvalues $\lambda_{2}, \ldots \lambda_{M}$ rule the exponential decay of correlations for our dynamical system. We observe that prefactors $\omega$ depend on the choice of functions, while the exponential decay rates (logarithms of $\lambda_{i}$ ) do not: the correlation spectrum is thus an universal property of the dynamics (once we fix the overall functional space our Perron-Frobenius operator acts on).

Example 16.8 Bernoulli shift eigenfunctions: So let us come back the Bernoulli shift example (16.6), on the space of analytic functions on a disk: apart from the origin we have only simple eigenvalues $\lambda_{k}=2^{-k} k=0,1, \ldots$. The eigenvalue $\lambda_{0}=1$ corresponds to probability conservation: the corresponding eigenfunction $B_{0}(x)=1$ indicates that the natural measure has a constant density over the unit interval. If we now take any analytic function $\eta(x)$ with zero average (with respect to the Lebesgue measure), we have that $\omega_{1}(\eta, \eta)=0$, and from (16.28) we have that the asymptotic decay of correlation function is (unless also $\omega_{1}(\eta, \eta)=0$ )

$$
\begin{equation*}
C_{\eta, \eta}(n) \sim \exp (-n \log 2) \tag{16.29}
\end{equation*}
$$

thus $-\log \lambda_{1}$ gives the exponential decay rate of correlations (with a prefactor that depends on the choice of the function). Actually the Bernoulli shift case may be treated exactly, as for analytic functions we can employ the Euler-MacLaurin summation formula

$$
\begin{equation*}
\eta(z)=\int_{0}^{1} d w \eta(w)+\sum_{m=1}^{\infty} \frac{\eta^{(m-1)}(1)-\eta^{(m-1)}(0)}{m!} B_{m}(z) . \tag{16.30}
\end{equation*}
$$

As we are considering zero-average functions, we have from (16.27), and the fact that Bernoulli polynomials are eigenvectors of the Perron-Frobenius operator

$$
C_{\eta, \eta}(n)=\sum_{m=1}^{\infty} \frac{\left(2^{-m}\right)^{n}\left(\eta^{(m)}(1)-\eta^{(m)}(0)\right)}{m!} \int_{0}^{1} d z \eta(z) B_{m}(z)
$$

The decomposition (16.30) is also useful to make us realize that the linear functionals $\psi_{i}^{*}$ are quite singular objects: if we write it as

$$
\eta(z)=\sum_{m=0}^{\infty} B_{m}(z) \psi_{m}^{*}[\eta]
$$

we see that these functionals are of the form

$$
\psi_{i}^{*}[\varepsilon]=\int_{0}^{1} d w \Psi_{i}(w) \varepsilon(w)
$$

where

$$
\begin{equation*}
\Psi_{i}(w)=\frac{(-1)^{i-1}}{i!}\left(\delta^{(i-1)}(w-1)-\delta^{(i-1)}(w)\right) \tag{16.31}
\end{equation*}
$$

when $i \geq 1$, while $\Psi_{0}(w)=1$. Such a representation is only meaningful when the function $\varepsilon$ is analytic in $w, w-1$ neighborhoods.

### 16.7 Troubles ahead

The above discussion confirms that for this simple example formal manipulations with traces and determinants are justified: the Perron-Frobenius operator has isolated eigenvalues, the trace formulas are explicitly verified, and the spectral determinant is an entire function whose zeroes yield the eigenvalues. Real life is harder, as we may appreciate through the following considerations:

- Our discussion tacitly assumed something that is physically entirely reasonable: our evolution operator is acting on the space of analytic functions, that is, we are allowed to represent the initial density $\rho(x)$ by its Taylor expansions in the neighborhoods of periodic points. This is however far from being the only possible choice: mathematicians often work with the function space $\mathbb{C}^{k+\alpha}$, that is, the space of $k$

Figure 16.4: Spectrum of the PerronFrobenius operator acting on the space of $\mathbb{C}^{k+\alpha}$ Hölder-continuous functions: only $k$ isolated eigenvalues remain between the spectral radius,
 and the essential spectral radius which bounds the "essential", continuous spectrum.
times differentiable functions whose $k$ 'th derivatives are Hölder continuous with an exponent $0<\alpha \leq 1$ : then every $y^{\eta}$ with $\operatorname{Re} \eta>k$ is an eigenfunction of Perron-Frobenius operator and we have

$$
\mathcal{L} y^{\eta}=\frac{1}{|\Lambda| \Lambda^{\eta}} y^{\eta}, \quad \eta \in \mathbb{C}
$$

This spectrum is quite different from the analytic case: only a small number of isolated eigenvalues remain, enclosed between the spectral radius and a smaller disk of radius $1 /|\Lambda|^{k+1}$, see figure 16.4. In literature the radius of this disk is called the essential spectral radius.

In sect. 16.4 we will discuss this point further, with the aid of a less trivial 1-dimensional example. The physical point of view is complementary to the standard setting of ergodic theory, where many chaotic properties of a dynamical system are encoded by the presence of a continuous spectrum, used to prove asymptotic decay of correlations in the space of $L^{2}$ square-integrable functions.

- A deceptively innocent assumption hides behind many features discussed so far: that (16.1) maps a given function space into itself. This is strictly related to the expanding property of the map: if $f(x)$ is smooth in a domain $D$ then $f(x / \Lambda)$ is smooth on a larger domain, provided $|\Lambda|>1$. This is not obviously the case for hyperbolic systems in higher dimensions, and, as we have seen in sect. 16.5, extensions of the results obtained for expanding maps will be highly nontrivial,
- It is not at all clear that the above analysis of a simple one-branch, one fixed point repeller can be extended to dynamical systems with a Cantor set infinity of periodic points: we show that in sect. 16.4.


## Commentary

Remark 16.1 Surveys of rigorous theory. We recommend references listed in sect. 1.8 for an introduction into the mathematical literature on this subject. For a physicist, Driebe's monograph [1.25] might be the most accessible introduction
into mathematics touched upon in this chapter. There are a number of reviews of the mathematical approach to dynamical zeta functions and spectral determinants, with pointers to the original references, such as refs. [16.1, 16.2]. An alternative approach to spectral properties of the Perron-Frobenius operator is illustrated in ref. [16.3].

The ergodic theory, as presented by Sinai [16.14] and others, tempts one to describe the densities that the evolution operator acts on in terms of either integrable or square-integrable functions. As we have already seen, for our purposes, this space is not suitable. An introduction to ergodic theory is given by Sinai, Kornfeld and Fomin [16.15]; more advanced and more old fashioned presentations are Walters [16.16] and Denker, Grillenberger and Sigmund [16.17]; and a more formal Peterson [16.18].

Remark 16.2 Fredholm theory. Our brief summary of Fredholm theory is based on the exposition in ref. [16.4]. A technical introduction of the theory from an operatorial point of view is contained in ref. [16.5]. The theory has been generalized in ref. [16.6].

Remark 16.3 Bernoulli shift. For a more detailed discussion, consult chaper 3 of ref. [1.25]. The extension of Fredholm theory to the case or Bernoulli shift on $\mathbb{C}^{k+\alpha}$ (in which the Perron-Frobenius operator is not compact - technically it is only quasi-compact, that is the essential spectral radius is strictly smaller than the spectral radius) has been given by Ruelle [16.7]: a concise and readable statement of the results is contained in ref. [16.8].

Remark 16.4 Higher dimensions and generalized Fredholm theory. When extending Bernoulli shift to higher dimensions, formulation of the Fredholm theory [16.6] which avoids problems with multi-dimensional residue calculus may be used: see ref. [16.9].

Remark 16.5 Hyperbolic dynamics. When dealing with hyperbolic systems one might try to reduce back to the expanding case by projecting the dynamics along the unstable directions. As mentioned in the text this might be technically quite involved, as usually such the unstable foliation is not characterized by very strong smoothness properties. For such an approach, see ref. [16.3].

Remark 16.6 Spectral determinants for smooth flows. The theorem on page 279 applies also to hyperbolic analytic maps in dimensions and smooth hyperbolic analytic flows in $(d+1)$ dimensions, provided that the flow can be reduced to a piecewise analytic map by suspension on a Poincaré section complemented by an analytic "ceiling" function (3.4) which accounts for a variation in the section return times. For example, if we take as the ceiling function $g(x)=e^{s T(x)}$, where $T(x)$ is the time of the next Poincaré section for a trajectory staring at $x$, we reproduce the flow spectral determinant (15.13). Proofs are getting too hard for the purposes of this chapter.

Remark 16.7 Explicit diagonalization. For 1-d repellers a diagonalization of an explicit truncated $L_{m n}$ matrix evaluated in a judiciously chosen basis may yield many more eigenvalues than a cycle expansion (see refs. [16.10, 16.11]). The reasons why one persists anyway in using the periodic orbit theory are partially aesthetic, and partially pragmatic. Explicit $L_{m n}$ demands explicit choice of a basis and is thus non-invariant, in contrast to cycle expansions which utilize only the invariant information about the flow. In addition, we usually do not know how to construct $L_{m n}$ for a realistic flow, such as the hyperbolic 3-disk game of pinball flow of sect. 1.3, whereas the periodic orbit formulas are general and straightforward to apply.

Remark 16.8 Perron-Frobenius theorem. A proof of the Perron-Frobenius theorem may be found in ref. [16.12]. For positive transfer operators such theorem has been generalized by Ruelle [16.13].

Remark 16.9 Axiom A systems. Proofs outlined in sect. 16.5 follow the thesis work of H.H. Rugh [16.9, 16.19, 16.20]. For mathematical introduction to the subject, consult the excellent review by V. Baladi [16.1]. Rigorous treatment is given in refs. [16.9, 16.19, 16.20]. It would take us too far to give and explain the definition of the Axiom A systems (see refs. [1.15, 1.16]). Axiom A implies, however, the existence of a Markov partition of the phase space from which the properties 2 and 3 assumed on page 268 follow.

Remark 16.10 Exponential mixing speed of the Bernoulli shift. We see from (16.29) that for the Bernoulli shift the exponential decay rate of correlations coincides with the Lyapunov exponent: while such an identity holds for a number of systems, it is by no means a general result, and there exist explicit counterexamples.

Remark 16.11 Left eigenfunctions. We shall never use explicit form of left eigenfunctions, corresponding to highly singular kernels like (16.31). Many details have been elaborated in a number of papers, like ref. [16.21], with a daring physical interpretation.

Remark 16.12 Ulam's idea. The approximation of Perron-Frobenius operator defined by (10.37) has been shown to reproduce correctly the spectrum for expanding maps, once finer and finer Markov partitions are used [16.22]. The subtle point of choosing a phase space partitioning for a "generic case" is discussed in ref. [16.23].

## Résumé

A theory of evaluation of dynamical averages by means of trace formulas and spectral determinants requires a deeper understanding of their analyticity and convergence. In the case of especially well-behaved "Axiom $A$ " systems where both the symbolic dynamics and hyperbolicity are under
control, it is possible to treat traces and determinants in a rigorous fashion, and strong results about analyticity properties of dynamical zeta functions and spectral determinants outlined above follow.

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## Exercises

Exercise 16.1 What space does $\mathcal{L}$ act on? Show that (16.2) is a complete basis on the space of analytic functions on a disk (and thus that we found the complete set of eigenvalues).

Exercise 16.2 What space does $\mathcal{L}$ act on? What can be said about the spectrum of (16.1) on $L^{1}[0,1]$ ? Compare the result with figure 16.4.

Exercise 16.3 Euler formula. Derive the Euler formula (16.5)

$$
\begin{align*}
\prod_{k=0}^{\infty}\left(1+t u^{k}\right) & =1+\frac{t}{1-u}+\frac{t^{2} u}{(1-u)\left(1-u^{2}\right)}+\frac{t^{3} u^{3}}{(1-u)\left(1-u^{2}\right)\left(1-u^{3}\right)} \cdots \\
& =\sum_{k=0}^{\infty} t^{k} \frac{u^{\frac{k(k-1)}{2}}}{(1-u) \cdots\left(1-u^{k}\right)}, \quad|u|<1 . \tag{16.32}
\end{align*}
$$

Exercise 16.4 2-d product expansion**. We conjecture that the expansion corresponding to (16.32) is in this case

$$
\begin{align*}
\prod_{k=0}^{\infty}\left(1+t u^{k}\right)^{k+1}= & \sum_{k=0}^{\infty} \frac{F_{k}(u)}{(1-u)^{2}\left(1-u^{2}\right)^{2} \cdots\left(1-u^{k}\right)^{2}} t^{k} \\
= & 1+\frac{1}{(1-u)^{2}} t+\frac{2 u}{(1-u)^{2}\left(1-u^{2}\right)^{2}} t^{2} \\
& +\frac{u^{2}\left(1+4 u+u^{2}\right)}{(1-u)^{2}\left(1-u^{2}\right)^{2}\left(1-u^{3}\right)^{2}} t^{3}+\cdots \tag{16.33}
\end{align*}
$$

$F_{k}(u)$ is a polynomial in $u$, and the coefficients fall off asymptotically as $C_{n} \approx u^{n^{3 / 2}}$. Verify; if you have a proof to all orders, e-mail it to the authors. (See also solution 16.3).

Exercise 16.5 Bernoulli shift on $L$ spaces. Check that the family (16.21) belongs to $L^{1}([0,1])$. What can be said about the essential spectral radius on $L^{2}([0,1])$ ? A useful reference is [16.25].

Exercise 16.6 Cauchy integrals. Rework all complex analysis steps used in the Bernoulli shift example on analytic functions on a disk.

Exercise 16.7 Escape rate. Consider the escape rate from a strange repeller: find a choice of trial functions $\xi$ and $\varphi$ such that (16.27) gives the fraction on particles surviving after $n$ iterations, if their initial density distribution is $\rho_{0}(x)$. Discuss the behavior of such an expression in the long time limit.

## Chapter 17

## Fixed points, and how to get them

(F. Christiansen)

Having set up the dynamical context, now we turn to the key and unavoidable piece of numerics in this subject; search for the solutions $(x, T)$, $x \in \mathbb{R}^{d}, T \in \mathbb{R}$ of the periodic orbit condition

$$
\begin{equation*}
f^{t+T}(x)=f^{t}(x), \quad T>0 \tag{17.1}
\end{equation*}
$$

for a given flow or mapping.
We know from chapter 14 that cycles are the necessary ingredient for evaluation of spectra of evolution operators. In chapter 11 we have developed a qualitative theory of how these cycles are laid out topologically. This chapter is intended as a hands-on guide to extraction of periodic orbits, and should be skipped on first reading - you can return to it whenever the need for finding actual cycles arises. A serious cyclist might want to also learn about the variational methods to find cycles, chapter 31. They are particularly useful when little is understood about the topology of a flow, such as in high-dimensional periodic orbit searches.

fast track:
chapter 18, p. 309
A prime cycle $p$ of period $T_{p}$ is a single traversal of the periodic orbit, so our task will be to find a cycle point $x \in p$ and the shortest time $T_{p}$ for which (17.1) has a solution. A cycle point of a flow $f^{t}$ which crosses a Poincaré section $n_{p}$ times is a fixed point of the $P^{n_{p}}$ iterate of the Poincaré section return map $P$, hence we shall refer to all cycles as "fixed points" in this chapter. By cyclic invariance, stability eigenvalues and the period of the cycle are independent of the choice of the initial point, so it will suffice to solve (17.1) at a single cycle point.

If the cycle is an attracting limit cycle with a sizable basin of attraction, it can be found by integrating the flow for sufficiently long time. If the cycle is unstable, simple integration forward in time will not reveal it, and methods to be described here need to be deployed. In essence, any method for finding a cycle is based on devising a new dynamical system which possesses the same cycle, but for which this cycle is attractive. Beyond that, there is a great freedom in constructing such systems, and many different methods are used in practice.

Due to the exponential divergence of nearby trajectories in chaotic dynamical systems, fixed point searches based on direct solution of the fixedpoint condition (17.1) as an initial value problem can be numerically very unstable. Methods that start with initial guesses for a number of points along the cycle, such as the multipoint shooting method described here in sect. 17.3, and the variational) methods of chapter 31, are considerably more robust and safer.

A prerequisite for any exhaustive cycle search is a good understanding of the topology of the flow: a preliminary step to any serious periodic orbit calculation is preparation of a list of all distinct admissible prime periodic symbol sequences, such as the list given in table 11.1. The relations between the temporal symbol sequences and the spatial layout of the topologically distinct regions of the phase space discussed in chapters 11 and 12 should enable us to guess location of a series of periodic points along a cycle. Armed with such informed guess we proceed to improve it by methods such as the Newton-Raphson iteration; we illustrate this by considering 1 -dimensional and $d$-dimensional maps.

### 17.1 Where are the cycles?

Ergodic exploration of recurrences that we turn to now sometimes performs admirably well in getting us started.

Both in the example of the Rössler flow and of the Kuramoto-Sivashinsky system we sketched the attractors by running a long chaotic trajectory, and noted that the attractors are very thin, but otherwise the return maps that we plotted were disquieting - neither figure 3.2 nor figure 3.3 appeared to be 1-to-1 maps. In this section we show how to use such information to approximately locate cycles. In the remainder of this chapter and in chapter 31 we shall learn how to turn such guesses into highly accurate cycles.

## Example 17.1 Rössler attractor

(G. Simon and P. Cvitanović)

Run a long simulation of the Rössler flow $f^{t}$, plot a Poincaré section, as in figure 3.1, and extract the corresponding Poincaré return map $P$, as in figure 3.2. Luck is with us; the figure 17.1(a) return map $y \rightarrow P_{1}(y, z)$ looks much like a parabola, so we take the unimodal map symbolic dynamics, sect. 11.3.1, as our guess for the covering dynamics. Strictly speaking, the attractor is "fractal", but for all practical purposes the return map


Figure 17.1: (a) $y \rightarrow P_{1}(y, z)$ return map for $x=0, y>0$ Poincaré section of the Rössler flow figure 2.3. (b) The $\overline{1}$-cycle found by taking the fixed point $y_{k+n}=$ $y_{k}$ together with the fixed point of the $z \rightarrow z$ return map (not shown) an initial guess $\left.\left(0, y^{(0)}, z^{( } 0\right)\right)$ for the Newton-Raphson search. (c) $y_{k+3}=P_{1}^{3}\left(x_{k}, y_{k}\right)$, the third iterate of Poincaré return map (3.1) together with the corresponding plot for $z_{k+3}=P_{2}^{3}\left(x_{k}, y_{k}\right)$, is used to pick starting guesses for the Newton-Raphson searches for the two 3-cycles: (d) the $\overline{001}$ cycle, and (e) the $\overline{011}$ cycle. (G. Simon)
is 1-dimensional; your printer will need a resolution better than $10^{14}$ dots per inch to start resolving its structure.

Periodic points of a prime cycle $p$ of cycle length $n_{p}$ for the $x=0, y>0$ Poincaré section of the Rössler flow figure 2.3 are fixed points $(y, z)=\operatorname{Pn}(y, z)$ of the $n$th Poincaré return map.

Using the fixed point $y_{k+1}=y_{k}$ in figure 17.1(a) together with the simultaneous fixed point of the $z \rightarrow P_{1}(y, z)$ return map (not shown) as a starting guess $\left(0, y^{(0)}, z^{(0)}\right)$ for the Newton-Raphson search for the cycle $p$ with symbolic dynamics label $\overline{1}$, we find the cycle figure $17.1(b)$ with the Poincaré section point $\left(0, y_{p}, z_{p}\right)$, period $T_{p}$, expanding, marginal, contracting stability eigenvalues $\left(\Lambda_{p, e}, \Lambda_{p, m}, \Lambda_{p, c}\right)$, and Lyapunov exponents $\left(\lambda_{p, e}, \lambda_{p, m}, \lambda_{p, c}\right)$ :

$$
\text { 1-cycle: } \begin{align*}
(x, y, z) & =(0,6.09176832,1.2997319) \\
T_{1} & =5.88108845586 \\
\left(\Lambda_{1, e}, \Lambda_{1, m}, \Lambda_{1, c}\right) & =\left(-2.40395353,1+10^{-14},-1.29 \times 10^{-14}\right) \\
\left(\lambda_{1, e}, \lambda_{1, m}, \lambda_{1, c}\right) & =\left(0.149141556,10^{-14},-5.44\right) \tag{17.2}
\end{align*}
$$

The Newton-Raphson method that we used is described in sect. 17.5.
As an example of a search for longer cycles, we use $y_{k+3}=P^{3}\left(y_{k}, z_{k}\right)$ the 17.7 third iterate of Poincaré return map (3.1) plotted in figure 17.1(c), together with a corresponding plot for $z_{k+3}=f^{3}\left(y_{k}, z_{k}\right)$, to pick starting guesses for the NewtonRaphson searches for the two 3-cycles plotted in figure 17.1(d), (e). For a listing of the short cycles of the Rössler flow, consult table 17.1.

The numerical evidence suggests (but a proof is lacking) that all cycles that comprise the strange attractor of the Rössler system are hyperbolic, each with an expanding eigenvalue $\left|\Lambda_{e}\right|>1$, a contracting eigenvalue $\left|\Lambda_{c}\right|<1$, and a marginal eigenvalue $\left|\Lambda_{m}\right|=1$ corresponding to displacements along the direction of the flow.

For the Rössler system the contracting eigenvalues turn out to be insanely contracting, a factor of $e^{-32}$ per one par-course of the attractor, so their numerical

Figure 17.2: The inverse time path to the $\overline{01}$ cycle of the logistic map $f(x)=4 x(1-x)$ from an initial guess of $x=0.2$. At each inverse iteration we chose the 0 , respectively 1 branch.

determination is quite difficult. Fortunately, they are irrelevant; for all practical purposes the strange attractor of the Rössler system is 1-dimensional, a very good realization of a horseshoe template.

### 17.2 One-dimensional mappings

### 17.2.1 Inverse iteration

Let us first consider a very simple method to find unstable cycles of a 1dimensional map such as the logistic map. Unstable cycles of 1-d maps are attracting cycles of the inverse map. The inverse map is not single valued, so at each backward iteration we have a choice of branch to make. By choosing branch according to the symbolic dynamics of the cycle we are trying to find, we will automatically converge to the desired cycle. The rate of convergence is given by the stability of the cycle, that is, the convergence is exponentially fast. Figure 17.2 shows such path to the $\overline{01}$-cycle of the logistic map.

The method of inverse iteration is fine for finding cycles for 1-d maps and some $2-d$ systems such as the repeller of exercise 17.11. It is not particularly fast, especially if the inverse map is not known analytically. However, it completely fails for higher dimensional systems where we have both stable and unstable directions. Inverse iteration will exchange these, but we will still be left with both stable and unstable directions. The best strategy is to directly attack the problem of finding solutions of $f^{T}(x)=x$.

### 17.2.2 Newton's method

Newton's method for determining a zero $x^{*}$ of a function $F(x)$ of one variable is based on a linearization around a starting guess $x_{0}$ :

$$
\begin{equation*}
F(x) \approx F\left(x_{0}\right)+F^{\prime}\left(x_{0}\right)\left(x-x_{0}\right) \tag{17.3}
\end{equation*}
$$

An approximate solution $x_{1}$ of $F(x)=0$ is

$$
\begin{equation*}
x_{1}=x_{0}-F\left(x_{0}\right) / F^{\prime}\left(x_{0}\right) \tag{17.4}
\end{equation*}
$$

Figure 17.3: Convergence of Newton's method $(\diamond)$ vs. inverse iteration $(+)$. The error after $n$ iterations searching for the $\overline{01}$-cycle of the logistic map $f(x)=4 x(1-x)$ with an initial starting guess of $x_{1}=0.2, x_{2}=0.8 . y$-axis is $\log _{10}$ of the error. The difference between the exponential convergence of the inverse iteration method and the super-exponential convergence of Newton's method is dramatic.


The approximate solution can then be used as a new starting guess in an iterative process. A fixed point of a map $f$ is a solution to $F(x)=x-f(x)=$ 0 . We determine $x$ by iterating

$$
\begin{align*}
x_{m} & =g\left(x_{m-1}\right)=x_{m-1}-F\left(x_{m-1}\right) / F^{\prime}\left(x_{m-1}\right) \\
& =x_{m-1}-\frac{1}{1-f^{\prime}\left(x_{m-1}\right)}\left(x_{m-1}-f\left(x_{m-1}\right)\right) . \tag{17.5}
\end{align*}
$$

Provided that the fixed point is not marginally stable, $f^{\prime}(x) \neq 1$ at the fixed point $x$, a fixed point of $f$ is a super-stable fixed point of the NewtonRaphson map $g, g^{\prime}(x)=0$, and with a sufficiently good initial guess, the Newton-Raphson iteration will converge super-exponentially fast.

To illustrate the efficiency of the Newton's method we compare it to the inverse iteration method in figure 17.3. Newton's method wins hands down: the number of significant digits of the accuracy of $x$ estimate doubles with each iteration.

In order to avoid jumping too far from the desired $x^{*}$ (see figure 17.4), one often initiates the search by the damped Newton's method,

$$
\Delta x_{m}=x_{m+1}-x_{m}=-\frac{F\left(x_{m}\right)}{F^{\prime}\left(x_{m}\right)} \Delta \tau, \quad 0<\Delta \tau \leq 1
$$

takes small $\Delta \tau$ steps at the beginning, reinstating to the full $\Delta \tau=1$ jumps only when sufficiently close to the desired $x^{*}$.

### 17.3 Multipoint shooting method

Periodic orbits of length $n$ are fixed points of $f^{n}$ so in principle we could use the simple Newton's method described above to find them. However, this is not an optimal strategy. $f^{n}$ will be a highly oscillating function with perhaps as many as $2^{n}$ or more closely spaced fixed points, and finding a specific periodic point, for example one with a given symbolic sequence, requires a very good starting guess. For binary symbolic dynamics we must expect to improve the accuracy of our initial guesses by at least a factor of $2^{n}$ to find orbits of length $n$. A better alternative is the multipoint shooting

Figure 17.4: Newton method: bad initial guess $x^{(b)}$ leads to the Newton estimate $x^{(b+1)}$ far away from the desired zero of $F(x)$. Sequence $\cdots, x^{(m)}, x^{(m+1)}, \cdots$, starting with a good guess converges super-exponentially to $x^{*}$. The method diverges if it iterates into the basin of attraction of a local minimum $x^{c}$.

method. While it might very hard to give a precise initial point guess for a long periodic orbit, if our guesses are informed by a good phase-space partition, a rough guess for each point along the desired trajectory might suffice, as for the individual short trajectory segments the errors have no time to explode exponentially.

A cycle of length $n$ is a zero of the $n$-dimensional vector function $F$ :

$$
F(x)=F\left(\begin{array}{c}
x_{1} \\
x_{2} \\
\cdot \\
x_{n}
\end{array}\right)=\left(\begin{array}{c}
x_{1}-f\left(x_{n}\right) \\
x_{2}-f\left(x_{1}\right) \\
\ldots \\
x_{n}-f\left(x_{n-1}\right)
\end{array}\right) .
$$

The relations between the temporal symbol sequences and the spatial layout of the topologically distinct regions of the phase space discussed in chapter 11 enable us to guess location of a series of periodic points along a cycle. Armed with such informed initial guesses we can initiate a NewtonRaphson iteration. The iteration in the Newton's method now takes the form of

$$
\begin{equation*}
\frac{d}{d x} F(x)\left(x^{\prime}-x\right)=-F(x), \tag{17.6}
\end{equation*}
$$

where $\frac{d}{d x} F(x)$ is an $[n \times n]$ matrix:

$$
\frac{d}{d x} F(x)=\left(\begin{array}{ccccc}
1 & & & & -f^{\prime}\left(x_{n}\right)  \tag{17.7}\\
-f^{\prime}\left(x_{1}\right) & 1 & & & \\
& \cdots & 1 & 1 & \\
& & \cdots & -f^{\prime}\left(x_{n-1}\right) & 1
\end{array}\right)
$$

This matrix can easily be inverted numerically by first eliminating the elements below the diagonal. This creates non-zero elements in the $n$th column. We eliminate these and are done. Let us take it step by step for a period 3 cycle. Initially the setup for the Newton step looks like this:

$$
\left(\begin{array}{ccc}
1 & 0 & -f^{\prime}\left(x_{3}\right)  \tag{17.8}\\
-f^{\prime}\left(x_{1}\right) & 1 & 0 \\
0 & -f^{\prime}\left(x_{2}\right) & 1
\end{array}\right)\left(\begin{array}{l}
\delta_{1} \\
\delta_{2} \\
\delta_{3}
\end{array}\right)=\left(\begin{array}{l}
-F_{1} \\
-F_{2} \\
-F_{3}
\end{array}\right)
$$

where $\delta_{i}=x_{i}^{\prime}-x_{i}$ is the correction of our guess for a solution and where $F_{i}=x_{i}-f\left(x_{i-1}\right)$. First we eliminate the below diagonal elements by adding $f^{\prime}\left(x_{1}\right)$ times the first row to the second row, then adding $f^{\prime}\left(x_{2}\right)$ times the second row to the third row. We then have

$$
\begin{align*}
& \left(\begin{array}{llc}
1 & 0 & -f^{\prime}\left(x_{3}\right) \\
0 & 1 & -f^{\prime}\left(x_{1}\right) f^{\prime}\left(x_{3}\right) \\
0 & 0 & 1-f^{\prime}\left(x_{2}\right) f^{\prime}\left(x_{1}\right) f^{\prime}\left(x_{3}\right)
\end{array}\right)\left(\begin{array}{l}
\delta_{1} \\
\delta_{2} \\
\delta_{3}
\end{array}\right)=  \tag{17.9}\\
& \\
& \\
& \left(\begin{array}{c}
-F_{1} \\
-F_{2}-f^{\prime}\left(x_{1}\right) F_{1} \\
-F_{3}-f^{\prime}\left(x_{2}\right) F_{2}-f^{\prime}\left(x_{2}\right) f^{\prime}\left(x_{1}\right) F_{1}
\end{array}\right)=
\end{align*}
$$

The next step is to invert the last element in the diagonal, that is, divide the third row by $1-f^{\prime}\left(x_{2}\right) f^{\prime}\left(x_{1}\right) f^{\prime}\left(x_{3}\right)$. It is clear that if this element is zero at the periodic orbit this step might lead to problems. In many cases this will just mean a slower convergence, but it might throw the Newton iteration completely off. We note that $f^{\prime}\left(x_{2}\right) f^{\prime}\left(x_{1}\right) f^{\prime}\left(x_{3}\right)$ is the stability of the cycle (when the Newton iteration has converged) and that this therefore is not a good method to find marginally stable cycles. We now have

$$
\begin{align*}
& \left(\begin{array}{ccc}
1 & 0 & -f^{\prime}\left(x_{3}\right) \\
0 & 1 & -f^{\prime}\left(x_{1}\right) f^{\prime}\left(x_{3}\right) \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{c}
\delta_{1} \\
\delta_{2} \\
\delta_{3}
\end{array}\right)= \\
&  \tag{17.10}\\
& \\
& \left(\begin{array}{c}
-F_{1} \\
-F_{2}-f^{\prime}\left(x_{1}\right) F_{1} \\
\frac{-F_{3}-f^{\prime}\left(x_{2}\right) F_{2}-f^{\prime}\left(x_{2}\right) f^{\prime}\left(x_{1}\right) F_{1}}{1-f^{\prime}\left(x_{2}\right) f^{\prime}\left(x_{1}\right) f^{\prime}\left(x_{3}\right)}
\end{array}\right)=
\end{align*}
$$

Finally we add $f^{\prime}\left(x_{3}\right)$ times the third row to the first row and $f^{\prime}\left(x_{1}\right) f^{\prime}\left(x_{3}\right)$ times the third row to the second row. On the left hand side the matrix is now the unit matrix, on the right hand side we have the corrections to our initial guess for the cycle, that is, we have gone through one step of the Newton iteration scheme.

When one sets up the Newton iteration on the computer it is not necessary to write the left hand side as a matrix. All one needs is a vector containing the $f^{\prime}\left(x_{i}\right)$ 's, a vector containing the $n$ 'th column, that is the cumulative product of the $f^{\prime}\left(x_{i}\right)$ 's and a vector containing the right hand side. After the iteration the vector containing the right hand side should be the correction to the initial guess.

## 17.4 d-dimensional mappings

Armed with symbolic dynamics informed initial guesses we can utilize the Newton-Raphson iteration in $d$-dimensions as well.

### 17.4.1 Newton's method for $d$-dimensional mappings

Newton's method for 1-dimensional mappings is easily extended to higher dimensions. In this case $f^{\prime}\left(x_{i}\right)$ is a $[d \times d]$ matrix. $\frac{d}{d x} F(x)$ is then an $[n d \times n d]$ matrix. In each of the steps that we went through above we are then manipulating $d$ rows of the left hand side matrix. (Remember that matrices do not commute - always multiply from the left.) In the inversion of the $n$ 'th element of the diagonal we are inverting a $[d \times d]$ matrix $\left(1-\prod f^{\prime}\left(x_{i}\right)\right)$ which can be done if none of the eigenvalues of $\prod f^{\prime}\left(x_{i}\right)$ equals 1 , that is, the cycle must not have any marginally stable directions.

Some $d$-dimensional mappings (such as the Hénon map (3.12)) can be written as 1-dimensional time delay mappings of the form

$$
\begin{equation*}
f\left(x_{i}\right)=f\left(x_{i-1}, x_{i-2}, \ldots, x_{i-d}\right) \tag{17.11}
\end{equation*}
$$

In this case $\frac{d}{d x} F(x)$ is an $[n \times n]$ matrix as in the case of usual 1-dimensional maps but with non-zero matrix elements on $d$ off-diagonals. In the elimination of these off-diagonal elements the last $d$ columns of the matrix will become non-zero and in the final cleaning of the diagonal we will need to invert a $[d \times d]$ matrix. In this respect, nothing is gained numerically by looking at such maps as 1 -dimensional time delay maps.

### 17.5 Flows

(F. Christiansen)

Further complications arise for flows due to the fact that for a periodic orbit the stability eigenvalue corresponding to the flow direction of necessity equals unity; the separation of any two points along a cycle remains unchanged after a completion of the cycle. More unit eigenvalues can arise if the flow satisfies conservation laws, such as the energy invariance for Hamiltonian systems. We now show how such problems are solved by increasing the number of fixed point conditions.

### 17.5.1 Newton's method for flows

A flow is equivalent to a mapping in the sense that one can reduce the flow to a mapping on the Poincaré surface of section. An autonomous flow (2.5) is given as

$$
\begin{equation*}
\dot{x}=v(x), \tag{17.12}
\end{equation*}
$$

The corresponding Jacobian matrix $\mathbf{J}$ (4.30) is obtained by integrating the linearized equation (4.32)

$$
\dot{\mathbf{J}}=\mathbf{A} \mathbf{J}, \quad A_{i j}(x)=\frac{\partial v_{i}(x)}{\partial x_{j}}
$$

along the trajectory. The flow and the corresponding Jacobian matrix are integrated simultaneously, by the same numerical routine. Integrating an initial condition on the Poincaré surface until a later crossing of the same and linearizing around the flow we can write

$$
\begin{equation*}
f\left(x^{\prime}\right) \approx f(x)+\mathbf{J}\left(x^{\prime}-x\right) . \tag{17.13}
\end{equation*}
$$

Notice here, that, even though all of $x^{\prime}, x$ and $f(x)$ are on the Poincaré surface, $f\left(x^{\prime}\right)$ is usually not. The reason for this is that $\mathbf{J}$ corresponds to a specific integration time and has no explicit relation to the arbitrary choice of Poincaré section. This will become important in the extended Newton's method described below.

To find a fixed point of the flow near a starting guess $x$ we must solve the linearized equation

$$
\begin{equation*}
(1-\mathbf{J})\left(x^{\prime}-x\right)=-(x-f(x))=-F(x) \tag{17.14}
\end{equation*}
$$

where $f(x)$ corresponds to integrating from one intersection of the Poincaré surface to another and $\mathbf{J}$ is integrated accordingly. Here we run into problems with the direction along the flow, since - as shown in sect. 8.2.1 - this corresponds to a unit eigenvector of $\mathbf{J}$. The matrix $(1-\mathbf{J})$ does therefore not have full rank. A related problem is that the solution $x^{\prime}$ of (17.14) is not guaranteed to be in the Poincaré surface of section. The two problems are solved simultaneously by adding a small vector along the flow plus an extra equation demanding that $x$ be in the Poincaré surface. Let us for the sake of simplicity assume that the Poincaré surface is a (hyper)-plane, that is, it is given by the linear equation

$$
\begin{equation*}
\left(x-x_{0}\right) \cdot a=0, \tag{17.15}
\end{equation*}
$$

where $a$ is a vector normal to the Poincaré section and $x_{0}$ is any point in the Poincaré section. (17.14) then becomes

$$
\left(\begin{array}{cc}
1-\mathbf{J} & v(x)  \tag{17.16}\\
a & 0
\end{array}\right)\binom{x^{\prime}-x}{\delta T}=\binom{-F(x)}{0} .
$$

The last row in this equation ensures that $x$ will be in the surface of section, and the addition of $v(x) \delta T$, a small vector along the direction of the flow,
ensures that such an $x$ can be found at least if $x$ is sufficiently close to a solution, that is, to a fixed point of $f$.

To illustrate this little trick let us take a particularly simple example; consider a 3 -d flow with the ( $x, y, 0$ )-plane as Poincaré section. Let all trajectories cross the Poincaré section perpendicularly, that is, with $v=\left(0,0, v_{z}\right)$, which means that the marginally stable direction is also perpendicular to the Poincaré section. Furthermore, let the unstable direction be parallel to the $x$-axis and the stable direction be parallel to the $y$-axis. In this case the Newton setup looks as follows

$$
\left(\begin{array}{cccc}
1-\Lambda & 0 & 0 & 0  \tag{17.17}\\
0 & 1-\Lambda_{s} & 0 & 0 \\
0 & 0 & 0 & v_{z} \\
0 & 0 & 1 & 0
\end{array}\right)\left(\begin{array}{c}
\delta_{x} \\
\delta_{y} \\
\delta_{z} \\
\delta t
\end{array}\right)=\left(\begin{array}{c}
-F_{x} \\
-F_{y} \\
-F_{z} \\
0
\end{array}\right) .
$$

If you consider only the upper-left $[3 \times 3]$ matrix (which is what we would have without the extra constraints that we have introduced) then this matrix is clearly not invertible and the equation does not have a unique solution. However, the full $[4 \times 4]$ matrix is invertible, as $\operatorname{det}(\cdot)=v_{z} \operatorname{det}\left(1-\mathbf{J}_{\perp}\right)$, where $\mathbf{J}_{\perp}$ is the monodromy matrix for a surface of section transverse to the orbit, see for ex. (30.2).

For periodic orbits (17.16) generalizes in the same way as (17.7), but with $n$ additional equations - one for each point on the Poincaré surface. The Newton setup looks like this

Solving this equation resembles the corresponding task for maps. However, in the process we will need to invert an $[(d+1) n \times(d+1) n]$ matrix rather than a $[d \times d]$ matrix. The task changes with the length of the cycle.

This method can be extended to take care of the same kind of problems if other eigenvalues of the Jacobian matrix equal 1. This happens if the flow has an invariant of motion, the most obvious example being energy conservation in Hamiltonian systems. In this case we add an extra equation for $x$ to be on the energy shell plus and extra variable corresponding to adding a small vector along the gradient of the Hamiltonian. We then have to solve

$$
\left(\begin{array}{ccc}
1-\mathbf{J} & v(x) & \nabla H(x)  \tag{17.18}\\
a & 0 & 0
\end{array}\right)\left(\begin{array}{c}
x^{\prime}-x \\
\delta t \\
\delta E
\end{array}\right)=\left(\begin{array}{c}
-(x-f(x)) \\
0 \\
0
\end{array}\right)
$$



Figure 17.5: Illustration of the optimal Poincaré surface. The original surface $y=0$ yields a large distance $x-f(x)$ for the Newton iteration. A much better choice is $y=0.7$.
simultaneously with

$$
\begin{equation*}
H\left(x^{\prime}\right)-H(x)=0 . \tag{17.19}
\end{equation*}
$$

This last equation is nonlinear. It is often best to treat this separately in the sense that we really solve this equation in each Newton step. This might mean putting in an additional Newton routine to solve the single step of (17.18) and (17.19) together. One might be tempted to linearize (17.19) and put it into (17.18) to do the two different Newton routines simultaneously, but this will not guarantee a solution on the energy shell. In fact, it may not even be possible to find any solution of the combined linearized equations, if the initial guess is not very good.

### 17.5.2 Newton's method with optimal surface of section

(F. Christiansen)


In some systems it might be hard to find a good starting guess for a fixed point, something that could happen if the topology and/or the symbolic dynamics of the flow is not well understood. By changing the Poincaré section one might get a better initial guess in the sense that $x$ and $f(x)$ are closer together. In figure 17.5 there is an illustration of this. The figure shows a Poincaré section, $y=0$, an initial guess $x$, the corresponding $f(x)$ and pieces of the trajectory near these two points.

If the Newton iteration does not converge for the initial guess $x$ we might have to work very hard to find a better guess, particularly if this is in a high-dimensional system (high-dimensional might in this context mean a Hamiltonian system with 3 degrees of freedom.) But clearly we could easily have a much better guess by simply shifting the Poincaré section to $y=0.7$ where the distance $x-f(x)$ would be much smaller. Naturally, one cannot see by eye the best surface in higher dimensional systems. The way to proceed is as follows: We want to have a minimal distance between our initial guess $x$ and the image of this $f(x)$. We therefore integrate the
flow looking for a minimum in the distance $d(t)=\left|f^{t}(x)-x\right| . d(t)$ is now a minimum with respect to variations in $f^{t}(x)$, but not necessarily with respect to $x$. We therefore integrate $x$ either forward or backward in time. Doing this we minimize $d$ with respect to $x$, but now it is no longer minimal with respect to $f^{t}(x)$. We therefore repeat the steps, alternating between correcting $x$ and $f^{t}(x)$. In most cases this process converges quite rapidly. The result is a trajectory for which the vector $(f(x)-x)$ connecting the two end points is perpendicular to the flow at both points. We can now choose to define a Poincaré surface of section as the hyper-plane that goes through $x$ and is normal to the flow at $x$. In other words the surface of section is determined by

$$
\begin{equation*}
\left(x^{\prime}-x\right) \cdot v(x)=0 . \tag{17.20}
\end{equation*}
$$

Note that $f(x)$ lies on this surface. This surface of section is optimal in the sense that a close return on the surface is really a local minimum of the distance between $x$ and $f^{t}(x)$. But more importantly, the part of the stability matrix that describes linearization perpendicular to the flow is exactly the stability of the flow in the surface of section when $f(x)$ is close to $x$. In this method, the Poincaré surface changes with each iteration of the Newton scheme. Should we later want to put the fixed point on a specific Poincaré surface it will only be a matter of moving along the trajectory.

## Commentary

Remark 17.1 Piece-wise linear maps. The Lozi map (3.14) is linear, and 100,000's of cycles can be easily computed by [2x2] matrix multiplication and inversion.

## Résumé

There is no general computational algorithm that is guaranteed to find all solutions (up to a given period $T_{\max }$ ) to the periodic orbit condition

$$
f^{t+T}(x)=f^{t}(x), \quad T>0
$$

for a general flow or mapping. Due to the exponential divergence of nearby trajectories in chaotic dynamical systems, direct solution of the periodic orbit condition can be numerically very unstable.

A prerequisite for a systematic and complete cycle search is a good (but hard to come by) understanding of the topology of the flow. Usually one starts by - possibly analytic - determination of the equilibria of the flow. Their locations, stabilities, stability eigenvectors and invariant manifolds offer skeletal information about the topology of the flow. Next step
is numerical long-time evolution of "typical" trajectories of the dynamical system under investigation. Such numerical experiments build up the "natural measure", and reveal regions most frequently visited. The periodic orbit searches can then be initialized by taking nearly recurring orbit segments and deforming them into a closed orbits. With a sufficiently good initial guess the Newton-Raphson formula (17.16)

$$
\left(\begin{array}{cc}
1-\mathbf{J} & v(x) \\
a & 0
\end{array}\right)\binom{\delta x}{\delta T}=\binom{f(x)-x}{0}
$$

yields improved estimate $x^{\prime}=x+\delta x, T^{\prime}=T+\delta T$. Iteration then yields the period $T$ and the location of a periodic point $x_{p}$ in the Poincaré surface $\left(x_{p}-x_{0}\right) \cdot a=0$, where $a$ is a vector normal to the Poincaré section at $x_{0}$.

The problem one faces with high-dimensional flows is that their topology is hard to visualize, and that even with a decent starting guess for a point on a periodic orbit, methods like the Newton-Raphson method are likely to fail. Methods that start with initial guesses for a number of points along the cycle, such as the multipoint shooting method of sect. 17.3, are more robust. The relaxation (or variational) methods take this strategy to its logical extreme, and start by a guess of not a few points along a periodic orbit, but a guess of the entire orbit. As these methods are intimately related to variational principles and path integrals, we postpone their introduction to chapter 31.

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## Exercises

Exercise 17.1 Cycles of the Ulam map. Test your cycle-searching routines by computing a bunch of short cycles and their stabilities for the Ulam map

$$
\begin{equation*}
f(x)=4 x(1-x) \tag{17.21}
\end{equation*}
$$

Exercise 17.2 Cycles stabilities for the Ulam map, exact. In exercise 17.1 you should have observed that the numerical results for the cycle stability eigenvalues (4.34) are exceptionally simple: the stability eigenvalue of the $x_{0}=0$ fixed point is 4 , while the eigenvalue of any other $n$-cycle is $\pm 2^{n}$. Prove this. (Hint: the Ulam map can be conjugated to the tent map (11.8). This problem is perhaps too hard, but give it a try - the answer is in many introductory books on nolinear dynamics.)

Exercise 17.3 Stability of billiard cycles. Compute stabilities of few simple cycles.
(a) A simple scattering billiard is the two-disk billiard. It consists of a disk of radius one centered at the origin and another disk of unit radius located at $L+2$. Find all periodic orbits for this system and compute their stabilities. (You might have done this already in exercise 1.2; at least now you will be able to see where you went wrong when you knew nothing about cycles and their extraction.)
(b) Find all periodic orbits and stabilities for a billiard ball bouncing between the diagonal $y=x$ and one of the hyperbola branches $y=1 / x$.


Exercise 17.4 Cycle stability. Add to the pinball simulator of exercise 6.1 a routine that evaluates the expanding eigenvalue for a given cycle.

Exercise 17.5 Pinball cycles. Determine the stability and length of all fundamental domain prime cycles of the binary symbol string lengths up to 5 (or longer) for $R: a=6$ 3-disk pinball.

Exercise 17.6 Newton-Raphson method. Implement the Newton-Raphson method in 2-d and apply it to determination of pinball cycles.

Exercise 17.7 Rössler system cycles. (continuation of exercise 2.7, and exercise 3.1) Determine all cycles up to 5 Poincaré sections returns for the Rössler system (2.14), as well as their stabilities. (Hint: implement (17.16), the multipoint shooting methods for flows; you can cross-check your shortest cycles against the ones listed in table 17.1.)

| $n_{p}$ | $p$ | $y_{p}$ | $z_{p}$ | $\Lambda_{e}$ |
| :---: | :--- | ---: | ---: | ---: |
| 1 | 1 | 6.091768319056803 | 1.299731937639821 | -2.4039535318268 |
| 2 | 01 | 3.915804049621049 | 3.692833386542665 | -3.5120069815161 |
| 3 | 001 | 2.278281031720258 | 7.416480984019008 | -2.3419235232340 |
|  | 011 | 2.932877559129124 | 5.670805943881501 | 5.3449081538885 |
| 4 | 0111 | 3.466758713211455 | 4.506217531477667 | -16.6967406980700 |
| 5 | 01011 | 4.162798782914948 | 3.303903338609633 | -23.1995830097831 |
|  | 01111 | 3.278914359770783 | 4.890452922955567 | 36.8863297988981 |
| 6 | 001011 | 2.122093931936202 | 7.886172854283211 | -6.8576654190825 |
|  | 010111 | 4.059210605826523 | 3.462265228606606 | 61.6490940089068 |
|  | 011111 | 3.361494458061049 | 4.718206217035575 | -92.0825560711089 |
| 7 | 0101011 | 3.842769382372052 | 3.815493592299824 | 77.7611048852412 |
|  | 0110111 | 3.025956697151134 | 5.451444475664179 | -95.1838846735358 |
|  | 0101111 | 4.102255295518855 | 3.395643547170646 | -142.2379888163439 |
|  | 0111111 | 3.327986189581191 | 4.787462810306583 | 218.0283602810993 |

Table 17.1: The Rössler system (2.14): The itinerary $p$, a periodic point $x_{p}=$ $\left(0, y_{p}, z_{p}\right)$ and the expanding eigenvalue $\Lambda_{p}$ for all cycles up to the topological length 7. (Joachim Mathiesen)

Exercise 17.8 Cycle stability, helium. Add to the helium integrator of exercise 2.10 a routine that evaluates the expanding eigenvalue for a given cycle.

Exercise 17.9 Colinear helium cycles. Determine the stability and length of all fundamental domain prime cycles up to symbol sequence length 5 or longer for collinear helium of figure 34.5.

Exercise 17.10 Uniqueness of unstable cycles***. Prove that there exists only one 3-disk prime cycle for a given finite admissible prime cycle symbol string. Hints: look at the Poincaré section mappings; can you show that there is exponential contraction to a unique periodic point with a given itinerary? Exercise 31.1 might be helpful in this effort.

## Exercise 17.11 Inverse iteration method for a Hamiltonian repeller.

Consider the Hénon map (3.12) for area-preserving ("Hamiltonian") parameter value $b=-1$. The coordinates of a periodic orbit of length $n_{p}$ satisfy the equation

$$
\begin{equation*}
x_{p, i+1}+x_{p, i-1}=1-a x_{p, i}^{2}, \quad i=1, \ldots, n_{p} \tag{17.22}
\end{equation*}
$$

with the periodic boundary condition $x_{p, 0}=x_{p, n_{p}}$. Verify that the itineraries and the stabilities of the short periodic orbits for the Hénon repeller (17.22) at $a=6$ are as listed in table 17.2.

Hint: you can use any cycle-searching routine you wish, but for the complete repeller case (all binary sequences are realized), the cycles can be evaluated simply by inverse iteration, using the inverse of (17.22)

$$
x_{p, i}^{\prime \prime}=S_{p, i} \sqrt{\frac{1-x_{p, i+1}^{\prime}-x_{p, i-1}^{\prime}}{a}}, \quad i=1, \ldots, n_{p}
$$

Here $S_{p, i}$ are the signs of the corresponding cycle point coordinates, $S_{p, i}=x_{p, i} /\left|x_{p, i}\right|$.

| $p$ | $\Lambda_{p}$ | $\sum x_{p, i}$ |
| :--- | ---: | ---: |
| 0 | $0.71516752438 \times 10^{1}$ | -0.6076252185107 |
| 1 | $-0.29528463259 \times 10^{1}$ | 0.2742918851774 |
| 10 | $-0.98989794855 \times 10^{1}$ | 0.3333333333333 |
| 100 | $-0.13190727397 \times 10^{3}$ | -0.2060113295833 |
| 110 | $0.55896964996 \times 10^{2}$ | 0.5393446629166 |
| 1000 | $-0.10443010730 \times 10^{4}$ | -0.8164965809277 |
| 1100 | $0.57799826989 \times 10^{4}$ | 0.0000000000000 |
| 1110 | $-0.10368832509 \times 10^{3}$ | 0.8164965809277 |
| 10000 | $-0.76065343718 \times 10^{4}$ | -1.4260322065792 |
| 11000 | $0.44455240007 \times 10^{4}$ | -0.6066540777738 |
| 10100 | $0.77020248597 \times 10^{3}$ | 0.1513755016405 |
| 11100 | $-0.71068835616 \times 10^{3}$ | 0.2484632276044 |
| 11010 | $-0.58949885284 \times 10^{3}$ | 0.8706954728949 |
| 11110 | $0.39099424812 \times 10^{3}$ | 1.0954854155465 |
| 100000 | $-0.54574527060 \times 10^{5}$ | -2.0341342556665 |
| 110000 | $0.32222060985 \times 10^{5}$ | -1.2152504370215 |
| 101000 | $0.51376165109 \times 10^{4}$ | -0.4506624359329 |
| 111000 | $-0.47846146631 \times 10^{4}$ | -0.3660254037844 |
| 110100 | $-0.63939998436 \times 10^{4}$ | 0.3333333333333 |
| 101100 | $-0.63939998436 \times 10^{4}$ | 0.3333333333333 |
| 111100 | $0.39019387269 \times 10^{4}$ | 0.5485837703548 |
| 111010 | $0.10949094597 \times 10^{4}$ | 1.1514633582661 |
| 111110 | $-0.10433841694 \times 10^{4}$ | 1.3660254037844 |

Table 17.2: All periodic orbits up to 6 bounces for the Hamiltonian Hénon mapping (17.22) with $a=6$. Listed are the cycle itinerary, its expanding eigenvalue $\Lambda_{p}$, and its "center of mass". The "center of mass" is listed because it turns out the "center of mass" is often a simple rational or a quadratic irrational.

Exercise 17.12 "Center of mass" puzzle**. Why is the "center of mass", listed in table 17.2, a simple rational every so often?

## Chapter 18

## Cycle expansions

Recycle... It's the Law!
Poster, New York City Department of Sanitation

The Euler product representations of spectral determinants (15.9) and dynamical zeta functions (15.15) are really only a shorthand notation - the zeros of the individual factors are not the zeros of the zeta function, and convergence of such objects is far from obvious. Now we shall give meaning to the dynamical zeta functions and spectral determinants by expanding them as cycle expansions, series representations ordered by increasing topological cycle length, with products in $(15.9),(15.15)$ expanded as sums over pseudocycles, products of $t_{p}$ 's. The zeros of correctly truncated cycle expansions yield the desired eigenvalues, and the expectation values of observables are given by the cycle averaging formulas obtained from the partial derivatives of dynamical zeta functions (or spectral determinants).

### 18.1 Pseudocycles and shadowing

How are periodic orbit formulas such as (15.15) evaluated? We start by computing the lengths and stability eigenvalues of the shortest cycles. This always requires numerical work, such as the Newton's method searches for periodic solutions; we shall assume that the numerics is under control, and that all short cycles up to a given (topological) length have been found. Examples of the data required for application of periodic orbit formulas are the lists of cycles given in tables 31.3 and 17.2. It is important not to miss any short cycles, as the calculation is as accurate as the shortest cycle dropped - including cycles longer than the shortest omitted does not improve the accuracy (more precisely, improves it, but painfully slowly).

Expand the dynamical zeta function (15.15) as a formal power series,

$$
\begin{equation*}
1 / \zeta=\prod_{p}\left(1-t_{p}\right)=1-\sum_{\left\{p_{1} p_{2} \ldots p_{k}\right\}}^{\prime}(-1)^{k+1} t_{p_{1}} t_{p_{2}} \ldots t_{p_{k}} \tag{18.1}
\end{equation*}
$$

where the prime on the sum indicates that the sum is over all distinct nonrepeating combinations of prime cycles. As we shall frequently use such sums, let us denote by $t_{\pi}=(-1)^{k+1} t_{p_{1}} t_{p_{2}} \ldots t_{p_{k}}$ an element of the set of all distinct products of the prime cycle weights $t_{p}$. The formal power series (18.1) is now compactly written as

$$
\begin{equation*}
1 / \zeta=1-\sum_{\pi}^{\prime} t_{\pi} \tag{18.2}
\end{equation*}
$$

For $k>1, \quad t_{\pi}$ are weights of pseudocycles; they are sequences of shorter cycles that shadow a cycle with the symbol sequence $p_{1} p_{2} \ldots p_{k}$ along segments $p_{1}, p_{2}, \ldots, p_{k}$. $\sum^{\prime}$ denotes the restricted sum, for which any given prime cycle $p$ contributes at most once to a given pseudocycle weight $t_{\pi}$.

The pseudocycle weight

$$
\begin{equation*}
t_{\pi}=(-1)^{k+1} \frac{1}{\left|\Lambda_{\pi}\right|} e^{\beta A_{\pi}-s T_{\pi}} z^{n_{\pi}} \tag{18.3}
\end{equation*}
$$

depends on the pseudocycle topological length, integrated observable, period, and stability

$$
\begin{array}{ll}
n_{\pi}=n_{p_{1}}+\ldots+n_{p_{k}}, & T_{\pi}=T_{p_{1}}+\ldots+T_{p_{k}} \\
A_{\pi}=A_{p_{1}}+\ldots+A_{p_{k}}, & \Lambda_{\pi}=\Lambda_{p_{1}} \Lambda_{p_{2}} \cdots \Lambda_{p_{k}} . \tag{18.4}
\end{array}
$$

Throughout this text, the terms "periodic orbit" and "cycle" are used interchangeably; while "periodic orbit" is more precise, "cycle" (which has many other uses in mathematics) is easier on the ear than "pseudo-periodicorbit." We never use acronynims such as UPOs (Unstable Periodic Orbits).

### 18.1.1 Curvature expansions

The simplest example is the pseudocycle sum for a system described by a complete binary symbolic dynamics. In this case the Euler product (15.15) is given by

$$
\begin{aligned}
1 / \zeta= & \left(1-t_{0}\right)\left(1-t_{1}\right)\left(1-t_{01}\right)\left(1-t_{001}\right)\left(1-t_{011}\right) \\
& \left(1-t_{0001}\right)\left(1-t_{0011}\right)\left(1-t_{0111}\right)\left(1-t_{00001}\right)\left(1-t_{00011}\right) \\
& \left(1-t_{00101}\right)\left(1-t_{00111}\right)\left(1-t_{01011}\right)\left(1-t_{01111}\right) \ldots
\end{aligned}
$$

(see table 11.1), and the first few terms of the expansion (18.2) ordered by increasing total pseudocycle length are:

$$
\begin{aligned}
1 / \zeta= & 1-t_{0}-t_{1}-t_{01}-t_{001}-t_{011}-t_{0001}-t_{0011}-t_{0111}-\ldots \\
& +t_{0} t_{1}+t_{0} t_{01}+t_{01} t_{1}+t_{0} t_{001}+t_{0} t_{011}+t_{001} t_{1}+t_{011} t_{1} \\
& -t_{0} t_{01} t_{1}-\ldots
\end{aligned}
$$

We refer to such series representation of a dynamical zeta function or a spectral determinant, expanded as a sum over pseudocycles, and ordered by increasing cycle length and instability, as a cycle expansion.

The next step is the key step: regroup the terms into the dominant fundamental contributions $t_{f}$ and the decreasing curvature corrections $\hat{c}_{n}$, each $\hat{c}_{n}$ split into prime cycles p of length $n_{p}=n_{1}$ grouped together with pseudocycles whose full itineraries build up the itinerary of p. For the binary case this regrouping is given by

$$
\begin{align*}
1 / \zeta= & 1-t_{0}-t_{1}-\left[\left(t_{01}-t_{1} t_{0}\right)\right]-\left[\left(t_{001}-t_{01} t_{0}\right)+\left(t_{011}-t_{01} t_{1}\right)\right] \\
& -\left[\left(t_{0001}-t_{0} t_{001}\right)+\left(t_{0111}-t_{011} t_{1}\right)\right. \\
& \left.+\left(t_{0011}-t_{001} t_{1}-t_{0} t_{011}+t_{0} t_{01} t_{1}\right)\right]-\ldots \\
= & 1-\sum_{f} t_{f}-\sum_{n} \hat{c}_{n} . \tag{18.5}
\end{align*}
$$

All terms in this expansion up to length $n_{p}=6$ are given in table 18.1. We refer to such regrouped series as curvature expansions.

Such separation into "fundamental" and "curvature" parts of cycle expansions is possible only for dynamical systems whose symbolic dynamics has finite grammar. The fundamental cycles $t_{0}, t_{1}$ have no shorter approximants; they are the "building blocks" of the dynamics in the sense that all longer orbits can be approximately pieced together from them. The fundamental part of a cycle expansion is given by the sum of the products of all non-intersecting loops of the associated Markov graph (see sect. 13.3 and examples in sect. 18.3). The terms grouped in brackets are the curvature corrections; the terms grouped in parenthesis are combinations of longer cycles and corresponding sequences of "shadowing" pseudocycles. If all orbits are weighted equally $\left(t_{p}=z^{n_{p}}\right)$, such combinations cancel exactly, and the dynamical zeta function reduces to the topological polynomial (13.21). If the flow is continuous and smooth, orbits of similar symbolic dynamics will traverse the same neighborhoods and will have similar weights, and the weights in such combinations will almost cancel. The utility of cycle expansions of dynamical zeta functions and spectral determinants, lies precisely in this organization into nearly cancelling combinations: cycle expansions are dominated by short cycles, with long cycles giving exponentially decaying corrections.

In the case where we know of no finite grammar symbolic dynamics that would help us organize the cycles, the best thing to use is a stability cutoff which we shall discuss in sect. 18.4. The idea is to truncate the cycle expansion by including only the pseudocycles such that $\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right| \leq$ $\Lambda_{\max }$, with the cutoff $\Lambda_{\max }$ equal to or greater than the most unstable $\Lambda_{p}$ in the data set.

| $-t_{0}$ |  |  |  |
| :--- | :--- | :--- | :--- |
| $-t_{1}$ |  |  |  |
| $-t_{10}$ | $+t_{1} t_{0}$ |  |  |
| $-t_{100}$ | $+t_{10} t_{0}$ |  |  |
| $-t_{101}$ | $+t_{10} t_{1}$ |  |  |
| $-t_{1000}$ | $+t_{100} t_{0}$ |  |  |
| $-t_{1001}$ | $+t_{100} t_{1}$ | $+t_{101} t_{0}$ | $-t_{1} t_{10} t_{0}$ |
| $-t_{1011}$ | $+t_{101} t_{1}$ |  |  |
| $-t_{10000}$ | $+t_{1000} t_{0}$ |  |  |
| $-t_{10001}$ | $+t_{1001} t_{0}$ | $+t_{1000} t_{1}$ | $-t_{0} t_{100} t_{1}$ |
| $-t_{10010}$ | $+t_{100} t_{10}$ |  |  |
| $-t_{10101}$ | $+t_{101} t_{10}$ |  |  |
| $-t_{10011}$ | $+t_{1011} t_{0}$ | $+t_{1001} t_{1}$ | $-t_{0} t_{101} t_{1}$ |
| $-t_{10111}$ | $+t_{1011} t_{1}$ |  |  |
| $-t_{100000}$ | $+t_{10000} t_{0}$ |  |  |
| $-t_{100001}$ | $+t_{10001} t_{0}$ | $+t_{10000} t_{1}$ | $-t_{0} t_{1000} t_{1}$ |
| $-t_{100010}$ | $+t_{10010} t_{0}$ | $+t_{1000} t_{10}$ | $-t_{0} t_{100} t_{10}$ |
| $-t_{100011}$ | $+t_{10011} t_{0}$ | $+t_{10001} t_{1}$ | $-t_{0} t_{1001} t_{1}$ |
| $-t_{100101}$ | $-t_{100110}$ | $+t_{10010} t_{1}$ | $+t_{10110} t_{0}$ |
|  | $+t_{10} t_{1001}$ | $+t_{100} t_{101}$ | $-t_{0} t_{10} t_{101}-t_{1} t_{10} t_{100}$ |
| $-t_{101110}$ | $+t_{10110} t_{1}$ | $+t_{1011} t_{10}$ | $-t_{1} t_{101} t_{10}$ |
| $-t_{100111}$ | $+t_{10011} t_{1}$ | $+t_{10111} t_{0}$ | $-t_{0} t_{1011} t_{1}$ |
| $-t_{101111}$ | $+t_{10111} t_{1}$ |  |  |

Table 18.1: The binary curvature expansion (18.5) up to length 6 , listed in such way that the sum of terms along the $p$ th horizontal line is the curvature $\hat{c}_{p}$ associated with a prime cycle $p$, or a combination of prime cycles such as the $t_{100101}+t_{100110}$ pair.

### 18.1.2 Evaluation of dynamical zeta functions

Cycle expansions of dynamical zeta functions are evaluated numerically by first computing the weights $t_{p}=t_{p}(\beta, s)$ of all prime cycles $p$ of topological length $n_{p} \leq N$ for given fixed $\beta$ and $s$. Denote by subscript ( $i$ ) the $i$ th prime cycle computed, ordered by the topological length $n_{(i)} \leq n_{(i+1)}$. The dynamical zeta function $1 / \zeta_{N}$ truncated to the $n_{p} \leq N$ cycles is computed recursively, by multiplying

$$
1 / \zeta_{(i)}=1 / \zeta_{(i-1)}\left(1-t_{(i)} z^{n_{(i)}}\right),
$$

and truncating the expansion at each step to a finite polynomial in $z^{n}$, $n \leq N$. The result is the $N$ th order polynomial approximation

$$
\begin{equation*}
1 / \zeta_{N}=1-\sum_{n=1}^{N} \hat{c}_{n} z^{n} . \tag{18.6}
\end{equation*}
$$

In other words, a cycle expansion is a Taylor expansion in the dummy variable $z$ raised to the topological cycle length. If both the number of cycles and their individual weights grow not faster than exponentially with the cycle length, and we multiply the weight of each cycle $p$ by a factor $z^{n_{p}}$, the cycle expansion converges for sufficiently small $|z|$.

If the dynamics is given by iterated mapping, the leading zero of (18.6) as function of $z$ yields the leading eigenvalue of the appropriate evolution operator. For continuous time flows, $z$ is a dummy variable that we set to $z=1$, and the leading eigenvalue of the evolution operator is given by the leading zero of (18.6) as function of $s$.

### 18.1.3 Evaluation of traces, spectral determinants

Due to the lack of factorization of the full pseudocycle weight, $\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p_{1} p_{2}}\right) \neq$ $\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p_{1}}\right) \operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p_{2}}\right)$, the cycle expansions for the spectral determinant (15.9) are somewhat less transparent than is the case for the dynamical zeta functions.

We commence the cycle expansion evaluation of a spectral determinant by computing recursively the trace formula (14.9) truncated to all prime cycles $p$ and their repeats such that $n_{p} r \leq N$ :

$$
\begin{align*}
& \left.\operatorname{tr} \frac{z \mathcal{L}}{1-z \mathcal{L}}\right|_{(i)}=\left.\operatorname{tr} \frac{z \mathcal{L}}{1-z \mathcal{L}}\right|_{(i-1)}+n_{(i)} \sum_{r=1}^{n_{(i)} r \leq N} \frac{e^{\left(\beta \cdot A_{(i)}-s T_{(i)}\right) r}}{\left|\prod\left(1-\Lambda_{(i), j}^{r}\right)\right|} z^{n_{(i)} r} \\
& \left.\operatorname{tr} \frac{z \mathcal{L}}{1-z \mathcal{L}}\right|_{N}=\sum_{n=1}^{N} C_{n} z^{n}, \quad C_{n}=\operatorname{tr} \mathcal{L}^{n} \tag{18.7}
\end{align*}
$$

This is done numerically: the periodic orbit data set consists of the list of the cycle periods $T_{p}$, the cycle stability eigenvalues $\Lambda_{p, 1}, \Lambda_{p, 2}, \ldots, \Lambda_{p, d}$, and the cycle averages of the observable $A_{p}$ for all prime cycles $p$ such that $n_{p} \leq N$. The coefficient of $z^{n_{p} r}$ is then evaluated numerically for the given $(\beta, s)$ parameter values. Now that we have an expansion for the trace formula (14.8) as a power series, we compute the $N$ th order approximation to the spectral determinant (15.3),

$$
\begin{equation*}
\left.\operatorname{det}(1-z \mathcal{L})\right|_{N}=1-\sum_{n=1}^{N} Q_{n} z^{n}, \quad Q_{n}=n \text {th cumulant }, \tag{18.8}
\end{equation*}
$$

as follows. The logarithmic derivative relation (15.4) yields

$$
\begin{aligned}
\left(\operatorname{tr} \frac{z \mathcal{L}}{1-z \mathcal{L}}\right) \operatorname{det}(1-z \mathcal{L}) & =-z \frac{d}{d z} \operatorname{det}(1-z \mathcal{L}) \\
\left(C_{1} z+C_{2} z^{2}+\cdots\right)\left(1-Q_{1} z-Q_{2} z^{2}-\cdots\right) & =Q_{1} z+2 Q_{2} z^{2}+3 Q_{3} z^{3} \cdots
\end{aligned}
$$

so the $n$th order term of the spectral determinant cycle (or in this case, the cumulant) expansion is given recursively by the trace formula expansion coefficients

$$
\begin{equation*}
Q_{n}=\frac{1}{n}\left(C_{n}-C_{n-1} Q_{1}-\cdots C_{1} Q_{n-1}\right), \quad Q_{1}=C_{1} \tag{18.9}
\end{equation*}
$$

| $R: a$ | $N$ | .$\quad \operatorname{det}(s-\mathcal{A})$ | $1 / \zeta(s)$ | $1 / \zeta(s)_{3}$-disk |
| ---: | ---: | :--- | :--- | :--- |
|  | 1 | 0.39 | 0.407 |  |
|  | 2 | 0.4105 | 0.41028 | 0.435 |
|  | 3 | 0.410338 | 0.410336 | 0.4049 |
|  | 4 | 0.4103384074 | 0.4103383 | 0.40945 |
|  | 5 | 0.4103384077696 | 0.4103384 | 0.410367 |
|  | 6 | 0.410338407769346482 | 0.4103383 | 0.410338 |
|  | 7 | 0.4103384077693464892 |  | 0.4103396 |
|  | 8 | 0.410338407769346489338468 |  |  |
|  | 9 | 0.4103384077693464893384613074 |  |  |
|  | 10 | 0.4103384077693464893384613078192 |  |  |
|  | 1 | 0.41 |  |  |
|  | 2 | 0.72 |  |  |
|  | 3 | 0.675 |  |  |
|  | 4 | 0.67797 |  |  |
|  | 5 | 0.677921 |  |  |
|  | 6 | 0.6779227 |  |  |
|  | 7 | 0.6779226894 |  |  |
|  | 8 | 0.6779226896002 | 0.677922689599532 |  |
|  | 10 | 0.67792268959953606 |  |  |

Table 18.2: 3-disk repeller escape rates computed from the cycle expansions of the spectral determinant (15.6) and the dynamical zeta function (15.15), as function of the maximal cycle length $N$. The first column indicates the disk-disk center separation to disk radius ratio $R: a$, the second column gives the maximal cycle length used, and the third the estimate of the classical escape rate from the fundamental domain spectral determinant cycle expansion. As for larger disk-disk separations the dynamics is more uniform, the convergence is better for $R: a=6$ than for $R: a=3$. For comparison, the fourth column lists a few estimates from from the fundamental domain dynamical zeta function cycle expansion (18.5), and the fifth from the full 3-disk cycle expansion (18.32). The convergence of the fundamental domain dynamical zeta function is significantly slower than the convergence of the corresponding spectral determinant, and the full (unfactorized) 3-disk dynamical zeta function has still poorer convergence. (P.E. Rosenqvist.)

Given the trace formula (18.7) truncated to $z^{N}$, we now also have the spectral determinant truncated to $z^{N}$.

The same program can also be reused to compute the dynamical zeta function cycle expansion (18.6), by replacing $\Pi\left(1-\Lambda_{(i), j}^{r}\right)$ in (18.7) by the product of expanding eigenvalues $\Lambda_{(i)}=\prod_{e} \Lambda_{(i), e}$, (see sect. 15.3).

The calculation of the leading eigenvalue of a given evolution operator is now straightforward. After the prime cycles and the pseudocycles have been grouped into subsets of equal topological length, the dummy variable can be set equal to $z=1$. With $z=1$, expansion (18.8) is the cycle expansion for (15.6), the spectral determinant $\operatorname{det}(s-\mathcal{A})$. We vary $s$ in cycle weights, and determine the eigenvalue $s_{\alpha}$ by finding $s=s_{\alpha}$ for which (18.8) vanishes. The convergence of a leading eigenvalue for a nice hyperbolic system is illustrated by the listing of pinball escape rate $\gamma$ estimates computed from truncations of (18.5) and (18.8) to different maximal cycle lengths, table 18.2.

The pleasant surprise is that the coefficients in these expansions can be proven to fall off exponentially or even faster, due to analyticity of $\operatorname{det}(s-$


Figure 18.1: Examples of the complex $s$ plane scans: contour plots of the logarithm of the absolute values of (a) $1 / \zeta(s)$, (b) spectral determinant $\operatorname{det}(s-\mathcal{A})$ for the 3-disk system, separation $a: R=6, A_{1}$ subspace are evaluated numerically. The eigenvalues of the evolution operator $\mathcal{L}$ are given by the centers of elliptic neighborhoods of the rapidly narrowing rings. While the dynamical zeta function is analytic on a strip $\operatorname{Im} s \geq-1$, the spectral determinant is entire and reveals further families of zeros. (P.E. Rosenqvist)
$\mathcal{A})$ or $1 / \zeta(s)$ for $s$ values well beyond those for which the corresponding trace formula diverges.

### 18.1.4 Newton algorithm for determination of the evolution operator eigenvalues

- 

The cycle expansions of spectral determinants yield the eigenvalues of the evolution operator beyond the leading one. A convenient way to search for these is by plotting either the absolute magnitude $\ln \mid \operatorname{det}(1-$ $\mathcal{L}) \mid$ or the phase of spectral determinants and dynamical zeta functions as functions of complex $s$. The eye is guided to the zeros of spectral determinants and dynamical zeta functions by means of complex $s$ plane contour plots, with different intervals of the absolute value of the function under investigation assigned different colors; zeros emerge as centers of elliptic neighborhoods of rapidly changing colors. Detailed scans of the whole area of the complex $s$ plane under investigation and searches for the zeros of spectral determinants, figure 18.1, reveal complicated patterns of resonances even for something so simple as the 3 -disk game of pinball. With a good starting guess (such as a location of a zero suggested by the complex $s$ scan of figure 18.1), a zero $1 / \zeta(s)=0$ can now be easily determined by standard numerical methods, such as the iterative Newton algorithm (17.4)

$$
\begin{equation*}
s_{n+1}=s_{n}-\left(\zeta\left(s_{n}\right) \frac{\partial}{\partial s} \zeta^{-1}\left(s_{n}\right)\right)^{-1}=s_{n}-\frac{1 / \zeta\left(s_{n}\right)}{\langle T\rangle_{\zeta}} . \tag{18.10}
\end{equation*}
$$

The derivative of $1 / \zeta(s)$ required for the Newton iteration is given by the cycle expansion (18.18) that we need to evaluate anyhow, as $\langle T\rangle_{\zeta}$ enters our cycle averaging formulas.

Figure 18.2: The eigenvalue condition is satisfied on the curve $F=0$ the $(\beta, s)$ plane. The expectation value of the observable (10.12) is given by the slope of the curve.


### 18.2 Cycle formulas for dynamical averages

The eigenvalue condition in any of the three forms that we have given so far - the level sum (19.18), the dynamical zeta function (18.2), the spectral determinant (18.8):

$$
\begin{align*}
& 1=\sum_{i}^{(n)} t_{i}, \quad t_{i}=t_{i}(\beta, s(\beta))=\frac{1}{\left|\Lambda_{i}\right|} e^{\beta \cdot A_{i}-s(\beta) T_{i}}  \tag{18.11}\\
& 0=1-\sum_{\pi}^{\prime} t_{\pi}, \quad t_{\pi}=t_{\pi}(z, \beta, s(\beta))  \tag{18.12}\\
& 0=1-\sum_{n=1}^{\infty} Q_{n}, \quad Q_{n}=Q_{n}(\beta, s(\beta)), \tag{18.13}
\end{align*}
$$

is an implicit equation for the eigenvalue $s=s(\beta)$ of form $F(\beta, s(\beta))=0$. The eigenvalue $s=s(\beta)$ as a function of $\beta$ is sketched in figure 18.2; the eigenvalue condition is satisfied on the curve $F=0$. The cycle averaging formulas for the slope and the curvature of $s(\beta)$ are obtained by taking derivatives of the eigenvalue condition. Evaluated along $F=0$, the first derivative leads to

$$
\begin{align*}
0 & =\frac{d}{d \beta} F(\beta, s(\beta)) \\
& =\frac{\partial F}{\partial \beta}+\left.\frac{d s}{d \beta} \frac{\partial F}{\partial s}\right|_{s=s(\beta)} \quad \Longrightarrow \quad \frac{d s}{d \beta}=-\frac{\partial F}{\partial \beta} / \frac{\partial F}{\partial s}, \tag{18.14}
\end{align*}
$$

and the second derivative of $F(\beta, s(\beta))=0$ yields

$$
\begin{equation*}
\frac{d^{2} s}{d \beta^{2}}=-\left[\frac{\partial^{2} F}{\partial \beta^{2}}+2 \frac{d s}{d \beta} \frac{\partial^{2} F}{\partial \beta \partial s}+\left(\frac{d s}{d \beta}\right)^{2} \frac{\partial^{2} F}{\partial s^{2}}\right] / \frac{\partial F}{\partial s} \tag{18.15}
\end{equation*}
$$

Denoting by

$$
\langle A\rangle_{F}=-\left.\frac{\partial F}{\partial \beta}\right|_{\beta, s=s(\beta)}, \quad\langle T\rangle_{F}=\left.\frac{\partial F}{\partial s}\right|_{\beta, s=s(\beta)}
$$

$$
\begin{equation*}
\left\langle(A-\langle A\rangle)^{2}\right\rangle_{F}=\left.\frac{\partial^{2} F}{\partial \beta^{2}}\right|_{\beta, s=s(\beta)} \tag{18.16}
\end{equation*}
$$

respectively the mean cycle expectation value of $A$ and the mean cycle period computed from the $F(\beta, s(\beta))=0$ condition, we obtain the cycle averaging formulas for the expectation value of the observable (10.12) and its variance

$$
\begin{align*}
\langle a\rangle & =\frac{\langle A\rangle_{F}}{\langle T\rangle_{F}} \\
\left\langle(a-\langle a\rangle)^{2}\right\rangle & =\frac{1}{\langle T\rangle_{F}}\left\langle(A-\langle A\rangle)^{2}\right\rangle_{F} . \tag{18.17}
\end{align*}
$$

These formulas are the central result of the periodic orbit theory. As we shall see below, for each choice of the eigenvalue condition function $F(\beta, s)$ in (19.18), (18.2) and (18.8), the above quantities have explicit cycle expansions.

### 18.2.1 Dynamical zeta function cycle expansions

For the dynamical zeta function condition (18.12), the cycle averaging formulas (18.14), (18.17) require evaluation of the derivatives of dynamical zeta function at a given eigenvalue. Substituting the cycle expansion (18.2) for dynamical zeta function we obtain

$$
\begin{align*}
\langle A\rangle_{\zeta} & :=-\frac{\partial}{\partial \beta} \frac{1}{\zeta}=\sum^{\prime} A_{\pi} t_{\pi}  \tag{18.18}\\
\langle T\rangle_{\zeta} & :=\frac{\partial}{\partial s} \frac{1}{\zeta}=\sum^{\prime} T_{\pi} t_{\pi}, \quad\langle n\rangle_{\zeta}:=-z \frac{\partial}{\partial z} \frac{1}{\zeta}=\sum^{\prime} n_{\pi} t_{\pi}
\end{align*}
$$

where the subscript in $\langle\cdots\rangle_{\zeta}$ stands for the dynamical zeta function average over prime cycles, $A_{\pi}, T_{\pi}$, and $n_{\pi}$ are evaluated on pseudocycles (18.4), and pseudocycle weights $t_{\pi}=t_{\pi}(z, \beta, s(\beta))$ are evaluated at the eigenvalue $s(\beta)$. In most applications, $s(\beta)$ is typically the leading eigenvalue.

For bounded flows the leading eigenvalue (the escape rate) vanishes, $s(0)=0$, so

$$
\begin{equation*}
\langle A\rangle_{\zeta}=\sum_{\pi}^{\prime}(-1)^{k+1} \frac{A_{p_{1}}+A_{p_{2}} \cdots+A_{p_{k}}}{\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right|}, \tag{18.19}
\end{equation*}
$$

and similarly for $\langle T\rangle_{\zeta},\langle n\rangle_{\zeta}$. For example, for the complete binary symbolic dynamics the mean cycle period $\langle T\rangle_{\zeta}$ is given by

$$
\begin{align*}
\langle T\rangle_{\zeta}= & \frac{T_{0}}{\left|\Lambda_{0}\right|}+\frac{T_{1}}{\left|\Lambda_{1}\right|}+\left(\frac{T_{01}}{\left|\Lambda_{01}\right|}-\frac{T_{0}+T_{1}}{\left|\Lambda_{0} \Lambda_{1}\right|}\right) \\
& +\left(\frac{T_{001}}{\left|\Lambda_{001}\right|}-\frac{T_{01}+T_{0}}{\left|\Lambda_{01} \Lambda_{0}\right|}\right)+\left(\frac{T_{011}}{\left|\Lambda_{011}\right|}-\frac{T_{01}+T_{1}}{\left|\Lambda_{01} \Lambda_{1}\right|}\right)+. \tag{18.20}
\end{align*}
$$

Note that the cycle expansions for averages are grouped into the same shadowing combinations as the dynamical zeta function cycle expansion (18.5), with nearby pseudocycles nearly cancelling each other.

The cycle averaging formulas for the expectation value of the observable $\langle a\rangle$ follow by substitution into (18.17). Assuming zero mean drift $\langle a\rangle=0$, the cycle expansion for the variance $\left\langle(A-\langle A\rangle)^{2}\right\rangle_{\zeta}$ is given by

$$
\begin{equation*}
\left\langle A^{2}\right\rangle_{\zeta}=\sum^{\prime}(-1)^{k+1} \frac{\left(A_{p_{1}}+A_{p_{2}} \cdots+A_{p_{k}}\right)^{2}}{\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right|} \tag{18.21}
\end{equation*}
$$

### 18.2.2 Spectral determinant cycle expansions

The dynamical zeta function cycle expansions have a particularly simple structure, with the shadowing apparent already by a term-by-term inspection of table 18.2. For "nice" hyperbolic systems the shadowing ensures exponential convergence of the dynamical zeta function cycle expansions. This, however, is not the best achievable convergence. As has been explained in chapter 16, for such systems the spectral determinant constructed from the same cycle data base is entire, and its cycle expansion converges faster than exponentially. Hence in practice, the best convergence is attained by the spectral determinant cycle expansion (18.13) and its derivatives.

The $\partial / \partial s, \partial / \partial \beta$ derivatives are in this case computed recursively, by taking derivatives of the spectral determinant cycle expansion contributions (18.9) and (18.7). The cycle averaging formulas formulas are exact, and highly convergent for nice hyperbolic dynamical systems. An example of its utility is the cycle expansion formula for the Lyapunov exponent of sect. 18.2.4. Further applications of cycle expansions will be discussed in chapter 19.

### 18.2.3 Continuous vs. discrete mean return time

The mean cycle period $\langle T\rangle_{\zeta}$ fixes the normalization of the unit of time; it can be interpreted as the average near recurrence or the average first return time. For example, if we have evaluated a billiard expectation value $\langle a\rangle$ in terms of continuous time, and would like to also have the corresponding average $\langle a\rangle_{\text {dscr }}$ measured in discrete time given by the number of reflections off billiard walls, the two averages are related by

$$
\begin{equation*}
\langle a\rangle_{\mathrm{dscr}}=\langle a\rangle\langle T\rangle_{\zeta} /\langle n\rangle_{\zeta}, \tag{18.22}
\end{equation*}
$$

where $\langle n\rangle_{\zeta}$ is the average of the number of bounces $n_{p}$ along the cycle $p$.

### 18.2.4 Cycle expansion formula for Lyapunov exponents

In sect. 10.3 we defined the Lyapunov exponent for a 1- $d$ mapping, related it to the leading eigenvalue of an evolution operator and promised to evaluate it. Now we are finally in position to deliver on our promise.

The cycle averaging formula (18.19) yields a closed expression for the Lyapunov exponent in terms of prime cycles:

$$
\begin{equation*}
\lambda=\frac{1}{\langle n\rangle_{\zeta}} \sum^{\prime}(-1)^{k+1} \frac{\log \left|\Lambda_{p_{1}}\right|+\cdots+\log \left|\Lambda_{p_{k}}\right|}{\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right|} \tag{18.23}
\end{equation*}
$$

For a repeller, the $1 /\left|\Lambda_{p}\right|$ weights are replaced by normalized measure (19.10) $\exp \left(\gamma n_{p}\right) /\left|\Lambda_{p}\right|$, where $\gamma$ is the escape rate.

We mention here without proof that for 2-d Hamiltonian flows such as our game of pinball there is only one expanding eigenvalue and (18.23) applies as it stands.
in depth:
chapter H.1, p. 689

### 18.3 Cycle expansions for finite alphabets

A finite Markov graph like the one given in figure 13.3(d) is a compact encoding of the transition or the Markov matrix for a given subshift. It is a sparse matrix, and the associated determinant (13.17) can be written down by inspection: it is the sum of all possible partitions of the graph into products of non-intersecting loops, with each loop carrying a minus sign:

$$
\operatorname{det}(1-T)=1-t_{0}-t_{0011}-t_{0001}-t_{00011}+t_{0} t_{0011}+t_{0011} t_{0001}(18.24)
$$

The simplest application of this determinant is to the evaluation of the topological entropy; if we set $t_{p}=z^{n_{p}}$, where $n_{p}$ is the length of the $p$ cycle, the determinant reduces to the topological polynomial (13.18).

The determinant (18.24) is exact for the finite graph figure 13.3(e), as well as for the associated transfer operator of example 10.1. For the associated (infinite dimensional) evolution operator, it is the beginning of the cycle expansion of the corresponding dynamical zeta function:

$$
\begin{align*}
& 1 / \zeta=1-t_{0}-t_{0011}-t_{0001}+t_{0001} t_{0011} \\
&-\left(t_{00011}-t_{0} t_{0011}+\ldots \text { curvatures }\right) \ldots \tag{18.25}
\end{align*}
$$

The cycles 0,0001 and 0011 are the fundamental cycles introduced in (18.5); they are not shadowed by any combinations of shorter cycles, and are the basic building blocks of the dynamics. All other cycles appear together with their shadows (for example, $t_{00011}-t_{0} t_{0011}$ combination is of that type) and yield exponentially small corrections for hyperbolic systems.

For the cycle counting purposes both $t_{a b}$ and the pseudocycle combination $t_{a+b}=t_{a} t_{b}$ in (18.2) have the same weight $z^{n_{a}+n_{b}}$, so all curvature combinations $t_{a b}-t_{a} t_{b}$ vanish exactly, and the topological polynomial (13.21) offers a quick way of checking the fundamental part of a cycle expansion.

Since for finite grammars the topological zeta functions reduce to polynomials, we are assured that there are just a few fundamental cycles and that all long cycles can be grouped into curvature combinations. For example, the fundamental cycles in exercise 11.5 are the three 2 -cycles which bounce back and forth between two disks and the two 3 -cycles which visit every disk. It is only after these fundamental cycles have been included that a cycle expansion is expected to start converging smoothly, that is, only for $n$ larger than the lengths of the fundamental cycles are the curvatures $\hat{c}_{n}$, a measure of the deviations between long orbits and their short cycle approximants, expected to fall off rapidly with $n$.

### 18.4 Stability ordering of cycle expansions

There is never a second chance. Most often there is not even the first chance.
John Wilkins
(C.P. Dettmann and P. Cvitanović)

Most dynamical systems of interest have no finite grammar, so at any order in $z$ a cycle expansion may contain unmatched terms which do not fit neatly into the almost cancelling curvature corrections. Similarly, for intermittent systems that we shall discuss in chapter 21, curvature corrections are in general not small, so again the cycle expansions may converge slowly. For such systems schemes which collect the pseudocycle terms according to some criterion other than the topology of the flow may converge more quickly than expansions based on the topological length.

All chaotic systems exhibit some degree of shadowing, and a good truncation criterion should do its best to respect the shadowing at least approximately. If a long cycle is shadowed by two or more shorter cycles and the flow is smooth, the period and the action will be additive in sense that the period of the longer cycle is approximately the sum of the shorter cycle periods. Similarly, stability is multiplicative, so shadowing is approximately preserved by including all terms with pseudocycle stability

$$
\begin{equation*}
\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right| \leq \Lambda_{\max } \tag{18.26}
\end{equation*}
$$

and ignoring all more unstable pseudocycles.
Two such schemes for ordering cycle expansions which approximately respect shadowing are truncations by the pseudocycle period (or action) and the stability ordering that we shall discuss here. In these schemes a dynamical zeta function or a spectral determinant is expanded keeping all terms for which the period, action or stability for a combination of cycles (pseudocycle) is less than a given cutoff.

The two settings in which the stability ordering may be preferable to the ordering by topological cycle length are the cases of bad grammar and of intermittency.

### 18.4.1 Stability ordering for bad grammars

For generic flows it is often not clear what partition of the phase space generates the "optimal" symbolic dynamics. Stability ordering does not require understanding dynamics in such detail: if you can find the cycles, you can use stability ordered cycle expansions. Stability truncation is thus easier to implement for a generic dynamical system than the curvature expansions (18.5) which rely on finite subshift approximations to a given flow.

Cycles can be detected numerically by searching a long trajectory for near recurrences. The long trajectory method for finding cycles preferentially finds the least unstable cycles, regardless of their topological length. Another practical advantage of the method (in contrast to the Newton method searches) is that it only finds cycles in a given connected ergodic component of phase space, even if isolated cycles or other ergodic regions exist elsewhere in the phase space.

Why should stability ordered cycle expansion of a dynamical zeta function converge better than the rude trace formula (19.9)? The argument has essentially already been laid out in sect. 13.7: in truncations that respect shadowing most of the pseudocycles appear in shadowning combinations and nearly cancel, and only the relatively small subset affected by the longer and longer pruning rules appears not shadowed. So the error is typically of the order of $1 / \Lambda$, smaller by factor $e^{h T}$ than the trace formula (19.9) error, where $h$ is the entropy and $T$ typical cycle length for cycles of stability $\Lambda$.

### 18.4.2 Smoothing



The breaking of exact shadowing cancellations deserves further comment. Partial shadowing which may be present can be (partially) restored by smoothing the stability ordered cycle expansions by replacing the $1 / \Lambda$ weigth for each term with pseudocycle stability $\Lambda=\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}$ by
$f(\Lambda) / \Lambda$. Here, $f(\Lambda)$ is a monotonically decreasing function from $f(0)=1$ to $f\left(\Lambda_{\max }\right)=0$. No smoothing corresponds to a step function.

A typical "shadowing error" induced by the cutoff is due to two pseudocycles of stability $\Lambda$ separated by $\Delta \Lambda$, and whose contribution is of opposite signs. Ignoring possible weighting factors the magnitude of the resulting term is of order $1 / \Lambda-1 /(\Lambda+\Delta \Lambda) \approx \Delta \Lambda / \Lambda^{2}$. With smoothing there is an extra term of the form $f^{\prime}(\Lambda) \Delta \Lambda / \Lambda$, which we want to minimise. A reasonable guess might be to keep $f^{\prime}(\Lambda) / \Lambda$ constant and as small as possible, that is

$$
f(\Lambda)=1-\left(\frac{\Lambda}{\Lambda_{\max }}\right)^{2}
$$

The results of a stability ordered expansion should always be tested for robustness by varying the cutoff. If this introduces significant variations, smoothing is probably necessary.

### 18.4.3 Stability ordering for intermittent flows

- 

Longer but less unstable cycles can give larger contributions to a cycle expansion than short but highly unstable cycles. In such situation truncation by length may require an exponentially large number of very unstable cycles before a significant longer cycle is first included in the expansion. This situation is best illustrated by intermittent maps that we shall study in detail in chapter 21, the simplest of which is the Farey map

$$
f(x)=\left\{\begin{array}{lll}
x /(1-x) & 0 \leq x \leq 1 / 2 & L  \tag{18.27}\\
(1-x) / x & 1 / 2 \leq x \leq 1 & R,
\end{array}\right.
$$

a map which will reappear in the intermittency chapter 21, and in chapter 24 , in context of circle maps.

For this map the symbolic dynamics is of complete binary type, so lack of shadowing is not due to lack of a finite grammar, but rather to the intermittency caused by the existence of the marginal fixed point $x_{L}=0$, for which the stability equals $\Lambda_{L}=1$. This fixed point does not participate directly in the dynamics and is omitted from cycle expansions. Its presence is felt in the stabilities of neighboring cycles with $n$ consecutive repeats of the symbol $L$ 's whose stability falls of only as $\Lambda \sim n^{2}$, in contrast to the most unstable cycles with $n$ consecutive $R$ 's which are exponentially unstable, $\left|\Lambda_{L R^{n}}\right| \sim[(\sqrt{5}+1) / 2]^{2 n}$.

The symbolic dynamics is of complete binary type, so a quick count in the style of sect. 13.5.2 leads to a total of $74,248,450$ prime cycles of length 30 or less, not including the marginal point $x_{L}=0$. Evaluating a cycle expansion to this order would be no mean computational feat. However, the


Figure 18.3: Comparison of cycle expansion truncation schemes for the Farey map (18.27); the deviation of the truncated cycles expansion for $\left|1 / \zeta_{N}(0)\right|$ from the exact flow conservation value $1 / \zeta(0)=0$ is a measure of the accuracy of the truncation. The jagged line is logarithm of the stability ordering truncation error; the smooth line is smoothed according to sect. 18.4.2; the diamonds indicate the error due the topological length truncation, with the maximal cycle length $N$ shown. They are placed along the stability cutoff axis at points determined by the condition that the total number of cycles is the same for both truncation schemes.
least unstable cycle omitted has stability of roughly $\Lambda_{R L^{30}} \sim 30^{2}=900$, and so amounts to a $0.1 \%$ correction. The situation may be much worse than this estimate suggests, because the next, $R L^{31}$ cycle contributes a similar amount, and could easily reinforce the error. Adding up all such omitted terms, we arrive at an estimated error of about $3 \%$, for a cycle-length truncated cycle expansion based on more than $10^{9}$ pseudocycle terms! On the other hand, truncating by stability at say $\Lambda_{\max }=3000$, only 409 prime cycles suffice to attain the same accuracy of about $3 \%$ error (see figure 18.3).

As the Farey map maps the unit interval onto itself, the leading eigenvalue of the Perron-Frobenius operator should equal $s_{0}=0$, so $1 / \zeta(0)=0$. Deviation from this exact result serves as an indication of the convergence of a given cycle expansion. The errors of different truncation schemes are indicated in figure 18.3. We see that topological length truncation schemes are hopelessly bad in this case; stability length truncations are somewhat better, but still rather bad. In simple cases like this one, where intermittency is caused by a single marginal fixed point, the convergence can be improved by going to infinite alphabets.

### 18.5 Dirichlet series

The most patient reader will thank me for compressing so much nonsense and falsehood into a few lines. Gibbon

$$
\begin{equation*}
f(s)=\sum_{j=1}^{\infty} a_{j} e^{-\lambda_{j} s} \tag{18.28}
\end{equation*}
$$

where $s, a_{j}$ are complex numbers, and $\left\{\lambda_{j}\right\}$ is a monotonically increasing series of real numbers $\lambda_{1}<\lambda_{2}<\cdots<\lambda_{j}<\cdots$. A classical example of a Dirichlet series is the Riemann zeta function for which $a_{j}=1, \lambda_{j}=\ln j$. In the present context, formal series over individual pseudocycles such as (18.2) ordered by the increasing pseudocycle periods are often Dirichlet series. For example, for the pseudocycle weight (18.3), the Dirichlet series is obtained by ordering pseudocycles by increasing periods $\lambda_{\pi}=T_{p_{1}}+T_{p_{2}}+$ $\ldots+T_{p_{k}}$, with the coefficients

$$
a_{\pi}=\frac{e^{\beta \cdot\left(A_{p_{1}}+A_{p_{2}}+\ldots+A_{p_{k}}\right)}}{\left|\Lambda_{p_{1}} \Lambda_{p_{2}} \ldots \Lambda_{p_{k}}\right|} d_{\pi},
$$

where $d_{\pi}$ is a degeneracy factor, in the case that $d_{\pi}$ pseudocycles have the same weight.

If the series $\sum\left|a_{j}\right|$ diverges, the Dirichlet series is absolutely convergent for $\operatorname{Re} s>\sigma_{a}$ and conditionally convergent for $\operatorname{Re} s>\sigma_{c}$, where $\sigma_{a}$ is the abscissa of absolute convergence

$$
\begin{equation*}
\sigma_{a}=\lim _{N \rightarrow \infty} \sup \frac{1}{\lambda_{N}} \ln \sum_{j=1}^{N}\left|a_{j}\right|, \tag{18.29}
\end{equation*}
$$

and $\sigma_{c}$ is the abscissa of conditional convergence

$$
\begin{equation*}
\sigma_{c}=\lim _{N \rightarrow \infty} \sup \frac{1}{\lambda_{N}} \ln \left|\sum_{j=1}^{N} a_{j}\right| . \tag{18.30}
\end{equation*}
$$

We shall encounter another example of a Dirichlet series in the semiclassical quantization chapter 28, where the inverse Planck constant is a complex variable $s=i / \hbar, \lambda_{\pi}=S_{p_{1}}+S_{p_{2}}+\ldots+S_{p_{k}}$ is the pseudocycle action, and $a_{\pi}=1 / \sqrt{\left|\Lambda_{p_{1}} \Lambda_{p_{2}} \ldots \Lambda_{p_{k}}\right|}$ (times possible degeneracy and topological phase factors). As the action is in general not a linear function of energy (except for billiards and for scaling potentials, where a variable $s$ can be extracted from $S_{p}$ ), semiclassical cycle expansions are Dirichlet series in variable $s=i / \hbar$ but not in $E$, the complex energy variable.

## Commentary

Remark 18.1 Pseudocycle expansions. Bowen's introduction of shadowing $\epsilon$ pseudoorbits [1.16] was a significant contribution to Smale's theory. Expression
"pseudoorbits" seems to have been introduced in the Parry and Pollicott's 1983 paper [15.5]. Following them M. Berry [18.9] had used the expression "pseudoorbits" in his 1986 paper on Riemann zeta and quantum chaos. Cycle and curvature expansions of dynamical zeta functions and spectral determinants were introduced in refs. [18.10, 18.2]. Some literature [15.14] refers to the pseudoorbits as "composite orbits", and to the cycle expansions as "Dirichlet series" (see also remark 18.6 and sect. 18.5).

Remark 18.2 Cumulant expansion. To statistical mechanician the curvature expansions are very reminiscent of cumulant expansions. Indeed, (18.9) is the standard Plemelj-Smithies cumulant formula (K.28) for the Fredholm determinant, discussed in more detail in appendix K.

Remark 18.3 Exponential growth of the number of cycles. Going from $N_{n} \approx$ $N^{n}$ periodic points of length $n$ to $M_{n}$ prime cycles reduces the number of computations from $N_{n}$ to $M_{n} \approx N^{n-1} / n$. Use of discrete symmetries (chapter 22) reduces the number of $n$th level terms by another factor. While the formulation of the theory from the trace (14.24) to the cycle expansion (18.5) thus does not eliminate the exponential growth in the number of cycles, in practice only the shortest cycles are used, and for them the computational labor saving can be significant.

Remark 18.4 Shadowing cycle-by-cycle. A glance at the low order curvatures in the table 18.1 leads to a temptation of associating curvatures to individual cycles, such as $\hat{c}_{0001}=t_{0001}-t_{0} t_{001}$. Such combinations tend to be numerically small (see for example ref. [18.3], table 1). However, splitting $\hat{c}_{n}$ into individual cycle curvatures is not possible in general [11.14]; the first example of such ambiguity in the binary cycle expansion is given by the $t_{001011}, t_{010011} 0 \leftrightarrow 1$ symmetric pair of 6 -cycles; the counterterm $t_{001} t_{011}$ in table 18.1 is shared by the two cycles.

Remark 18.5 Stability ordering. The stability ordering was introduced by Dahlqvist and Russberg [18.12] in a study of chaotic dynamics for the $\left(x^{2} y^{2}\right)^{1 / a}$ potential. The presentation here runs along the lines of Dettmann and Morriss [18.13] for the Lorentz gas which is hyperbolic but the symbolic dynamics is highly pruned, and Dettmann and Cvitanović [18.14] for a family of intermittent maps. In the applications discussed in the above papers, the stability ordering yields a considerable improvement over the topological length ordering.

Remark 18.6 Are cycle expansions Dirichlet series? Even though some literature [15.14] refers to cycle expansions as "Dirichlet series", they are not Dirichlet series. Cycle expansions collect contributions of individual cycles into groups that correspond to the coefficients in cumulant expansions of spectral determinants, and the convergence of cycle expansions is controlled by general properties of spectral determinants. Dirichlet series order cycles by their periods or actions, and are only conditionally convergent in regions of interest. The abscissa of absolute convergence is in this context called the "entropy barrier"; contrary to the frequently voiced anxieties, this number does not necessarily have much to do with the actual convergence of the theory.

## Résumé

A cycle expansion is a series representation of a dynamical zeta function, trace formula or a spectral determinant, with products in (15.15), (30.18) expanded as sums over pseudocycles, products of the prime cycle weigths $t_{p}$.

If a flow is hyperbolic and has a topology of a Smale horseshoe, the associated zeta functions have nice analytic structure: the dynamical zeta functions are holomorphic, the spectral determinants are entire, and the spectrum of the evolution operator is discrete. The situation is considerably more reassuring than what practitioners of quantum chaos fear; there is no "abscissa of absolute convergence" and no "entropy barier", the exponential proliferation of cycles is no problem, spectral determinants are entire and converge everywhere, and the topology dictates the choice of cycles to be used in cycle expansion truncations.

The basic observation is that the motion in dynamical systems of few degrees of freedom is in this case organized around a few fundamental cycles, with the cycle expansion of the Euler product

$$
1 / \zeta=1-\sum_{f} t_{f}-\sum_{n} \hat{c}_{n}
$$

regrouped into dominant fundamental contributions $t_{f}$ and decreasing curvature corrections $\hat{c}_{n}$. The fundamental cycles $t_{f}$ have no shorter approximants; they are the "building blocks" of the dynamics in the sense that all longer orbits can be approximately pieced together from them. A typical curvature contribution to $\hat{c}_{n}$ is a difference of a long cycle $\{a b\}$ minus its shadowing approximation by shorter cycles $\{a\}$ and $\{b\}$ :

$$
t_{a b}-t_{a} t_{b}=t_{a b}\left(1-t_{a} t_{b} / t_{a b}\right)
$$

The orbits that follow the same symbolic dynamics, such as $\{a b\}$ and a "pseudocycle" $\{a\}\{b\}$, lie close to each other, have similar weights, and for longer and longer orbits the curvature corrections fall off rapidly. Indeed, for systems that satisfy the "axiom A" requirements, such as the open disks billiards, curvature expansions converge very well.

Once a set of the shortest cycles has been found, and the cycle periods, stabilities and integrated observable computed, the cycle averaging formulas

$$
\begin{aligned}
\langle a\rangle & =\langle A\rangle_{\zeta} /\langle T\rangle_{\zeta} \\
\langle A\rangle_{\zeta} & =-\frac{\partial}{\partial \beta} \frac{1}{\zeta}=\sum^{\prime} A_{\pi} t_{\pi}, \quad\langle T\rangle_{\zeta}=\frac{\partial}{\partial s} \frac{1}{\zeta}=\sum^{\prime} T_{\pi} t_{\pi}
\end{aligned}
$$

yield the expectation value (the chaotic, ergodic average over the nonwandering set) of the observable $a(x)$.

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## Exercises

Exercise 18.1 Cycle expansions. Write programs that implement binary symbolic dynamics cycle expansions for (a) dynamical zeta functions, (b) spectral determinants. Combined with the cycles computed for a 2-branch repeller or a 3-disk system they will be useful in problem that follow.

Exercise 18.2 Escape rate for a 1-d repeller. (Continuation of exercise 15.1 - easy, but long)
Consider again the quadratic map (15.31)

$$
f(x)=A x(1-x)
$$

on the unit interval, for definitivness take either $A=9 / 2$ or $A=6$. Describing the itinerary of any trajectory by the binary alphabet $\{0,1\}$ ('0' if the iterate is in the first half of the interval and '1' if is in the second half), we have a repeller with a complete binary symbolic dynamics.
(a) Sketch the graph of $f$ and determine its two fixed points $\overline{0}$ and $\overline{1}$, together with their stabilities.
(b) Sketch the two branches of $f^{-1}$. Determine all the prime cycles up to topological length 4 using your pocket calculator and backwards iteration of $f$ (see sect. 17.2.1).
(c) Determine the leading zero of the zeta function (15.15) using the weigths $t_{p}=z^{n_{p}} /\left|\Lambda_{p}\right|$ where $\Lambda_{p}$ is the stability of the $p$ cycle.
(d) Show that for $A=9 / 2$ the escape rate of the repeller is $0.361509 \ldots$ using the spectral determinant, with the same cycle weight. If you have taken $A=6$, the escape rate is in $0.83149298 \ldots$, as shown in solution 18.2. Compare the coefficients of the spectral determinant and the zeta function cycle expansions. Which expansion converges faster?
(Per Rosenqvist)
Exercise 18.3 Escape rate for the Ulam map. (medium) We will try to compute the escape rate for the Ulam map (17.21)

$$
f(x)=4 x(1-x),
$$

using the method of cycle expansions. The answer should be zero, as nothing escapes.
(a) Compute a few of the stabilities for this map. Show that $\Lambda_{0}=4, \Lambda_{1}=-2$, $\Lambda_{01}=-4, \Lambda_{001}=-8$ and $\Lambda_{011}=8$.
(b) Show that

$$
\Lambda_{\epsilon_{1} \ldots \epsilon_{n}}= \pm 2^{n}
$$

and determine a rule for the sign.
(c) (hard) Compute the dynamical zeta function for this system

$$
\zeta^{-1}=1-t_{0}-t_{1}-\left(t_{01}-t_{0} t_{1}\right)-\cdots
$$

You might note that the convergence as function of the truncation cycle length is slow. Try to fix that by treating the $\Lambda_{0}=4$ cycle separately.

Exercise 18.4 Pinball escape rate, semi-analytical. Estimate the 3disk pinball escape rate for $R: a=6$ by substituting analytical cycle stabilities and periods (exercise 8.1 and exercise 8.2) into the appropriate binary cycle expansion. Compare with the numerical estimate exercise 10.3
Exercise 18.5 Pinball escape rate, from numerical cycles. Compute the escape rate for $R: a=6$ 3-disk pinball by substituting list of numerically computed cycle stabilities of exercise 17.5 into the binary cycle expansion.

Exercise 18.6 Pinball resonances, in the complex plane. Plot the logarithm of the absolute value of the dynamical zeta function and/or the spectral determinant cycle expansion (18.5) as contour plots in the complex s plane. Do you find zeros other than the one corresponding to the complex one? Do you see evidence for a finite radius of convergence for either cycle expansion?
Exercise 18.7 Counting the 3-disk pinball counterterms. Verify that the number of terms in the 3-disk pinball curvature expansion (18.31) is given by

$$
\begin{aligned}
\prod_{p}\left(1+t_{p}\right) & =\frac{1-3 z^{4}-2 z^{6}}{1-3 z^{2}-2 z^{3}}=1+3 z^{2}+2 z^{3}+\frac{z^{4}\left(6+12 z+2 z^{2}\right)}{1-3 z^{2}-2 z^{3}} \\
& =1+3 z^{2}+2 z^{3}+6 z^{4}+12 z^{5}+20 z^{6}+48 z^{7}+84 z^{8}+184 z^{9}+\ldots
\end{aligned}
$$

This means that, for example, $c_{6}$ has a total of 20 terms, in agreement with the explicit 3-disk cycle expansion (18.32).
Exercise 18.8 3-disk unfactorized zeta cycle expansions. Check that the curvature expansion (18.2) for the 3 -disk pinball, assuming no symmetries between disks, is given by

$$
\begin{align*}
1 / \zeta= & \left(1-z^{2} t_{12}\right)\left(1-z^{2} t_{13}\right)\left(1-z^{2} t_{23}\right)\left(1-z^{3} t_{123}\right)\left(1-z^{3} t_{132}\right) \\
& \left(1-z^{4} t_{1213}\right)\left(1-z^{4} t_{1232}\right)\left(1-z^{4} t_{1323}\right)\left(1-z^{5} t_{12123}\right) \cdots \\
= & 1-z^{2} t_{12}-z^{2} t_{23}-z^{2} t_{31}-z^{3} t_{123}-z^{3} t_{132} \\
& -z^{4}\left[\left(t_{1213}-t_{12} t_{13}\right)+\left(t_{1232}-t_{12} t_{23}\right)+\left(t_{1323}-t_{13} t_{23}\right)\right] \\
& -z^{5}\left[\left(t_{12123}-t_{12} t_{123}\right)+\cdots\right]-\cdots \tag{18.31}
\end{align*}
$$

The symmetrically arranged 3-disk pinball cycle expansion of the Euler product (18.2) (see table 13.4 and figure 22.2) is given by:

$$
\begin{align*}
1 / \zeta= & \left(1-z^{2} t_{12}\right)^{3}\left(1-z^{3} t_{123}\right)^{2}\left(1-z^{4} t_{1213}\right)^{3} \\
& \left(1-z^{5} t_{12123}\right)^{6}\left(1-z^{6} t_{1212133}\right)^{6}\left(1-z^{6} t_{121323}\right)^{3} \ldots \\
= & 1-3 z^{2} t_{12}-2 z^{3} t_{123}-3 z^{4}\left(t_{1213}-t_{12}^{2}\right)-6 z^{5}\left(t_{12123}-t_{12} t_{123}\right) \\
& -z^{6}\left(6 t_{121213}+3 t_{121323}+t_{12}^{3}-9 t_{12} t_{1213}-t_{123}^{2}\right) \\
& -6 z^{7}\left(t_{1212123}+t_{1212313}+t_{1213123}+t_{12}^{2} t_{123}-3 t_{12} t_{12123}-t_{123} t_{1213}\right) \\
& -3 z^{8}\left(2 t_{12121213}+t_{12121313}+2 t_{12121323}+2 t_{12123123}\right. \\
& \quad+2 t_{12123213}+t_{12132123}+3 t_{12}^{2} t_{1213}+t_{12} t_{123}^{2} \\
& \left.\quad-6 t_{12} t_{121213}-3 t_{12} t_{121323}-4 t_{123} t_{12123}-t_{1213}^{2}\right)-\cdots \quad(18.32) \tag{18.32}
\end{align*}
$$

Remark 18.7 Unsymmetrized cycle expansions. The above 3-disk cycle expansions might be useful for cross-checking purposes, but, as we shall see in chapter 22, they are not recommended for actual computations, as the factorized zeta functions yield much better convergence.

Exercise 18.9 4-disk unfactorized dynamical zeta function cycle expansions For the symmetriclly arranged 4-disk pinball the symmetry group is $C_{4 v}$, of order 8. The degenerate cycles can have multiplicities 2, 4 or 8 (see table 13.2):

$$
\begin{align*}
1 / \zeta= & \left(1-z^{2} t_{12}\right)^{4}\left(1-z^{2} t_{13}\right)^{2}\left(1-z^{3} t_{123}\right)^{8}\left(1-z^{4} t_{1213}\right)^{8}\left(1-z^{4} t_{1214}\right)^{4} \\
& \left(1-z^{4} t_{1234}\right)^{2}\left(1-z^{4} t_{1243}\right)^{4}\left(1-z^{5} t_{12123}\right)^{8}\left(1-z^{5} t_{12124}\right)^{8}\left(1-z^{5} t_{12134}\right)^{8} \\
& \left(1-z^{5} t_{12143}\right)^{8}\left(1-z^{5} t_{12313}\right)^{8}\left(1-z^{5} t_{12413}\right)^{8} \cdots \tag{18.33}
\end{align*}
$$

and the cycle expansion is given by

$$
\begin{align*}
1 / \zeta= & 1-z^{2}\left(4 t_{12}+2 t_{13}\right)-8 z^{3} t_{123} \\
& -z^{4}\left(8 t_{1213}+4 t_{1214}+2 t_{1234}+4 t_{1243}-6 t_{12}^{2}-t_{13}^{2}-8 t_{12} t_{13}\right) \\
& -8 z^{5}\left(t_{12123}+t_{12124}+t_{12134}+t_{12143}+t_{12313}+t_{12413}-4 t_{12} t_{123}-2 t_{13} t_{123}\right) \\
& -4 z^{6}\left(2 S_{8}+S_{4}+t_{12}^{3}+3 t_{12}^{2} t_{13}+t_{12} t_{13}^{2}-8 t_{12} t_{1213}-4 t_{12} t_{1214}\right. \\
& -2 t_{12} t_{1234}-4 t_{12} t_{1243}-4 t_{13} t_{1213}-2 t_{13} t_{1214}-t_{13} t_{1234} \\
& \left.-2 t_{13} t_{1243}-7 t_{123}^{2}\right)-\cdots \tag{18.34}
\end{align*}
$$

where in the coefficient to $z^{6}$ the abbreviations $S_{8}$ and $S_{4}$ stand for the sums over the weights of the 12 orbits with multiplicity 8 and the 5 orbits of multiplicity 4 , respectively; the orbits are listed in table 13.4.

Exercise 18.10 Tail resummations. A simple illustration of such tail resummation is the $\zeta$ function for the Ulam map (17.21) for which the cycle structure is exceptionally simple: the eigenvalue of the $x_{0}=0$ fixed point is 4 , while the eigenvalue of any other $n$-cycle is $\pm 2^{n}$. Typical cycle weights used in thermodynamic averaging are $t_{0}=4^{\tau} z, t_{1}=t=2^{\tau} z, t_{p}=t^{n_{p}}$ for $p \neq 0$. The simplicity of the cycle eigenvalues enables us to evaluate the $\zeta$ function by a simple trick: we note that if the value of any $n$-cycle eigenvalue were $t^{n}$, (15.21) would yield $1 / \zeta=1-2 t$. There is only one cycle, the $x_{0}$ fixed point, that has a different weight $\left(1-t_{0}\right)$, so we factor it out, multiply the rest by $(1-t) /(1-t)$, and obtain a rational $\zeta$ function

$$
\begin{equation*}
1 / \zeta(z)=\frac{(1-2 t)\left(1-t_{0}\right)}{(1-t)} \tag{18.35}
\end{equation*}
$$

Consider how we would have detected the pole at $z=1 / t$ without the above trick. As the $\overline{0}$ fixed point is isolated in its stability, we would have kept the factor $\left(1-t_{0}\right)$ in (18.5) unexpanded, and noted that all curvature combinations in (18.5) which include the $t_{0}$ factor are unbalanced, so that the cycle expansion is an infinite series:

$$
\begin{equation*}
\prod_{p}\left(1-t_{p}\right)=\left(1-t_{0}\right)\left(1-t-t^{2}-t^{3}-t^{4}-\ldots\right) \tag{18.36}
\end{equation*}
$$

(we shall return to such infinite series in chapter 21). The geometric series in the brackets sums up to (18.35). Had we expanded the $\left(1-t_{0}\right)$ factor, we would have noted that the ratio of the successive curvatures is exactly $c_{n+1} / c_{n}=t$; summing we would recover the rational $\zeta$ function (18.35).

## Chapter 19

## Why cycle?

"Progress was a labyrinth ... people plunging blindly in and then rushing wildly back, shouting that they had found it ... the invisible king the élan vital the principle of evolution ... writing a book, starting a war, founding a school...."
F. Scott Fitzgerald, This Side of Paradise

In the preceding chapters we have moved rather briskly through the evolution operator formalism. Here we slow down in order to develop some fingertip feeling for the traces of evolution operators. We start out by explaining how qualitatively how local exponential instability and exponential growth in topologically distinct trajectories lead to a global exponential instability.

### 19.1 Escape rates

We start by verifying the claim (10.11) that for a nice hyperbolic flow the trace of the evolution operator grows exponentially with time. Consider again the game of pinball of figure 1.1. Designate by $\mathcal{M}$ a phase space region that encloses the three disks, say the surface of the table $\times$ all pinball directions. The fraction of initial points whose trajectories start out within the phase space region $\mathcal{M}$ and recur within that region at the time $t$ is given by

$$
\begin{equation*}
\hat{\Gamma}_{\mathcal{M}}(t)=\frac{1}{|\mathcal{M}|} \iint_{\mathcal{M}} d x d y \delta\left(y-f^{t}(x)\right) \tag{19.1}
\end{equation*}
$$

This quantity is eminently measurable and physically interesting in a variety of problems spanning from nuclear physics to celestial mechanics. The integral over $x$ takes care of all possible initial pinballs; the integral over $y$ checks whether they are still within $\mathcal{M}$ by the time $t$. If the dynamics is bounded, and $\mathcal{M}$ envelops the entire accessible phase space, $\hat{\Gamma}_{\mathcal{M}}(t)=1$
for all $t$. However, if trajectories exit $\mathcal{M}$ the recurrence fraction decreases with time. For example, any trajectory that falls off the pinball table in figure 1.1 is gone for good.

These observations can be made more concrete by examining the pinball phase space of figure 1.8. With each pinball bounce the initial conditions that survive get thinned out, each strip yielding two thiner strips within it. The total fraction of survivors (1.2) after $n$ bounces is given by

$$
\begin{equation*}
\hat{\Gamma}_{n}=\frac{1}{|\mathcal{M}|} \sum_{i}^{(n)}\left|\mathcal{M}_{i}\right| \tag{19.2}
\end{equation*}
$$

where $i$ is a binary label of the $i$ th strip, and $\left|\mathcal{M}_{i}\right|$ is the area of the $i$ th strip. The phase space volume is preserved by the flow, so the strips of survivors are contracted along the stable eigendirections, and ejected along the unstable eigendirections. As a crude estimate of the number of survivors in the $i$ th strip, assume that the spreading of a ray of trajectories per bounce is given by a factor $\Lambda$, the mean value of the expanding eigenvalue of the corresponding Jacobian matrix of the flow, and replace $\left|\mathcal{M}_{i}\right|$ by the phase space strip width estimate $\left|\mathcal{M}_{i}\right| /|\mathcal{M}| \sim 1 / \Lambda_{i}$. This estimate of a size of a neighborhood (given already on p. 112) is right in spirit, but not without drawbacks. One problem is that in general the eigenvalues of a Jacobian matrix have no invariant meaning; they depend on the choice of coordinates. However, we saw in chapter 14 that the sizes of neighborhoods are determined by stability eigenvalues of periodic points, and those are invariant under smooth coordinate transformations.

In this approximation $\hat{\Gamma}_{n}$ receives $2^{n}$ contributions of equal size

$$
\begin{equation*}
\hat{\Gamma}_{1} \sim \frac{1}{\Lambda}+\frac{1}{\Lambda}, \cdots \quad, \hat{\Gamma}_{n} \sim \frac{2^{n}}{\Lambda^{n}}=e^{-n(\lambda-h)}:=e^{-n \gamma} \tag{19.3}
\end{equation*}
$$

up to preexponential factors. We see here the interplay of the two key ingredients of chaos first alluded to in sect. 1.3.1: the escape rate $\gamma$ equals local expansion rate (the Lyapunov exponent $\lambda=\ln \Lambda$ ), minus the rate of global reinjection back into the system (the topological entropy $h=\ln 2$ ). As we shall see in (20.16), with correctly defined "entropy" this result is exact.

As at each bounce one loses routinely the same fraction of trajectories, one expects the sum (19.2) to fall off exponentially with $n$. More precisely, by the hyperbolicity assumption of sect. 14.1.1 the expanding eigenvalue of the Jacobian matrix of the flow is exponentially bounded from both above and below,

$$
\begin{equation*}
1<\left|\Lambda_{\min }\right| \leq|\Lambda(x)| \leq\left|\Lambda_{\max }\right|, \tag{19.4}
\end{equation*}
$$

and the area of each strip in (19.2) is bounded by $\left|\Lambda_{\text {max }}^{-n}\right| \leq\left|\mathcal{M}_{i}\right| \leq$ $\left|\Lambda_{\text {min }}^{-n}\right|$. Replacing $\left|\mathcal{M}_{i}\right|$ in (19.2) by its over (under) estimates in terms
of $\left|\Lambda_{\text {max }}\right|,\left|\Lambda_{\text {min }}\right|$ immediately leads to exponential bounds $\left(2 /\left|\Lambda_{\max }\right|\right)^{n} \leq$ $\hat{\Gamma}_{n} \leq\left(2 /\left|\Lambda_{\text {min }}\right|\right)^{n}$, that is,

$$
\begin{equation*}
\ln \left|\Lambda_{\max }\right| \geq-\frac{1}{n} \ln \hat{\Gamma}_{n}+\ln 2 \geq \ln \left|\Lambda_{\min }\right| \tag{19.5}
\end{equation*}
$$

The argument based on (19.5) establishes only that the sequence $\gamma_{n}=$ $-\frac{1}{n} \ln \Gamma_{n}$ has a lower and an upper bound for any $n$. In order to prove that $\gamma_{n}$ converge to the limit $\gamma$, we first show that for hyperbolic systems the sum over survivor intervals (19.2) can be replaced by the sum over periodic orbit stabilities. By (19.4) the size of $\mathcal{M}_{i}$ strip can be bounded by the stability $\Lambda_{i}$ of $i$ th periodic point:

$$
\begin{equation*}
\mathcal{C}_{1} \frac{1}{\left|\Lambda_{i}\right|}<\frac{\left|\mathcal{M}_{i}\right|}{|\mathcal{M}|}<\mathcal{C}_{2} \frac{1}{\left|\Lambda_{i}\right|}, \tag{19.6}
\end{equation*}
$$

for any periodic point $i$ of period $n$, with constants $\mathcal{C}_{j}$ dependent on the dynamical system but independent of $n$. The meaning of these bounds is that for longer and longer cycles in a system of bounded hyperbolicity, the shrinking of the $i$ th strip is better and better approximated by by the derivaties evaluated on the periodic point within the strip. Hence the survival probability can be bounded close to the cycle point stability sum

$$
\begin{equation*}
\hat{\mathcal{C}}_{1} \Gamma_{n}<\sum_{i}^{(n)} \frac{\left|\mathcal{M}_{i}\right|}{|\mathcal{M}|}<\hat{\mathcal{C}}_{2} \Gamma_{n} \tag{19.7}
\end{equation*}
$$

where $\Gamma_{n}=\sum_{i}^{(n)} 1 /\left|\Lambda_{i}\right|$ is the asymptotic trace sum (14.22). In this way we have established that for hyperbolic systems the survival probability sum (19.2) can be replaced by the periodic orbit sum (14.22).

We conclude that for hyperbolic, locally unstable flows the fraction (19.1) of initial $x$ whose trajectories remain trapped within $\mathcal{M}$ up to time $t$ is expected to decay exponentially,

$$
\Gamma_{\mathcal{M}}(t) \propto e^{-\gamma t}
$$

where $\gamma$ is the asymptotic escape rate defined by

$$
\begin{equation*}
\gamma=-\lim _{t \rightarrow \infty} \frac{1}{t} \ln \Gamma_{\mathcal{M}}(t) \tag{19.8}
\end{equation*}
$$

### 19.1.1 Periodic orbit averages

We now refine the reasoning of sect. 19.1. Consider the trace (14.6) in the asymptotic limit (14.21):

$$
\operatorname{tr} \mathcal{L}^{n}=\int d x \delta\left(x-f^{n}(x)\right) e^{\beta A^{n}(x)} \approx \sum_{i}^{(n)} \frac{e^{\beta A^{n}\left(x_{i}\right)}}{\left|\Lambda_{i}\right|}
$$

The factor $1 /\left|\Lambda_{i}\right|$ was interpreted in (19.2) as the area of the $i$ th phase space strip. Hence $\operatorname{tr} \mathcal{L}^{n}$ is a discretization of the integral $\int d x e^{\beta A^{n}(x)}$ approximated by a tesselation into strips centered on periodic points $x_{i}$, figure 1.9, with the volume of the $i$ th neighborhood given by estimate $\left|\mathcal{M}_{i}\right| \sim 1 /\left|\Lambda_{i}\right|$, and $e^{\beta A^{n}(x)}$ estimated by $e^{\beta A^{n}\left(x_{i}\right)}$, its value at the $i$ th periodic point. If the symbolic dynamics is a complete, any rectangle $\left[s_{-m} \cdots s_{0} . s_{1} s_{2} \cdots s_{n}\right.$ ] of sect. 12.3.1 always contains the cycle point $\overline{s_{-m} \cdots s_{0} s_{1} s_{2} \cdots s_{n}}$; hence even though the periodic points are of measure zero (just like rationals in the unit interval), they are dense on the non-wandering set. Equiped with a measure for the associated rectangle, periodic orbits suffice to cover the entire non-wandering set. The average of $e^{\beta A^{n}}$ evaluated on the non-wandering set is therefore given by the trace, properly normalized so $\langle 1\rangle=1$ :

$$
\begin{equation*}
\left\langle e^{\beta A^{n}}\right\rangle_{n} \approx \frac{\sum_{i}^{(n)} e^{\beta A^{n}\left(x_{i}\right)} /\left|\Lambda_{i}\right|}{\sum_{i}^{(n)} 1 /\left|\Lambda_{i}\right|}=\sum_{i}^{(n)} \mu_{i} e^{\beta A^{n}\left(x_{i}\right)} . \tag{19.9}
\end{equation*}
$$

Here $\mu_{i}$ is the normalized natural measure

$$
\begin{equation*}
\sum_{i}^{(n)} \mu_{i}=1, \quad \quad \mu_{i}=e^{n \gamma} /\left|\Lambda_{i}\right| \tag{19.10}
\end{equation*}
$$

correct both for the closed systems as well as the open systems of sect. 10.1.3.
Unlike brute numerical slicing of the integration space into an arbitrary lattice (for a critique, see sect. 10.4), the periodic orbit theory is smart, as it automatically partitions integrals by the intrinsic topology of the flow, and assigns to each tile the invariant natural measure $\mu_{i}$.

### 19.1.2 Unstable periodic orbits are dense

(L. Rondoni and P. Cvitanović)

Our goal in sect. 10.1 was to evaluate the space and time averaged expectation value (10.9). An average over all periodic orbits can accomplish the job only if the periodic orbits fully explore the asymptotically accessible phase space.

Why should the unstable periodic points end up being dense? The cycles are intuitively expected to be dense because on a connected chaotic set a typical trajectory is expected to behave ergodically, and pass infinitely many times arbitrarily close to any point on the set, including the initial point of the trajectory itself. The argument is more or less the following. Take a partition of $\mathcal{M}$ in arbitrarily small regions, and consider particles that start out in region $\mathcal{M}_{i}$, and return to it in $n$ steps after some peregrination in phase space. In particular, a particle might return a little to the left of its original position, while a close neighbor might return a little to
the right of its original position. By assumption, the flow is continuous, so generically one expects to be able to gently move the initial point in such a way that the trajectory returns precisely to the initial point, that is, one expects a periodic point of period $n$ in cell $i$. (This is by no means guaranteed to always work, and it must be checked for the particular system at hand. A variety of ergodic but insufficiently mixing counter-examples can be constructed - the most familiar being a quasiperiodic motion on a torus.) As we diminish the size of regions $\mathcal{M}_{i}$, aiming a trajectory that returns to $\mathcal{M}_{i}$ becomes increasingly difficult. Therefore, we are guaranteed that unstable (because of the expansiveness of the map) orbits of larger and larger period are densely interspersed in the asymptotic non-wandering set.

### 19.2 Flow conservation sum rules

If the dynamical system is bounded, all trajectories remain confined for all times, escape rate (19.8) equals $\gamma=-s_{0}=0$, and the leading eigenvalue (J.2) of the Perron-Frobenius operator (9.10) is simply $\exp (-t \gamma)=1$. Conservation of material flow thus implies that for bound flows cycle expansions of dynamical zeta functions and spectral determinants satisfy exact flow conservation sum rules:

$$
\begin{align*}
1 / \zeta(0,0) & =1+\sum_{\pi}^{\prime} \frac{(-1)^{k}}{\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right|}=0 \\
F(0,0) & =1-\sum_{n=1}^{\infty} c_{n}(0,0)=0 \tag{19.11}
\end{align*}
$$

obtained by setting $s=0$ in (18.12), (18.13) cycle weights $t_{p}=e^{-s T_{p}} /\left|\Lambda_{p}\right| \rightarrow$ $1 /\left|\Lambda_{p}\right|$. These sum rules depend neither on the cycle periods $T_{p}$ nor on the observable $a(x)$ under investigation, but only on the cycle stabilities $\Lambda_{p, 1}$, $\Lambda_{p, 2}, \cdots, \Lambda_{p, d}$, and their significance is purely geometric: they are a measure of how well periodic orbits tesselate the phase space. Conservation of material flow provides the first and very useful test of the quality of finite cycle length truncations, and is something that you should always check first when constructing a cycle expansion for a bounded flow.

The trace formula version of the flow conservation flow sum rule comes in two varieties, one for the maps, and another for the flows. By flow conservation the leading eigenvalue is $s_{0}=0$, and for maps (18.11) yields

$$
\begin{equation*}
\operatorname{tr} \mathcal{L}^{n}=\sum_{i \in \operatorname{Fix} f^{n}} \frac{1}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}^{n}\left(x_{i}\right)\right)\right|}=1+e^{s_{1} n}+\ldots \tag{19.12}
\end{equation*}
$$

For flows one can apply this rule by grouping together cycles from $t=T$ to $t=T+\Delta T$

$$
\frac{1}{\Delta T} \sum_{p, r}^{T \leq r T_{p} \leq T+\Delta T} \quad \frac{T_{p}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|}=\frac{1}{\Delta T} \int_{T}^{T+\Delta T} d t\left(1+e^{s_{1} t}+\ldots\right)
$$

$$
=1+\frac{1}{\Delta T} \sum_{\alpha=1}^{\infty} \frac{e^{s_{\alpha} T}}{s_{\alpha}}\left(e^{s_{\alpha} \Delta T}-1\right) \approx 1+e^{s_{1} T}(\notin \leq
$$

As is usual for the the fixed level trace sums, the convergence of (19.12) is controled by the gap between the leading and the next-to-leading eigenvalues of the evolution operator.

### 19.3 Correlation functions

The time correlation function $C_{A B}(t)$ of two observables $A$ and $B$ along the trajectory $x(t)=f^{t}\left(x_{0}\right)$ is defined as

$$
\left.C_{A B}\left(t ; x_{0}\right)=\lim _{T \rightarrow \infty} \frac{1}{T} \int_{0}^{T} d \tau A(x(\tau+t)) B(x(\tau)), \quad x_{0}=\chi(0) 14\right)
$$

If the system is ergodic, with invariant continuous measure $\varrho(x) d x$, then correlation functions do not depend on $x_{0}$ (apart from a set of zero measure), and may be computed by a phase average as well

$$
\begin{equation*}
C_{A B}(t)=\int_{\mathcal{M}} d x_{0} \varrho\left(x_{0}\right) A\left(f^{t}\left(x_{0}\right)\right) B\left(x_{0}\right) . \tag{19.15}
\end{equation*}
$$

For a chaotic system we expect that time evolution will loose the information contained in the initial conditions, so that $C_{A B}(t)$ will approach the uncorrelated limit $\langle A\rangle \cdot\langle B\rangle$. As a matter of fact the asymptotic decay of correlation functions

$$
\begin{equation*}
\hat{C}_{A B}:=C_{A B}-\langle A\rangle\langle B\rangle \tag{19.16}
\end{equation*}
$$

for any pair of observables coincides with the definition of mixing, a fundamental property in ergodic theory. We now assume $\langle B\rangle=0$ (otherwise we may define a new observable by $B(x)-\langle B\rangle$ ). Our purpose is now to connect the asymptotic behavior of correlation functions with the spectrum of $\mathcal{L}$. We can write (19.15) as

$$
\tilde{C}_{A B}(t)=\int_{\mathcal{M}} d x \int_{\mathcal{M}} d y A(y) B(x) \varrho(x) \delta\left(y-f^{t}(x)\right)
$$

and recover the evolution operator

$$
\tilde{C}_{A B}(t)=\int_{\mathcal{M}} d x \int_{\mathcal{M}} d y A(y) \mathcal{L}^{t}(y, x) B(x) \varrho(x)
$$

We also recall that in sect. 9.1 we showed that $\rho(x)$ is the eigenvector of $\mathcal{L}$ corresponding to probability conservation

$$
\int_{\mathcal{M}} d y \mathcal{L}^{t}(x, y) \rho(y)=\rho(x)
$$

Now, we can expand the $x$ dependent part in terms of the eigenbasis of $\mathcal{L}$ :

$$
B(x) \varrho(x)=\sum_{\alpha=0}^{\infty} c_{\alpha} \varphi_{\alpha}(x),
$$

where $\varphi_{0}=\varrho(x)$. Since the average of the left hand side is zero the coefficient $c_{0}$ must vanish. The action of $\mathcal{L}$ then can be written as

$$
\begin{equation*}
\tilde{C}_{A B}(t)=\sum_{\alpha \neq 0} e^{-s_{\alpha} t} c_{\alpha} \int_{\mathcal{M}} d y A(y) \varphi_{\alpha}(y) \tag{19.17}
\end{equation*}
$$

We see immediately that if the spectrum has a gap, that is, the second largest leading eigenvalue is isolated from the largest eigenvalue ( $s_{0}=0$ ) then (19.17) implies an exponential decay of correlations

$$
\tilde{C}_{A B}(t) \sim e^{-\nu t} .
$$

The correlation decay rate $\nu=s_{1}$ then depends only on intrinsic properties of the dynamical system (the position of the next-to-leading eigenvalue of the Perron-Frobenius operator), while the choice of particular observables influences just the prefactor.

The importance of correlation functions, beyond the mentioned theoretical features, is that they are often accessible from time series measurable in laboratory experiments and numerical simulations: moreover they are linked to transport exponents.

### 19.4 Trace formulas $v s$. level sums

Trace formulas (14.9) and (14.19) diverge precisely where one would like to use them, at $s$ equal to eigenvalues $s_{\alpha}$. Instead, one can proceed as follows; according to (14.23) the "level" sums (all symbol strings of length $n$ ) are asymptotically going like $e^{s_{0} n}$

$$
\sum_{i \in \operatorname{Fix} f^{n}} \frac{e^{\beta A^{n}\left(x_{i}\right)}}{\left|\Lambda_{i}\right|}=e^{s_{0} n}
$$

so an $n$th order estimate $s_{(n)}$ is given by

$$
\begin{equation*}
1=\sum_{i \in \operatorname{Fix}^{n} n} \frac{e^{\beta A^{n}\left(x_{i}\right)} e^{-s_{(n)} n}}{\left|\Lambda_{i}\right|} \tag{19.18}
\end{equation*}
$$

which generates a "normalized measure". The difficulty with estimating this $n \rightarrow \infty$ limit is at least twofold:

1. due to the exponential growth in number of intervals, and the exponential decrease in attainable accuracy, the maximal $n$ attainable experimentally or numerically is in practice of order of something between 5 to 20.
2. the preasymptotic sequence of finite estimates $s_{(n)}$ is not unique, because the sums $\Gamma_{n}$ depend on how we define the escape region, and because in general the areas $\mathcal{M}_{i}$ in the sum (19.2) should be weighted by the density of initial conditions $x_{0}$. For example, an overall measuring unit rescaling $\mathcal{M}_{i} \rightarrow \alpha \mathcal{M}_{i}$ introduces $1 / n$ corrections in $s_{(n)}$ defined by the log of the sum (19.8): $s_{(n)} \rightarrow s_{(n)}-\ln \alpha / n$. This can be partially fixed by defining a level average

$$
\begin{equation*}
\left\langle e^{\beta A(s)}\right\rangle_{(n)}:=\sum_{i \in \operatorname{Fix} f^{n}} \frac{e^{\beta A^{n}\left(x_{i}\right)} e^{s n}}{\left|\Lambda_{i}\right|} \tag{19.19}
\end{equation*}
$$

and requiring that the ratios of successive levels satisfy

$$
1=\frac{\left\langle e^{\beta A\left(s_{(n)}\right)}\right\rangle_{(n+1)}}{\left\langle e^{\beta A\left(s_{(n)}\right)}\right\rangle_{(n)}}
$$

This avoids the worst problem with the formula (19.18), the inevitable $1 / n$ corrections due to its lack of rescaling invariance. However, even though much published pondering of "chaos" relies on it, there is no need for such gymnastics: the dynamical zeta functions and spectral determinants are already invariant under all smooth nonlinear conjugacies $x \rightarrow h(x)$, not only linear rescalings, and require no $n \rightarrow \infty$ extrapolations. Comparing with the cycle expansions (18.5) we see what the difference is; while in the level sum approach we keep increasing exponentially the number of terms with no reference to the fact that most are already known from shorter estimates, in the cycle expansions short terms dominate, longer ones enter only as exponentially small corrections.

The beauty of the trace formulas is that they are coordinatization independent: both $\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}\right)\right|=\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}^{T_{p}}(x)\right)\right|$ and $e^{\beta A_{p}}=e^{\beta A^{T_{p}}(x)}$ contribution to the cycle weight $t_{p}$ are independent of the starting periodic point point $x$. For the Jacobian matrix $\mathbf{J}_{p}$ this follows from the chain rule for derivatives, and for $e^{\beta A_{p}}$ from the fact that the integral over $e^{\beta A^{t}(x)}$ is evaluated along a closed loop. In addition, $\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}\right)\right|$ is invariant under smooth coordinate transformations.

### 19.4.1 Equipartition measures



There exist many strange sets which cannot be partitioned by the topology of a dynamical flow: some well known examples are the Mandelbrot set, the period doubling repeller and the probabilistically generated
fractal aggregates. In such cases the choice of measure is wide open. One easy choice is the equipartition or cylinder measure: given a symbolic dynamics partition, weigh all symbol sequences of length $n$ equally. Given a symbolic dynamics, the equipartition measure is easy to implement: the rate of growth of the number of admissible symbol sequences $K_{n}$ with the sequence length $n$ is given by the topological entropy $h$ (discussed in sect. 13.1) and the equipartition measure for the $i$ th region $\mathcal{M}_{i}$ is simply

$$
\begin{equation*}
\Delta \mu_{i}=1 / K_{n} \rightarrow e^{-n h} \tag{19.20}
\end{equation*}
$$

The problem with the equipartition measure is twofold: it usually has no physical basis, and it is not an intrinsic invariant property of the strange set, as it depends on the choice of a partition. One is by no means forced to use either the natural or the equipartition measure; there is a variety of other choices, depending on the problem. Also the stability eigenvalues $\Lambda_{i}$ need not refer to motion in the dynamical space; in more general settings it can be a renormalization scaling function, or even a scaling function describing a non-wandering set in the parameter space (sect. 24.3).

## Commentary

Remark 19.1 Nonhyperbolic measures. $\quad \mu_{i}=1 /\left|\Lambda_{i}\right|$ is the natural measure only for the strictly hyperbolic systems. For non-hyperbolic systems, the measure develops folding cusps. For example, for Ulam type maps (unimodal maps with quadratic critical point mapped onto the "left" unstable fixed point $x_{0}$, discussed in more detail in chapter 21), the measure develops a square-root singularity on the $\overline{0}$ cycle:

$$
\begin{equation*}
\mu_{0}=\frac{1}{\left|\Lambda_{0}\right|^{1 / 2}} . \tag{19.21}
\end{equation*}
$$

The thermodynamics averages are still expected to converge in the "hyperbolic" phase where the positive entropy of unstable orbits dominates over the marginal orbits, but they fail in the "non-hyperbolic" phase. The general case remains unclear, and we refer the reader to the literature [H.18, H.14, H.11, 12.21].

Remark 19.2 Trace formula periodic orbit averaging. The cycle averaging formulas are not the first thing that one would intuitively write down; the approximate trace formulas are more accessibly heuristically. The trace formula averaging (19.13) seems to have be discussed for the first time by Hannay and Ozorio de Almeida [H.1, 7.11]. Another novelty of the cycle averaging formulas and one of their main virtues, in contrast to the explicit analytic results such as those of ref. [18.4], is that their evaluation does not require any explicit construction of the (coordinate dependent) eigenfunctions of the Perron-Frobenius operator (that is, the natural measure $\rho_{0}$ ).

Remark 19.3 The choice of observables We have been quite sloppy on the mathematical side, as in discussing the spectral features of $\mathcal{L}$ the choice of the function space is crucial (especially when one is looking beyond the dominant eigenvalue). As a matter of fact in the function space where usually ergodic properties are defined, $L^{2}(d \mu)$ there is $\underline{\text { no }}$ gap, due to unitarity property of the Koopman operator: this means that there exist (ugly yet summable) functions for which no exponential decay is present even if the Fredholm determinant has isolated zeroes. A particularly nice example is worked out in [H.20], and a more mathematical argument is presented in [H.21].

Remark 19.4 Lattice models The relationship between the spectral gap and exponential decay properties is very well known in the statistical mechanical framework, where one deals with spatial correlations in lattice systems and links them to the gap of the transfer matrix.

Remark 19.5 Role of noise in dynamical systems. In most practical applications in addition to the chaotic deterministic dynamics there is always an additional external noise. The noise can be characterized by its strength $\sigma$ and distribution. Lyapunov exponents, correlation decay and dynamo rate can be defined in this case the same way as in the deterministic case. We can think that noise completely destroys the results derived here. However, one can show that the deterministic formulas remain valid until the noise level is small. A small level of noise even helps as it makes the dynamics ergodic. Non-communicating parts of the phase space become weakly connected due to the noise. This is a good argument to explain why periodic orbit theory works in non-ergodic systems. For small amplitude noise one can make a noise expansion

$$
\bar{\lambda}=\bar{\lambda}_{0}+\bar{\lambda}_{1} \sigma^{2}+\overline{\lambda_{2}} \sigma^{4}+\ldots,
$$

around the deterministic averages $\lambda_{0}$. The expansion coefficients $\overline{\lambda_{1}}, \overline{\lambda_{2}}, \ldots$ can also be expressed via periodic orbit formulas. The calculation of these coefficients is one of the challenges facing periodic orbit theory today.

## Résumé

We conclude this chapter by a general comment on the relation of the finite trace sums such as (19.2) to the spectral determinants and dynamical zeta functions. One might be tempted to believe that given a deterministic rule, a sum like (19.2) could be evaluated to any desired precision. For short finite times this is indeed true: every region $\mathcal{M}_{i}$ in (19.2) can be accurately delineated, and there is no need for fancy theory. However, if the dynamics is unstable, local variations in initial conditions grow exponentially and in finite time attain the size of the system. The difficulty with estimating the $n \rightarrow \infty$ limit from (19.2) is then at least twofold:

1. due to the exponential growth in number of intervals, and the exponential decrease in attainable accuracy, the maximal $n$ attainable experimentally or numerically is in practice of order of something between 5 to 20;
2. the preasymptotic sequence of finite estimates $\gamma_{n}$ is not unique, because the sums $\Gamma_{n}$ depend on how we define the escape region, and because in general the areas $\mathcal{M}_{i}$ in the sum (19.2) should be weighted by the density of initial $x_{0}$.

In contrast, the dynamical zeta functions and spectral determinants are already invariant under all smooth nonlinear conjugacies $x \rightarrow h(x)$, not only linear rescalings, and require no $n \rightarrow \infty$ extrapolations.

## References

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## Exercises

## Exercise 19.1 Escape rate of the logistic map.

(a) Calculate the fraction of trajectories remaining trapped in the interval $[0,1]$ for the logistic map

$$
\begin{equation*}
f(x)=a\left(1-4(x-0.5)^{2}\right), \tag{19.22}
\end{equation*}
$$

and determine the $a$ dependence of the escape rate $\gamma(a)$ numerically.
(b) Work out a numerical method for calculating the lengths of intervals of trajectories remaining stuck for $n$ iterations of the map.
(c) What is your expectation about the a dependence near the critical value $a_{c}=1$ ?

Exercise 19.2 Four scale map decay. Compute the second largest eigenvalue of the Perron-Frobenius operator for the four scale map

$$
f(x)= \begin{cases}a_{1} x & \text { if } 0<x<b / a_{1},  \tag{19.23}\\ (1-b)\left(\left(x-b / a_{1}\right) /\left(b-b / a_{1}\right)\right)+b & \text { if } b / a_{1}<x<b, \\ a_{2}(x-b) & \text { if } b<x<b+b / a_{2}, \\ (1-b)\left(\left(x-b-b / a_{2}\right) /\left(1-b-b / a_{2}\right)\right)+b & \text { if } b+b / a_{2}<x<1 .\end{cases}
$$

Exercise 19.3 Lyapunov exponents for 1-dimensional maps. Extend your cycle expansion programs so that the first and the second moments of observables can be computed. Use it to compute the Lyapunov exponent for some or all of the following maps:
(a) the piecewise-linear flow conserving map, the skew tent map

$$
f(x)= \begin{cases}a x & \text { if } 0 \leq x \leq a^{-1} \\ \frac{a}{a-1}(1-x) & \text { if } a^{-1} \leq x \leq 1\end{cases}
$$

(b) the Ulam map $f(x)=4 x(1-x)$
(c) the skew Ulam map

$$
f(x)=0.1218 x(1-x)(1-0.6 x)
$$

with a peak at 0.7.
(d) the repeller of $f(x)=A x(1-x)$, for either $A=9 / 2$ or $A=6$ (this is a continuation of exercise 18.2).
(e) for the 2-branch flow conserving map

$$
\begin{array}{ll}
f_{0}(x)=\frac{h-p+\sqrt{(h-p)^{2}+4 h x}}{2 h}, \quad x \in[0, p] & \text { (19.24) }  \tag{19.24}\\
f_{1}(x)=\frac{h+p-1+\sqrt{(h+p-1)^{2}+4 h(x-p)}}{2 h}, & x \in[p, 1]
\end{array}
$$

This is a nonlinear perturbation of $(h=0)$ Bernoulli map (16.6); the first 15 eigenvalues of the Perron-Frobenius operator are listed in ref. [19.1] for $p=0.8, h=0.1$. Use these parameter values when computing the Lyapunov exponent.

Cases (a) and (b) can be computed analytically; cases (c), (d) and (e) require numerical computation of cycle stabilities. Just to see whether the theory is worth the trouble, also cross check your cycle expansions results for cases (c) and (d) with Lyapunov exponent computed by direct numerical averaging along trajectories of randomly chosen initial points:
(f) trajectory-trajectory separation (10.27) (hint: rescale $\delta x$ every so often, to avoid numerical overflows),
(g) iterated stability (10.31).

How good is the numerical accuracy compared with the periodic orbit theory predictions?
oo

## Chapter 20

## Thermodynamic formalism

So, naturalists observe, a flea hath smaller fleas that on him prey; and those have smaller still to bite 'em; and so proceed ad infinitum.
Jonathan Swift
(G. Vattay)

In the preceding chapters we characterized chaotic systems via global quantities such as averages. It turned out that these are closely related to very fine details of the dynamics like stabilities and time periods of individual periodic orbits. In statistical mechanics a similar duality exists. Macroscopic systems are characterized with thermodynamic quantities (pressure, temperature and chemical potential) which are averages over fine details of the system called microstates. One of the greatest achievements of the theory of dynamical systems was when in the sixties and seventies Bowen, Ruelle and Sinai made the analogy between these two subjects explicit. Later this "Thermodynamic Formalism" of dynamical systems became widely used when the concept of fractals and multifractals has been introduced. The formalism made it possible to calculate various fractal dimensions in an elegant way and become a standard instrument in a wide range of scientific fields. Next we sketch the main ideas of this theory and show how periodic orbit theory helps to carry out calculations.

### 20.1 Rényi entropies

As we have already seen trajectories in a dynamical system can be characterized by their symbolic sequences from a generating Markov partition. We can locate the set of starting points $\mathcal{M}_{s_{1} s_{2} \ldots s_{n}}$ of trajectories whose symbol sequence starts with a given set of $n$ symbols $s_{1} s_{2} \ldots s_{n}$. We can associate many different quantities to these sets. There are geometric measures such as the volume $V\left(s_{1} s_{2} \ldots s_{n}\right)$, the area $A\left(s_{1} s_{2} \ldots s_{n}\right)$ or the length $l\left(s_{1} s_{2} \ldots s_{n}\right)$ of this set. Or in general we can have some measure $\mu\left(\mathcal{M}_{s_{1} s_{2} \ldots s_{n}}\right)=\mu\left(s_{1} s_{2} \ldots s_{n}\right)$ of this set. As we have seen in (19.10) the
most important is the natural measure, which is the probability that a non-periodic trajectory visits the set $\mu\left(s_{1} s_{2} \ldots s_{n}\right)=P\left(s_{1} s_{2} \ldots s_{n}\right)$. The natural measure is additive. Summed up for all possible symbol sequences of length $n$ it gives the measure of the whole phase space:

$$
\begin{equation*}
\sum_{s_{1} s_{2} \ldots s_{n}} \mu\left(s_{1} s_{2} \ldots s_{n}\right)=1 \tag{20.1}
\end{equation*}
$$

expresses probability conservation. Also, summing up for the last symbol we get the measure of a one step shorter sequence

$$
\sum_{s_{n}} \mu\left(s_{1} s_{2} \ldots s_{n}\right)=\mu\left(s_{1} s_{2} \ldots s_{n-1}\right) .
$$

As we increase the length $(n)$ of the sequence the measure associated with it decreases typically with an exponential rate. It is then useful to introduce the exponents

$$
\begin{equation*}
\lambda\left(s_{1} s_{2} \ldots s_{n}\right)=-\frac{1}{n} \log \mu\left(s_{1} s_{2} \ldots s_{n}\right) . \tag{20.2}
\end{equation*}
$$

To get full information on the distribution of the natural measure in the symbolic space we can study the distribution of exponents. Let the number of symbol sequences of length $n$ with exponents between $\lambda$ and $\lambda+d \lambda$ be given by $N_{n}(\lambda) d \lambda$. For large $n$ the number of such sequences increases exponentially. The rate of this exponential growth can be characterized by $g(\lambda)$ such that

$$
N_{n}(\lambda) \sim \exp (n g(\lambda)) .
$$

The knowledge of the distribution $N_{n}(\lambda)$ or its essential part $g(\lambda)$ fully characterizes the microscopic structure of our dynamical system.

As a natural next step we would like to calculate this distribution. However it is very time consuming to calculate the distribution directly by making statistics for millions of symbolic sequences. Instead, we introduce auxiliary quantities which are easier to calculate and to handle. These are called partition sums

$$
\begin{equation*}
Z_{n}(\beta)=\sum_{s_{1} s_{2} \ldots s_{n}} \mu^{\beta}\left(s_{1} s_{2} \ldots s_{n}\right), \tag{20.3}
\end{equation*}
$$

as they are obviously motivated by Gibbs type partition sums of statistical mechanics. The parameter $\beta$ plays the role of inverse temperature $1 / k_{B} T$ and $E\left(s_{1} s_{2} \ldots s_{n}\right)=-\log \mu\left(s 1 s_{2} \ldots s_{n}\right)$ is the energy associated with the microstate labeled by $s_{1} s_{2} \ldots s_{n}$ We are tempted also to introduce something analogous with the Free energy. In dynamical systems this is called the Rényi entropy [H.5] defined by the growth rate of the partition sum

$$
\begin{equation*}
K_{\beta}=\lim _{n \rightarrow \infty} \frac{1}{n} \frac{1}{1-\beta} \log \left(\sum_{s_{1} s_{2} \ldots s_{n}} \mu^{\beta}\left(s_{1} s_{2} \ldots s_{n}\right)\right) . \tag{20.4}
\end{equation*}
$$

In the special case $\beta \rightarrow 1$ we get Kolmogorov's entropy

$$
K_{1}=\lim _{n \rightarrow \infty} \frac{1}{n} \sum_{s_{1} s_{2} \ldots s_{n}}-\mu\left(s_{1} s_{2} \ldots s_{n}\right) \log \mu\left(s_{1} s_{2} \ldots s_{n}\right),
$$

while for $\beta=0$ we recover the topological entropy

$$
h_{\text {top }}=K_{0}=\lim _{n \rightarrow \infty} \frac{1}{n} \log N(n)
$$

where $N(n)$ is the number of existing length $n$ sequences. To connect the partition sums with the distribution of the exponents, we can write them as averages over the exponents

$$
Z_{n}(\beta)=\int d \lambda N_{n}(\lambda) \exp (-n \lambda \beta)
$$

where we used the definition (20.2). For large $n$ we can replace $N_{n}(\lambda)$ with its asymptotic form

$$
Z_{n}(\beta) \sim \int d \lambda \exp (n g(\lambda)) \exp (-n \lambda \beta)
$$

For large $n$ this integral is dominated by contributions from those $\lambda^{*}$ which maximize the exponent

$$
g(\lambda)-\lambda \beta
$$

The exponent is maximal when the derivative of the exponent vanishes

$$
\begin{equation*}
g^{\prime}\left(\lambda^{*}\right)=\beta \tag{20.5}
\end{equation*}
$$

From this equation we can determine $\lambda^{*}(\beta)$. Finally the partition sum is

$$
Z_{n}(\beta) \sim \exp \left(n\left[g\left(\lambda^{*}(\beta)\right)-\lambda^{*}(\beta) \beta\right]\right)
$$

Using the definition (20.4) we can now connect the Rényi entropies and $g(\lambda)$

$$
\begin{equation*}
(\beta-1) K_{\beta}=\lambda^{*}(\beta) \beta-g\left(\lambda^{*}(\beta)\right) \tag{20.6}
\end{equation*}
$$

Equations (20.5) and (20.6) define the Legendre transform of $g(\lambda)$. This equation is analogous with the thermodynamic equation connecting the entropy and the free energy. As we know from thermodynamics we can invert the Legendre transform. In our case we can express $g(\lambda)$ from the Rényi entropies via the Legendre transformation

$$
\begin{equation*}
g(\lambda)=\lambda \beta^{*}(\lambda)-\left(\beta^{*}(\lambda)-1\right) K_{\beta^{*}(\lambda)}, \tag{20.7}
\end{equation*}
$$

where now $\beta^{*}(\lambda)$ can be determined from

$$
\begin{equation*}
\frac{d}{d \beta^{*}}\left[\left(\beta^{*}-1\right) K_{\beta^{*}}\right]=\lambda . \tag{20.8}
\end{equation*}
$$

Obviously, if we can determine the Rényi entropies we can recover the distribution of probabilities from (20.7) and (20.8).

The periodic orbit calculation of the Rényi entropies can be carried out by approximating the natural measure corresponding to a symbol sequence by the expression (19.10)

$$
\begin{equation*}
\mu\left(s_{1}, \ldots, s_{n}\right) \approx \frac{e^{n \gamma}}{\mid \Lambda_{\overline{s_{1} s_{2} \ldots s_{n}}}} . \tag{20.9}
\end{equation*}
$$

The partition sum (20.3) now reads

$$
\begin{equation*}
Z_{n}(\beta) \approx \sum_{i} \frac{e^{n \beta \gamma}}{\left|\Lambda_{i}\right|^{\beta}}, \tag{20.10}
\end{equation*}
$$

where the summation goes for periodic orbits of length $n$. We can define the characteristic function

$$
\begin{equation*}
\Omega(z, \beta)=\exp \left(-\sum_{n} \frac{z^{n}}{n} Z_{n}(\beta)\right) . \tag{20.11}
\end{equation*}
$$

According to (20.4) for large $n$ the partition sum behaves as

$$
\begin{equation*}
Z_{n}(\beta) \sim e^{-n(\beta-1) K_{\beta}} \tag{20.12}
\end{equation*}
$$

Substituting this into (20.11) we can see that the leading zero of the characteristic function is

$$
z_{0}(\beta)=e^{(\beta-1) K_{\beta}}
$$

On the other hand substituting the periodic orbit approximation (20.10) into (20.11) and introducing prime and repeated periodic orbits as usual we get

$$
\Omega(z, \beta)=\exp \left(-\sum_{p, r} \frac{z^{n_{p} r} e^{\beta \gamma n_{p} r}}{r\left|\Lambda_{p}^{r}\right|^{\beta}}\right) .
$$

We can see that the characteristic function is the same as the zeta function we introduced for Lyapunov exponents (H.14) except we have $z e^{\beta \gamma}$ instead of $z$. Then we can conclude that the Rényi entropies can be expressed with the pressure function directly as

$$
\begin{equation*}
P(\beta)=(\beta-1) K_{\beta}+\beta \gamma, \tag{20.13}
\end{equation*}
$$

since the leading zero of the zeta function is the pressure. The Rényi entropies $K_{\beta}$, hence the distribution of the exponents $g(\lambda)$ as well, can be calculated via finding the leading eigenvalue of the operator (H.4).

## Figure 20.1:



Figure 20.2: $g(\lambda)$ and $P(\beta)$ for the map of Exercise 20.4 at $a=3$ and $b=3 / 2$. See Solutions N for calculation details.


From (20.13) we can get all the important quantities of the thermodynamic formalism. For $\beta=0$ we get the topological entropy

$$
\begin{equation*}
P(0)=-K_{0}=-h_{t o p} \tag{20.14}
\end{equation*}
$$

For $\beta=1$ we get the escape rate

$$
\begin{equation*}
P(1)=\gamma \tag{20.15}
\end{equation*}
$$

Taking the derivative of (20.13) in $\beta=1$ we get Pesin's formula [H.2] connecting Kolmogorov's entropy and the Lyapunov exponent

$$
\begin{equation*}
P^{\prime}(1)=\bar{\lambda}=K_{1}+\gamma \tag{20.16}
\end{equation*}
$$

It is important to note that, as always, these formulas are strictly valid for nice hyperbolic systems only. At the end of this Chapter we discuss the important problems we are facing in non-hyperbolic cases.

On figure 20.2 we show a typical pressure and $g(\lambda)$ curve computed for the two scale tent map of Exercise 20.4. We have to mention, that all typical hyperbolic dynamical system produces a similar parabola like curve. Although this is somewhat boring we can interpret it like a sign of a high level of universality: The exponents $\lambda$ have a sharp distribution around the most probable value. The most probable value is $\lambda=P^{\prime}(0)$ and $g(\lambda)=h_{\text {top }}$ is the topological entropy. The average value in closed systems is where $g(\lambda)$ touches the diagonal: $\bar{\lambda}=g(\bar{\lambda})$ and $1=g^{\prime}(\bar{\lambda})$.

Next, we are looking at the distribution of trajectories in real space.

### 20.2 Fractal dimensions

By looking at the repeller we can recognize an interesting spatial structure. In the 3-disk case the starting points of trajectories not leaving the system after the first bounce form two strips. Then these strips are subdivided into an infinite hierarchy of substrips as we follow trajectories which do not leave the system after more and more bounces. The finer strips are similar to strips on a larger scale. Objects with such self similar properties are called fractals.

We can characterize fractals via their local scaling properties. The first step is to draw a uniform grid on the surface of section. We can look at various measures in the square boxes of the grid. The most interesting measure is again the natural measure located in the box. By decreasing the size of the grid $\epsilon$ the measure in a given box will decrease. If the distribution of the measure is smooth then we expect that the measure of the $i$ th box is proportional with the dimension of the section

$$
\mu_{i} \sim \epsilon^{d} .
$$

If the measure is distributed on a hairy object like the repeller we can observe unusual scaling behavior of type

$$
\mu_{i} \sim \epsilon^{\alpha_{i}},
$$

where $\alpha_{i}$ is the local "dimension" or Hölder exponent of the the object. As $\alpha$ is not necessarily an integer here we are dealing with objects with fractional dimensions. We can study the distribution of the measure on the surface of section by looking at the distribution of these local exponents. We can define

$$
\alpha_{i}=\frac{\log \mu_{i}}{\log \epsilon}
$$

the local Hölder exponent and then we can count how many of them are between $\alpha$ and $\alpha+d \alpha$. This is $N_{\epsilon}(\alpha) d \alpha$. Again, in smooth objects this function scales simply with the dimension of the system

$$
N_{\epsilon}(\alpha) \sim \epsilon^{-d}
$$

while for hairy objects we expect an $\alpha$ dependent scaling exponent

$$
N_{\epsilon}(\alpha) \sim \epsilon^{-f(\alpha)} .
$$

$f(\alpha)$ can be interpreted [H.7] as the dimension of the points on the surface of section with scaling exponent $\alpha$. We can calculate $f(\alpha)$ with the help of partition sums as we did for $g(\lambda)$ in the previous section. First we define

$$
\begin{equation*}
Z_{\epsilon}(q)=\sum_{i} \mu_{i}^{q} . \tag{20.17}
\end{equation*}
$$

Then we would like to determine the asymptotic behavior of the partition sum characterized by the $\tau(q)$ exponent

$$
Z_{\epsilon}(q) \sim \epsilon^{-\tau(q)} .
$$

The partition sum can be written in terms of the distribution function of $\alpha$-s

$$
Z_{\epsilon}(q)=\int d \alpha N_{\epsilon}(\alpha) \epsilon^{q \alpha} .
$$

Using the asymptotic form of the distribution we get

$$
Z_{\epsilon}(q) \sim \int d \alpha \epsilon^{q \alpha-f(\alpha)}
$$

As $\epsilon$ goes to zero the integral is dominated by the term maximizing the exponent. This $\alpha^{*}$ can be determined from the equation

$$
\frac{d}{d \alpha^{*}}\left(q \alpha^{*}-f\left(\alpha^{*}\right)\right)=0,
$$

leading to

$$
q=f^{\prime}\left(\alpha^{*}\right) .
$$

Finally we can read off the scaling exponent of the partition sum

$$
\tau(q)=\alpha^{*} q-f\left(\alpha^{*}\right)
$$

In a uniform fractal characterized by a single dimension both $\alpha$ and $f(\alpha)$ collapse to $\alpha=f(\alpha)=D$. The scaling exponent then has the form $\tau(q)=$ $(q-1) D$. In case of non uniform fractals we can introduce generalized dimensions [H.9] $D_{q}$ via the definition

$$
D_{q}=\tau(q) /(q-1) .
$$

Some of these dimensions have special names. For $q=0$ the partition sum (20.17) counts the number of non empty boxes $\bar{N}_{\epsilon}$. Consequently

$$
D_{0}=-\lim _{\epsilon \rightarrow 0} \frac{\log \bar{N}_{\epsilon}}{\log \epsilon}
$$

is called the box counting dimension. For $q=1$ the dimension can be determined as the limit of the formulas for $q \rightarrow 1$ leading to

$$
D_{1}=\lim _{\epsilon \rightarrow 0} \sum_{i} \mu_{i} \log \mu_{i} / \log \epsilon .
$$

This is the scaling exponent of the Shannon information entropy [H.16] of the distribution, hence its name is information dimension.

Using equisize grids is impractical in most of the applications. Instead, we can rewrite (20.17) into the more convenient form

$$
\begin{equation*}
\sum_{i} \frac{\mu_{i}^{q}}{\epsilon^{\tau}(q)} \sim 1 . \tag{20.18}
\end{equation*}
$$

If we cover the $i$ th branch of the fractal with a grid of size $l_{i}$ instead of $\epsilon$ we can use the relation [H.8]

$$
\begin{equation*}
\sum_{i} \frac{\mu_{i}^{q}}{l_{i}^{\tau}(q)} \sim 1 \tag{20.19}
\end{equation*}
$$

the non-uniform grid generalization of 20.18. Next we show how can we use the periodic orbit formalism to calculate fractal dimensions. We have already seen that the width of the strips of the repeller can be approximated with the stabilities of the periodic orbits situating in them

$$
l_{i} \sim \frac{1}{\left|\Lambda_{i}\right|}
$$

Then using this relation and the periodic orbit expression of the natural measure we can write (20.19) into the form

$$
\begin{equation*}
\sum_{i} \frac{e^{q \gamma n}}{\left|\Lambda_{i}\right|^{q-\tau(q)}} \sim 1, \tag{20.20}
\end{equation*}
$$

where the summation goes for periodic orbits of length $n$. The sum for stabilities can be expressed with the pressure function again

$$
\sum_{i} \frac{1}{\left|\Lambda_{i}\right|^{q-\tau(q)}} \sim e^{-n P(q-\tau(q))}
$$

and (20.20) can be written as

$$
e^{q \gamma n} e^{-n P(q-\tau(q))} \sim 1,
$$

for large $n$. Finally we get an implicit formula for the dimensions

$$
\begin{equation*}
P\left(q-(q-1) D_{q}\right)=q \gamma . \tag{20.21}
\end{equation*}
$$

Solving this equation directly gives us the partial dimensions of the multifractal repeller along the stable direction. We can see again that the pressure function alone contains all the relevant information. Setting $q=0$ in (20.21) we can prove that the zero of the pressure function is the boxcounting dimension of the repeller

$$
P\left(D_{0}\right)=0 .
$$

Taking the derivative of (20.21) in $q=1$ we get

$$
P^{\prime}(1)\left(1-D_{1}\right)=\gamma .
$$

This way we can express the information dimension with the escape rate and the Lyapunov exponent

$$
\begin{equation*}
D_{1}=1-\gamma / \bar{\lambda} . \tag{20.22}
\end{equation*}
$$

If the system is bound $(\gamma=0)$ the information dimension and all other dimensions are $D_{q}=1$. Also since $D_{1} 0$ is positive (20.22) proves that the Lyapunov exponent must be larger than the escape rate $\bar{\lambda}>\gamma$ in general.

## Commentary

Remark 20.1 Mild phase transition. In non-hyperbolic systems the formulas derived in this chapter should be modified. As we mentioned in 19.1 in nonhyperbolic systems the periodic orbit expression of the measure can be

$$
\mu_{0}=e^{\gamma n} /\left|\Lambda_{0}\right|^{\delta},
$$

where $\delta$ can differ from 1 . Usually it is $1 / 2$. For sufficiently negative $\beta$ the corresponding term $1 /\left|\Lambda_{0}\right|^{\beta}$ can dominate (20.10) while in (20.3) $e^{\gamma n} /\left|\Lambda_{0}\right|^{\mid \beta \beta}$ plays no dominant role. In this case the pressure as a function of $\beta$ can have a kink at the critical point $\beta=\beta_{c}$ where $\beta_{c} \log \left|\Lambda_{0}\right|=\left(\beta_{c}-1\right) K_{\beta_{c}}+\beta_{c} \gamma$. For $\beta<\beta_{c}$ the pressure and the Rényi entropies differ

$$
P(\beta) \neq(\beta-1) K_{\beta}+\beta \gamma
$$

This phenomena is called phase transition. This is however not a very deep problem. We can fix the relation between pressure and the entropies by replacing $1 /\left|\Lambda_{0}\right|$ with $1 /\left|\Lambda_{0}\right|^{\delta}$ in (20.10).

Remark 20.2 Hard phase transition The really deep trouble of thermodynamics is caused by intermittency. In that case we have periodic orbits with $\left|\Lambda_{0}\right| \rightarrow 1$ as $n \rightarrow \infty$. Then for $\beta>1$ the contribution of these orbits dominate both (20.10) and (20.3). Consequently the partition sum scales as $Z_{n}(\beta) \rightarrow 1$ and both the pressure and the entropies are zero. In this case quantities connected with $\beta \leq 1$ make sense only. These are for example the topological entropy, Kolmogorov entropy, Lyapunov exponent, escape rate, $D_{0}$ and $D_{1}$. This phase transition cannot be fixed. It is probably fair to say that quantities which depend on this phase transition are only of mathematical interest and not very useful for characterization of realistic dynamical systems.

Remark 20.3 Multifractals. For reasons that remain mysterious to the authors perhaps so that Mandelbrot can refer to himself both as the mother of fractals and the grandmother of multifractals - some physics literature referes to any fractal generated by more than one scale as a "multi" -fractal. This usage seems to divide fractals into 2 classes; one consisting essentially of the above Cantor set and the Serapinski gasket, and the second consisting of anything else, including all cases of physical interest.

## Résumé

In this chapter we have shown that thermodynamic quantities and various fractal dimensions can be expressed in terms of the pressure function. The pressure function is the leading eigenvalue of the operator which generates the Lyapunov exponent. In the Lyapunov case $\beta$ is just an auxiliary variable. In thermodynamics it plays an essential role. The good news of the chapter is that the distribution of locally fluctuating exponents should not be computed via making statistics. We can use cyclist formulas for determining the pressure. Then the pressure can be found using short cycles +
curvatures. Here the head reach the tail of the snake. We just argued that the statistics of long trajectories coded in $g(\lambda)$ and $P(\beta)$ can be calculated from short cycles. To use this intimate relation between long and short trajectories effectively is still a research level problem.

## References

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[20.2] Ya.B. Pesin, Uspekhi Mat. Nauk 32, 55 (1977), [Russian Math. Surveys 32, 55 (1977)].
[20.3] Even though the thermodynamic formalism is of older vintage (we refer the reader to ref. [20.4] for a comprehensive overview), we adhere here to the notational conventions of ref. [20.5] which are more current in the physics literature: we strongly recommend also ref. [20.6], dealing with period doubling universality.
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## Exercises

Exercise 20.1 Thermodynamics in higher dimensions Introduce the time averages of the eigenvalues of the Jacobian

$$
\begin{equation*}
\lambda_{i}=\lim _{t \rightarrow \infty} \frac{1}{t} \log \left|\Lambda_{i}^{t}\left(x_{0}\right)\right| \tag{20.23}
\end{equation*}
$$

as a generalization of (10.31).
Show that in higher dimensions Pesin's formula is

$$
\begin{equation*}
K_{1}=\sum_{i} \lambda_{i}-\gamma, \tag{20.24}
\end{equation*}
$$

where the summation goes for the positive $\lambda_{i}$-s only. (Hint: Use the higher dimensional generalization of (19.10)

$$
\mu_{i}=e^{n \gamma} /\left|\prod_{j} \Lambda_{i, j}\right|
$$

where the product goes for the expanding eigenvalues of the Jacobian of the periodic orbit.

## Exercise 20.2 Bunimovich stadium Kolmogorov entropy. Take for defini-

 tiveness $a=1.6$ and $d=1$ in the Bunimovich stadium of exercise 6.4,
estimate the Lyapunov exponent by averaging over a very long trajectory. Biham and Kvale [20.14] estimate the discrete time Lyapunov to $\lambda \approx 1.0 \pm .1$, the continuous time Lyapunov to $\lambda \approx 0.43 \pm .02$, the topological entropy (for their symbolic dynamics) $h \approx 1.15 \pm .03$.

Exercise 20.3 Entropy of rugged-edge billiards. Take a semi-circle of diameter $\varepsilon$ and replace the sides of a unit square by $\lfloor 1 / \varepsilon\rfloor$ catenated copies of the semi-circle.

(a) Is the billiard ergodic as $\varepsilon \rightarrow 0$ ?
(b) (hard) Show that the entropy of the billiard map is

$$
K_{1} \rightarrow-\frac{2}{\pi} \ln \varepsilon+\text { const }
$$

as $\varepsilon \rightarrow 0$. (Hint: do not write return maps.)
(c) (harder) Show that when the semi-circles of the Bunimovich stadium are far apart, say $L$, the entropy for the flow decays as

$$
K_{1} \rightarrow \frac{2 \ln L}{\pi L}
$$

Exercise 20.4 Two scale map Compute all those quantities - dimensions, escape rate, entropies, etc. - for the repeller of the one dimensional map

$$
f(x)=\left\{\begin{array}{lll}
1+a x & \text { if } & x<0  \tag{20.25}\\
1-b x & \text { if } & x>0
\end{array}\right.
$$

where $a$ and $b$ are larger than 2. Compute the fractal dimension, plot the pressure and compute the $f(\alpha)$ spectrum of singularities.

Exercise 20.5 Four scale map Compute the Rényi entropies and $g(\lambda)$ for the four scale map

$$
f(x)= \begin{cases}a_{1} x & \text { if } 0<x<b / a_{1}  \tag{20.26}\\ (1-b)\left(\left(x-b / a_{1}\right) /\left(b-b / a_{1}\right)\right)+b & \text { if } b / a_{1}<x<b \\ a_{2}(x-b) & \text { if } b<x<b+b / a_{2} \\ (1-b)\left(\left(x-b-b / a_{2}\right) /\left(1-b-b / a_{2}\right)\right)+b & \text { if } b+b / a_{2}<x<1\end{cases}
$$

Hint: Calculate the pressure function and use (20.13).

Exercise 20.6 Transfer matrix Take the unimodal map $f(x)=\sin (\pi x)$ of the interval $I=[0,1]$. Calculate the four preimages of the intervals $I_{0}=[0,1 / 2]$ and $I_{1}=[1 / 2,1]$. Extrapolate $f(x)$ with piecewise linear functions on these intervals. Find $a_{1}, a_{2}$ and $b$ of the previous exercise. Calculate the pressure function of this linear extrapolation. Work out higher level approximations by linearly extrapolating the map on the $2^{n}$-th preimages of $I$.

## Chapter 21

## Intermittency

Sometimes They Come Back<br>Stephen King<br>(R. Artuso, P. Dahlqvist, G. Tanner and P. Cvitanović)

In the theory of chaotic dynamics developed so far we assumed that the evolution operators have discrete spectra $\left\{z_{0}, z_{1}, z_{2}, \ldots\right\}$ given by the zeros of

$$
1 / \zeta(z)=(\cdots) \prod_{k}\left(1-z / z_{k}\right)
$$

The assumption was based on the tacit premise that the dynamics is everywhere exponentially unstable. Real life is nothing like that - phase spaces are generically infinitely interwoven patterns of stable and unstable behaviors. The stable (in the case of Hamiltonian flows, integrable) orbits do not communicate with the ergodic components of the phase space, and can be treated by classical methods. In general, one is able to treat the dynamics near stable orbits as well as chaotic components of the phase space dynamics well within a periodic orbit approach. Problems occur at the broderline between chaos and regular dynamics where marginally stable orbits and manifolds present difficulties and still unresolved challenges.

We shall use the simplest example of such behavior - intermittency in 1-dimensional maps - to illustrate effects of marginal stability. The main message will be that spectra of evolution operators are no longer discrete, dynamical zeta functions exhibit branch cuts of the form

$$
1 / \zeta(z)=(\cdots)+(1-z)^{\alpha}(\cdots),
$$

and correlations decay no longer exponentially, but as power laws.


Figure 21.1: Typical phase space for an area-preserving map with mixed phase space dynamics; here the standard map for $k=1.2$.

### 21.1 Intermittency everywhere

In many fluid dynamics experiments one observes transitions from regular behaviors to behaviors where long time intervals of regular behavior ("laminar phases") are interrupted by fast irregular bursts. The closer the parameter is to the onset of such bursts, the longer are the intervals of regular behavior. The distributions of laminar phase intervals are well described by power laws.

This phenomenon is called intermittency, and it is a very general aspect of dynamics, a shadow cast by non-hyperbolic, marginally stable phase space regions. Complete hyperbolicity assumed in (14.5) is the exception rather than the rule, and for almost any dynamical system of interest (dynamics in smooth potentials, billiards with smooth walls, the infinite horizon Lorentz gas, etc.) one encounters mixed phase spaces with islands of stability coexisting with hyperbolic regions, see figure 21.1. Wherever stable islands are interspersed with chaotic regions, trajectories which come close to the stable islands can stay 'glued' for arbitrarily long times. These intervals of regular motion are interrupted by irregular bursts as the trajectory is re-injected into the chaotic part of the phase space. How the trajectories are precisely 'glued' to the marginally stable region is often hard to describe. What coarsely looks like a border of an island will under magnification dissolve into infinities of island chains of decreasing sizes, broken tori and bifurcating orbits, as illustrated in figure 21.1.

Intermittency is due to the existence of fixed points and cycles of marginal stability (8.2), or (in studies of the onset of intermittency) to the proximity of a nearly marginal complex or unstable orbits. In Hamiltonian systems intermittency goes hand in hand with the existence of (marginally stable) KAM tori. In more general settings, the existence of marginal or nearly marginal orbits is due to incomplete intersections of stable and unstable manifolds in a Smale horseshoe type dynamics (see figure 13.2). Following the stretching and folding of the invariant manifolds in time one will inevitably find phase space points at which the stable and unstable manifolds

Figure 21.2: A complete binary repeller with a marginal fixed point.

are almost or exactly tangential to each other, implying non-exponential separation of nearby points in phase space or, in other words, marginal stability. Under small parameter perturbations such neighborhoods undergo tangent bifurcations - a stable/unstable pair of periodic orbits is destroyed or created by coalescing into a marginal orbit, so the pruning which we shall encounter in chapter 12, and the intermittency discussed here are two sides of the same coin.

How to deal with the full complexity of a typical Hamiltonian system with mixed phase space is a very difficult, still open problem. Nevertheless, it is possible to learn quite a bit about intermittency by considering rather simple examples. Here we shall restrict our considerations to 1-dimensional maps which in the neighborhood of a single marginally stable fixed point at $x=0$ take the form

$$
\begin{equation*}
x \mapsto f(x)=x+O\left(x^{1+s}\right) \tag{21.1}
\end{equation*}
$$

and are expanding everywhere else. Such a map may allow for escape, like the map shown in figure 21.2 or the dynamics may be bounded, like the Farey map (18.27) 163,164c153,154

$$
x \mapsto f(x)= \begin{cases}x /(1-x) & x \in[0,1 / 2[ \\ (1-x) / x & x \in[1 / 2,1]\end{cases}
$$

introduced in sect. 18.4.
Figure 21.3 compares a trajectory of the tent map (11.8) side by side with a trajectory of the Farey map. In a stark contrast to the uniformly chaotic trajectory of the tent map, the Farey map trajectory alternates intermittently between slow regular motion close to the marginally stable fixed point, and chaotic bursts.

The presence of marginal stability has striking dynamical consequences: correlation decay may exhibit long range power law asymptotic behavior and diffusion processes can assume anomalous character. Escape from a repeller of the form figure 21.2 may be algebraic rather than exponential. In long time explorations of the dynamics intermittency manifests itself


Figure 21.3: (a) A tent map trajectory. (b) A Farey map trajectory.
by enhancement of natural measure in the proximity of marginally stable cycles.

The questions we shall address here are: how does marginal stability affect zeta functions or spectral determinants? And, can we deduce power law decays of correlations from cycle expansions?

In example 16.5 we saw that marginal stability violates one of the conditions which ensure that the spectral determinant is an entire function. Already the simple fact that the cycle weight $1 /\left|1-\Lambda_{p}^{r}\right|$ in the trace (14.3) or the spectral determinant (15.3) diverges for marginal orbits with $\left|\Lambda_{p}\right|=1$ tells us that we have to treat these orbits with care.

In the following we will incorporate marginal stability orbits into cycleexpansions in a systematic manner. To get to know the difficulties lying ahead, we will start in sect. 21.2 with a piecewise linear map, with the asymptotics (21.1). We will construct a dynamical zeta function in the usual way without worrying too much about its justification and show that it has a branch cut singularity. We will calculate the rate of escape from our piecewise linear map and find that it is characterized by decay, rather than exponential decay, a power law. We will show that dynamical zeta functions in the presence of marginal stability can still be written in terms of periodic orbits, exactly as in chapters 10 and 19 , with one exception: the marginally stable orbits have to be explicitly excluded. This innocent looking step has far reaching consequences; it forces us to change the symbolic dynamics from a finite to an infinite alphabet, and entails a reorganization of the order of summations in cycle expansions, sect. 21.2.4.

Branch cuts are typical also for smooth intermittent maps with isolated marginally stable fixed points and cycles. In sect. 21.3, we discuss the
cycle expansions and curvature combinations for zeta functions of smooth maps tailored to intermittency. The knowledge of the type of singularity one encounters enables us to develop the efficient resummation method presented in sect. 21.3.1.

Finally, in sect. 21.4, we discuss a probabilistic approach to intermittency that yields approximate dynamical zeta functions and provides valuable information about more complicated systems, such as billiards.

### 21.2 Intermittency for pedestrians

Intermittency does not only present us with a large repertoire of interesting dynamics, it is also at the root of many sorrows such as slow convergence of cycle expansions. In order to get to know the kind of problems which arise when studying dynamical zeta functions in the presence of marginal stability we will consider an artfully concocted piecewise linear model first. From there we will move on to the more general case of smooth intermittant maps, sect. 21.3.

### 21.2.1 A toy map

The Bernoulli shift map (16.6) is an idealized, but highly instructive, example of a hyperbolic map. To study intermittency we will now construct a likewise piecewise linear model, an intermittent map stripped down to its bare essentials.

Consider a map $x \mapsto f(x)$ on the unit interval $\mathcal{M}=[0,1]$ with two monotone branches

$$
f(x)=\left\{\begin{array}{ll}
f_{0}(x) & \text { for } x \in \mathcal{M}_{0}=[0, a]  \tag{21.2}\\
f_{1}(x) & \text { for } x \in \mathcal{M}_{1}=[b, 1]
\end{array} .\right.
$$

The two branches are assumed complete, that is $f_{0}\left(\mathcal{M}_{0}\right)=f_{1}\left(\mathcal{M}_{1}\right)=\mathcal{M}$. The map allows escape if $a<b$ and is bounded if $a=b$ (see figure 21.2 and figure 21.4). We take the right branch to be expanding and linear:

$$
f_{1}(x)=\frac{1}{1-b}(x-b) .
$$

Next, we will construct the left branch in a way, which will allow us to model the intermittent behavior (21.1) near the origin. We chose a monotonically decreasing sequence of points $q_{n}$ in $[0, a]$ with $q_{1}=a$ and $q_{n} \rightarrow 0$ as $n \rightarrow \infty$. This sequence defines a partition of the left interval $\mathcal{M}_{0}$ into an infinite number of connected intervals $\mathcal{M}_{n}, n \geq 2$ with

$$
\begin{equation*}
\left.\left.\mathcal{M}_{n}=\right] q_{n}, q_{n-1}\right] \quad \text { and } \quad \mathcal{M}_{0}=\bigcup_{n=2}^{\infty} \mathcal{M}_{n} \tag{21.3}
\end{equation*}
$$

Figure 21.4: A piecewise linear intermittent map of (21.2) type: more specifically, the map piecewise linear over intervals (21.8) of the toy example studied below, $a=.5, b=.6, s=1.0$.


The map $f_{0}(x)$ is now specified by the following requirements

- $f_{0}(x)$ is continuous.
- $f_{0}(x)$ is linear on the intervals $\mathcal{M}_{n}$ for $n \geq 2$.
- $f_{0}\left(q_{n}\right)=q_{n-1}$, that is $\mathcal{M}_{n}=f_{0}^{-n+1}([a, 1])$.

This fixes the map for any given sequence $\left\{q_{n}\right\}$. The last condition ensures the existence of a simple Markov partition. The slopes of the various linear segments are

$$
\begin{array}{lll}
f_{0}^{\prime}(x)=\frac{f_{0}\left(q_{n-1}\right)-f_{0}\left(q_{n}\right)}{q_{n}-1-q_{n}} & =\frac{\left|\mathcal{M}_{n-1}\right|}{\left|\mathcal{M}_{n}\right|} & \text { for } x \in \mathcal{M}_{n}, n \geq 3 \\
f_{0}^{\prime}(x)=\frac{f_{0}\left(q_{1}-f_{0}\left(q_{2}\right)\right.}{q_{1}-q_{2}} & =\frac{1-a}{\left|\mathcal{M}_{2}\right|} & \text { for } x \in \mathcal{M}_{2}  \tag{21.4}\\
f_{0}^{\prime}(x)=\quad \frac{1-1}{1-b} & =\frac{|\mathcal{M}|}{\left|\mathcal{M}_{1}\right|} & \text { for } x \in \mathcal{M}_{1}
\end{array}
$$

with $\left|\mathcal{M}_{n}\right|=q_{n-1}-q_{n}$ for $n \geq 2$. Note that we do not require as yet that the map exhibit intermittent behavior.

We will see that the family of periodic orbits with code $10^{n}$ plays a key role for intermittent maps of the form (21.1). An orbit $10^{n}$ enters the intervals $\mathcal{M}_{1} \rightarrow \mathcal{M}_{n+1} \rightarrow \mathcal{M}_{n} \rightarrow \ldots \rightarrow \mathcal{M}_{2}$ successively and the family approaches the marginal stable fixed point at $x=0$ for $n \rightarrow \infty$. The stability of a cycle $10^{n}$ for $n \geq 1$ is given by the chain rule (4.34),

$$
\begin{equation*}
\Lambda_{10^{n}}=f_{0}^{\prime}\left(x_{n+1}\right) f_{0}^{\prime}\left(x_{n}\right) \ldots f_{0}^{\prime}\left(x_{2}\right) f_{1}^{\prime}\left(x_{1}\right)=\frac{1}{\left|\mathcal{M}_{n+1}\right|} \frac{1-a}{1-b}, \tag{21.5}
\end{equation*}
$$

with $x_{i} \in \mathcal{M}_{i}$.
The properties of the map (21.2) are completely determined by the sequence $\left\{q_{n}\right\}$. By choosing $q_{n}=2^{-n}$, for example, we recover the uniformly hyperbolic Bernoulli shift map (16.6). An intermittent map of the form (21.3) having the asymptotic behavior (21.1) can be constructed by choosing an algebraically decaying sequence $\left\{q_{n}\right\}$ behaving asymptotically like

$$
\begin{equation*}
q_{n} \sim \frac{1}{n^{1 / s}}, \tag{21.6}
\end{equation*}
$$

where $s$ is the intermittency exponent in (21.1). Such a partition leads to intervals whose length decreases asymptotically like a power-law, that is,

$$
\begin{equation*}
\left|\mathcal{M}_{n}\right| \sim \frac{1}{n^{1+1 / s}} \tag{21.7}
\end{equation*}
$$

As can be seen from (21.5), the stability eigenvalues of periodic orbit families approaching the marginal fixed point, such as the $10^{n}$ family increase in turn only algebraically with the cycle length.

It may now seem natural to construct an intermittent toy map in terms of a partition $\left|\mathcal{M}_{n}\right|=1 / n^{1+1 / s}$, that is, a partition which follows (21.7) exactly. Such a choice leads to a dynamical zeta function which can be written in terms of so-called Jonquière functions (or polylogarithms) which arise naturally also in the context of the Farey map (18.27), and the anomalous diffusion of sect. 23.3. We will, however, not go along this route here; instead, we will engage in a bit of reverse engineering and construct a less obvious partition which will simplify the algebra considerably later without loosing any of the key features typical for intermittent systems. We fix the intermittent toy map by specifying the intervals $\mathcal{M}_{n}$ in terms of Gamma functions according to

$$
\begin{equation*}
\left|\mathcal{M}_{n}\right|=\mathcal{C} \frac{\Gamma(n+m-1 / s-1)}{\Gamma(n+m)} \quad \text { for } \quad n \geq 2 \tag{21.8}
\end{equation*}
$$

where $m=[1 / s]$ denotes the integer part of $1 / s$ and $\mathcal{C}$ is a normalization constant fixed by the condition $\sum_{n=2}^{\infty}\left|\mathcal{M}_{n}\right|=q_{1}=a$, that is,

$$
\begin{equation*}
\mathcal{C}=a\left[\sum_{n=m+1}^{\infty} \frac{\Gamma(n-1 / s)}{\Gamma(n+1)}\right]^{-1} . \tag{21.9}
\end{equation*}
$$

Using Stirling's formula for the Gamma function

$$
\Gamma(z) \sim e^{-z} z^{z-1 / 2} \sqrt{2 \pi}(1+1 / 12 z+\ldots)
$$

we verify that the intervals decay asymptotically like $n^{-(1+1 / s)}$, as required by the condition (21.7).

Next, let us write down the dynamical zeta function of the toy map in terms of its periodic orbits, that is

$$
1 / \zeta(z)=\prod_{p}\left(1-\frac{z^{n_{p}}}{\left|\Lambda_{p}\right|}\right)
$$

One may be tempted to expand the dynamical zeta function in terms of the binary symbolic dynamics of the map; we saw, however, in sect. 18.4 that
such cycle expansion converges extremely slowly. The shadowing mechanism between orbits and pseudo-orbits fails for orbits of the form $10^{n}$ with stabilities given by (21.5), due to the marginal stability of the fixed point $\overline{0}$. It is therefore advantageous to choose as the fundamental cycles the family of orbits with code $10^{n}$ or, equivalently, switch from the finite (binary) alphabet to an infinite alphabet given by

$$
10^{n-1} \rightarrow n .
$$

Due to the piecewise-linear form of the map which maps intervals $\mathcal{M}_{n}$ exactly onto $\mathcal{M}_{n-1}$, all periodic orbits entering the left branch at least twice are canceled exactly by pseudo cycles, and the cycle expanded dynamical zeta function depends only on the fundamental series $1,10,100, \ldots$ :

$$
\begin{align*}
1 / \zeta(z) & =\prod_{p \neq 0}\left(1-\frac{z^{n_{p}}}{\left|\Lambda_{p}\right|}\right)=1-\sum_{n=1}^{\infty} \frac{z^{n}}{\left|\Lambda_{10^{n-1}}\right|} \\
& =1-(1-b) z-\mathcal{C} \frac{1-b}{1-a} \sum_{n=2}^{\infty} \frac{\Gamma(n+m-1 / s-1)}{\Gamma(n+m)} z^{n} . \tag{21.10}
\end{align*}
$$

The fundamental term (18.5) consists here of an infinite sum over algebraically decaying cycle weights. The sum is divergent for $|z| \geq 1$. We will see that this behavior is due to a branch cut of $1 / \zeta$ starting at $z=1$. We need to find analytic continuations of sums over algebraically decreasing terms in (21.10). Note also that we omitted the fixed point $\overline{0}$ in the above Euler product; we will discussed this point as well as a proper derivation of the zeta function in more detail in sect. 21.2.4.

### 21.2.2 Branch cuts

Starting from the dynamical zeta function (21.10), we first have to worry about finding an analytical continuation of the sum for $|z| \geq 1$. We do, however, get this part for free here due to the particular choice of interval lengths made in (21.8). The sum over ratios of Gamma functions in (21.10) can be evaluated analytically by using the following identities valid for $1 / s=\alpha>0$ (the famed binomial theorem in disguise),

- $\alpha$ non-integer

$$
\begin{equation*}
(1-z)^{\alpha}=\sum_{n=0}^{\infty} \frac{\Gamma(n-\alpha)}{\Gamma(-\alpha) \Gamma(n+1)} z^{n} \tag{21.11}
\end{equation*}
$$

- $\alpha$ integer

$$
\begin{align*}
(1-z)^{\alpha} \log (1-z) & =\sum_{n=1}^{\alpha}(-1)^{n} c_{n} z^{n}  \tag{21.12}\\
& +(-1)^{\alpha+1} \alpha!\sum_{n=\alpha+1}^{\infty} \frac{(n-\alpha-1)!}{n!} z^{n}
\end{align*}
$$

with

$$
c_{n}=\binom{\alpha}{n} \sum_{k=0}^{n-1} \frac{1}{\alpha-k} .
$$

In order to simplify the notation, we restrict the intermittency parameter to the range $1 \leq 1 / s<2$ with $[1 / s]=m=1$. All what follows can easily be generalized to arbitrary $s>0$ using equations (21.11) and (21.12). The infinite sum in (21.10) can now be evaluated with the help of (21.11) or (21.12), that is,

$$
\sum_{n=2}^{\infty} \frac{\Gamma(n-1 / s)}{\Gamma(n+1)} z^{n}= \begin{cases}\Gamma\left(-\frac{1}{s}\right)\left[(1-z)^{1 / s}-1+\frac{1}{s} z\right] & \text { for } \quad 1<1 / s<2 \\ (1-z) \log (1-z)+z & \text { for } \quad s=1\end{cases}
$$

The normalization constant $\mathcal{C}$ in (21.8) can be evaluated explicitly using (21.9) and the dynamical zeta function can be given in closed form. We obtain for $1<1 / s<2$

$$
\begin{equation*}
1 / \zeta(z)=1-(1-b) z-\frac{a}{1 / s-1} \frac{1-b}{1-a}\left((1-z)^{1 / s}-1+\frac{1}{s} z\right) . \tag{21.13}
\end{equation*}
$$

and for $s=1$,

$$
\begin{equation*}
1 / \zeta(z)=1-(1-b) z-a \frac{1-b}{1-a}((1-z) \log (1-z)+z) . \tag{21.14}
\end{equation*}
$$

It now becomes clear why the particular choice of intervals $\mathcal{M}_{n}$ made in the last section is useful; by summing over the infinite family of periodic orbits $0^{n} 1$ explicitly, we have found the desired analytical continuation for the dynamical zeta function for $|z| \geq 1$. The function has a branch cut starting at the branch point $z=1$ and running along the positive real axis. That means, the dynamical zeta function takes on different values when approaching the positive real axis for $\operatorname{Re} z>1$ from above and below. The dynamical zeta function for general $s>0$ takes on the form

$$
\begin{equation*}
1 / \zeta(z)=1-(1-b) z-\frac{a}{g_{s}(1)} \frac{1-b}{1-a} \frac{1}{z^{m-1}}\left((1-z)^{1 / s}-g_{s}(z)\right) \tag{21.15}
\end{equation*}
$$

for non-integer $s$ with $m=[1 / s]$ and

$$
\begin{equation*}
1 / \zeta(z)=1-(1-b) z-\frac{a}{g_{m}(1)} \frac{1-b}{1-a} \frac{1}{z^{m-1}}\left((1-z)^{m} \log (1-z)-g_{m}(z)\right) \tag{21.16}
\end{equation*}
$$

for $1 / s=m$ integer and $g_{s}(z)$ are polynomials of order $m=[1 / s]$ which can be deduced from (21.11) or (21.12). We thus find algebraic branch cuts
for non integer intermittency exponents $1 / s$ and logarithmic branch cuts for $1 / s$ integer. We will see in sect. 21.3 that branch cuts of that form are generic for 1-dimensional intermittent maps.

Branch cuts are the all important new feature of dynamical zeta functions due to intermittency. So, how do we calculate averages or escape rates of the dynamics of the map from a dynamical zeta function with branch cuts? We take 'a learning by doing' approach and calculate the escape from our toy map for $a<b$.

### 21.2.3 Escape rate

Our starting point for the calculation of the fraction of survivors after $n$ time steps, is the integral representation (15.19)

$$
\begin{equation*}
\Gamma_{n}=\frac{1}{2 \pi i} \oint_{\gamma_{r}^{-}} z^{-n}\left(\frac{d}{d z} \log \zeta^{-1}(z)\right) d z \tag{21.17}
\end{equation*}
$$

where the contour encircles the origin in the clockwise direction. If the contour lies inside the unit circle $|z|=1$, we may expand the logarithmic derivative of $\zeta^{-1}(z)$ as a convergent sum over all periodic orbits. Integrals and sums can be interchanged, the integrals can be solved term by term, and the formula (14.22) is recovered. For hyperbolic maps, cycle expansion methods or other techniques may provide an analytic extension of the dynamical zeta function beyond the leading zero; we may therefore deform the original contour into a larger circle with radius $R$ which encircles both poles and zeros of $\zeta^{-1}(z)$, see figure 21.5(a). Residue calculus turns this into a sum over the zeros $z_{\alpha}$ and poles $z_{\beta}$ of the dynamical zeta function, that is

$$
\begin{equation*}
\Gamma_{n}=\sum_{\left|z_{\alpha}\right|<R}^{\text {zeros }} \frac{1}{z_{\alpha}^{n}}-\sum_{\left|z_{\beta}\right|<R}^{\text {poles }} \frac{1}{z_{\beta}^{n}}+\frac{1}{2 \pi i} \oint_{\gamma_{R}^{-}} d z z^{-n} \frac{d}{d z} \log \zeta^{-1} \tag{21.18}
\end{equation*}
$$

where the last term gives a contribution from a large circle $\gamma_{R}^{-}$. We thus find exponential decay of $\Gamma_{n}$ dominated by the leading zero or pole of $\zeta^{-1}(z)$, see chapter 20.1 for more details.

Things change considerably in the intermittent case. The point $z=1$ is a branch cut singularity and there exists no Taylor series expansion of $\zeta^{-1}$ around $z=1$. Secondly, the path deformation that led us to (21.18) requires more care, as it must not cross the branch cut. When expanding the contour to large $|z|$ values, we have to deform it along the branch $\operatorname{Re}(z) \geq 1, \operatorname{Im}(z)=0$ encircling the branch cut in anti-clockwise direction, see figure $21.5(\mathrm{~b})$. We will denote the detour around the cut as $\gamma_{c u t}$. We may write symbolically

$$
\oint_{\gamma_{r}}=\sum^{\text {zeros }}-\sum^{\text {poles }}+\oint_{\gamma_{R}}+\oint_{\gamma_{c u t}}
$$



Figure 21.5: The survival probability $\Gamma_{n}$ calculated by contour integration; integrating (21.17) inside the domain of convergence $|z|<1$ (shaded area) of $1 / \zeta(z)$ in periodic orbit representation yields (14.22). A deformation of the contour $\gamma_{r}^{-}$(dashed line) to a larger circle $\gamma_{R}^{-}$gives contributions from the poles and zeros $(\times)$of $1 / \zeta(z)$ between the two circles. These are the only contributions for hyperbolic maps (a), for intermittent systems additional contributions arise, given by the contour $\gamma_{c u t}$ running along the branch cut (b).
where the sums include only the zeros and the poles in the area enclosed by the contours. The asymptotics is controlled by the zero, pole or cut closest to the origin.

Let us now go back to our intermittent toy map. The asymptotics of the survival probability of the map is here governed by the behavior of the integrand $\frac{d}{d z} \log \zeta^{-1}$ in (21.17) at the branch point $z=1$. We restrict ourselves again to the case $1<1 / s<2$ first and write the dynamical zeta function (21.13) in the form

$$
1 / \zeta(z)=a_{0}+a_{1}(1-z)+b_{0}(1-z)^{1 / s} \equiv G(1-z)
$$

and

$$
a_{0}=\frac{b-a}{1-a}, \quad b_{0}=\frac{a}{1-1 / s} \frac{1-b}{1-a} .
$$

Setting $u=1-z$, we need to evaluate

$$
\begin{equation*}
\frac{1}{2 \pi i} \oint_{\gamma_{\text {cut }}}(1-u)^{-n} \frac{d}{d u} \log G(u) d u \tag{21.19}
\end{equation*}
$$

where $\gamma_{\text {cut }}$ goes around the cut (that is, the negative $u$ axis). Expanding the integrand $\frac{d}{d u} \log G(u)=G^{\prime}(u) / G(u)$ in powers of $u$ and $u^{1 / s}$ at $u=0$, one obtains

$$
\begin{equation*}
\frac{d}{d u} \log G(u)=\frac{a_{1}}{a_{0}}+\frac{1}{s} \frac{b_{0}}{a_{0}} u^{1 / s-1}+O(u) . \tag{21.20}
\end{equation*}
$$

Figure 21.6: The asymptotic escape from an intermittent repeller is a power law. Normally it is preceded by an exponential, which can be related to zeros close to the cut but beyond the branch point $z=1$, as in figure 21.5(b).


The integrals along the cut may be evaluated using the general formula

$$
\begin{equation*}
\frac{1}{2 \pi i} \oint_{\gamma_{c u t}} u^{\alpha}(1-u)^{-n} d u=\frac{\Gamma(n-\alpha-1)}{\Gamma(n) \Gamma(-\alpha)} \sim \frac{1}{n^{\alpha+1}}(1+O(1 / n))( \tag{21.21}
\end{equation*}
$$

which can be obtained by deforming the contour back to a loop around the point $u=1$, now in positive (anti-clockwise) direction. The contour integral then picks up the $(n-1)$ st term in the Taylor expansion of the function $u^{\alpha}$ at $u=1$, cf. (21.11). For the continuous time case the corresponding formula is

$$
\begin{equation*}
\frac{1}{2 \pi i} \oint_{\gamma_{c u t}} z^{\alpha} e^{z t} d z=\frac{1}{\Gamma(-\alpha)} \frac{1}{t^{\alpha+1}} \tag{21.22}
\end{equation*}
$$

Plugging (21.20) into (21.19) and using (21.21) we get the asymptotic result

$$
\begin{equation*}
\Gamma_{n} \sim \frac{b_{0}}{a_{0}} \frac{1}{s} \frac{1}{\Gamma(1-1 / s)} \frac{1}{n^{1 / s}}=\frac{a}{s-1} \frac{1-b}{b-a} \frac{1}{\Gamma(1-1 / s)} \frac{1}{n^{1 / s}} \tag{21.23}
\end{equation*}
$$

We see that, asymptotically, the escape from an intermittent repeller is described by power law decay rather than the exponential decay we are familiar with for hyperbolic maps; a numerical simulation of the power-law escape from an intermittent repeller is shown in figure 21.6.

For general non-integer $1 / s>0$, we write

$$
1 / \zeta(z)=A(u)+(u)^{1 / s} B(u) \equiv G(u)
$$

with $u=1-z$ and $A(u), B(u)$ are functions analytic in a disc of radius 1 around $u=0$. The leading terms in the Taylor series expansions of $A(u)$ and $B(u)$ are

$$
a_{0}=\frac{b-a}{1-a}, \quad b_{0}=\frac{a}{g_{s}(1)} \frac{1-b}{1-a}
$$

see (21.15). Expanding $\frac{d}{d u} \log G(u)$ around $u=0$, one again obtains leading order contributions according to (21.20) and the general result follows immediately using (21.21), that is,

$$
\begin{equation*}
\Gamma_{n} \sim \frac{a}{s g_{s}(1)} \frac{1-b}{b-a} \frac{1}{\Gamma(1-1 / s)} \frac{1}{n^{1 / s}} . \tag{21.24}
\end{equation*}
$$

Applying the same arguments for integer intermittency exponents $1 / s=m$, one obtains

$$
\begin{equation*}
\Gamma_{n} \sim(-1)^{m+1} \frac{a}{s g_{m}(1)} \frac{1-b}{b-a} \frac{m!}{n^{m}} . \tag{21.25}
\end{equation*}
$$

So far, we have considered the survival probability for a repeller, that is we assumed $a<b$. The formulas (21.24) and (21.25) do obviously not apply for the case $a=b$, that is, for the bounded map. The coefficient $a_{0}=(b-a) /(1-a)$ in the series representation of $G(u)$ is zero, and the expansion of the logarithmic derivative of $G(u)(21.20)$ is no longer valid. We get instead

$$
\frac{d}{d u} \log G(u)=\left\{\begin{array}{cl}
\frac{1}{u}\left(1+O\left(u^{1 / s-1}\right)\right) & s<1 \\
\frac{1}{u}\left(\frac{1}{s}+O\left(u^{1-1 / s}\right)\right) & s>1
\end{array},\right.
$$

assuming non-integer $1 / s$ for convenience. One obtains for the survival probability.

$$
\Gamma_{n} \sim\left\{\begin{array}{cc}
1+O\left(n^{1-1 / s}\right) & s<1 \\
1 / s+O\left(n^{1 / s-1}\right) & s>1
\end{array} .\right.
$$

For $s>1$, this is what we expect. There is no escape, so the survival probability is equal to 1 , which we get as an asymptotic result here. The result for $s>1$ is somewhat more worrying. It says that $\Gamma_{n}$ defined as sum over the instabilities of the periodic orbits as in (19.12) does not tend to unity for large $n$. However, the case $s>1$ is in many senses anomalous. For instance, the invariant density cannot be normalized. It is therefore not reasonable to expect that periodic orbit theories will work without complications.

### 21.2.4 Why does it work (anyway)?

Due to the piecewise linear nature of the map constructed in the previous section, we had the nice property that interval lengths did exactly coincide with the inverse of the stability of periodic orbits of the system, that is

$$
\left|\mathcal{M}_{n}\right|=\mid / \Lambda_{10}^{n-1} .
$$

There is thus no problem in replacing the survival probability $\Gamma_{n}$ given by (1.2), (19.2), that is the fraction of phase space $\mathcal{M}$ surviving $n$ iterations of the map,

$$
\Gamma_{n}=\frac{1}{|\mathcal{M}|} \sum_{i}^{(n)}\left|\mathcal{M}_{i}\right|
$$

by a sum over periodic orbits of the form (14.22). The only orbit to worry about is the marginal fixed point $\overline{0}$ itself which we excluded from the zeta function (21.10).

For smooth intermittent maps, things are less clear and the fact that we had to prune the marginal fixed point is a warning sign that interval estimates by periodic orbit stabilities might go horribly wrong. The derivation of the survival probability in terms of cycle stabilities in chapter 19 did indeed rely heavily on a hyperbolicity assumption which is clearly not fulfilled for intermittent maps. We therefore have to carefully reconsider this derivation in order to show that periodic orbit formulas are actually valid for intermittent systems in the first place.

We will for simplicity consider maps, which have a finite number of say $s$ branches defined on intervals $\mathcal{M}_{s}$ and we assume that the map maps each interval $\mathcal{M}_{s}$ onto $\mathcal{M}$, that is $f\left(\mathcal{M}_{s}\right)=\mathcal{M}$. This ensures the existence of a complete symbolic dynamics - just to make things easy (see figure 21.2).

The generating partition is composed of the domains $\mathcal{M}_{s}$. The $n$th level partition $\mathcal{C}^{(n)}=\left\{\mathcal{M}_{i}\right\}$ can be constructed iteratively. Here $i$ 's are words $i=s_{2} s_{2} \ldots s_{n}$ of length $n$, and the intervals $\mathcal{M}_{i}$ are constructed recursively

$$
\begin{equation*}
\mathcal{M}_{s j}=f_{s}^{-1}\left(\mathcal{M}_{j}\right), \tag{21.26}
\end{equation*}
$$

where $s j$ is the concatenation of letter $s$ with word $j$ of length $n_{j}<n$.
In what follows we will concentrate on the survival probability $\Gamma_{n}$, postponing other quantities of interest, such as averages, to later considerations. In establishing the equivalence of the survival probability and the periodic orbit formula for the escape rate for hyperbolic systems we have assumed that the map is expanding, with a minimal expansion rate $\left|f^{\prime}(x)\right| \geq \Lambda_{\min }>1$. This enabled us to bound the size of every survivor strip $\mathcal{M}_{i}$ by (19.6), the stability $\Lambda_{i}$ of the periodic orbit $i$ within the $\mathcal{M}_{i}$, and bound the survival probability by the periodic orbit sum (19.7).

The bound (19.6)

$$
\mathcal{C}_{1} \frac{1}{\left|\Lambda_{i}\right|}<\frac{\left|\mathcal{M}_{i}\right|}{|\mathcal{M}|}<\mathcal{C}_{2} \frac{1}{\left|\Lambda_{i}\right|}
$$

relies on hyperbolicity, and is thus indeed violated for intermittent systems. The problem is that now there is no lower bound on the expansion rate, the
minimal expansion rate is $\Lambda_{\text {min }}=1$. The survivor strip $\mathcal{M}_{0^{n}}$ which includes the marginal fixed point is thus completely overestimated by $1 /\left|\Lambda_{0^{n}}\right|=1$ which is constant for all $n$.

However, bounding survival probability strip by strip is not what is required for establishing the bound (19.7). For intermittent systems a somewhat weaker bound can be established, saying that the average size of intervals along a periodic orbit can be bounded close to the stability of the periodic orbit for all but the interval $\mathcal{M}_{0^{n}}$. The weaker bound applies to averaging over each prime cycle $p$ separately

$$
\begin{equation*}
\mathcal{C}_{1} \frac{1}{\left|\Lambda_{p}\right|}<\frac{1}{n_{p}} \sum_{i \in p} \frac{\left|\mathcal{M}_{i}\right|}{|\mathcal{M}|}<\mathcal{C}_{2} \frac{1}{\left|\Lambda_{p}\right|}, \tag{21.27}
\end{equation*}
$$

where the word $i$ represents a code of the periodic orbit $p$ and all its cyclic permutations. It can be shown that one can find positive constants $\mathcal{C}_{1}$, $\mathcal{C}_{2}$ independent of $p$. Summing over all periodic orbits leads then again to (19.7).

To study averages of multiplicative weights we follow sect. 10.1 and introduce a phase space observable $a(x)$ and the integrated quantity

$$
A^{n}(x)=\sum_{k=0}^{n-1} a\left(f^{k}(x)\right)
$$

This leads us to introduce the generating function (10.10)

$$
\left\langle e^{\beta A^{n}(x)}\right\rangle
$$

where $\langle$.$\rangle denote some averaging over the distribution of initial points, which$ we choose to be uniform (rather than the a priori unknown invariant density). Again, all we have to show is, that constants $\mathcal{C}_{1}, \mathcal{C}_{2}$ exist, such that

$$
\begin{equation*}
\mathcal{C}_{1} \frac{e^{\beta A_{p}}}{\left|\Lambda_{p}\right|}<\frac{1}{n_{p}} \sum_{i \in p} \frac{1}{|\mathcal{M}|} \int_{\mathcal{M}_{Q}} e^{\beta A^{n}(x)} d x<\mathcal{C}_{2} \frac{e^{\beta A_{p}}}{\left|\Lambda_{p}\right|} \tag{21.28}
\end{equation*}
$$

is valid for all $p$. After performing the above average one gets

$$
\begin{equation*}
\mathcal{C}_{1} \Gamma_{n}(\beta)<\frac{1}{|\mathcal{M}|} \int_{\mathcal{M}} e^{\beta A(x, n)} d x<\mathcal{C}_{2} \Gamma_{n}(\beta) \tag{21.29}
\end{equation*}
$$

with

$$
\begin{equation*}
\Gamma_{n}(\beta)=\sum_{p}^{n} \frac{e^{\beta A_{p}}}{\left|\Lambda_{p}\right|} . \tag{21.30}
\end{equation*}
$$

Figure 21.7: Markov graph corresponding to the alphabet $\left\{0^{k-1} 1 ; \overline{0}, k \geq 1\right\}$

and a dynamical zeta function can be derived. In the intermittent case one can expect that the bound (21.28) holds using an averaging argument similar to the one discussed in (21.27). This justifies the use of dynamical zeta functions for intermittent systems.

One lesson we should have learned so far is that the natural alphabet to use is not $\{0,1\}$ but rather the infinite alphabet $\left\{0^{k-1} 1, \overline{0} ; k \geq 1\right\}$. The symbol 0 occurs unaccompanied by any 1's only in the $\overline{0}$ marginal fixed point which is disconnected from the rest of the Markov graph see figure 21.7.

What happens if we remove a single prime cycle from a dynamical zeta function? In the hyperbolic case such a removal introduces a pole in the $1 / \zeta$ and slows down the convergence of cycle expansions. The heuristic interpretation of such a pole is that for a subshift of finite type removal of a single prime cycle leads to unbalancing of cancellations within the infinity of of shadowing pairs. Nevertheless, removal of a single prime cycle is an exponentially small perturbation of the trace sums, and the asymptotics of the associated trace formulas is unaffected.

In the intermittent case, the fixed point $\overline{0}$ does not provide any shadowing (cf. sect. J.1), and a statement such as

$$
\Lambda_{1 \cdot 0^{k+1}} \approx \Lambda_{1 \cdot 0^{k}} \Lambda_{0},
$$

is meaningless. It seems therefore sensible to take out the factor $\left(1-t_{0}\right)=$ $1-z$ from the product representation of the dynamical zeta function (15.15), that is, to consider a pruned dynamical zeta function $1 / \zeta_{\text {inter }}(z)$ defined by

$$
1 / \zeta(z)=(1-z) 1 / \zeta_{\text {inter }}(z) .
$$

We saw in the last sections, that the zeta function $1 / \zeta_{\text {inter }}(z)$ has all the nice properties we know from the hyperbolic case, that is, we can find a cycle expansion with - in the toy model case - vanishing curvature contributions and we can calculate dynamical properties like escape after having understood, how to handle the branch cut. But you might still be worried about leaving out the extra factor $1-z$ all together. It turns out, that this is not only a matter of convenience, omitting the marginal $\overline{0}$ cycle is a dire necessity. The cycle weight $\Lambda_{0}^{n}=1$ overestimates the corresponding interval length of $\mathcal{M}_{0^{n}}$ in the partition of the phase space $\mathcal{M}$ by an increasing amount thus leading to wrong results when calculating escape. By leaving out the $\overline{0}$ cycle (and thus also the $\mathcal{M}_{0^{n}}$ contribution), we are guaranteed to get at least the right asymptotical behavior.

Note also, that if we are working with the spectral determinant (15.3), given in product form as

$$
\operatorname{det}(1-z \mathcal{L})=\prod_{p} \prod_{m=0}^{\infty}\left(1-\frac{z^{n_{p}}}{\left|\Lambda_{p}\right| \Lambda_{p}^{m}}\right)
$$

for intermittent maps the marginal stable cycle has to be excluded. It introduces an (unphysical) essential singularity at $z=1$ due the presence of a factor $(1-z)^{\infty}$ stemming from the $\overline{0}$ cycle.

### 21.3 Intermittency for cyclists

Admittedly, the toy map is what is says - a toy model. The piece wise linearity of the map led to exact cancellations of the curvature contributions leaving only the fundamental terms. There are still infinitely many orbits included in the fundamental term, but the cycle weights were chosen in such a way that the zeta function could be written in closed form. For a smooth intermittent map this all will not be the case in general; still, we will argue that we have already seen almost all the fundamentally new features due to intermittency. What remains are technicalities - not necessarily easy to handle, but nothing very surprise any more.

In the following we will sketch, how to make cycle expansion techniques work for general 1-dimensional maps with a single isolated marginal fixed point. To keep the notation simple, we will consider two-branch maps with a complete binary symbolic dynamics as for example the Farey map, figure 21.3, or the repeller depicted in figure 21.2. We again assume that the behavior near the fixed point is given by (21.1). This implies that the stability of a family of periodic orbits approaching the marginally stable orbit, as for example the family $10^{n}$, will increase only algebraically, that is we find again for large $n$

$$
\frac{1}{\Lambda_{10^{n}}} \sim \frac{1}{n^{1+1 / s}}
$$

where $s$ denotes the intermittency exponent.
When considering zeta functions or trace formulas, we again have to take out the marginal orbit $\overline{0}$; periodic orbit contributions of the form $t_{0^{n} 1}$ are now unbalanced and we arrive at a cycle expansion in terms of infinitely many fundamental terms as for our toy map. This corresponds to moving from our binary symbolic dynamics to an infinite symbolic dynamics by making the identification

$$
10^{n-1} \rightarrow n ; \quad 10^{n-1} 10^{m-1} \rightarrow n m ; \quad 10^{n-1} 10^{m-1} 10^{k-1} \rightarrow n m k ; \ldots
$$

| $\infty$ - alphabet | binary alphabet |  |  |  |  |  |
| ---: | ---: | ---: | ---: | ---: | ---: | ---: |
|  |  | $n=1$ | $n=2$ | $n=3$ | $n=4$ | $n=5$ |
| 1 -cycles | $n$ | 1 | 10 | 100 | 1000 | 10000 |
| 2 -cycles | $m n$ |  |  |  |  |  |
|  | $1 n$ | 11 | 110 | 1100 | 11000 | 110000 |
|  | $2 n$ | 101 | 0101 | 10100 | 101000 | 1010000 |
|  | $3 n$ | 1001 | 10010 | 100100 | 1001000 | 10010000 |
| 3 -cycles | $4 m n$ | 10001 | 100010 | 1000100 | 10001000 | 100010000 |
|  | $11 n$ | 111 | 1110 | 11100 | 111000 | 1110000 |
|  | $12 n$ | 1101 | 11010 | 110100 | 1101000 | 11010000 |
|  | $13 n$ | 11001 | 110010 | 1100100 | 11001000 | 110010000 |
|  | $21 n$ | 1011 | 10110 | 101100 | 1011000 | 10110000 |
|  | $22 n$ | 10101 | 101010 | 1010100 | 10101000 | 101010000 |
|  | $23 n$ | 101001 | 1010010 | 10100100 | 101001000 | 1010010000 |
|  | $31 n$ | 10011 | 100110 | 1001100 | 10011000 | 100110000 |
|  | $32 n$ | 100101 | 1001010 | 10010100 | 100101000 | 1001010000 |
| $33 n$ | 1001001 | 10010010 | 100100100 | 1001001000 | 10010010000 |  |

Table 21.1: Infinite alphabet versus the original binary alphabet for the shortest periodic orbit families. Repetitions of prime cycles $\left(11=1^{2}, 0101=01^{2}, \ldots\right)$ and their cyclic repeats $(110=101,1110=1101, \ldots)$ are accounted for by cancellations and combination factors in the cycle expansion (21.31).
see also table 21.1. The topological length of the orbit is thus no longer determined by the iterations of our two-branch map, but by the number of times the cycle goes from the right to the left branch. Equivalently, one may define a new map, for which all the iterations on the left branch are done in one step. Such a map is called an induced map and the topological length of orbits in the infinite alphabet corresponds to the iterations of this induced map.

For generic intermittent maps, curvature contributions in the cycle expanded zeta function will not vanish exactly. The most natural way to organize the cycle expansion is to collect orbits and pseudo orbits of the same topological length with respect to the infinite alphabet. Denoting cycle weights in the new alphabet as $t_{n m \ldots}=t_{10^{n-1} 10^{m-1} \ldots}$, one obtains

$$
\begin{align*}
\zeta^{-1} & =\prod_{p \neq 0}\left(1-t_{p}\right)=1-\sum_{n=1}^{\infty} c e  \tag{21.31}\\
& =1-\sum_{n=1}^{\infty} t_{n}-\sum_{m=1}^{\infty} \sum_{n=1}^{\infty} \frac{1}{2}\left(t_{m n}-t_{m} t_{n}\right) \\
& -\sum_{k=1}^{\infty} \sum_{m=1}^{\infty} \sum_{n=1}^{\infty}\left(\frac{1}{3} t_{k m n}-\frac{1}{2} t_{k m} t_{n}+\frac{1}{6} t_{k} t_{m} t_{n}\right)-\sum_{l=1}^{\infty} \sum_{k=1}^{\infty} \sum_{m=1}^{\infty} \sum_{n=1}^{\infty} \ldots
\end{align*}
$$

The first sum is the fundamental term, which we have already seen in the toy model, (21.10). The curvature terms $c_{n}$ in the expansion are now $e$-fold infinite sums where the prefactors take care of double counting of prime periodic orbits.

Let us consider the fundamental term first. For generic intermittent maps, we can not expect to obtain an analytic expression for the infinite sum of the form

$$
\begin{equation*}
f(z)=\sum_{n=0}^{\infty} h_{n} z^{n} . \tag{21.32}
\end{equation*}
$$

with algebraically decreasing coefficients

$$
h_{n} \sim \frac{1}{n^{\alpha}} \text { with } \quad \alpha>0
$$

To evaluate the sum, we face the same problem as for our toy map: the power series diverges for $z>1$, that is, exactly in the 'interesting' region where poles, zeros or branch cuts of the zeta function are to be expected. By carefully subtracting the asymptotic behavior with the help of (21.11) or (21.12), one can in general construct an analytic continuation of $f(z)$ around $z=1$ of the form

$$
\begin{array}{ll}
f(z) \sim A(z)+(1-z)^{\alpha-1} B(z) & \alpha \notin \mathbb{N}  \tag{21.33}\\
f(z) \sim A(z)+(1-z)^{\alpha-1} \ln (1-z) & \alpha \in \mathbb{N}
\end{array}
$$

where $A(z)$ and $B(z)$ are functions analytic in a disc around $z=1$. We thus again find that the zeta function (21.31) has a branch cut along the real axis $\operatorname{Re} z \geq 1$. From here on we can switch to auto-pilot and derive algebraic escape, decay of correlation and all the rest. We find in particular that the asymptotic behavior derived in (21.24) and (21.25) is a general result, that is, the survival probability is given asymptotically by

$$
\begin{equation*}
\Gamma_{n} \sim C \frac{1}{n^{1 / s}} \tag{21.34}
\end{equation*}
$$

for all 1-dimensional maps of the form (21.1). We have to work a bit harder if we want more detailed information like the prefactor $C$, exponential precursors given by zeros or poles of the dynamical zeta function or higher order corrections. This information is buried in the functions $A(z)$ and $B(z)$ or more generally in the analytically continued zeta function. To get this analytic continuation, one may follow either of the two different strategies which we will sketch next.

### 21.3.1 Resummation

One way to get information about the zeta function near the branch cut is to derive the leading coefficients in the Taylor series of the functions
$A(z)$ and $B(z)$ in (21.33) at $z=1$. This can be done in principle, if the coefficients $h_{n}$ in sums like (21.32) are known (as for our toy model). One then considers a resummation of the form

$$
\begin{equation*}
\sum_{j=0}^{\infty} h_{j} z^{j}=\sum_{j=0}^{\infty} a_{j}(1-z)^{j}+(1-z)^{\alpha-1} \sum_{j=0}^{\infty} b_{j}(1-z)^{j}, \tag{21.35}
\end{equation*}
$$

and the coefficients $a_{j}$ and $b_{j}$ are obtained in terms of the $h_{j}$ 's by expanding $(1-z)^{j}$ and $(1-z)^{j+\alpha-1}$ on the right hand side around $z=0$ using (21.11) and equating the coefficients.

In practical calculations one often has only a finite number of coefficients $h_{j}, 0 \leq j \leq N$, which may have been obtained by finding periodic orbits and their stabilities numerically. One can still design a resummation scheme for the computation of the coefficients $a_{j}$ and $b_{j}$ in (21.35). We replace the infinite sums in (21.35) by finite sums of increasing degrees $n_{a}$ and $n_{b}$, and require that

$$
\begin{equation*}
\sum_{i=0}^{n_{a}} a_{i}(1-z)^{i}+(1-z)^{\alpha-1} \sum_{i=0}^{n_{b}} b_{i}(1-z)^{i}=\sum_{i=0}^{N} h_{i} z^{i}+O\left(z^{N+1}\right) . \tag{21.36}
\end{equation*}
$$

One proceeds again by expanding the right hand side around $z=0$, skipping all powers $z^{N+1}$ and higher, and then equating coefficients. It is natural to require that $\left|n_{b}+\alpha-1-n_{a}\right|<1$, so that the maximal powers of the two sums in (21.36) are adjacent. If one chooses $n_{a}+n_{b}+2=N+1$, then, for each cutoff length $N$, the integers $n_{a}$ and $n_{b}$ are uniquely determined from a linear system of equations. The price we pay is that the so obtained coefficients depend on the cutoff $N$. One can now study convergence of the coefficients $a_{j}$, and $b_{j}$, with respect to increasing values of $N$, or various quantities derived from $a_{j}$ and $b_{j}$. Note that the leading coefficients $a_{0}$ and $b_{0}$ determine the prefactor $C$ in (21.34), cf. (21.23). The resummed expression can also be used to compute zeros, inside or outside the radius of convergence of the cycle expansion $\sum h_{j} z^{j}$.

The scheme outlined in this section tacitly assumes that a representation of form (21.33) holds in a disc of radius 1 around $z=1$. Convergence is improved further if additional information about the asymptotics of sums like (21.32) is used to improve the ansatz (21.35).

### 21.3.2 Analytical continuation by integral transformations

We will now introduce a method which provides an analytic continuation of sums of the form (21.32) without explicitly relying on an ansatz (21.35). The main idea is to rewrite the sum (21.32) as a sum over integrals with the help of the Poisson summation formula and find an analytic continuation of each integral by contour deformation. In order to do so, we need to know the $n$ dependence of the coefficients $h_{n} \equiv h(n)$ explicitly for all $n$. If the
coefficients are not known analytically, one may proceed by approximating the large $n$ behavior in the form

$$
h(n)=n^{-\alpha}\left(C_{1}+C_{2} n^{-1}+\ldots\right), \quad n \neq 0
$$

and determine the constants $C_{i}$ numerically from periodic orbit data. By using the Poisson resummation identity

$$
\begin{equation*}
\sum_{n=-\infty}^{\infty} \delta(x-n)=\sum_{m=-\infty}^{\infty} \exp (2 \pi i m x) \tag{21.37}
\end{equation*}
$$

we may write the sum as (21.32)

$$
\begin{equation*}
f(z)=\frac{1}{2} h(0)+\sum_{m=-\infty}^{\infty} \int_{0}^{\infty} d x e^{2 \pi i m x} h(x) z^{x} . \tag{21.38}
\end{equation*}
$$

The continuous variable $x$ corresponds to the discrete summation index $n$ and it is convenient to write $z=r \exp (i \sigma)$ from now on. The integrals are still not convergent for $r>0$, but an analytical continuation can be found by considering the contour integral, where the contour goes out along the real axis, makes a quarter circle to either the positive or negative imaginary axis and goes back to zero. By letting the radius of the circle go to infinity, we essentially rotate the line of integration from the real onto the imaginary axis. For the $m=0$ term in (21.38), we transform $x \rightarrow i x$ and the integral takes on the form

$$
\int_{0}^{\infty} d x h(x) r^{x} e^{i x \sigma}=i \int_{0}^{\infty} d x h(i x) r^{i x} e^{-x \sigma} .
$$

The integrand is now exponentially decreasing for all $r>0$ and $\sigma \neq 0$ or $2 \pi$. The last condition reminds us again of the existence of a branch cut at $\operatorname{Re} z \geq 1$. By the same technique, we find the analytic continuation for all the other integrals in (21.38). The real axis is then rotated according to $x \rightarrow \operatorname{sign}(m) i x$ where $\operatorname{sign}(m)$ refers to the sign of $m$.

$$
\int_{0}^{\infty} d x e^{ \pm 2 \pi i|m| x} h(x) r^{x} e^{i x \sigma}= \pm \mathrm{i} \int_{0}^{\infty} d x h( \pm i x) r^{ \pm i x} e^{-x(2 \pi|m| \pm \sigma)}
$$

Changing summation and integration, we can carry out the sum over $|m|$ explicitly and one finally obtains the compact expression

$$
\begin{align*}
f(z) & =\frac{1}{2} h(0)+i \int_{0}^{\infty} d x h(i x) r^{i x} e^{-x \sigma}  \tag{21.39}\\
& +i \int_{0}^{\infty} d x \frac{e^{-2 \pi x}}{1-e^{-2 \pi x}}\left[h(i x) r^{i x} e^{-x \sigma}-h(-i x) r^{-i x} e^{x \sigma}\right]
\end{align*}
$$

The transformation from the original sum to the two integrals in (21.39) is exact for $r \leq 1$, and provides an analytic continuation for $r>0$. The expression (21.39) is especially useful for an efficient numerical calculations of a dynamical zeta function for $|z|>1$, which is essential when searching for its zeros and poles.

### 21.3.3 Curvature contributions

So far, we have discussed only the fundamental term $\sum_{n=1}^{\infty} t_{n}$ in (21.31), and showed how to deal with such power series with algebraically decreasing coefficients. The fundamental term determines the main structure of the zeta function in terms of the leading order branch cut. Corrections to both the zeros and poles of the dynamical zeta function as well as the leading and subleading order terms in expansions like (21.33) are contained in the curvature terms in (21.31). The first curvature correction is the 2 -cycle sum

$$
\sum_{m=1}^{\infty} \sum_{n=1}^{\infty} \frac{1}{2}\left(t_{m n}-t_{m} t_{n}\right),
$$

with algebraically decaying coefficients which again diverge for $|z|>1$. The analytically continued curvature terms have as usual branch cuts along the positive real $z$ axis. Our ability to calculate the higher order curvature terms depends on how much we know about the cycle weights $t_{m n}$. The form of the cycle stability (21.5) suggests that $t_{m n}$ decrease asymptotically as

$$
\begin{equation*}
t_{m n} \sim \frac{1}{(n m)^{1+1 / s}} \tag{21.40}
\end{equation*}
$$

for 2-cycles, and in general for $n$-cycles as

$$
t_{m_{1} m_{2} \ldots m_{n}} \sim \frac{1}{\left(m_{1} m_{2} \ldots m_{n}\right)^{1+1 / s}}
$$

If we happen to know the cycle weights $t_{m_{1} m_{2} \ldots m_{n}}$ analytically, we may proceed as in sect. 21.3.2, transform the multiple sums into multiple integrals and rotate the integration contours.

We have reached the edge of what has been accomplished so far in computing and what is worth the dynamical zeta functions from periodic orbit data. In the next section, we describe a probabilistic method applicable to intermittent maps which does not rely on periodic orbits.

### 21.4 BER zeta functions

$\circ$So far we have focused on 1-d models as the simplest setting in which to investigate dynamical implications of marginal fixed points. We now take an altogether different track and describe how probabilistic methods may be employed in order to write down approximate dynamical zeta functions for intermittent systems.

We will discuss the method in a very general setting, for a flow in arbitrary dimension. The key idea is to introduce a surface of section $\mathcal{P}$ such that all trajectories traversing this section will have spent some time both near the marginal stable fixed point and in the chaotic phase. An important quantity in what follows is (3.4), the first return time $\tau(x)$, or the time of flight of a trajectory starting in $x$ to the next return to the surface of section $\mathcal{P}$. The period of a periodic orbit $p$ intersecting the $\mathcal{P}$ section $n_{p}$ times is

$$
T_{p}=\sum_{k=0}^{n_{p}-1} \tau\left(f^{k}\left(x_{p}\right)\right)
$$

where $f(x)$ is the Poincaré map, and $x_{p} \in \mathcal{P}$ is a cycle point. The dynamical zeta function (15.15)

$$
\begin{equation*}
1 / \zeta(z, s, \beta)=\prod_{p}\left(1-\frac{z^{n_{p}} e^{\beta A_{p}-s T_{p}}}{\left|\Lambda_{p}\right|}\right), \quad A_{p}=\sum_{k=0}^{n_{p}-1} a\left(f^{k}\left(x_{p}\right)\right), \tag{21.41}
\end{equation*}
$$

associated with the observable $a(x)$ captures the dynamics of both the flow and the Poincaré map. The dynamical zeta function for the flow is obtained as $1 / \zeta(s, \beta)=1 / \zeta(1, s, \beta)$, and the dynamical zeta function for the discrete time Poincaré map is $1 / \zeta(z, \beta)=1 / \zeta(z, 0, \beta)$.

Our basic assumption will be probabilistic. We assume that the chaotic interludes render the consecutive return (or recurrence) times $T\left(x_{i}\right)$, $T\left(x_{i+1}\right)$ and observables $a\left(x_{i}\right), a\left(x_{i+1}\right)$ effectively uncorrelated. Consider the quantity $e^{\beta A\left(x_{0}, n\right)-s T\left(x_{0}, n\right)}$ averaged over the surface of section $\mathcal{P}$. With the above probabilistic assumption the large $n$ behavior is

$$
\left\langle e^{\beta A\left(x_{0}, n\right)-s T\left(x_{0}, n\right)}\right\rangle_{\mathcal{P}} \sim\left(\int_{\mathcal{P}} e^{\beta a(x)-s \tau} \rho(x) d x\right)^{n}
$$

where $\rho(x)$ is the invariant density of the Poincaré map. This type of behavior is equivalent to there being only one zero $z_{0}(s, \beta)=\int e^{\beta a(x)-s \tau(x)} \rho(x) d x$ of $1 / \zeta(z, s, \beta)$ in the $z-\beta$ plane. In the language of Ruelle-Pollicott resonances this means that there is an infinite gap to the first resonance. This in turn implies that $1 / \zeta(z, s, \beta)$ may be written as

$$
\begin{equation*}
1 / \zeta(z, s, \beta)=z-\int_{\mathcal{P}} e^{\beta a(x)-s \tau(x)} \rho(x) d x \tag{21.42}
\end{equation*}
$$

where we have neglected a possible analytic and non-zero prefactor. The dynamical zeta function of the flow is now

$$
\begin{equation*}
1 / \zeta(s, \beta)=1 / \zeta(1, s, \beta)=1-\int_{\mathcal{P}} e^{\beta a(x)} \rho(x) e^{-s \tau(x)} d x \tag{21.43}
\end{equation*}
$$

Normally, the best one can hope for is a finite gap to the leading resonance of the Poincaré map. with the above dynamical zeta function only approximatively valid. As it is derived from an approximation due to Baladi, Eckmann, and Ruelle, we shall refer to it as the BER zeta function $1 / \zeta_{\mathrm{BER}}(s, \beta)$ in what follows.

A central role is played by the probability distribution of return times

$$
\begin{equation*}
\psi(\tau)=\int_{\mathcal{P}} \delta(\tau-\tau(x)) \rho(x) d x \tag{21.44}
\end{equation*}
$$

The BER zeta function at $\beta=0$ is then given in terms of the Laplace transform of this distribution

$$
1 / \zeta_{\mathrm{BER}}(s)=1-\int_{0}^{\infty} \psi(\tau) e^{-s \tau} d \tau .
$$

Example 21.1 Return times for the Bernoulli map. For the Bernoulli shift map (16.6)

$$
x \mapsto f(x)=2 x \bmod 1,
$$

one easily derives the distribution of return times

$$
\psi_{n}=\frac{1}{2^{n}} \quad n \geq 1
$$

The BER zeta function becomes (by the discrete Laplace transform (14.8))

$$
\begin{align*}
1 / \zeta_{B E R}(z) & =1-\sum_{n=1}^{\infty} \psi_{n} z^{n}=1-\sum_{n=1}^{\infty} \frac{z^{n}}{2^{n}} \\
& =\frac{1-z}{1-z / 2}=\zeta^{-1}(z) /\left(1-z / \Lambda_{0}\right) \tag{21.45}
\end{align*}
$$

Thanks to the uniformity of the piecewise linear map measure (10.19) the "approximate" zeta function is in this case the exact dynamical zeta function, with the cycle point $\overline{0}$ pruned.

Example 21.2 Return times for the model of sect. 21.2.1. For the toy model of sect. 21.2.1 one gets $\psi_{1}=\left|\mathcal{M}_{1}\right|$, and $\psi_{n}=\left|\mathcal{M}_{n}\right|(1-b) /(1-a)$, for $n \geq 2$, leading to a $B E R$ zeta function

$$
1 / \zeta_{B E R}(z)=1-z\left|\mathcal{M}_{1}\right|-\sum_{n=2}^{\infty}\left|\mathcal{M}_{n}\right| z^{n}
$$

which again coincides with the exact result, (21.10).

It may seem surprising that the BER approximation produces exact results in the two examples above. The reason for this peculiarity is that both these systems are piecewise linear and have complete Markov partitions. As long as the map is piecewise linear and complete, and the probabilistic approximation is exactly fulfilled, the cycle expansion curvature terms vanish. The BER zeta function and the fundamental part of a cycle expansion discussed in sect. 18.1.1 are indeed intricately related, but not identical in general. In particular, note that the BER zeta function obeys the flow conservation sum rule (19.11) by construction, whereas the fundamental part of a cycle expansion as a rule does not.

## Commentary

Remark 21.1 What about the evolution operator formalism? The main virtue of evolution operators was their semigroup property (10.25). This was natural for hyperbolic systems where instabilities grow exponentially, and evolution operators capture this behavior due to their multiplicative nature. Whether the evolution operator formalism is a good way to capture the slow, power law instabilities of intermittent dynamics is less clear. The approach taken here leads us to a formulation in terms of dynamical zeta functions rather than spectral determinants, circumventing evolution operators altogether. It is not known if the spectral determinants formulation would yield any benefits when applied to intermittent chaos. Some results on spectral determinants and intermittency can be found in [21.2]. A useful mathematical technique to deal with isolated marginally stable fixed point is that of inducing, that is, replacing the intermittent map by a completely hyperbolic map with infinite alphabet and redefining the discrete time; we have used this method implicitly by changing from a finite to an infinite alphabet. We refer to refs. [21.3, 21.19] for detailed discussions of this technique, as well as applications to 1-dimensional maps.

Remark 21.2 Intermittency. Intermittency was discovered by Manneville and Pomeau [21.1] in their study of the Lorentz system. They demonstrated that in neighborhood of parameter value $r_{c}=166.07$ the mean duration of the periodic motion scales as $\left(r-r_{c}\right)^{1 / 2}$. In ref. [21.5] they explained this phenomenon in terms of a 1-dimensional map (such as (21.1)) near tangent bifurcation, and classified possible types of intermittency.

Piecewise linear models like the one considered here have been studied by Gaspard and Wang [21.6]. The escape problem has here been treated following
ref. [21.7], resummations following ref. [21.8]. The proof of the bound (21.27) can be found in P. Dahlqvist's notes on ChaosBook.org/PDahlqvistEscape.ps.gz.

Farey map (18.27) has been studied widely in the context of intermittent dynamics, for example in refs. [21.16, 21.17, 18.3, 21.18, L.23, 18.14, 21.2]. The Fredholm determinant and the dynamical zeta functions for the Farey map (18.27) and the related Gauss shift map (24.38) have been studied by Mayer [21.16]. He relates the continued fraction transformation to the Riemann zeta function, and constructs a Hilbert space on which the evolution operator is self-adjoint, and its eigenvalues are exponentially spaced, just as for the dynamical zeta functions [21.23] for "Axiom A" hyperbolic systems.

Remark 21.3 Tauberian theorems. In this chapter we used Tauberian theorems for power series and Laplace transforms: Feller's monograph [21.9] is a highly recommended introduction to these methods.

Remark 21.4 Probabilistic methods, BER zeta functions. Probabilistic description of intermittent chaos was introduced by Geisal and Thomae [21.10]. The BER approximation studied here is inspired by Baladi, Eckmann and Ruelle [21.14], with further developments in refs. [21.13, 21.15].

## Résumé

The presence of marginally stable fixed points and cycles changes the analytic structure of dynamical zeta functions and the rules for constructing cycle expansions. The marginal orbits have to be omitted, and the cycle expansions now need to include families of infinitely many longer and longer unstable orbits which accumulate toward the marginally stable cycles. Correlations for such non-hyperbolic systems may decay algebraically with the decay rates controlled by the branch cuts of dynamical zeta functions. Compared to pure hyperbolic systems, the physical consequences are drastic: exponential decays are replaced by slow power-law decays, and transport properties, such as the diffusion may become anomalous.

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## Exercises

Exercise 21.1 Integral representation of Jonquière functions. Check the integral representation

$$
\begin{equation*}
J(z, \alpha)=\frac{z}{\Gamma(\alpha)} \int_{0}^{\infty} d \xi \frac{\xi^{\alpha-1}}{e^{\xi}-z} \quad \text { for } \quad \alpha>0 . \tag{21.46}
\end{equation*}
$$

Note how the denominator is connected to Bose-Einstein distribution. Compute $J(x+$ $i \epsilon)-J(x-i \epsilon)$ for a real $x>1$.

Exercise 21.2 Power law correction to a power law. Expand (21.20) further and derive the leading power law correction to (21.23).

Exercise 21.3 Power-law fall off. In cycle expansions the stabilities of orbits do not always behave in a geometric fashion. Consider the map $f$


This map behaves as $f \rightarrow x$ as $x \rightarrow 0$. Define a symbolic dynamics for this map by assigning 0 to the points that land on the interval $[0,1 / 2)$ and 1 to the points that land on $(1 / 2,1]$. Show that the stability of orbits that spend a long time on the 0 side goes as $n^{2}$. In particular, show that

$$
\underbrace{\Lambda_{00 \ldots 0}}_{n} \sim n^{2}
$$

Exercise 21.4 Power law fall-off of stability eigenvalues in the stadium billiard**. From the cycle expansions point of view, the most important consequence of the shear in $\mathbf{J}^{n}$ for long sequences of rotation bounces $n_{k}$ in (6.13) is that the $\Lambda_{n}$ grows only as a power law in number of bounces:

$$
\begin{equation*}
\Lambda_{n} \propto n_{k}^{2} \tag{21.47}
\end{equation*}
$$

Check.
Exercise 21.5 Probabilistic zeta function for maps. Derive the probabilistic zeta function for a map with recurrence distribution $\psi_{n}$.

Exercise 21.6 Accelerated diffusion. Consider a map $h$, such that $\hat{h}=\hat{f}$, but now running branches are turner into standing branches and vice versa, so that $1,2,3,4$ are standing while 0 leads to both positive and negative jumps. Build the corresponding dynamical zeta function and show that

$$
\sigma^{2}(t) \sim \begin{cases}t & \text { for } \alpha>2 \\ t \ln t & \text { for } \alpha=2 \\ t^{3-\alpha} & \text { for } \alpha \in(1,2) \\ t^{2} / \ln t & \text { for } \alpha=1 \\ t^{2} & \text { for } \alpha \in(0,1)\end{cases}
$$

Exercise 21.7 Anomalous diffusion (hyperbolic maps). Anomalous diffusive properties are associated to deviations from linearity of the variance of the phase variable we are looking at: this means the the diffusion constant (10.13) either vanishes or diverges. We briefly illustrate in this exercise how the local local properties of a map are crucial to account for anomalous behavior even for hyperbolic systems.

Consider a class of piecewise linear maps, relevant to the problem of the onset of diffusion, defined by

$$
f_{\epsilon}(x)= \begin{cases}\Lambda x & \text { for } x \in\left[0, x_{1}^{+}\right]  \tag{21.48}\\ a-\Lambda_{\epsilon, \gamma}\left|x-x^{+}\right| & \text {for } x \in\left[x_{1}^{+}, x_{2}^{+}\right] \\ 1-\Lambda^{\prime}\left(x-x_{2}^{+}\right) & \text {for } x \in\left[x_{2}^{+}, x_{1}^{-}\right] \\ 1-a+\Lambda_{\epsilon, \gamma}\left|x-x^{-}\right| & \text {for } x \in\left[x_{1}^{-}, x_{2}^{-}\right] \\ 1+\Lambda(x-1) & \text { for } x \in\left[x_{2}^{-}, 1\right]\end{cases}
$$

where $\Lambda=\left(1 / 3-\epsilon^{1 / \gamma}\right)^{-1}, \Lambda^{\prime}=\left(1 / 3-2 \epsilon^{1 / \gamma}\right), \Lambda_{\epsilon, \gamma}=\epsilon^{1-1 / \gamma}, a=1+\epsilon, x^{+}=1 / 3$, $x_{1}^{+}=x^{+}-\epsilon^{1 / \gamma}, x_{2}^{+}=x^{+}+\epsilon^{1 / \gamma}$, and the usual symmetry properties (23.11) are satisfied.

Thus this class of maps is characterized by two escaping windows (through which the diffusion process may take place) of size $2 \epsilon^{1 / \gamma}$ : the exponent $\gamma$ mimicks the order of the maximum for a continuous map, while piecewise linearity, besides making curvatures vanish and leading to finite cycle expansions, prevents the appearance of stable cycles. The symbolic dynamics is easily described once we consider a sequence of parameter values $\left\{\epsilon_{m}\right\}$, where $\epsilon_{m}=\Lambda^{-(m+1)}$ : we then partition the unit interval though the sequence of points $0, x_{1}^{+}, x^{+}, x_{2}^{+}, x_{1}^{-}, x^{-}, x_{2}^{-}, 1$ and label the corresponding sub-intervals $1, s_{a}, s_{b}, 2, d_{b}, d_{a}, 3$ : symbolic dynamics is described by an unrestricted grammar over the following set of symbols

$$
\left\{1,2,3, s_{\#} \cdot 1^{i}, d_{\#} \cdot 3^{k}\right\} \quad \#=a, b \quad i, k=m, m+1, m+2, \ldots
$$

This leads to the following dynamical zeta function:

$$
\zeta_{0}^{-1}(z, \alpha)=1-\frac{2 z}{\Lambda}-\frac{z}{\Lambda^{\prime}}-4 \cosh (\alpha) \epsilon_{m}^{1 / \gamma-1} \frac{z^{m+1}}{\Lambda^{m}}\left(1-\frac{z}{\Lambda}\right)^{-1}
$$

from which, by (23.8) we get

$$
\begin{equation*}
D=\frac{2 \epsilon_{m}^{1 / \gamma-1} \Lambda^{-m}(1-1 / \Lambda)^{-1}}{1-\frac{2}{\Lambda}-\frac{1}{\Lambda^{\prime}}-4 \epsilon_{m}^{1 / \gamma-1}\left(\frac{m+1}{\Lambda^{m}(1-1 / \Lambda)}+\frac{1}{\Lambda^{m+1}(1-1 / \Lambda)^{2}}\right)} \tag{21.49}
\end{equation*}
$$

The main interest in this expression is that it allows exploring how $D$ vanishes in the $\epsilon \mapsto 0(m \mapsto \infty)$ limit: as a matter of fact, from (21.49) we get the asymptotic behavior $D \sim \epsilon^{1 / \gamma}$, which shows how the onset of diffusion is governed by the order of the map at its maximum.

Remark 21.5 Onset of diffusion for continuous maps. The zoology of behavior for continuous maps at the onset of diffusion is described in refs. [23.11, 23.12, 21.24]: our treatment for piecewise linear maps was introduced in ref. [21.25].

## Chapter 22

## Discrete symmetries

Utility of discrete symmetries in reducing spectrum calculations is familiar from quantum mechanics. Here we show that the classical spectral determinants factor in essentially the same way as in quantum mechanics. In the process we also learn how to simplify the classical dynamics. The main result of this chapter can be stated as follows:

If the dynamics possesses a discrete symmetry, the contribution of a cycle $p$ of multiplicity $m_{p}$ to a dynamical zeta function factorizes into a product over the $d_{\alpha}$-dimensional irreducible representations $D_{\alpha}$ of the symmetry group,

$$
\left(1-t_{p}\right)^{m_{p}}=\prod_{\alpha} \operatorname{det}\left(1-D_{\alpha}\left(h_{\tilde{p}}\right) t_{\tilde{p}}\right)^{d_{\alpha}}, \quad t_{p}=t_{\tilde{p}}^{g / m_{p}},
$$

where $t_{\tilde{p}}$ is the cycle weight evaluated on the fundamental domain, $g$ is the dimension of the group, $h_{\tilde{p}}$ is the group element relating the fundamental domain cycle $\tilde{p}$ to a segment of the full space cycle $p$, and $m_{p}$ is the multiplicity of the $p$ cycle. As the dynamical zeta functions have particularly simple cycle expansions, a simple geometrical shadowing interpretation of their convergence, and as they suffice for determination of leading eigenvalues, we shall concentrate in this chapter on their factorizations; the full spectral determinants can be factorized by the same techniques. To emphasize the group theoretic structure of zeta functions, we shall combine all the non-group-theory dependence of a $p$-cycle into a cycle weight $t_{p}$.

This chapter is meant to serve as a detailed guide to computation of dynamical zeta functions and spectral determinants for systems with discrete symmetries. Familiarity with basic group-theoretic notions is assumed, with the definitions relegated to appendix I.1. We develop here the cycle expansions for factorized determinants, and exemplify them by working out a series of cases of physical interest: $C_{2}, C_{3 v}$ symmetries in this chapter, and $C_{2 v}, C_{4 v}$ symmetries in appendix I below.

### 22.1 Preview

Dynamical systems often come equipped with discrete symmetries, such as the reflection and the rotation symmetries of various potentials. Such symmetries simplify and improve the cycle expansions in a rather beautiful way; they can be exploited to relate classes of periodic orbits and reduce dynamics to a fundamental domain. Furthermore, in classical dynamics, just as in quantum mechanics, the symmetrized subspaces can be probed by linear operators of different symmetries. If a linear operator commutes with the symmetry, it can be block-diagonalized, and, as we shall now show, the associated spectral determinants and dynamical zeta functions factorize.

Invariance of a system under symmetries means that the symmetry image of a cycle is again a cycle, with the same weight. The new orbit may be topologically distinct (in which case it contributes to the multiplicity of the cycle) or it may be the same cycle, shifted in time. A cycle is symmetric if some symmetry operations act on it like a shift in time, advancing the starting point to the starting point of a symmetry related segment. A symmetric cycle can thus be subdivided into a sequence of repeats of an irreducible segment. The period or any average evaluated along the full orbit is given by the sum over the segments, whereas the stability is given by the product of the stability matrices of the individual segments.

Cycle degeneracies induced by the symmetry are removed by desymmetrization, reduction of the full dynamics to the dynamics on a fundamental domain. The phase space can be completely tiled by a fundamental domain and its symmetry images. The irreducible segments of cycles in the full space, folded back into the fundamental domain, are closed orbits in the reduced space.

### 22.1.1 3-disk game of pinball

We have already exploited a discrete symmetry in our introduction to the 3disk game of pinball, sect. 1.3. As the three disks are equidistantly spaced, our game of pinball has a sixfold symmetry. The symmetry group of relabeling the 3 disks is the permutation group $S_{3}$; however, it is better to think of this group geometrically, as $\mathrm{C}_{3 v}$, the group of rotations by $\pm 2 \pi / 3$ and reflections across the three symmetry axes. Applying an element (identity, rotation by $\pm 2 \pi / 3$, or one of the three possible reflections) of this symmetry group to any trajectory yields another trajectory. For instance, the cycles $\overline{12}, \overline{23}$, and $\overline{13}$, are related to each other by rotation by $\pm 2 \pi / 3$, or, equivalently, by a relabeling of the disks.

An irreducible segment corresponds to a periodic orbit in the fundamental domain, a one-sixth slice of the full 3 -disk system, with the symmetry axes acting as reflecting mirrors, see figure 11.6. A set of orbits related in the full space by discrete symmetries maps onto a single fundamental domain orbit. The reduction to the fundamental domain desymmetrizes the dynamics and removes all global discrete symmetry induced degeneracies:
rotationally symmetric global orbits (such as the 3 -cycles $\overline{123}$ and $\overline{132}$ ) have degeneracy 2 , reflection symmetric ones (such as the 2 -cycles $\overline{12}, \overline{13}$ and $\overline{23}$ ) have degeneracy 3 , and global orbits with no symmetry are 6 -fold degenerate. Table 11.2 lists some of the shortest binary symbols strings, together with the corresponding full 3 -disk symbol sequences and orbit symmetries. Some examples of such orbits are shown in figure 22.3.

We shall return to the 3-disk game of pinball desymmetrization in sects. 22.2.2 and 22.6, but first we develop a feeling for discrete symmetries by working out a simple 1- $d$ example.

### 22.1.2 Reflection symmetric 1-d maps

Consider $f$, a map on the interval with reflection symmetry $f(-x)=-f(x)$. A simple example is the piecewise-linear sawtooth map of figure 22.1. Denote the reflection operation by $\mathbf{C} x=-x$. The symmetry of the map implies that if $\left\{x_{n}\right\}$ is a trajectory, than also $\left\{\mathbf{C} x_{n}\right\}$ is a trajectory because $\mathbf{C} x_{n+1}=\mathbf{C} f\left(x_{n}\right)=f\left(\mathbf{C} x_{n}\right)$. The dynamics can be restricted to a fundamental domain, in this case to one half of the original interval; every time a trajectory leaves this interval, it can be mapped back using C. Furthermore, the evolution operator commutes with $\mathbf{C}, \mathcal{L}(y, x)=\mathcal{L}(\mathbf{C} y, \mathbf{C} x)$. $\mathbf{C}$ satisfies $\mathbf{C}^{2}=\mathbf{e}$ and can be used to decompose the phase space into mutually orthogonal symmetric and antisymmetric subspaces by means of projection operators

$$
\begin{align*}
P_{A_{1}} & =\frac{1}{2}(\mathbf{e}+\mathbf{C}), \quad P_{A_{2}}=\frac{1}{2}(\mathbf{e}-\mathbf{C}) \\
\mathcal{L}_{A_{1}}(y, x) & =P_{A_{1}} \mathcal{L}(y, x)=\frac{1}{2}(\mathcal{L}(y, x)+\mathcal{L}(-y, x)) \\
\mathcal{L}_{A_{2}}(y, x) & =P_{A_{2}} \mathcal{L}(y, x)=\frac{1}{2}(\mathcal{L}(y, x)-\mathcal{L}(-y, x)) . \tag{22.1}
\end{align*}
$$

To compute the traces of the symmetrization and antisymmetrization projection operators (22.1), we have to distinguish three kinds of cycles: asymmetric cycles $a$, symmetric cycles $s$ built by repeats of irreducible segments $\tilde{s}$, and boundary cycles $b$. Now we show that the spectral determinant can be written as the product over the three kinds of cycles: $\operatorname{det}(1-\mathcal{L})=\operatorname{det}(1-\mathcal{L})_{a} \operatorname{det}(1-\mathcal{L})_{\tilde{s}} \operatorname{det}(1-\mathcal{L})_{b}$.

Asymmetric cycles: A periodic orbits is not symmetric if $\left\{x_{a}\right\} \cap\left\{\mathbf{C} x_{a}\right\}=$ $\emptyset$, where $\left\{x_{a}\right\}$ is the set of periodic points belonging to the cycle $a$. Thus C generates a second orbit with the same number of points and the same stability properties. Both orbits give the same contribution to the first term and no contribution to the second term in (22.1); as they are degenerate, the prefactor $1 / 2$ cancels. Resuming as in the derivation of (15.15) we find


Figure 22.1: The Ulam sawtooth map with the $C_{2}$ symmetry $f(-x)=-f(x)$. (a) boundary fixed point $\bar{C}$, (b) symmetric 2-cycle $\overline{L R}$, (c) asymmetric 2-cycles pair $\{\overline{L C}, \overline{C R}\}$. The Ulam sawtooth map restricted to the fundamental domain; pieces of the global map (a) are reflected into the upper right quadrant. (d) Boundary fixed point $\bar{C}$ maps into the fixed point $\bar{c}$, symmetric 2 -cycle $\overline{L R}$ maps into fixed point $\bar{s}$, and the asymmetric fixed point pair $\{\bar{L}, \bar{R}\}$ maps into a single fixed point $\bar{r}$, (e) the asymmetric 2 -cycles pair $\{\overline{L C}, \overline{C R}\}$ maps into a single 2 -cycle $\overline{c r}$.
that asymmetric orbits yield the same contribution to the symmetric and the antisymmetric subspaces:

$$
\operatorname{det}\left(1-\mathcal{L}_{ \pm}\right)_{a}=\prod_{a} \prod_{k=0}^{\infty}\left(1-\frac{t_{a}}{\Lambda_{a}^{k}}\right), \quad t_{a}=\frac{z^{n_{a}}}{\left|\Lambda_{a}\right|}
$$

Symmetric cycles: A cycle $s$ is reflection symmetric if operating with $\mathbf{C}$ on the set of cycle points reproduces the set. The period of a symmetric cycle is always even $\left(n_{s}=2 n_{\tilde{s}}\right)$ and the mirror image of the $x_{s}$ cycle point is reached by traversing the irreducible segment $\tilde{s}$ of length $n_{\tilde{s}}, f^{n_{\tilde{s}}}\left(x_{s}\right)=\mathbf{C} x_{s}$. $\delta\left(x-f^{n}(x)\right)$ picks up $2 n_{\tilde{s}}$ contributions for every even traversal, $n=r n_{\tilde{s}}, r$ even, and $\delta\left(x+f^{n}(x)\right)$ for every odd traversal, $n=r n_{\tilde{s}}$, $r$ odd. Absorb the group-theoretic prefactor in the stability eigenvalue by defining the stability computed for a segment of length $n_{\tilde{s}}$,

$$
\Lambda_{\tilde{s}}=-\left.\frac{\partial f^{n_{\tilde{s}}}(x)}{\partial x}\right|_{x=x_{s}}
$$

Restricting the integration to the infinitesimal neighborhood $\mathcal{M}_{s}$ of the $s$ cycle, we obtain the contribution to $\operatorname{tr} \mathcal{L}_{ \pm}^{n}$ :

$$
z^{n} \operatorname{tr} \mathcal{L}_{ \pm}^{n} \rightarrow \int_{\mathcal{M}_{s}} d x z^{n} \frac{1}{2}\left(\delta\left(x-f^{n}(x)\right) \pm \delta\left(x+f^{n}(x)\right)\right)
$$

$$
\begin{aligned}
& =n_{\tilde{s}}\left(\sum_{r=2}^{\text {even }} \delta_{n, r n_{\tilde{s}}} \frac{t_{\tilde{s}}^{r}}{1-1 / \Lambda_{\tilde{s}}^{r}} \pm \sum_{r=1}^{\text {odd }} \delta_{n, r n_{\tilde{s}}} \frac{t_{\tilde{s}}^{r}}{1-1 / \Lambda_{\tilde{s}}^{r}}\right) \\
& =n_{\tilde{s}} \sum_{r=1}^{\infty} \delta_{n, r n_{\tilde{s}}} \frac{\left( \pm t_{\tilde{s}}\right)^{r}}{1-1 / \Lambda_{\tilde{s}}^{r}} .
\end{aligned}
$$

Substituting all symmetric cycles $s$ into $\operatorname{det}\left(1-\mathcal{L}_{ \pm}\right)$and resuming we obtain:

$$
\operatorname{det}\left(1-\mathcal{L}_{ \pm}\right)_{\tilde{s}}=\prod_{\tilde{s}} \prod_{k=0}^{\infty}\left(1 \mp \frac{t_{\tilde{s}}}{\Lambda_{\tilde{s}}^{k}}\right)
$$

Boundary cycles: In the example at hand there is only one cycle which is neither symmetric nor antisymmetric, but lies on the boundary of the fundamental domain, the fixed point at the origin. Such cycle contributes simultaneously to both $\delta\left(x-f^{n}(x)\right)$ and $\delta\left(x+f^{n}(x)\right)$ :

$$
\begin{aligned}
z^{n} \operatorname{tr} \mathcal{L}_{ \pm}^{n} & \rightarrow \int_{\mathcal{M}_{b}} d x z^{n} \frac{1}{2}\left(\delta\left(x-f^{n}(x)\right) \pm \delta\left(x+f^{n}(x)\right)\right) \\
& =\sum_{r=1}^{\infty} \delta_{n, r} t_{b}^{r} \frac{1}{2}\left(\frac{1}{1-1 / \Lambda_{b}^{r}} \pm \frac{1}{1+1 / \Lambda_{b}^{r}}\right) \\
z^{n} \operatorname{tr} \mathcal{L}_{+}^{n} & \rightarrow \sum_{r=1}^{\infty} \delta_{n, r} \frac{t_{b}^{r}}{1-1 / \Lambda_{b}^{2 r}} ; \quad z^{n} \operatorname{tr} \mathcal{L}_{-}^{n} \rightarrow \sum_{r=1}^{\infty} \delta_{n, r} \frac{1}{\Lambda_{b}^{r}} \frac{t_{b}^{r}}{1-1 / \Lambda_{b}^{2 r}} .
\end{aligned}
$$

Boundary orbit contributions to the factorized spectral determinants follow by resummation:

$$
\operatorname{det}\left(1-\mathcal{L}_{+}\right)_{b}=\prod_{k=0}^{\infty}\left(1-\frac{t_{b}}{\Lambda_{b}^{2 k}}\right), \quad \operatorname{det}\left(1-\mathcal{L}_{-}\right)_{b}=\prod_{k=0}^{\infty}\left(1-\frac{t_{b}}{\Lambda_{b}^{2 k+1}}\right)
$$

Only even derivatives contribute to the symmetric subspace (and odd to the antisymmetric subspace) because the orbit lies on the boundary.

Finally, the symmetry reduced spectral determinants follow by collecting the above results:

$$
\begin{align*}
& F_{+}(z)=\prod_{a} \prod_{k=0}^{\infty}\left(1-\frac{t_{a}}{\Lambda_{a}^{k}}\right) \prod_{\tilde{s}} \prod_{k=0}^{\infty}\left(1-\frac{t_{\tilde{s}}}{\Lambda_{\tilde{s}}^{k}}\right) \prod_{k=0}^{\infty}\left(1-\frac{t_{b}}{\Lambda_{b}^{2 k}}\right) \\
& F_{-}(z)=\prod_{a} \prod_{k=0}^{\infty}\left(1-\frac{t_{a}}{\Lambda_{a}^{k}}\right) \prod_{\tilde{s}} \prod_{k=0}^{\infty}\left(1+\frac{t_{\tilde{s}}}{\Lambda_{\tilde{s}}^{k}}\right) \prod_{k=0}^{\infty}\left(1-\frac{t_{b}}{\Lambda_{b}^{2 k+1}}\right) \tag{22.2}
\end{align*}
$$

We shall work out the symbolic dynamics of such reflection symmetric systems in some detail in sect. 22.5. As reflection symmetry is essentially the only discrete symmetry that a map of the interval can have, this example completes the group-theoretic factorization of determinants and zeta functions for $1-d$ maps. We now turn to discussion of the general case.

### 22.2 Discrete symmetries

A dynamical system is invariant under a symmetry group $G=\left\{e, g_{2}, \ldots, g_{|G|}\right\}$ if the equations of motion are invariant under all symmetries $g \in G$. For a map $x_{n+1}=f\left(x_{n}\right)$ and the evolution operator $\mathcal{L}(y, x)$ defined by (10.23) this means

$$
\begin{align*}
f(x) & =\mathbf{g}^{-1} f(\mathbf{g} x) \\
\mathcal{L}(y, x) & =\mathcal{L}(\mathbf{g} y, \mathbf{g} x) \tag{22.3}
\end{align*}
$$

Bold face letters for group elements indicate a suitable representation on phase space. For example, if a 2 -dimensional map has the symmetry $x_{1} \rightarrow$ $-x_{1}, x_{2} \rightarrow-x_{2}$, the symmetry group $G$ consists of the identity and $C$, a rotation by $\pi$ around the origin. The map $f$ must then commute with rotations by $\pi, f(\mathbf{C} x)=\mathbf{C} f(x)$, with $\mathbf{C}$ given by the $[2 \times 2]$ matrix

$$
\mathbf{C}=\left(\begin{array}{cc}
-1 & 0  \tag{22.4}\\
0 & -1
\end{array}\right) .
$$

C satisfies $\mathbf{C}^{2}=e$ and can be used to decompose the phase space into mutually orthogonal symmetric and antisymmetric subspaces by means of projection operators (22.1). More generally the projection operator onto the $\alpha$ irreducible subspace of dimension $d_{\alpha}$ is given by $P_{\alpha}=\left(d_{\alpha} /|G|\right) \sum \chi_{\alpha}(h) \mathbf{h}^{-1}$, where $\chi_{\alpha}(h)=\operatorname{tr} D_{\alpha}(h)$ are the group characters, and the transfer operator $\mathcal{L}$ splits into a sum of inequivalent irreducible subspace contributions $\sum_{\alpha} \operatorname{tr} \mathcal{L}_{\alpha}$,

$$
\begin{equation*}
\mathcal{L}_{\alpha}(y, x)=\frac{d_{\alpha}}{|G|} \sum_{h \in G} \chi_{\alpha}(h) \mathcal{L}\left(\mathbf{h}^{-1} y, x\right) \tag{22.5}
\end{equation*}
$$

The prefactor $d_{\alpha}$ in the above reflects the fact that a $d_{\alpha}$-dimensional representation occurs $d_{\alpha}$ times.

### 22.2.1 Cycle degeneracies

If $g \in G$ is a symmetry of the dynamical problem, the weight of a cycle $p$ and the weight of its image under a symmetry transformation $g$ are equal, $t_{g p}=t_{p}$. The number of degenerate cycles (topologically distinct, but

Figure 22.2: The symmetries of three disks on an equilateral triangle. The fundamental domain is indicated by the shaded wedge.

mapped into each other by symmetry transformations) depends on the cycle symmetries. Associated with a given cycle $p$ is a maximal subgroup $\mathcal{H}_{p} \subseteq G$, $\mathcal{H}_{p}=\left\{e, b_{2}, b_{3}, \ldots, b_{h}\right\}$ of order $h_{p}$, whose elements leave $p$ invariant. The elements of the quotient space $b \in G / \mathcal{H}_{p}$ generate the degenerate cycles $b p$, so the multiplicity of a degenerate cycle is $m_{p}=g / h_{p}$.

Taking into account these degeneracies, the Euler product (15.15) takes the form

$$
\begin{equation*}
\prod_{p}\left(1-t_{p}\right)=\prod_{\hat{p}}\left(1-t_{\hat{p}}\right)^{m_{\hat{p}}} . \tag{22.6}
\end{equation*}
$$

Here $\hat{p}$ is one of the $m_{p}$ degenerate cycles, picked to serve as the label for the entire class. Our labeling convention is usually lexical, that is, we label a cycle $p$ by the cycle point whose label has the lowest value, and we label a class of degenerate cycles by the one with the lowest label $\hat{p}$. In what follows we shall drop the hat in $\hat{p}$ when it is clear from the context that we are dealing with symmetry distinct classes of cycles.

### 22.2.2 Example: $C_{3 v}$ invariance

An illustration of the above is afforded by $C_{3 v}$, the group of symmetries of a game of pinball with three equal size, equally spaced disks, figure 22.2. The group consists of the identity element $e$, three reflections across axes $\left\{\sigma_{12}, \sigma_{23}, \sigma_{13}\right\}$, and two rotations by $2 \pi / 3$ and $4 \pi / 3$ denoted $\left\{C_{3}, C_{3}^{2}\right\}$, so its dimension is $g=6$. On the disk labels $\{1,2,3\}$ these symmetries act as permutations which map cycles into cycles. For example, the flip across the symmetry axis going through disk 1 interchanges the symbols 2 and 3; it maps the cycle $\overline{12123}$ into $\overline{13132}$, figure 22.3a.

The subgroups of $C_{3 v}$ are $C_{v}$, consisting of the identity and any one of the reflections, of dimension $h=2$, and $C_{3}=\left\{e, C_{3}, C_{3}^{2}\right\}$, of dimension $h=3$, so possible cycle multiplicities are $g / h=2,3$ or 6 .

The $C_{3}$ subgroup invariance is exemplified by the cycles $\overline{123}$ and $\overline{132}$ which are invariant under rotations by $2 \pi / 3$ and $4 \pi / 3$, but are mapped

Figure 22.3: Some examples of 3-disk cycles: (a) $\overline{12123}$ and $\overline{13132}$ are mapped into each other by $\sigma_{23}$, the flip across 1 axis; this cycle has degeneracy 6 under $C_{3 v}$ symmetries. ( $C_{3 v}$ is the symmetry group of the equilateral triangle.) Similarly (b) $\overline{123}$ and $\overline{132}$ and (c) $\overline{1213}$, $\overline{1232}$ and $\overline{1323}$ are degenerate under $C_{3 v}$. (d) The cycles $\overline{121212313}$ and $\overline{121212323}$ are related by time reversal but not by any $C_{3 v}$ symmetry. (from ref. [1.2])
a)

b)

c)

into each other by any reflection, figure $22.3 \mathrm{~b} ; \mathcal{H}_{p}=\left\{e, C_{3}, C_{3}^{2}\right\}$, and the degeneracy is $g / h_{c_{3}}=2$.

The $C_{v}$ type of a subgroup is exemplified by the invariances of $\hat{p}=1213$. This cycle is invariant under reflection $\sigma_{23}\{\overline{1213}\}=\overline{1312}=\overline{1213}$, so the invariant subgroup is $\mathcal{H}_{\hat{p}}=\left\{e, \sigma_{23}\right\}$. Its order is $h_{C_{v}}=2$, so the degeneracy is $m_{\hat{p}}=g / h_{C_{v}}=3$; the cycles in this class, $\overline{1213}, \overline{1232}$ and $\overline{1323}$, are related by $2 \pi / 3$ rotations, figure 22.3 (c).

A cycle of no symmetry, such as $\overline{12123}$, has $\mathcal{H}_{p}=\{e\}$ and contributes in all six terms (the remaining cycles in the class are $\overline{12132}, \overline{12313}, \overline{12323}$, $\overline{13132}$ and $\overline{13232}$ ), figure 22.3a.

Besides the above discrete symmetries, for Hamiltonian systems cycles may be related by time reversal symmetry. An example are the cycles $\overline{121212313}$ and $\overline{121212323}=\overline{313212121}$ which are related by no space symmetry (figure $22.3(\mathrm{~d})$ ).

The Euler product (15.15) for the $C_{3 v}$ symmetric 3-disk problem is given in (18.32).

### 22.3 Dynamics in the fundamental domain

So far we have used the discrete symmetry to effect a reduction in the number of independent cycles in cycle expansions. The next step achieves much more: the symmetries can be used to restrict all computations to a fundamental domain. We show here that to each global cycle $p$ corresponds a fundamental domain cycle $\tilde{p}$. Conversely, each fundamental domain cycle $\tilde{p}$ traces out a segment of the global cycle $p$, with the end point of the cycle $\tilde{p}$ mapped into the irreducible segment of $p$ with the group element $h_{\tilde{p}}$.

An important effect of a discrete symmetry is that it tessellates the phase space into copies of a fundamental domain, and thus induces a natural partition of phase space. The group elements $g=\{a, b, \cdots, d\}$ which map the fundamental domain $\tilde{M}$ into its copies $g \tilde{M}$, can double in function as letters of a symbolic dynamics alphabet. If the dynamics is symmetric under interchanges of disks, the absolute disk labels $\epsilon_{i}=1,2, \cdots, N$ can be replaced by the symmetry-invariant relative disk $\rightarrow$ disk increments $g_{i}$, where $g_{i}$ is the discrete group element that maps disk $i-1$ into disk $i$. We demonstrate the reduction for a series of specific examples in sect. 22.4. An immediate gain arising from symmetry invariant relabeling is that $N$ disk symbolic dynamics becomes ( $N-1$ )-nary, with no restrictions on the admissible sequences. However, the main gain is in the close connection between the symbol string symmetries and the phase space symmetries which will aid us in the dynamical zeta function factorizations. Once the connection between the full space and the reduced space is established, working in the fundamental domain (ie., with irreducible segments) is so much simpler that we never use the full space orbits in actual computations.

If the dynamics is invariant under a discrete symmetry, the phase space $M$ can be completely tiled by the fundamental domain $\tilde{M}$ and its images $a \tilde{M}, b \tilde{M}, \ldots$ under the action of the symmetry group $G=\{e, a, b, \ldots\}$,

$$
M=\sum_{a \in G} M_{a}=\sum_{a \in G} a \tilde{M} .
$$

In the above example (22.4) with symmetry group $G=\{e, C\}$, the phase space $M=\left\{x_{1}-x_{2}\right.$ plane $\}$ can be tiled by a fundamental domain $\tilde{M}=\{$ halfplane $\left.x_{1} \geq 0\right\}$, and $\mathbf{C} \tilde{M}=\left\{\right.$ half-plane $\left.x_{1} \leq 0\right\}$, its image under rotation by $\pi$.

If the space $M$ is decomposed into $g$ tiles, a function $\phi(x)$ over $M$ splits into a $g$-dimensional vector $\phi_{a}(x)$ defined by $\phi_{a}(x)=\phi(x)$ if $x \in$ $M_{a}, \phi_{a}(x)=0$ otherwise. Let $h=a b^{-1}$ conflicts with be the symmetry operation that maps the endpoint domain $M_{b}$ into the starting point domain $M_{a}$, and let $D(h)_{b a}$, the left regular representation, be the $[g \times g]$ matrix whose $b, a$-th entry equals unity if $a=h b$ and zero otherwise; $D(h)_{b a}=$ $\delta_{b h, a}$. Since the symmetries act on phase space as well, the operation $h$ enters in two guises: as a $[g \times g]$ matrix $D(h)$ which simply permutes the domain labels, and as a $[d \times d]$ matrix representation $\mathbf{h}$ of a discrete symmetry operation on the $d$ phase-space coordinates. For instance, in the above example (22.4) $h \in C_{2}$ and $D(h)$ can be either the identity or the interchange of the two domain labels,

$$
D(e)=\left(\begin{array}{ll}
1 & 0  \tag{22.7}\\
0 & 1
\end{array}\right), \quad D(C)=\left(\begin{array}{ll}
0 & 1 \\
1 & 0
\end{array}\right) .
$$

Note that $D(h)$ is a permutation matrix, mapping a tile $M_{a}$ into a different tile $M_{h a} \neq M_{a}$ if $h \neq e$. Consequently only $D(e)$ has diagonal elements, and $\operatorname{tr} D(h)=g \delta_{h, e}$. However, the phase-space transformation $\mathbf{h} \neq \mathbf{e}$ leaves
invariant sets of boundary points; for example, under reflection $\sigma$ across a symmetry axis, the axis itself remains invariant. The boundary periodic orbits that belong to such point-wise invariant sets will require special care in $\operatorname{tr} \mathcal{L}$ evaluations.

One can associate to the evolution operator (10.23) a $[g \times g]$ matrix evolution operator defined by

$$
\mathcal{L}_{b a}(y, x)=D(h)_{b a} \mathcal{L}(y, x),
$$

if $x \in M_{a}$ and $y \in M_{b}$, and zero otherwise. Now we can use the invariance condition (22.3) to move the starting point $x$ into the fundamental domain $x=\mathbf{a} \tilde{x}, \mathcal{L}(y, x)=\mathcal{L}\left(\mathbf{a}^{-1} y, \tilde{x}\right)$, and then use the relation $a^{-1} b=h^{-1}$ to also relate the endpoint $y$ to its image in the fundamental domain, $\tilde{\mathcal{L}}(\tilde{y}, \tilde{x}):=$ $\mathcal{L}\left(\mathbf{h}^{-1} \tilde{y}, \tilde{x}\right)$. With this operator which is restricted to the fundamental domain, the global dynamics reduces to

$$
\mathcal{L}_{b a}(y, x)=D(h)_{b a} \tilde{\mathcal{L}}(\tilde{y}, \tilde{x}) .
$$

While the global trajectory runs over the full space $M$, the restricted trajectory is brought back into the fundamental domain $\tilde{M}$ any time it crosses into adjoining tiles; the two trajectories are related by the symmetry operation $h$ which maps the global endpoint into its fundamental domain image.

Now the traces (15.3) required for the evaluation of the eigenvalues of the transfer operator can be evaluated on the fundamental domain alone

$$
\begin{equation*}
\operatorname{tr} \mathcal{L}=\int_{M} d x \mathcal{L}(x, x)=\int_{\tilde{M}} d \tilde{x} \sum_{h} \operatorname{tr} D(h) \mathcal{L}\left(\mathbf{h}^{-1} \tilde{x}, \tilde{x}\right) \tag{22.8}
\end{equation*}
$$

The fundamental domain integral $\int d \tilde{x} \mathcal{L}\left(\mathbf{h}^{-1} \tilde{x}, \tilde{x}\right)$ picks up a contribution from every global cycle (for which $h=e$ ), but it also picks up contributions from shorter segments of global cycles. The permutation matrix $D(h)$ guarantees by the identity $\operatorname{tr} D(h)=0, h \neq e$, that only those repeats of the fundamental domain cycles $\tilde{p}$ that correspond to complete global cycles $p$ contribute. Compare, for example, the contributions of the $\overline{12}$ and $\overline{0}$ cycles of figure 11.6. $\operatorname{tr} D(h) \tilde{\mathcal{L}}$ does not get a contribution from the $\overline{0}$ cycle, as the symmetry operation that maps the first half of the $\overline{12}$ into the fundamental domain is a reflection, and $\operatorname{tr} D(\sigma)=0$. In contrast, $\sigma^{2}=e, \operatorname{tr} D\left(\sigma^{2}\right)=6$ insures that the repeat of the fundamental domain fixed point $\operatorname{tr}(D(h) \tilde{\mathcal{L}})^{2}=6 t_{0}^{2}$, gives the correct contribution to the global trace $\operatorname{tr} \mathcal{L}^{2}=3 \cdot 2 t_{12}$.

Let $p$ be the full orbit, $\tilde{p}$ the orbit in the fundamental domain and $h_{\tilde{p}}$ an element of $\mathcal{H}_{p}$, the symmetry group of $p$. Restricting the volume integrations to the infinitesimal neighborhoods of the cycles $p$ and $\tilde{p}$, respectively, and performing the standard resummations, we obtain the identity

$$
\begin{equation*}
\left(1-t_{p}\right)^{m_{p}}=\operatorname{det}\left(1-D\left(h_{\tilde{p}}\right) t_{\tilde{p}}\right), \tag{22.9}
\end{equation*}
$$

valid cycle by cycle in the Euler products $(15.15)$ for $\operatorname{det}(1-\mathcal{L})$. Here "det" refers to the $[g \times g]$ matrix representation $D\left(h_{\tilde{p}}\right)$; as we shall see, this determinant can be evaluated in terms of standard characters, and no explicit representation of $D\left(h_{\tilde{p}}\right)$ is needed. Finally, if a cycle $p$ is invariant under the symmetry subgroup $\mathcal{H}_{p} \subseteq G$ of order $h_{p}$, its weight can be written as a repetition of a fundamental domain cycle

$$
\begin{equation*}
t_{p}=t_{\tilde{p}}^{h_{p}} \tag{22.10}
\end{equation*}
$$

computed on the irreducible segment that corresponds to a fundamental domain cycle. For example, in figure 11.6 we see by inspection that $t_{12}=t_{0}^{2}$ and $t_{123}=t_{1}^{3}$.

### 22.3.1 Boundary orbits

Before we can turn to a presentation of the factorizations of dynamical zeta functions for the different symmetries we have to discuss an effect that arises for orbits that run on a symmetry line that borders a fundamental domain. In our 3 -disk example, no such orbits are possible, but they exist in other systems, such as in the bounded region of the Hénon-Heiles potential and in 1-d maps. For the symmetrical 4-disk billiard, there are in principle two kinds of such orbits, one kind bouncing back and forth between two diagonally opposed disks and the other kind moving along the other axis of reflection symmetry; the latter exists for bounded systems only. While there are typically very few boundary orbits, they tend to be among the shortest orbits, and their neglect can seriously degrade the convergence of cycle expansions, as those are dominated by the shortest cycles.

While such orbits are invariant under some symmetry operations, their neighborhoods are not. This affects the stability matrix $\mathbf{J}_{p}$ of the linearization perpendicular to the orbit and thus the eigenvalues. Typically, e.g. if the symmetry is a reflection, some eigenvalues of $\mathbf{J}_{p}$ change sign. This means that instead of a weight $1 / \operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}\right)$ as for a regular orbit, boundary cycles also pick up contributions of form $1 / \operatorname{det}\left(\mathbf{1}-\mathbf{h J} \mathbf{J}_{p}\right)$, where $\mathbf{h}$ is a symmetry operation that leaves the orbit pointwise invariant; see for example sect. 22.1.2.

Consequences for the dynamical zeta function factorizations are that sometimes a boundary orbit does not contribute. A derivation of a dynamical zeta function (15.15) from a determinant like (15.9) usually starts with an expansion of the determinants of the Jacobian. The leading order terms just contain the product of the expanding eigenvalues and lead to the dynamical zeta function (15.15). Next to leading order terms contain products of expanding and contracting eigenvalues and are sensitive to their signs. Clearly, the weights $t_{p}$ in the dynamical zeta function will then be affected by reflections in the Poincaré surface of section perpendicular to the orbit. In all our applications it was possible to implement these effects by the following simple prescription.

If an orbit is invariant under a little group $\mathcal{H}_{p}=\left\{e, b_{2}, \ldots, b_{h}\right\}$, then the corresponding group element in (22.9) will be replaced by a projector. If the weights are insensitive to the signs of the eigenvalues, then this projector is

$$
\begin{equation*}
g_{p}=\frac{1}{h} \sum_{i=1}^{h} b_{i} . \tag{22.11}
\end{equation*}
$$

In the cases that we have considered, the change of sign may be taken into account by defining a sign function $\epsilon_{p}(g)= \pm 1$, with the "-" sign if the symmetry element $g$ flips the neighborhood. Then (22.11) is replaced by

$$
\begin{equation*}
g_{p}=\frac{1}{h} \sum_{i=1}^{h} \epsilon\left(b_{i}\right) b_{i} . \tag{22.12}
\end{equation*}
$$

We have illustrated the above in sect. 22.1.2 by working out the full factorization for the 1 -dimensional reflection symmetric maps.

### 22.4 Factorizations of dynamical zeta functions

In the above we have shown that a discrete symmetry induces degeneracies among periodic orbits and decomposes periodic orbits into repetitions of irreducible segments; this reduction to a fundamental domain furthermore leads to a convenient symbolic dynamics compatible with the symmetry, and, most importantly, to a factorization of dynamical zeta functions. This we now develop, first in a general setting and then for specific examples.

### 22.4.1 Factorizations of dynamical dynamical zeta functions

According to (22.9) and (22.10), the contribution of a degenerate class of global cycles (cycle $p$ with multiplicity $m_{p}=g / h_{p}$ ) to a dynamical zeta function is given by the corresponding fundamental domain cycle $\tilde{p}$ :

$$
\begin{equation*}
\left(1-t_{\tilde{p}}^{h_{p}}\right)^{g / h_{p}}=\operatorname{det}\left(1-D\left(h_{\tilde{p}}\right) t_{\tilde{p}}\right) \tag{22.13}
\end{equation*}
$$

Let $D(h)=\bigoplus_{\alpha} d_{\alpha} D_{\alpha}(h)$ be the decomposition of the matrix representation $D(h)$ into the $d_{\alpha}$ dimensional irreducible representations $\alpha$ of a finite group $G$. Such decompositions are block-diagonal, so the corresponding contribution to the Euler product (15.9) factorizes as

$$
\begin{equation*}
\operatorname{det}(1-D(h) t)=\prod_{\alpha} \operatorname{det}\left(1-D_{\alpha}(h) t\right)^{d_{\alpha}} \tag{22.14}
\end{equation*}
$$

where now the product extends over all distinct $d_{\alpha}$-dimensional irreducible representations, each contributing $d_{\alpha}$ times. For the cycle expansion purposes, it has been convenient to emphasize that the group-theoretic factorization can be effected cycle by cycle, as in (22.13); but from the transfer operator point of view, the key observation is that the symmetry reduces the transfer operator to a block diagonal form; this block diagonalization implies that the dynamical zeta functions (15.15) factorize as

$$
\begin{equation*}
\frac{1}{\zeta}=\prod_{\alpha} \frac{1}{\zeta_{\alpha}^{d_{\alpha}}}, \quad \frac{1}{\zeta_{\alpha}}=\prod_{\tilde{p}} \operatorname{det}\left(1-D_{\alpha}\left(h_{\tilde{p}}\right) t_{\tilde{p}}\right) \tag{22.15}
\end{equation*}
$$

Determinants of $d$-dimensional irreducible representations can be evaluated using the expansion of determinants in terms of traces,

$$
\begin{align*}
\operatorname{det}(1+M)= & 1+\operatorname{tr} M+\frac{1}{2}\left((\operatorname{tr} M)^{2}-\operatorname{tr} M^{2}\right) \\
& +\frac{1}{6}\left((\operatorname{tr} M)^{3}-3(\operatorname{tr} M)\left(\operatorname{tr} M^{2}\right)+2 \operatorname{tr} M^{3}\right) \\
& +\cdots+\frac{1}{d!}\left((\operatorname{tr} M)^{d}-\cdots\right), \tag{22.16}
\end{align*}
$$

(see (K.26), for example) and each factor in (22.14) can be evaluated by looking up the characters $\chi_{\alpha}(h)=\operatorname{tr} D_{\alpha}(h)$ in standard tables [22.14]. In terms of characters, we have for the 1-dimensional representations

$$
\operatorname{det}\left(1-D_{\alpha}(h) t\right)=1-\chi_{\alpha}(h) t
$$

for the 2-dimensional representations

$$
\operatorname{det}\left(1-D_{\alpha}(h) t\right)=1-\chi_{\alpha}(h) t+\frac{1}{2}\left(\chi_{\alpha}(h)^{2}-\chi_{\alpha}\left(h^{2}\right)\right) t^{2},
$$

and so forth.
In the fully symmetric subspace $\operatorname{tr} D_{A_{1}}(h)=1$ for all orbits; hence a straightforward fundamental domain computation (with no group theory weights) always yields a part of the full spectrum. In practice this is the most interesting subspectrum, as it contains the leading eigenvalue of the transfer operator.

### 22.4.2 Factorizations of spectral determinants

Factorization of the full spectral determinant (15.3) proceeds in essentially the same manner as the factorization of dynamical zeta functions outlined
above. By (22.5) and (22.8) the trace of the transfer operator $\mathcal{L}$ splits into the sum of inequivalent irreducible subspace contributions $\sum_{\alpha} \operatorname{tr} \mathcal{L}_{\alpha}$, with

$$
\operatorname{tr} \mathcal{L}_{\alpha}=d_{\alpha} \sum_{h \in G} \chi_{\alpha}(h) \int_{\tilde{M}} d \tilde{x} \mathcal{L}\left(\mathbf{h}^{-1} \tilde{x}, \tilde{x}\right) .
$$

This leads by standard manipulations to the factorization of (15.9) into

$$
\begin{align*}
F(z) & =\prod_{\alpha} F_{\alpha}(z)^{d_{\alpha}} \\
F_{\alpha}(z) & =\exp \left(-\sum_{\tilde{p}} \sum_{r=1}^{\infty} \frac{1}{r} \frac{\chi_{\alpha}\left(h_{\tilde{p}}^{r}\right) z^{n_{\tilde{p}} r}}{\left|\operatorname{det}\left(\mathbf{1}-\tilde{\mathbf{J}}_{\tilde{p}}^{r}\right)\right|}\right) \tag{22.17}
\end{align*}
$$

where $\tilde{\mathbf{J}}_{\tilde{p}}=\mathbf{h}_{\tilde{p}} \mathbf{J}_{\tilde{p}}$ is the fundamental domain Jacobian. Boundary orbits require special treatment, discussed in sect. 22.3.1, with examples given in the next section as well as in the specific factorizations discussed below.

The factorizations (22.15), (22.17) are the central formulas of this chapter. We now work out the group theory factorizations of cycle expansions of dynamical zeta functions for the cases of $C_{2}$ and $C_{3 v}$ symmetries. The cases of the $C_{2 v}, C_{4 v}$ symmetries are worked out in appendix I below.

## $22.5 C_{2}$ factorization

As the simplest example of implementing the above scheme consider the $C_{2}$ symmetry. For our purposes, all that we need to know here is that each orbit or configuration is uniquely labeled by an infinite string $\left\{s_{i}\right\}, s_{i}=+,-$ and that the dynamics is invariant under the $+\leftrightarrow-$ interchange, that is, it is $C_{2}$ symmetric. The $C_{2}$ symmetry cycles separate into two classes, the selfdual configurations,,,,$+-++--+++---+--+-++- \cdots$, with multiplicity $m_{p}=1$, and the asymmetric configurations,,+-++- , ,$--+ \cdots$, with multiplicity $m_{p}=2$. For example, as there is no absolute distinction between the "up" and the "down" spins, or the "left" or the "right" lobe, $t_{+}=t_{-}, t_{++-}=t_{+--}$, and so on.

The symmetry reduced labeling $\rho_{i} \in\{0,1\}$ is related to the standard $s_{i} \in\{+,-\}$ Ising spin labeling by

$$
\begin{align*}
& \text { If } s_{i}=s_{i-1} \text { then } \rho_{i}=1 \\
& \text { If } s_{i} \neq s_{i-1} \text { then } \rho_{i}=0 \tag{22.18}
\end{align*}
$$

For example, $\bar{\mp}=\cdots++++\cdots$ maps into $\cdots 111 \cdots=\overline{1}$ (and so does $\overline{=}$ ), $\overline{-+}=\cdots-+-+\cdots$ maps into $\cdots 000 \cdots=\overline{0}, \overline{-++-}=\cdots--++--+$

| $\tilde{p}$ | $p$ | $m_{p}$ |
| :--- | :--- | :--- |
| 1 | + | 2 |
| 0 | -+ | 1 |
| 01 | --++ | 1 |
| 001 | -++ | 2 |
| 011 | ---+++ | 1 |
| 0001 | -+--+-++ | 2 |
| 0011 | -+++ | 1 |
| 0111 | ----++++ | 2 |
| 00001 | -+-+- | 1 |
| 00011 | -+---+-+++ | 1 |
| 00101 | -++--+--++ | 2 |
| 00111 | -+---+-+++ | 1 |
| 01011 | --+++ | 1 |
| 01111 | -----+++++ | 1 |
| 001011 | -++---+--+++ | 1 |
| 001101 | -+++--+---++ | 1 |

Table 22.1: Correspondence between the $C_{2}$ symmetry reduced cycles $\tilde{p}$ and the standard Ising model periodic configurations $p$, together with their multiplicities $m_{p}$. Also listed are the two shortest cycles (length 6) related by time reversal, but distinct under $C_{2}$.
$+\cdots$ maps into $\cdots 0101 \cdots=\overline{01}$, and so forth. A list of such reductions is given in table 22.1.

Depending on the maximal symmetry group $\mathcal{H}_{p}$ that leaves an orbit $p$ invariant (see sects. 22.2 and 22.3 as well as sect. 22.1.2), the contributions to the dynamical zeta function factor as

$$
\begin{array}{cc} 
& A_{1}
\end{array} A_{2},
$$

For example:

$$
\begin{aligned}
& \mathcal{H}_{++-}=\{e\}: \quad\left(1-t_{++-}\right)^{2}=\left(1-t_{001}\right)\left(1-t_{001}\right) \\
& \mathcal{H}_{+-}=\{e, \sigma\}: \quad\left(1-t_{+-}\right)=\left(1-t_{0}\right)\left(1+t_{0}\right), \quad t_{+-}=t_{0}^{2}
\end{aligned}
$$

This yields two binary cycle expansions. The $A_{1}$ subspace dynamical zeta function is given by the standard binary expansion (18.5). The antisymmetric $A_{2}$ subspace dynamical zeta function $\zeta_{A_{2}}$ differs from $\zeta_{A_{1}}$ only by a minus sign for cycles with an odd number of 0 's:

$$
\begin{align*}
1 / \zeta_{A_{2}}= & \left(1+t_{0}\right)\left(1-t_{1}\right)\left(1+t_{10}\right)\left(1-t_{100}\right)\left(1+t_{101}\right)\left(1+t_{1000}\right) \\
& \left(1-t_{1001}\right)\left(1+t_{1011}\right)\left(1-t_{10000}\right)\left(1+t_{10001}\right) \\
& \left(1+t_{10010}\right)\left(1-t_{10011}\right)\left(1-t_{10101}\right)\left(1+t_{10111}\right) \ldots \\
= & 1+t_{0}-t_{1}+\left(t_{10}-t_{1} t_{0}\right)-\left(t_{100}-t_{10} t_{0}\right)+\left(t_{101}-t_{10} t_{1}\right) \\
& -\left(t_{1001}-t_{1} t_{001}-t_{101} t_{0}+t_{10} t_{0} t_{1}\right)-\ldots \ldots \tag{22.20}
\end{align*}
$$

Note that the group theory factors do not destroy the curvature corrections (the cycles and pseudo cycles are still arranged into shadowing combinations).

If the system under consideration has a boundary orbit ( $c f$. sect. 22.3.1) with group-theoretic factor $\mathbf{h}_{p}=(\mathbf{e}+\sigma) / 2$, the boundary orbit does not contribute to the antisymmetric subspace

$$
\begin{gather*}
\\
\text { boundary: }  \tag{22.21}\\
A_{1} \quad A_{2} \\
\left(1-t_{p}\right)= \\
\left(1-t_{\tilde{p}}\right)\left(1-0 t_{\tilde{p}}\right)
\end{gather*}
$$

This is the $1 / \zeta$ part of the boundary orbit factorization of sect. 22.1.2.

## 22.6 $C_{3 v}$ factorization: 3-disk game of pinball

The next example, the $C_{3 v}$ symmetry, can be worked out by a glance at figure 11.6 a . For the symmetric 3 -disk game of pinball the fundamental domain is bounded by a disk segment and the two adjacent sections of the symmetry axes that act as mirrors (see figure 11.6b). The three symmetry axes divide the space into six copies of the fundamental domain. Any trajectory on the full space can be pieced together from bounces in the fundamental domain, with symmetry axes replaced by flat mirror reflections. The binary $\{0,1\}$ reduction of the ternary three disk $\{1,2,3\}$ labels has a simple geometric interpretation: a collision of type 0 reflects the projectile to the disk it comes from (back-scatter), whereas after a collision of type 1 projectile continues to the third disk. For example, $\overline{23}=\cdots 232323 \cdots$ maps into $\cdots 000 \cdots=\overline{0}$ (and so do $\overline{12}$ and $\overline{13}$ ), $\overline{123}=\cdots 12312 \cdots$ maps into $\cdots 111 \cdots=\overline{1}$ (and so does $\overline{132}$ ), and so forth. A list of such reductions for short cycles is given in table 11.2.
$C_{3 v}$ has two 1-dimensional irreducible representations, symmetric and antisymmetric under reflections, denoted $A_{1}$ and $A_{2}$, and a pair of degenerate 2-dimensional representations of mixed symmetry, denoted $E$. The contribution of an orbit with symmetry $g$ to the $1 / \zeta$ Euler product (22.14) factorizes according to

$$
\begin{equation*}
\operatorname{det}(1-D(h) t)=\left(1-\chi_{A_{1}}(h) t\right)\left(1-\chi_{A_{2}}(h) t\right)\left(1-\chi_{E}(h) t+\chi_{A_{2}}(h) t^{2}\right)^{2} \tag{22.22}
\end{equation*}
$$

with the three factors contributing to the $C_{3 v}$ irreducible representations $A_{1}, A_{2}$ and $E$, respectively, and the 3 -disk dynamical zeta function factorizes into $\zeta=\zeta_{A_{1}} \zeta_{A_{2}} \zeta_{E}^{2}$. Substituting the $C_{3 v}$ characters [22.14]

| $C_{3 v}$ | $A_{1}$ | $A_{2}$ | $E$ |
| :---: | :---: | ---: | ---: |
| $e$ | 1 | 1 | 2 |
| $C_{3}, C_{3}^{2}$ | 1 | 1 | -1 |
| $\sigma_{v}$ | 1 | -1 | 0 |

into (22.22), we obtain for the three classes of possible orbit symmetries (indicated in the first column)

$$
\left.\begin{array}{rl}
\mathbf{h}_{\tilde{p}} & A_{1} A_{2} E \\
e:\left(1-t_{\tilde{p}}\right)^{6} & =\left(1-t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1-2 t_{\tilde{p}}+t_{\tilde{p}}^{2}\right)^{2} \\
C_{3}, C_{3}^{2}: & \left(1-t_{\tilde{p}}^{3}\right)^{2}
\end{array}=\left(1-t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1+t_{\tilde{p}}+t_{\tilde{p}}^{2}\right)^{2}\right) .
$$

where $\sigma_{v}$ stands for any one of the three reflections.
The Euler product (15.15) on each irreducible subspace follows from the factorization (22.23). On the symmetric $A_{1}$ subspace the $\zeta_{A_{1}}$ is given by the standard binary curvature expansion (18.5). The antisymmetric $A_{2}$ subspace $\zeta_{A_{2}}$ differs from $\zeta_{A_{1}}$ only by a minus sign for cycles with an odd number of 0 's, and is given in (22.20). For the mixed-symmetry subspace $E$ the curvature expansion is given by

$$
\begin{aligned}
1 / \zeta_{E}= & \left(1+z t_{1}+z^{2} t_{1}^{2}\right)\left(1-z^{2} t_{0}^{2}\right)\left(1+z^{3} t_{100}+z^{6} t_{100}^{2}\right)\left(1-z^{4} t_{10}^{2}\right) \\
& \left(1+z^{4} t_{1001}+z^{8} t_{1001}^{2}\right)\left(1+z^{5} t_{10000}+z^{10} t_{10000}^{2}\right) \\
& \left(1+z^{5} t_{10101}+z^{10} t_{10101}^{2}\right)\left(1-z^{5} t_{10011}\right)^{2} \cdots \\
= & 1+z t_{1}+z^{2}\left(t_{1}^{2}-t_{0}^{2}\right)+z^{3}\left(t_{001}-t_{1} t_{0}^{2}\right) \\
& +z^{4}\left[t_{0011}+\left(t_{001}-t_{1} t_{0}^{2}\right) t_{1}-t_{01}^{2}\right] \\
& +z^{5}\left[t_{00001}+t_{01011}-2 t_{00111}+\left(t_{0011}-t_{01}^{2}\right) t_{1}+\left(t_{1}^{2}-t_{0}^{2}\right) t_{100}\right](2 z .24)
\end{aligned}
$$

We have reinserted the powers of $z$ in order to group together cycles and pseudocycles of the same length. Note that the factorized cycle expansions retain the curvature form; long cycles are still shadowed by (somewhat less obvious) combinations of pseudocycles.

Referring back to the topological polynomial (13.31) obtained by setting $t_{p}=1$, we see that its factorization is a consequence of the $C_{3 v}$ factorization of the $\zeta$ function:

$$
\begin{equation*}
1 / \zeta_{A_{1}}=1-2 z, \quad 1 / \zeta_{A_{2}}=1, \quad 1 / \zeta_{E}=1+z \tag{22.25}
\end{equation*}
$$

as obtained from (18.5), (22.20) and (22.24) for $t_{p}=1$.
Their symmetry is $K=\{\mathbf{e}, \sigma\}$, so according to (22.11), they pick up the group-theoretic factor $\mathbf{h}_{p}=(\mathbf{e}+\sigma) / 2$. If there is no sign change in $t_{p}$, then evaluation of $\operatorname{det}\left(1-\frac{\mathrm{e}+\sigma}{2} t_{\tilde{p}}\right)$ yields

$$
A_{1} \quad A_{2} \quad E
$$

boundary: $\left(1-t_{p}\right)^{3}=\left(1-t_{\tilde{p}}\right)\left(1-0 t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)^{2}, \quad t_{p}=t_{\tilde{p}}(22.26)$

However, if the cycle weight changes sign under reflection, $t_{\sigma \tilde{p}}=-t_{\tilde{p}}$, the boundary orbit does not contribute to the subspace symmetric under reflection across the orbit;

$$
\begin{gathered}
A_{1} \quad A_{2} \quad E
\end{gathered}
$$

## Commentary

Remark 22.1 Some examples of systems with discrete symmetries. This chapter is based on ref. [22.1]. One has a $C_{2}$ symmetry in the Lorenz system [2.1, 22.15], the Ising model, and in the 3 -dimensional anisotropic Kepler potential [30.6, 30.19, 30.20 ], a $C_{3 v}$ symmetry in Hénon-Heiles type potentials [22.2, 22.6, 22.7, 22.5], a $C_{4 v}$ symmetry in quartic oscillators [22.9, 22.10], in the pure $x^{2} y^{2}$ potential [22.11, 22.12] and in hydrogen in a magnetic field [22.13], and a $C_{2 v}=C_{2} \times C_{2}$ symmetry in the stadium billiard [22.4]. A very nice application of the symmetry factorization is carried out in ref. [22.8].

Remark 22.2 Who did it? This chapter is based on long collaborative effort with B. Eckhardt, ref. [22.1]. The group-theoretic factorizations of dynamical zeta functions that we develop here were first introduced and applied in ref. [6.3]. They are closely related to the symmetrizations introduced by Gutzwiller [30.6] in the context of the semiclassical periodic orbit trace formulas, put into more general group-theoretic context by Robbins [22.4], whose exposition, together with Lauritzen's [22.5] treatment of the boundary orbits, has influenced the presentation given here. A related group-theoretic decomposition in context of hyperbolic billiards was utilized in ref. [22.8].

Remark 22.3 Computations The techniques of this chapter have been applied to computations of the 3 -disk classical and quantum spectra in refs. [1.2, 32.15], and to a "Zeeman effect" pinball and the $x^{2} y^{2}$ potentials in refs. [22.3, 18.12]. In a larger perspective, the factorizations developed above are special cases of a general approach to exploiting the group-theoretic invariances in spectra computations, such as those used in enumeration of periodic geodesics [22.8, 15.4, 15.14] for hyperbolic billiards [30.4] and Selberg zeta functions [25.2].

Remark 22.4 Other symmetries. In addition to the symmetries exploited here, time reversal symmetry and a variety of other non-trivial discrete symmetries can induce further relations among orbits; we shall point out several of examples of cycle degeneracies under time reversal. We do not know whether such symmetries can be exploited for further improvements of cycle expansions.

Remark 22.5 Cycles and symmetries. We conclude this section with a few comments about the role of symmetries in actual extraction of cycles. In the example
at hand, the $N$-disk billiard systems, a fundamental domain is a sliver of the $N$ disk configuration space delineated by a pair of adjoining symmetry axes, with the directions of the momenta indicated by arrows. The flow may further be reduced to a return map on a Poincaré surface of section, on which an appropriate transfer operator may be constructed. While in principle any Poincaré surface of section will do, a natural choice in the present context are crossings of symmetry axes.

In actual numerical integrations only the last crossing of a symmetry line needs to be determined. The cycle is run in global coordinates and the group elements associated with the crossings of symmetry lines are recorded; integration is terminated when the orbit closes in the fundamental domain. Periodic orbits with non-trivial symmetry subgroups are particularly easy to find since their points lie on crossings of symmetry lines.

Remark $22.6 \quad C_{2}$ symmetry The $C_{2}$ symmetry arises, for example, in the Lorenz system [22.15], in the 3-dimensional anisotropic Kepler problem [30.6, $30.19,30.20$ ] or in the cycle expansions treatments of the Ising model [12.64].

Remark 22.7 Hénon-Heiles potential An example of a system with $C_{3 v}$ symmetry is provided by the motion of a particle in the Hénon-Heiles potential [22.2]

$$
V(r, \theta)=\frac{1}{2} r^{2}+\frac{1}{3} r^{3} \sin (3 \theta) .
$$

Our coding is not directly applicable to this system because of the existence of elliptic islands and because the three orbits that run along the symmetry axis cannot be labeled in our code. However, since these orbits run along the boundary of the fundamental domain, they require the special treatment discussed in sect. 22.3.1.

## Résumé

If a dynamical system has a discrete symmetry, the symmetry should be exploited; much is gained, both in understanding of the spectra and ease of their evaluation. Once this is appreciated, it is hard to conceive of a calculation without factorization; it would correspond to quantum mechanical calculations without wave-function symmetrizations.

Reduction to the fundamental domain simplifies symbolic dynamics and eliminates symmetry induced degeneracies. While the resummation of the theory from the trace sums to the cycle expansions does not reduce the exponential growth in number of cycles with the cycle length, in practice only the short orbits are used, and for them the labor saving is dramatic. For example, for the 3 -disk game of pinball there are 256 periodic points of length 8 , but reduction to the fundamental domain non-degenerate prime cycles reduces the number of the distinct cycles of length 8 to 30 .

In addition, cycle expansions of the symmetry reduced dynamical zeta functions converge dramatically faster than the unfactorized dynamical zeta
functions. One reason is that the unfactorized dynamical zeta function has many closely spaced zeros and zeros of multiplicity higher than one; since the cycle expansion is a polynomial expansion in topological cycle length, accommodating such behavior requires many terms. The dynamical zeta functions on separate subspaces have more evenly and widely spaced zeros, are smoother, do not have symmetry-induced multiple zeros, and fewer cycle expansion terms (short cycle truncations) suffice to determine them. Furthermore, the cycles in the fundamental domain sample phase space more densely than in the full space. For example, for the 3 -disk problem, there are 9 distinct (symmetry unrelated) cycles of length 7 or less in full space, corresponding to 47 distinct periodic points. In the fundamental domain, we have 8 (distinct) periodic orbits up to length 4 and thus 22 different periodic points in $1 / 6$-th the phase space, that is, an increase in density by a factor 3 with the same numerical effort.

We emphasize that the symmetry factorization (22.23) of the dynamical zeta functionis intrinsic to the classical dynamics, and not a special property of quantal spectra. The factorization is not restricted to the Hamiltonian systems, or only to the configuration space symmetries; for example, the discrete symmetry can be a symmetry of the Hamiltonian phase space [22.4]. In conclusion, the manifold advantages of the symmetry reduced dynamics should thus be obvious; full space cycle expansions, such as those of exercise 18.8, are useful only for cross checking purposes.

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## Exercises

Exercise 22.1 Sawtooth map desymmetrization. Work out the some of the shortest global cycles of different symmetries and fundamental domain cycles for the sawtooth map of figure 22.1. Compute the dynamical zeta function and the spectral determinant of the Perron-Frobenius operator for this map; check explicitely the factorization (22.2).

Exercise 22.2 2-d asymmetric representation. The above expressions can sometimes be simplified further using standard group-theoretical methods. For example, the $\frac{1}{2}\left((\operatorname{tr} M)^{2}-\operatorname{tr} M^{2}\right)$ term in (22.16) is the trace of the antisymmetric part of the $M \times M$ Kronecker product; if $\alpha$ is a 2-dimensional representation, this is the $A_{2}$ antisymmetric representation, so

$$
\begin{equation*}
2-\operatorname{dim}: \quad \operatorname{det}\left(1-D_{\alpha}(h) t\right)=1-\chi_{\alpha}(h) t+\chi_{A_{2}}(h) t^{2} . \tag{22.28}
\end{equation*}
$$

## Exercise 22.3 3-disk desymmetrization.

a) Work out the 3-disk symmetry factorization for the 0 and 1 cycles, i.e. which symmetry do they have, what is the degeneracy in full space and how do they factorize (how do they look in the $A_{1}, A_{2}$ and the $E$ representations).
b) Find the shortest cycle with no symmetries and factorize it like in a)
c) Find the shortest cycle that has the property that its time reversal is not described by the same symbolic dynamics.
d) Compute the dynamical zeta functions and the spectral determinants (symbolically) in the three representations; check the factorizations (22.15) and (22.17).

Exercise 22.4 The group $C_{3 v}$. We will compute a few of the properties of the group $C_{3 v}$, the group of symmetries of an equilateral triangle
(a) All discrete groups are isomorphic to a permutation group or one of its subgroups, and elements of the permutation group can be expressed as cycles. Express the elements of the group $C_{3 v}$ as cycles. For example, one of the rotations is (123), meaning that vertex 1 maps to 2 and 2 to 3 and 3 to 1 .
(b) Find the subgroups of the group $C_{3 v}$.
(c) Find the classes of $C_{3 v}$ and the number of elements in them.
(d) Their are three irreducible representations for the group. Two are one dimensional and the other one is formed by $2 \times 2$ matrices of the form

$$
\left[\begin{array}{cc}
\cos \theta & \sin \theta \\
-\sin \theta & \cos \theta
\end{array}\right] .
$$

Find the matrices for all six group elements.
(e) Use your representation to find the character table for the group.

Exercise $22.5 \quad C_{2}$ factorizations: the Lorenz and Ising systems. In the Lorenz system [2.1, 22.15] the labels + and - stand for the left or the right lobe of the attractor and the symmetry is a rotation by $\pi$ around the $z$-axis. Similarly, the Ising Hamiltonian (in the absence of an external magnetic field) is invariant under spin flip. Work out the factorizations for some of the short cycles in either system.

Exercise 22.6 Ising model. The Ising model with two states $\epsilon_{i}=\{+,-\}$ per site, periodic boundary condition, and Hamiltonian

$$
H(\epsilon)=-J \sum_{i} \delta_{\epsilon_{i}, \epsilon_{i+1}},
$$

is invariant under spin-flip: $+\leftrightarrow-$. Take advantage of that symmetry and factorize the dynamical zeta function for the model, that is, find all the periodic orbits that contribute to each factor and their weights.

Exercise 22.7 One orbit contribution. If $p$ is an orbit in the fundamental domain with symmetry $h$, show that it contributes to the spectral determinant with a factor

$$
\operatorname{det}\left(1-D(h) \frac{t_{p}}{\lambda_{p}^{k}}\right),
$$

where $D(h)$ is the representation of $h$ in the regular representation of the group.

## Chapter 23

## Deterministic diffusion

This is a bizzare and discordant situation. M.V. Berry

## (R. Artuso and P. Cvitanović)

The advances in the theory of dynamical systems have brought a new life to Boltzmann's mechanical formulation of statistical mechanics. Sinai, Ruelle and Bowen (SRB) have generalized Boltzmann's notion of ergodicity for a constant energy surface for a Hamiltonian system in equilibrium to dissipative systems in nonequilibrium stationary states. In this more general setting the attractor plays the role of a constant energy surface, and the SRB measure of sect. 9.1 is a generalization of the Liouville measure. Such measures are purely microscopic and indifferent to whether the system is at equilibrium, close to equilibrium or far from it. "Far for equilibrium" in this context refers to systems with large deviations from Maxwell's equilibrium velocity distribution. Furthermore, the theory of dynamical systems has yielded new sets of microscopic dynamics formulas for macroscopic observables such as diffusion constants and the pressure, to which we turn now.

We shall apply cycle expansions to the analysis of transport properties of chaotic systems.

The resulting formulas are exact; no probabilistic assumptions are made, and the all correlations are taken into account by the inclusion of cycles of all periods. The infinite extent systems for which the periodic orbit theory yields formulas for diffusion and other transport coefficients are spatially periodic, the global phase space being tiled with copies of a elementary cell. The motivation are physical problems such as beam defocusing in particle accelerators or chaotic behavior of passive tracers in 2-d rotating flows, problems which can be described as deterministic diffusion in periodic arrays.

In sect. 23.1 we derive the formulas for diffusion coefficients in a simple physical setting, the $2-d$ periodic Lorentz gas. This system, however, is

Figure 23.1: Deterministic diffusion in a finite horizon periodic Lorentz gas. (Courtesy of T. Schreiber)

not the best one to exemplify the theory, due to its complicated symbolic dynamics. Therefore we apply the theory first to diffusion induced by a 1-d maps in sect. 23.2.

### 23.1 Diffusion in periodic arrays

The 2-d Lorentz gas is an infinite scatterer array in which diffusion of a light molecule in a gas of heavy scatterers is modelled by the motion of a point particle in a plane bouncing off an array of reflecting disks. The Lorentz gas is called "gas" as one can equivalently think of it as consisting of any number of pointlike fast "light molecules" interacting only with the stationary "heavy molecules" and not among themselves. As the scatterer array is built up from only defocusing concave surfaces, it is a pure hyperbolic system, and one of the simplest nontrivial dynamical systems that exhibits deterministic diffusion, figure 23.1. We shall now show that the periodic Lorentz gas is amenable to a purely deterministic treatment. In this class of open dynamical systems quantities characterizing global dynamics, such as the Lyapunov exponent, pressure and diffusion constant, can be computed from the dynamics restricted to the elementary cell. The method applies to any hyperbolic dynamical system that is a periodic tiling $\hat{\mathcal{M}}=\bigcup_{\hat{n} \in T} \mathcal{M}_{\hat{n}}$ of the dynamical phase space $\hat{\mathcal{M}}$ by translates $\mathcal{M}_{\hat{n}}$ of an elementary cell $\mathcal{M}$, with $T$ the Abelian group of lattice translations. If the scattering array has further discrete symmetries, such as reflection symmetry, each elementary cell may be built from a fundamental domain $\widetilde{\mathcal{M}}$ by the action of a discrete (not necessarily Abelian) group $G$. The symbol $\hat{\mathcal{M}}$ refers here to the full phase space, that is,, both the spatial coordinates and the momenta. The spatial component of $\hat{\mathcal{M}}$ is the complement of the disks in the whole space.

We shall now relate the dynamics in $\mathcal{M}$ to diffusive properties of the Lorentz gas in $\hat{\mathcal{M}}$.

These concepts are best illustrated by a specific example, a Lorentz gas

Figure 23.2: Tiling of $\hat{\mathcal{M}}$, a periodic lattice of reflecting disks, by the fundamental domain $\widetilde{\mathcal{M}}$. Indicated is an example of a global trajectory $\hat{x}(t)$ together with the corresponding elementary cell trajectory $x(t)$ and the fundamental domain trajectory $\tilde{x}(t)$. (Courtesy of J.-P. Eckmann)

based on the hexagonal lattice Sinai billiard of figure 23.2. We distinguish two types of diffusive behavior; the infinite horizon case, which allows for infinite length flights, and the finite horizon case, where any free particle trajectory must hit a disk in finite time. In this chapter we shall restrict our consideration to the finite horizon case, with disks sufficiently large so that no infinite length free flight is possible. In this case the diffusion is normal, with $\hat{x}(t)^{2}$ growing like $t$. We shall return to the anomalous diffusion case in sect. 23.3.

As we will work with three kinds of phase spaces, good manners require that we repeat what hats, tildas and nothings atop symbols signify:
~ fundamental domain, triangle in figure 23.2
elementary cell, hexagon in figure 23.2

- full phase space, lattice in figure 23.2

It is convenient to define an evolution operator for each of the 3 cases of figure 23.2. $\hat{x}(t)=\hat{f}^{t}(\hat{x})$ denotes the point in the global space $\hat{\mathcal{M}}$ reached by the flow in time $t . x(t)=f^{t}\left(x_{0}\right)$ denotes the corresponding flow in the elementary cell; the two are related by

$$
\begin{equation*}
\hat{n}_{t}\left(x_{0}\right)=\hat{f}^{t}\left(x_{0}\right)-f^{t}\left(x_{0}\right) \in T, \tag{23.2}
\end{equation*}
$$

the translation of the endpoint of the global path into the elementary cell $\mathcal{M}$. The quantity $\tilde{x}(t)=\tilde{f}^{t}(\tilde{x})$ denotes the flow in the fundamental domain $\widetilde{\mathcal{M}} ; \tilde{f}^{t}(\tilde{x})$ is related to $f^{t}(\tilde{x})$ by a discrete symmetry $g \in G$ which maps $\tilde{x}(t) \in \widetilde{\mathcal{M}}$ to $x(t) \in \mathcal{M}$.

Fix a vector $\beta \in \mathbb{R}^{d}$, where $d$ is the dimension of the phase space. We will compute the diffusive properties of the Lorentz gas from the leading eigenvalue of the evolution operator (10.11)

$$
\begin{equation*}
s(\beta)=\lim _{t \rightarrow \infty} \frac{1}{t} \log \left\langle e^{\beta \cdot(\hat{x}(t)-x)}\right\rangle_{\mathcal{M}} \tag{23.3}
\end{equation*}
$$

where the average is over all initial points in the elementary cell, $x \in \mathcal{M}$.
If all odd derivatives vanish by symmetry, there is no drift and the second derivatives

$$
\left.\frac{\partial}{\partial \beta_{i}} \frac{\partial}{\partial \beta_{j}} s(\beta)\right|_{\beta=0}=\lim _{t \rightarrow \infty} \frac{1}{t}\left\langle(\hat{x}(t)-x)_{i}(\hat{x}(t)-x)_{j}\right\rangle_{\mathcal{M}}
$$

yield a (generally anisotropic) diffusion matrix. The spatial diffusion constant is then given by the Einstein relation (10.13)

$$
D=\left.\frac{1}{2 d} \sum_{i} \frac{\partial^{2}}{\partial \beta_{i}^{2}} s(\beta)\right|_{\beta=0}=\lim _{t \rightarrow \infty} \frac{1}{2 d t}\left\langle(\hat{q}(t)-q)^{2}\right\rangle_{\mathcal{M}}
$$

where the $i$ sum is restricted to the spatial components $q_{i}$ of the phase space vectors $x=(q, p)$, that is, if the dynamics is Hamiltonian to the number of the degrees of freedom.

We now turn to the connection between (23.3) and periodic orbits in the elementary cell. As the full $\hat{\mathcal{M}} \rightarrow \widetilde{\mathcal{M}}$ reduction is complicated by the nonabelian nature of $G$, we shall introduce the main ideas in the abelian $\hat{\mathcal{M}} \rightarrow \mathcal{M}$ context.

### 23.1.1 Reduction from $\hat{\mathcal{M}}$ to $\mathcal{M}$

The key idea follows from inspection of the relation

$$
\left\langle e^{\beta \cdot(\hat{x}(t)-x)}\right\rangle_{\mathcal{M}}=\frac{1}{|\mathcal{M}|} \int_{\substack{x \in \mathcal{M} \\ \hat{y} \in \mathcal{M}}} d x d \hat{y} e^{\beta \cdot(\hat{y}-x)} \delta\left(\hat{y}-\hat{f}^{t}(x)\right) .
$$

$|\mathcal{M}|=\int_{\mathcal{M}} d x$ is the volume of the elementary cell $\mathcal{M}$. As in sect. 10.2, we have used the identity $1=\int_{\mathcal{M}} d y \delta(y-\hat{x}(t))$ to motivate the introduction of the evolution operator $\mathcal{L}^{t}(\hat{y}, x)$. There is a unique lattice translation $\hat{n}$ such that $\hat{y}=y-\hat{n}$, with $y \in \mathcal{M}$, and $f^{t}(x)$ given by (23.2). The difference is a translation by a constant, and the Jacobian for changing integration from $d \hat{y}$ to $d y$ equals unity. Therefore, and this is the main point, translation invariance can be used to reduce this average to the elementary cell:

$$
\begin{equation*}
\left\langle e^{\beta \cdot(\hat{x}(t)-x)}\right\rangle_{\mathcal{M}}=\frac{1}{|\mathcal{M}|} \int_{x, y \in \mathcal{M}} d x d y e^{\beta \cdot\left(f^{t}(x)-x\right)} \delta\left(y-f^{t}(x)\right) . \tag{23.4}
\end{equation*}
$$

As this is a translation, the Jacobian is $\delta \hat{y} / \delta y=1$. In this way the global $\hat{f}^{t}(x)$ flow averages can be computed by following the flow $f^{t}\left(x_{0}\right)$ restricted to the elementary cell $\mathcal{M}$. The equation (23.4) suggests that we study the evolution operator

$$
\begin{equation*}
\mathcal{L}^{t}(y, x)=e^{\beta \cdot(\hat{x}(t)-x)} \delta\left(y-f^{t}(x)\right), \tag{23.5}
\end{equation*}
$$

where $\hat{x}(t)=\hat{f}^{t}(x) \in \hat{\mathcal{M}}$, but $x, f^{t}(x), y \in \mathcal{M}$. It is straightforward to check that this operator satisfies the semigroup property (10.25), $\int_{\mathcal{M}} d z \mathcal{L}^{t_{2}}(y, z) \mathcal{L}^{t_{1}}(z, x)=$ $\mathcal{L}^{t_{2}+t_{1}}(y, x)$. For $\beta=0$, the operator (23.5) is the Perron-Frobenius operator (9.10), with the leading eigenvalue $e^{s_{0}}=1$ because there is no escape from this system (this will lead to the flow conservation sum rule (19.11) later on).

The rest is old hat. The spectrum of $\mathcal{L}$ is evaluated by taking the trace

$$
\operatorname{tr} \mathcal{L}^{t}=\int_{\mathcal{M}} d x e^{\beta \cdot \hat{n}_{t}(x)} \delta(x-x(t))
$$

Here $\hat{n}_{t}(x)$ is the discrete lattice translation defined in (23.2). Two kinds of orbits periodic in the elementary cell contribute. A periodic orbit is called standing if it is also periodic orbit of the infinite phase space dynamics, $\hat{f}^{T_{p}}(x)=x$, and it is called running if it corresponds to a lattice translation in the dynamics on the infinite phase space, $\hat{f}^{T_{p}}(x)=x+\hat{n}_{p}$. In the theory of area-preserving maps such orbits are called accelerator modes, as the diffusion takes place along the momentum rather than the position coordinate. The travelled distance $\hat{n}_{p}=\hat{n}_{T_{p}}\left(x_{0}\right)$ is independent of the starting point $x_{0}$, as can be easily seen by continuing the path periodically in $\hat{\mathcal{M}}$.

The final result is the spectral determinant (15.6)

$$
\begin{equation*}
\operatorname{det}(s(\beta)-\mathcal{A})=\prod_{p} \exp \left(-\sum_{r=1}^{\infty} \frac{1}{r} \frac{e^{\left(\beta \cdot \hat{n}_{p}-s T_{p}\right) r}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|}\right) \tag{23.6}
\end{equation*}
$$

or the corresponding dynamical zeta function (15.15)

$$
\begin{equation*}
1 / \zeta(\beta, s)=\prod_{p}\left(1-\frac{e^{\left(\beta \cdot \hat{n}_{p}-s T_{p}\right)}}{\left|\Lambda_{p}\right|}\right) . \tag{23.7}
\end{equation*}
$$

The dynamical zeta function cycle averaging formula (18.17) for the diffusion constant (10.13), zero mean drift $\left\langle\hat{x}_{i}\right\rangle=0$, is given by

$$
\begin{equation*}
D=\frac{1}{2 d} \frac{\left\langle\hat{x}^{2}\right\rangle_{\zeta}}{\langle T\rangle_{\zeta}}=\frac{1}{2 d} \frac{1}{\langle T\rangle_{\zeta}} \sum^{\prime} \frac{(-1)^{k+1}\left(\hat{n}_{p_{1}}+\cdots+\hat{n}_{p_{k}}\right)^{2}}{\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right|} \tag{23.8}
\end{equation*}
$$

where the sum is over all distinct non-repeating combination of prime cycles. The derivation is standard, still the formula is strange. Diffusion is unbounded motion accross an infinite lattice; nevertheless, the reduction to the elementary cell enables us to compute relevant quantities in the usual way, in terms of periodic orbits.

A sleepy reader might protest that $x_{p}=x\left(T_{p}\right)-x(0)$ is manifestly equal to zero for a periodic orbit. That is correct; $\hat{n}_{p}$ in the above formula refers
to a displacement on the infinite periodic lattice, while $p$ refers to closed orbit of the dynamics reduced to the elementary cell, with $x_{p}$ belonging to the closed prime cycle $p$.

Even so, this is not an obvious formula. Globally periodic orbits have $\hat{x}_{p}^{2}=0$, and contribute only to the time normalization $\langle T\rangle_{\zeta}$. The mean square displacement $\left\langle\hat{x}^{2}\right\rangle_{\zeta}$ gets contributions only from the periodic runaway trajectories; they are closed in the elementary cell, but on the periodic lattice each one grows like $\hat{x}(t)^{2}=\left(\hat{n}_{p} / T_{p}\right)^{2}=v_{p}^{2} t^{2}$. So the orbits that contribute to the trace formulas and spectral determinants exhibit either ballistic transport or no transport at all: diffusion arises as a balance between the two kinds of motion, weighted by the $1 /\left|\Lambda_{p}\right|$ measure. If the system is not hyperbolic such weights may be abnormally large, with $1 /\left|\Lambda_{p}\right| \approx 1 / T_{p}{ }^{\alpha}$ rather than $1 /\left|\Lambda_{p}\right| \approx e^{-T_{p} \lambda}$, where $\lambda$ is the Lyapunov exponent, and they may lead to anomalous diffusion - accelerated or slowed down depending on whether the probabilities of the running or the standing orbits are enhanced.

We illustrate the main idea, tracking of a globally diffusing orbit by the associated confined orbit restricted to the elementary cell, with a class of simple 1-d dynamical systems where all transport coefficients can be evaluated analytically.

### 23.2 Diffusion induced by chains of 1- $d$ maps

In a typical deterministic diffusive process, trajectories originating from a given scatterer reach a finite set of neighboring scatterers in one bounce, and then the process is repeated. As was shown in chapter 11, the essential part of this process is the stretching along the unstable directions of the flow, and in the crudest approximation the dynamics can be modelled by $1-d$ expanding maps. This observation motivates introduction of a class of particularly simple $1-d$ systems, chains of piecewise linear maps.

We start by defining the map $\hat{f}$ on the unit interval as

$$
\hat{f}(\hat{x})=\left\{\begin{array}{ll}
\Lambda \hat{x} & \hat{x} \in[0,1 / 2)  \tag{23.9}\\
\Lambda \hat{x}+1-\Lambda & \hat{x} \in(1 / 2,1]
\end{array}, \quad \Lambda>2,\right.
$$

and then extending the dynamics to the entire real line, by imposing the translation property

$$
\begin{equation*}
\hat{f}(\hat{x}+\hat{n})=\hat{f}(\hat{x})+\hat{n} \quad \hat{n} \in \mathbb{Z} \tag{23.10}
\end{equation*}
$$

As the map is dicontinuous at $\hat{x}=1 / 2, \hat{f}(1 / 2)$ is undefined, and the $x=$ $1 / 2$ point has to be excluded from the Markov partition. The map is antisymmetric under the $\hat{x}$-coordinate flip

$$
\begin{equation*}
\hat{f}(\hat{x})=-\hat{f}(-\hat{x}), \tag{23.11}
\end{equation*}
$$



Figure 23.3: (a) $\hat{f}(\hat{x})$, the full space sawtooth map (23.9), $\Lambda>2$. (b) $f(x)$, the sawtooth map restricted to the unit circle (23.12), $\Lambda=6$.
so the dynamics will exhibit no mean drift; all odd derivatives of the generating function (10.11) with respect to $\beta$, evaluated at $\beta=0$, will vanish.

The map (23.9) is sketched in figure 23.3(a). Initial points sufficiently close to either of the fixed points in the initial unit interval remain in the elementary cell for one iteration; depending on the slope $\Lambda$, other points jump $\hat{n}$ cells, either to the right or to the left. Repetition of this process generates a random walk for almost every initial condition.

The translational symmetry (23.10) relates the unbounded dynamics on the real line to dynamics restricted to the elementary cell - in the example at hand, the unit interval curled up into a circle. Associated to $\hat{f}(\hat{x})$ we thus also consider the circle map

$$
\begin{equation*}
f(x)=\hat{f}(\hat{x})-[\hat{f}(\hat{x})], \quad x=\hat{x}-[\hat{x}] \in[0,1] \tag{23.12}
\end{equation*}
$$

figure 23.3(b), where $[\cdots]$ stands for the integer part (23.2). As noted above, the elementary cell cycles correspond to either standing or running orbits for the map on the full line: we shall refer to $\hat{n}_{p} \in \mathbb{Z}$ as the jumping number of the $p$ cycle, and take as the cycle weight

$$
\begin{equation*}
t_{p}=z^{n_{p}} e^{\beta \hat{n}_{p}} /\left|\Lambda_{p}\right| . \tag{23.13}
\end{equation*}
$$

For the piecewise linear map of figure 23.3 we can evaluate the dynamical zeta function in closed form. Each branch has the same value of the slope, and the map can be parametrized by a single parameter, for example its critical value $a=\hat{f}(1 / 2)$, the absolute maximum on the interval $[0,1]$ related to the slope of the map by $a=\Lambda / 2$. The larger $\Lambda$ is, the stronger is the stretching action of the map.

The diffusion constant formula (23.8) for 1-d maps is

$$
\begin{equation*}
D=\frac{1}{2} \frac{\left\langle\hat{n}^{2}\right\rangle_{\zeta}}{\langle n\rangle_{\zeta}} \tag{23.14}
\end{equation*}
$$

where the "mean cycle time" is given by (18.18)

$$
\begin{equation*}
\langle n\rangle_{\zeta}=\left.z \frac{\partial}{\partial z} \frac{1}{\zeta(0, z)}\right|_{z=1}=-\sum^{\prime}(-1)^{k} \frac{n_{p_{1}}+\cdots+n_{p_{k}}}{\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right|}, \tag{23.15}
\end{equation*}
$$

and the "mean cycle displacement squared" by (20.1)

$$
\begin{equation*}
\left\langle\hat{n}^{2}\right\rangle_{\zeta}=\left.\frac{\partial^{2}}{\partial \beta^{2}} \frac{1}{\zeta(\beta, 1)}\right|_{\beta=0}=-\sum^{\prime}(-1)^{k} \frac{\left(\hat{n}_{p_{1}}+\cdots+\hat{n}_{p_{k}}\right)^{2}}{\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right|}, \tag{23.16}
\end{equation*}
$$

the primed sum indicating all distinct non-repeating combinations of prime cycles. The evaluation of these formulas in this simple system will require nothing more than pencil and paper.

### 23.2.1 Case of unrestricted symbolic dynamics

Whenever $\Lambda$ is an integer number, the symbolic dynamics is exceedingly simple. For example, for the case $\Lambda=6$ illustrated in figure 23.3(b), the elementary cell map consists of 6 full branches, with uniform stretching factor $\Lambda=6$. The branches have different jumping numbers: for branches 1 and 2 we have $\hat{n}=0$, for branch 3 we have $\hat{n}=+1$, for branch $4 \hat{n}=-1$, and finally for branches 5 and 6 we have respectively $\hat{n}=+2$ and $\hat{n}=-2$. The same structure reappears whenever $\Lambda$ is an even integer $\Lambda=2 a$ : all branches are mapped onto the whole unit interval and we have two $\hat{n}=0$ branches, one branch for which $\hat{n}=+1$ and one for which $\hat{n}=-1$, and so on, up to the maximal jump $|\hat{n}|=$ $a-1$. The symbolic dynamics is thus full, unrestricted shift in $2 a$ symbols $\left\{0_{+}, 1_{+}, \ldots,(a-1)_{+},(a-1)_{-}, \ldots, 1_{-}, 0_{-}\right\}$, where the symbol indicates both the length and the direction of the corresponding jump.

For the piecewise linear maps with uniform stretching the weight associated with a given symbol sequence is a product of weights for individual steps, $t_{s q}=t_{s} t_{q}$. For the map of figure 23.3 there are 6 distinct weigths (23.13):

$$
\begin{aligned}
& t_{1}=t_{2}=z / \Lambda \\
& t_{3}=e^{\beta} z / \Lambda, \quad t_{4}=e^{-\beta} z / \Lambda, \quad t_{5}=e^{2 \beta} z / \Lambda, \quad t_{6}=e^{-2 \beta} z / \Lambda
\end{aligned}
$$

The piecewise linearity and the simple symbolic dynamics lead to the full cancellation of all curvature corrections in (18.5). The exact dynamical zeta
function (13.13) is given by the fixed point contributions:

$$
\begin{align*}
1 / \zeta(\beta, z) & =1-t_{0_{+}}-t_{0-}-\cdots-t_{(a-1)_{+}}-t_{(a-1)_{-}} \\
& =1-\frac{z}{a}\left(1+\sum_{j=1}^{a-1} \cosh (\beta j)\right) . \tag{23.17}
\end{align*}
$$

The leading (and only) eigenvalue of the evolution operator (23.5) is

$$
\begin{equation*}
s(\beta)=\log \left\{\frac{1}{a}\left(1+\sum_{j=1}^{a-1} \cosh (\beta j)\right)\right\}, \quad \Lambda=2 a, a \text { integer } . \tag{23.18}
\end{equation*}
$$

The flow conservation (19.11) sum rule is manifestly satisfied, so $s(0)=0$. The first derivative $s(0)^{\prime}$ vanishes as well by the left/right symmetry of the dynamics, implying vanishing mean drift $\langle\hat{x}\rangle=0$. The second derivative $s(\beta)^{\prime \prime}$ yields the diffusion constant (23.14):

$$
\langle n\rangle_{\zeta}=2 a \frac{1}{\Lambda}=1, \quad\left\langle\hat{x}^{2}\right\rangle_{\zeta}=2 \frac{0^{2}}{\Lambda}+2 \frac{1^{2}}{\Lambda}+2 \frac{2^{2}}{\Lambda}+\cdots+2 \frac{(a-1)^{2}}{\Lambda}(23.19)
$$

Using the identity $\sum_{k=1}^{n} k^{2}=n(n+1)(2 n+1) / 6$ we obtain

$$
\begin{equation*}
D=\frac{1}{24}(\Lambda-1)(\Lambda-2), \quad \Lambda \text { even integer } . \tag{23.20}
\end{equation*}
$$

Similar calculation for odd integer $\Lambda=2 k-1$ yields

$$
\begin{equation*}
D=\frac{1}{24}\left(\Lambda^{2}-1\right), \quad \Lambda \text { odd integer } \tag{23.21}
\end{equation*}
$$

### 23.2.2 Higher order transport coefficients

The same approach yields higher order transport coefficients

$$
\begin{equation*}
\mathcal{B}_{k}=\left.\frac{1}{k!} \frac{d^{k}}{d \beta^{k}} s(\beta)\right|_{\beta=0}, \quad \mathcal{B}_{2}=D \tag{23.22}
\end{equation*}
$$

known for $k>2$ as the Burnett coefficients. The behavior of the higher order coefficients yields information on the relaxation to the asymptotic distribution function generated by the diffusive process. Here $\hat{x}_{t}$ is the relevant dynamical variable and $\mathcal{B}_{k}$ 's are related to moments $\left\langle\hat{x}_{t}^{k}\right\rangle$ of arbitrary order.


Figure 23.4: (a) A partition of the unit interval into six intervals, labeled by the jumping number $\hat{n}(x) I=\left\{0_{+}, 1_{+}, 2_{+}, 2_{-}, 1_{-}, 0_{-}\right\}$. The partition is Markov, as the critical point is mapped onto the right border of $\mathcal{M}_{1_{+}}$. (b) The Markov graph for this partition. (c) The Markov graph in the compact notation of (23.26) (introduced by Vadim Moroz).

Were the diffusive process purely Gaussian

$$
\begin{equation*}
e^{t s(\beta)}=\frac{1}{\sqrt{4 \pi D t}} \int_{-\infty}^{+\infty} d \hat{x} e^{\beta \hat{x}} e^{-\hat{x}^{2} /(4 D t)}=e^{\beta^{2} D t} \tag{23.23}
\end{equation*}
$$

the only $\mathcal{B}_{k}$ coefficient different from zero would be $\mathcal{B}_{2}=D$. Hence, nonvanishing higher order coefficients signal deviations of deterministic diffusion from a Gaussian stochastic process.

For the map under consideration the first Burnett coefficient coefficient $\mathcal{B}_{4}$ is easily evaluated. For example, using (23.18) in the case of even integer slope $\Lambda=2 a$ we obtain

$$
\begin{equation*}
\mathcal{B}_{4}=-\frac{1}{4!\cdot 60}(a-1)(2 a-1)\left(4 a^{2}-9 a+7\right) . \tag{23.24}
\end{equation*}
$$

We see that deterministic diffusion is not a Gaussian stochastic process. Higher order even coefficients may be calculated along the same lines.

### 23.2.3 Case of finite Markov partitions

For piecewise-linear maps exact results may be obtained whenever the critical points are mapped in finite numbers of iterations onto partition boundary points, or onto unstable periodic orbits. We will work out here an example for which this occurs in two iterations, leaving other cases as exercises.

The key idea is to construct a Markov partition (11.4), with intervals mapped onto unions of intervals. As an example we determine a value of the parameter $4 \leq \Lambda \leq 6$ for which $f(f(1 / 2))=0$. As in the integer $\Lambda$ case, we
partition the unit interval into six intervals, labeled by the jumping number $\hat{n}(x) \in\left\{\mathcal{M}_{0_{+}}, \mathcal{M}_{1_{+}}, \mathcal{M}_{2_{+}}, \mathcal{M}_{2_{-}}, \mathcal{M}_{1_{-}}, \mathcal{M}_{0_{-}}\right\}$, ordered by their placement along the unit interval, figure 23.4(a).

In general the critical value $a=\hat{f}(1 / 2)$ will not correspond to an interval border, but now we choose $a$ such that the critical point is mapped onto the right border of $\mathcal{M}_{1_{+}}$. Equating $f(1 / 2)$ with the right border of $\mathcal{M}_{1_{+}}, x=1 / \Lambda$, we obtain a quadratic equation with the expanding solution $\Lambda=2(\sqrt{2}+1)$. For this parameter value $f\left(\mathcal{M}_{1_{+}}\right)=\mathcal{M}_{0_{+}} \cup \mathcal{M}_{1_{+}}$, $f\left(\mathcal{M}_{2_{-}}\right)=\mathcal{M}_{0_{-}} \cup \mathcal{M}_{1_{-}}$, while the remaining intervals map onto the whole unit interval $\mathcal{M}$. The transition matrix (11.2) is given by

$$
\phi^{\prime}=T \phi=\left(\begin{array}{cccccc}
1 & 1 & 1 & 0 & 1 & 1  \tag{23.25}\\
1 & 1 & 1 & 0 & 1 & 1 \\
1 & 1 & 0 & 0 & 1 & 1 \\
1 & 1 & 0 & 0 & 1 & 1 \\
1 & 1 & 0 & 1 & 1 & 1 \\
1 & 1 & 0 & 1 & 1 & 1
\end{array}\right)\left(\begin{array}{l}
\phi_{0_{+}} \\
\phi_{1_{+}} \\
\phi_{2_{+}} \\
\phi_{2_{-}} \\
\phi_{1_{-}} \\
\phi_{0_{-}}
\end{array}\right)
$$

One could diagonalize (23.25) on a computer, but, as we saw in sect. 11.5, the Markov graph figure 23.4(b) corresponding to figure 23.4(a) offers more insight into the dynamics. The graph figure 23.4(b) can be redrawn more compactly as Markov graph figure 23.4(c) by replacing parallel lines in a graph by their sum

The dynamics is unrestricted in the alphabet

$$
\mathcal{A}=\left\{0_{+}, 1_{+}, 2_{+} 0_{+}, 2_{+} 1_{+}, 2_{-} 1_{-}, 2_{-} 0_{-}, 1_{-}, 0_{-}\right\} .
$$

Applying the loop expansion (13.13) of sect. 13.3, we are led to the dynamical zeta function

$$
\begin{aligned}
1 / \zeta(\beta, z) & =1-t_{0_{+}}-t_{1_{+}}-t_{2_{+} 0_{+}}-t_{2_{+} 1_{+}}-t_{2_{-} 1_{-}}-t_{2_{-} 0_{-}}-t_{1_{-}}-t_{0_{-}} \\
& =1-\frac{2 z}{\Lambda}(1+\cosh (\beta))-\frac{2 z^{2}}{\Lambda^{2}}(\cosh (2 \beta)+\cosh (3 \beta))(23.27)
\end{aligned}
$$

For grammar as simple as this one, the dynamical zeta function is the sum over fixed points of the unrestricted alphabet. As the first check of this expression for the dynamical zeta function we verify that

$$
1 / \zeta(0,1)=1-\frac{4}{\Lambda}-\frac{4}{\Lambda^{2}}=0,
$$

as required by the flow conservation (19.11). Conversely, we could have started by picking the desired Markov partition, writing down the corresponding dynamical zeta function, and then fixing $\Lambda$ by the $1 / \zeta(0,1)=0$
condition. For more complicated Markov graphs this approach, together with the factorization (23.35), is helpful in reducing the order of the polynomial condition that fixes $\Lambda$.

The diffusion constant follows from (23.14)

$$
\begin{align*}
\langle n\rangle_{\zeta} & =4 \frac{1}{\Lambda}+4 \frac{2}{\Lambda^{2}}, \quad\left\langle\hat{n}^{2}\right\rangle_{\zeta}=2 \frac{1^{2}}{\Lambda}+2 \frac{2^{2}}{\Lambda^{2}}+2 \frac{3^{2}}{\Lambda^{2}} \\
D & =\frac{15+2 \sqrt{2}}{16+8 \sqrt{2}} . \tag{23.28}
\end{align*}
$$

It is by now clear how to build an infinite hierarchy of finite Markov partitions: tune the slope in such a way that the critical value $f(1 / 2)$ is mapped into the fixed point at the origin in a finite number of iterations $p$ $f^{P}(1 / 2)=0$. By taking higher and higher values of $p$ one constructs a dense set of Markov parameter values, organized into a hierarchy that resembles the way in which rationals are densely embedded in the unit interval. For example, each of the 6 primary intervals can be subdivided into 6 intervals obtained by the 2-nd iterate of the map, and for the critical point mapping into any of those in 2 steps the grammar (and the corresponding cycle expansion) is finite. So, if we can prove continuity of $D=D(\Lambda)$, we can apply the periodic orbit theory to the sawtooth map (23.9) for a random "generic" value of the parameter $\Lambda$, for example $\Lambda=4.5$. The idea is to bracket this value of $\Lambda$ by a sequence of nearby Markov values, compute the exact diffusion constant for each such Markov partition, and study their convergence toward the value of $D$ for $\Lambda=4.5$. Judging how difficult such problem is already for a tent map (see sect. 13.6 and appendix E.1), this is not likely to take only a week of work.

Expressions like (23.20) may lead to an expectation that the diffusion coefficient (and thus transport properties) are smooth functions of parameters controling the chaoticity of the system. For example, one might expect that the diffusion coefficient increases smoothly and monotonically as the slope $\Lambda$ of the map (23.9) is increased, or, perhaps more physically, that the diffusion coefficient is a smooth function of the Lyapunov exponent $\lambda$. This turns out not to be true: $D$ as a function of $\Lambda$ is a fractal, nowhere differentiable curve illustrated in figure 23.5. The dependence of $D$ on the map parameter $\Lambda$ is rather unexpected - even though for larger $\Lambda$ more points are mapped outside the unit cell in one iteration, the diffusion constant does not necessarily grow.

This is a consequence of the lack of structural stability, even of purely hyperbolic systems such as the Lozi map and the 1-d diffusion map (23.9). The trouble arises due to non-smooth dependence of the topological entropy on system parameters - any parameter change, no mater how small, leads to creation and destruction of ininitely many periodic orbits. As far as diffusion is concerned this means that even though local expansion rate is a smooth function of $\Lambda$, the number of ways in which the trajectory can re-enter the the initial cell is an irregular function of $\Lambda$.


Figure 23.5: The dependence of $D$ on the map parameter $a$ is continuous, but not monotone. (From ref. [23.7]). Here $a$ stands for the slope $\Lambda$ in (23.9).

The lesson is that lack of structural stabily implies lack of spectral stability, and no global observable is expected to depend smoothly on the system parameters. If you want to master the material, working through the project 0.1 and/or project O .2 is strongly recommended.

### 23.3 Marginal stability and anomalous diffusion

What effect does the intermittency of chapter 21 have on transport properties of $1-d$ maps? Consider a $1-d$ map of the real line on itself with the same properties as in sect. 23.2, except for a marginal fixed point at $x=0$.

A marginal fixed point unbalances the role of running and standing orbits, thus generating a mechanism that may result in anomalous diffusion.


Figure 23.6: (a) A map with marginal fixed point. (b) The map restricted to the unit circle.
[穿 remark 23.8
where $\alpha=1 / s$ is determined by the intermittency exponent (21.1), while $q$ is to be determined by the flow conservation (19.11) for $\hat{f}$ : -PCdefine $R$

$$
\frac{4}{\Lambda}+2 q \zeta(\alpha+1)=1
$$

so that $q=(\Lambda-4) / 2 \Lambda \zeta(\alpha+1)$. The dynamical zeta function picks up contributions just by the alphabet's letters, as we have imposed piecewise linearity, and can be expressed in terms of a Jonguiere function (23.29):

$$
\begin{equation*}
1 / \zeta_{0}(z, \beta)=1-\frac{4}{\Lambda} z \cosh \beta-\frac{\Lambda-4}{\Lambda \zeta(1+\alpha)} z \cosh \beta \cdot J(z, \alpha+1) . \tag{23.31}
\end{equation*}
$$

Its first zero $z(\beta)$ is determined by

$$
\frac{4}{\Lambda} z+\frac{\Lambda-4}{\Lambda \zeta(1+\alpha)} z \cdot J(z, \alpha+1)=\frac{1}{\cosh \beta} .
$$

By using implicit function derivation we see that $D$ vanishes (that is, $\left.z^{\prime \prime}(\beta)\right|_{\beta=1}=0$ ) when $\alpha \leq 1$. The physical interpretion is that a typical orbit will stick for long times near the $\overline{0}$ marginal fixed point, and the 'trapping time' will be larger for higher values of the intermittency parameter $s$ (recall $\alpha=s^{-1}$ ). Hence, we need to look more closely at the behavior of traces of high powers of the transfer operator.

The evaluation of transport coefficient requires one more derivative with respect to expectation values of phase functions (see sect. 23.1): if we use the diffusion dynamical zeta function (23.7), we may write the diffusion coefficient as an inverse Laplace transform, in such a way that any distinction between maps and flows has vanished. In the case of 1- $d$ diffusion we thus have

$$
\begin{equation*}
D=\left.\lim _{t \rightarrow \infty} \frac{d^{2}}{d \beta^{2}} \frac{1}{2 \pi i} \int_{a-i \infty}^{a+i \infty} d s e^{s t} \frac{\zeta^{\prime}(\beta, s)}{\zeta(\beta, s)}\right|_{\beta=0} \tag{23.32}
\end{equation*}
$$

where the $\zeta^{\prime}$ refers to the derivative with respect to $s$.
The evaluation of inverse Laplace transforms for high values of the argument is most conveniently performed using Tauberian theorems. We shall take

$$
\omega(\lambda)=\int_{0}^{\infty} d x e^{-\lambda x} u(x)
$$

with $u(x)$ monotone as $x \rightarrow \infty$; then, as $\lambda \mapsto 0$ and $x \mapsto \infty$ respectively (and $\rho \in(0, \infty)$,

$$
\omega(\lambda) \sim \frac{1}{\lambda^{\rho}} L\left(\frac{1}{\lambda}\right)
$$

if and only if

$$
u(x) \sim \frac{1}{\Gamma(\rho)} x^{\rho-1} L(x)
$$

where $L$ denotes any slowly varying function with $\lim _{t \rightarrow \infty} L(t y) / L(t)=1$. Now

$$
\frac{1 / \zeta_{0}{ }^{\prime}\left(e^{-s}, \beta\right)}{1 / \zeta_{0}\left(e^{-s}, \beta\right)}=\frac{\left(\frac{4}{\Lambda}+\frac{\Lambda-4}{\Lambda \zeta(1+\alpha)}\left(J\left(e^{-s}, \alpha+1\right)+J\left(e^{-s}, \alpha\right)\right)\right) \cosh \beta}{1-\frac{4}{\Lambda} e^{-s} \cosh \beta-\frac{\Lambda-4}{\Lambda \zeta(1+\alpha)} e^{-s}\left(e^{-s}, \alpha+1\right) \cosh \beta J}
$$

We then take the double derivative with respect to $\beta$ and obtain

$$
\begin{align*}
\frac{d^{2}}{d \beta^{2}} & \left(1 / \zeta_{0}{ }^{\prime}\left(e^{-s}, \beta\right) / \zeta^{-1}\left(e^{-s}, \beta\right)\right)_{\beta=0} \\
& =\frac{\frac{4}{\Lambda}+\frac{\Lambda-4}{\Lambda \zeta(1+\alpha)}\left(J\left(e^{-s}, \alpha+1\right)+J\left(e^{-s}, \alpha\right)\right)}{\left(1-\frac{4}{\Lambda} e^{-s}-\frac{\Lambda-4}{\Lambda \zeta(1+\alpha)} e^{-s} J\left(e^{-s}, \alpha+1\right)\right)^{2}}=g_{\alpha}(s) \tag{23.33}
\end{align*}
$$

The asymptotic behavior of the inverse Laplace transform (23.32) may then be evaluated via Tauberian theorems, once we use our estimate for the behavior of Jonquière functions near $z=1$. The deviations from normal behavior correspond to an explicit dependence of $D$ on time. Omitting prefactors (which can be calculated by the same procedure) we have

$$
g_{\alpha}(s) \sim \begin{cases}s^{-2} & \text { for } \alpha>1 \\ s^{-(\alpha+1)} & \text { for } \alpha \in(0,1) \\ 1 /\left(s^{2} \ln s\right) & \text { for } \alpha=1\end{cases}
$$

The anomalous diffusion exponents follow:

$$
\left\langle\left(x-x_{0}\right)^{2}\right\rangle_{t} \sim \begin{cases}t & \text { for } \alpha>1  \tag{23.34}\\ t^{\alpha} & \text { for } \alpha \in(0,1) \\ t / \ln t & \text { for } \alpha=1\end{cases}
$$

## Commentary

Remark 23.1 Lorentz gas. The original pinball model proposed by Lorentz [23.3] consisted of randomly, rather than regularly placed scatterers.

Remark 23.2 Who's dun it? Cycle expansions for the diffusion constant of a particle moving in a periodic array have been introduced independently by R. Artuso [23.4] (exact dynamical zeta function for 1-d chains of maps (23.8)), by W.N. Vance [23.5], and by P. Cvitanović, J.-P. Eckmann, and P. Gaspard [23.6] (the dynamical zeta function cycle expansion (23.8) applied to the Lorentz gas).

Remark 23.3 Lack of structural stability for D. Expressions like (23.20) may lead to an expectation that the diffusion coefficient (and thus transport properties) are smooth functions of the chaoticity of the system (parametrized, for example, by the Lyapunov exponent $\lambda=\ln \Lambda$ ). This turns out not to be true: $D$ as a function of $\Lambda$ is a fractal, nowhere differentiable curve shown in figure 23.5. The dependence of $D$ on the map parameter $\Lambda$ is rather unexpected - even though for larger $\Lambda$ more points are mapped outside the unit cell in one iteration, the diffusion constant does not necessarily grow. The fractal dependence of diffusion constant on the map parameter is discussed in refs. [23.7, 23.8, 23.9]. Statistical mechanicians tend to believe that such complicated behavior is not to be expected in systems with very many degrees of freedom, as the addition to a large integer dimension of a number smaller than 1 should be as unnoticeable as a microscopic perturbation of a macroscopic quantity. No fractal-like behavior of the conductivity for the Lorentz gas has been detected so far [23.10].

Remark 23.4 Diffusion induced by 1- $d$ maps. We refer the reader to refs. [23.11, 23.12] for early work on the deterministic diffusion induced by 1-dimenional maps. The sawtooth map (23.9) was introduced by Grossmann and Fujisaka [23.13] who derived the integer slope formulas (23.20) for the diffusion constant. The sawtooth map is also discussed in refs. [23.14].

Remark 23.5 Symmetry factorization in one dimension. In the $\beta=0$ limit the dynamics (23.11) is symmetric under $x \rightarrow-x$, and the zeta functions factorize into products of zeta functions for the symmetric and antisymmetric subspaces, as described in sect. 22.1.2:

$$
\begin{equation*}
\frac{1}{\zeta(0, z)}=\frac{1}{\zeta_{s}(0, z)} \frac{1}{\zeta_{a}(0, z)}, \quad \frac{\partial}{\partial z} \frac{1}{\zeta}=\frac{1}{\zeta_{s}} \frac{\partial}{\partial z} \frac{1}{\zeta_{a}}+\frac{1}{\zeta_{a}} \frac{\partial}{\partial z} \frac{1}{\zeta_{s}} \tag{23.35}
\end{equation*}
$$

The leading (material flow conserving) eigenvalue $z=1$ belongs to the symmetric subspace $1 / \zeta_{s}(0,1)=0$, so the derivatives (23.15) also depend only on the symmetric subspace:

$$
\begin{equation*}
\langle n\rangle_{\zeta}=\left.z \frac{\partial}{\partial z} \frac{1}{\zeta(0, z)}\right|_{z=1}=\left.\frac{1}{\zeta_{a}(0, z)} z \frac{\partial}{\partial z} \frac{1}{\zeta_{s}(0, z)}\right|_{z=1} \tag{23.36}
\end{equation*}
$$

Implementing the symmetry factorization is convenient, but not essential, at this level of computation.

Remark 23.6 Lorentz gas in the fundamental domain. The vector valued nature of the generating function (23.3) in the case under consideration makes it difficult to perform a calculation of the diffusion constant within the fundamental domain. Yet we point out that, at least as regards scalar quantities, the full reduction to $\tilde{\mathcal{M}}$ leads to better estimates. A proper symbolic dynamics in the fundamental domain has been introduced in ref. [23.15], numerical estimates for scalar quantities are reported in table 23.1, taken from ref. [23.16].

In order to perform the full reduction for diffusion one should express the dynamical zeta function (23.7) in terms of the prime cycles of the fundamental domain

| length | \# cycles | $\zeta(0,0)$ | $\lambda$ |
| ---: | ---: | ---: | ---: |
| 1 | 5 | -1.216975 | - |
| 2 | 10 | -0.024823 | 1.745407 |
| 3 | 32 | -0.021694 | 1.719617 |
| 4 | 104 | 0.000329 | 1.743494 |
| 5 | 351 | 0.002527 | 1.760581 |
| 6 | 1243 | 0.000034 | 1.756546 |

Table 23.1: Fundamental domain, $w=0.3$.
$\tilde{\mathcal{M}}$ of the lattice (see figure 23.2 ) rather than those of the elementary (WignerSeitz) cell $\mathcal{M}$. This problem is complicated by the breaking of the rotational symmetry by the auxilliary vector $\beta$, or, in other words, the non-commutativity of translations and rotations: see ref. [23.6].

Remark 23.7 Anomalous diffusion. Anomalous diffusion for 1-d intermittent maps was studied in the continuous time random walk approach in refs. [21.10, 21.11]. The first approach within the framework of cycle expansions (based on truncated dynamical zeta functions) was proposed in ref. [21.12]. Our treatment follows methods introduced in ref. [21.13], applied there to investigate the behavior of the Lorentz gas with unbounded horizon.

Remark 23.8 Jonquière functions. In statistical mechanics Jonquière functions

$$
\begin{equation*}
J(z, s)=\sum_{k=1}^{\infty} z^{k} / k^{s} \tag{23.37}
\end{equation*}
$$

appear in the theory of free Bose-Einstein gas, see refs. [21.21, 21.22].

## Résumé

The classical Boltzmann equation for evolution of 1-particle density is based on stosszahlansatz, neglect of particle correlations prior to, or after a 2particle collision. It is a very good approximate description of dilute gas dynamics, but a difficult starting point for inclusion of systematic corrections. In the theory developed here, no correlations are neglected - they are all included in the cycle averaging formula such as the cycle expansion for the diffusion constant

$$
D=\frac{1}{2 d} \frac{1}{\langle T\rangle_{\zeta}} \sum^{\prime}(-1)^{k+1} \frac{\left(\hat{n}_{p}+\cdots\right)}{\left|\Lambda_{p} \cdots\right|} \frac{\left(\hat{n}_{p_{1}}+\cdots+\hat{n}_{p_{k}}\right)^{2}}{\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right|} .
$$

Such formulas are exact; the issue in their applications is what are the most effective schemes of estimating the infinite cycle sums required for their evaluation. Unlike most statistical mechanics, here there are no phenomenological macroscopic parameters; quantities such as transport coefficients are calculable to any desired accuracy from the microscopic dynamics.

Though superficially indistinguishable from the probabilistic random walk diffusion, deterministic diffusion is quite recognizable, at least in low dimensional settings, through fractal dependence of the diffusion constant on the system parameters, and through non-Gaussion relaxation to equilibrium (non-vanishing Burnett coefficients).

For systems of a few degrees of freedom these results are on rigorous footing, but there are indications that they capture the essential dynamics of systems of many degrees of freedom as well.

Actual evaluation of transport coefficients is a test of the techniques developped above in physical settings. In cases of severe pruning the trace formulas and ergodic sampling of dominant cycles might be more effective strategy than the cycle expansions of dynamical zeta functions and systematic enumeration of all cycles.

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## Exercises

Exercise 23.1 Diffusion for odd integer $\Lambda$. Show that when the slope $\Lambda=2 k-1$ in (23.9) is an odd integer, the diffusion constant is given by $D=\left(\Lambda^{2}-1\right) / 24$, as stated in (23.21).

Exercise 23.2 Fourth-order transport coefficient. Verify (23.24). You will need the identity

$$
\sum_{k=1}^{n} k^{4}=\frac{1}{30} n(n+1)(2 n+1)\left(3 n^{2}+3 n-1\right) .
$$

## Exercise 23.3 Finite Markov partitions. Verify (23.28).

## Exercise 23.4 Maps with variable peak shape:

Consider the following piecewise linear map

$$
f_{\delta}(x)= \begin{cases}\frac{3 x}{1-\delta} & \text { for } x \in\left[0, \frac{1}{3}(1-\delta)\right]  \tag{23.38}\\ \frac{3}{2}-\left(\frac{2}{\delta}\left|\frac{4-\delta}{12}-x\right|\right) & \text { for } x \in\left[\frac{1}{3}(1-\delta), \frac{1}{6}(2+\delta)\right] \\ 1-\frac{3}{1-\delta}\left(x-\frac{1}{6}(2+\delta)\right) & \text { for } x \in\left[\frac{1}{6}(2+\delta), \frac{1}{2}\right]\end{cases}
$$

and the map in $[1 / 2,1]$ is obtained by antisymmetry with respect to $x=1 / 2, y=1 / 2$. Write the corresponding dynamical zeta function relevant to diffusion and then show that

$$
D=\frac{\delta(2+\delta)}{4(1-\delta)}
$$

See refs. [23.17, 23.18] for further details.

Exercise 23.5 Two-symbol cycles for the Lorentz gas. Write down all cycles labeled by two symbols, such as (06), (17), (15) and (05).

Appendix O contains several project-length deterministic diffusion exercises.

Exercise 23.6 Accelerated diffusion. Consider a map $h$, such that $\hat{h}=\hat{f}$ of figure 23.6(b), but now running branches are turner into standing branches and vice versa, so that $1,2,3,4$ are standing while 0 leads to both positive and negative jumps. Build the corresponding dynamical zeta function and show that

$$
\sigma^{2}(t) \sim \begin{cases}t & \text { for } \alpha>2 \\ t \ln t & \text { for } \alpha=2 \\ t^{3-\alpha} & \text { for } \alpha \in(1,2) \\ t^{2} / \ln t & \text { for } \alpha=1 \\ t^{2} & \text { for } \alpha \in(0,1)\end{cases}
$$

Exercise 23.7 Recurrence times for Lorentz gas with infinite horizon. Consider the Lorentz gas with unbounded horizon with a square lattice geometry, with disk radius $R$ and unit lattice spacing. Label disks according to the (integer) coordinates of their center: the sequence of recurrence times $\left\{t_{j}\right\}$ is given by the set of collision times. Consider orbits that leave the disk sitting at the origin and hit a disk far away after a free flight (along the horizontal corridor). Initial conditions are characterized by coordinates $(\phi, \alpha)$ ( $\phi$ determines the initial position along the disk, while $\alpha$ gives the angle of the initial velocity with respect to the outward normal: the appropriate measure is then $d \phi \cos \alpha d \alpha(\phi \in[0,2 \pi), \alpha \in[-\pi / 2, \pi / 2]$. Find how $\psi(T)$ scales for large values of $T$ : this is equivalent to investigating the scaling of portions of the phase space that lead to a first collision with disk $(n, 1)$, for large values of $n$ (as $n \mapsto \infty n \simeq T)$.

## Suggested steps

(a) Show that the condition assuring that a trajectory indexed by $(\phi, \alpha)$ hits the ( $m, n$ ) disk (all other disks being transparent) is written as

$$
\begin{equation*}
\left|\frac{d_{m, n}}{R} \sin \left(\phi-\alpha-\theta_{m, n}\right)+\sin \alpha\right| \leq 1 \tag{23.39}
\end{equation*}
$$

where $d_{m, n}=\sqrt{m^{2}+n^{2}}$ and $\theta_{m, n}=\arctan (n / m)$. You can then use a small $R$ expansion of (23.39).
(b) Now call $j_{n}$ the portion of the phase space leading to a first collision with disk $(n, 1)$ (take into account screening by disks $(1,0)$ or $(n-1,1)$ ). Denote by $J_{n}=\bigcup_{k=n+1}^{\infty} j_{k}$ and show that $J_{n} \sim 1 / n^{2}$, from which the result for the distribution function follows.

## Chapter 24

## Irrationally winding

I don't care for islands, especially very small ones.
D.H. Lawrence

## (R. Artuso and P. Cvitanović)

This chapter is concerned with the mode locking problems for circle maps: besides its physical relevance it nicely illustrates the use of cycle expansions away from the dynamical setting, in the realm of renormalization theory at the transition to chaos.

The physical significance of circle maps is connected with their ability to model the two-frequencies mode-locking route to chaos for dissipative systems. In the context of dissipative dynamical systems one of the most common and experimentally well explored routes to chaos is the two-frequency mode-locking route. Interaction of pairs of frequencies is of deep theoretical interest due to the generality of this phenomenon; as the energy input into a dissipative dynamical system (for example, a Couette flow) is increased, typically first one and then two of intrinsic modes of the system are excited. After two Hopf bifurcations (a fixed point with inward spiralling stability has become unstable and outward spirals to a limit cycle) a system lives on a two-torus. Such systems tend to mode-lock: the system adjusts its internal frequencies slightly so that they fall in step and minimize the internal dissipation. In such case the ratio of the two frequencies is a rational number. An irrational frequency ratio corresponds to a quasiperiodic motion - a curve that never quite repeats itself. If the mode-locked states overlap, chaos sets in. The likelihood that a mode-locking occurs depends on the strength of the coupling of the two frequencies.

Our main concern in this chapter is to illustrate the "global" theory of circle maps, connected with universality properties of the whole irrational winding set. We shall see that critical global properties may be expressed via cycle expansions involving "local" renormalization critical exponents. The renormalization theory of critical circle maps demands rather tedious numerical computations, and our intuition is much facilitated by approximating circle maps by number-theoretic models. The models that arise in
this way are by no means mathematically trivial, they turn out to be related to number-theoretic abysses such as the Riemann conjecture, already in the context of the "trivial" models.

### 24.1 Mode locking

The simplest way of modeling a nonlinearly perturbed rotation on a circle is by 1 -dimensional circle maps $x \rightarrow x^{\prime}=f(x)$, restricted to the one dimensional torus, such as the sine map

$$
\begin{equation*}
x_{n+1}=f\left(x_{n}\right)=x_{n}+\Omega-\frac{k}{2 \pi} \sin \left(2 \pi x_{n}\right) \quad \bmod 1 . \tag{24.1}
\end{equation*}
$$

$f(x)$ is assumed to be continuous, have a continuous first derivative, and a continuous second derivative at the inflection point (where the second derivative vanishes). For the generic, physically relevant case (the only one considered here) the inflection is cubic. Here $k$ parametrizes the strength of the nonlinear interaction, and $\Omega$ is the bare frequency.

The phase space of this map, the unit interval, can be thought of as the elementary cell of the map

$$
\begin{equation*}
\hat{x}_{n+1}=\hat{f}\left(\hat{x}_{n}\right)=\hat{x}_{n}+\Omega-\frac{k}{2 \pi} \sin \left(2 \pi \hat{x}_{n}\right) . \tag{24.2}
\end{equation*}
$$

where ${ }^{\wedge}$ is used in the same sense as in chapter 23.
The winding number is defined as

$$
\begin{equation*}
W(k, \Omega)=\lim _{n \rightarrow \infty}\left(\hat{x}_{n}-\hat{x}_{0}\right) / n . \tag{24.3}
\end{equation*}
$$

and can be shown to be independent of the initial value $\hat{x}_{0}$.
For $k=0$, the map is a simple rotation (the shift map) see figure 24.1

$$
\begin{equation*}
x_{n+1}=x_{n}+\Omega \quad \bmod 1, \tag{24.4}
\end{equation*}
$$

and the rotation number is given by the parameter $\Omega$.

$$
W(k=0, \Omega)=\Omega .
$$

For given values of $\Omega$ and $k$ the winding number can be either rational or irrational. For invertible maps and rational winding numbers $W=P / Q$


Figure 24.1: Unperturbed circle map ( $k=0$ in (24.1)) with golden mean rotation number.
the asymptotic iterates of the map converge to a unique attractor, a stable periodic orbit of period $Q$

$$
\hat{f}^{Q}\left(\hat{x}_{i}\right)=\hat{x}_{i}+P, \quad i=0,1,2, \cdots, Q-1 .
$$

This is a consequence of the independence of $\hat{x}_{0}$ previously mentioned. There is also an unstable cycle, repelling the trajectory. For any rational winding number, there is a finite interval of values of $\Omega$ values for which the iterates of the circle map are attracted to the $P / Q$ cycle. This interval is called the $P / Q$ mode-locked (or stability) interval, and its width is given by

$$
\begin{equation*}
\Delta_{P / Q}=Q^{-2 \mu_{P / Q}}=\Omega_{P / Q}^{\text {right }}-\Omega_{P / Q}^{\text {left }} . \tag{24.5}
\end{equation*}
$$

where $\Omega_{P / Q}^{\text {right }}\left(\Omega_{P / Q}^{\text {left }}\right)$ denote the biggest (smallest) value of $\Omega$ for which $W(k, \Omega)=P / Q$. Parametrizing mode lockings by the exponent $\mu$ rather than the width $\Delta$ will be convenient for description of the distribution of the mode-locking widths, as the exponents $\mu$ turn out to be of bounded variation. The stability of the $P / Q$ cycle is

$$
\Lambda_{P / Q}=\frac{\partial x_{Q}}{\partial x_{0}}=f^{\prime}\left(x_{0}\right) f^{\prime}\left(x_{1}\right) \cdots f^{\prime}\left(x_{Q-1}\right)
$$

For a stable cycle $\left|\Lambda_{P / Q}\right|$ lies between 0 (the superstable value, the "center" of the stability interval) and 1 (the $\Omega_{P / Q}^{\text {right }}, \Omega_{P / Q}^{\text {left }}$ endpoints of (24.5)).


Figure 24.2: The critical circle map ( $k=1$ in (24.1)) devil's staircase [24.3]; the winding number $W$ as function of the parameter $\Omega$.

For the shift map (24.4), the stability intervals are shrunk to points. As $\Omega$ is varied from 0 to 1 , the iterates of a circle map either mode-lock, with the winding number given by a rational number $P / Q \in(0,1)$, or do not mode-lock, in which case the winding number is irrational. A plot of the winding number $W$ as a function of the shift parameter $\Omega$ is a convenient visualization of the mode-locking structure of circle maps. It yields a monotonic "devil's staircase" of figure 24.2 whose self-similar structure we are to unravel. Circle maps with zero slope at the inflection point $x_{c}$ (see figure 24.3)

$$
f^{\prime}\left(x_{c}\right)=0, \quad f^{\prime \prime}\left(x_{c}\right)=0
$$

( $k=1, x_{c}=0$ in (24.1)) are called critical: they delineate the borderline of chaos in this scenario.

As the nonlinearity parameter $k$ increases, the mode-locked intervals become wider, and for the critical circle maps $(k=1)$ they fill out the whole interval. A critical map has a superstable $P / Q$ cycle for any rational $P / Q$, as the stability of any cycle that includes the inflection point equals zero. If the map is non-invertible ( $k>1$ ), it is called supercritical; the bifurcation structure of this regime is extremely rich and beyond the scope of this exposition.

The physically relevant transition to chaos is connected with the critical case, however the apparently simple "free" shift map limit is quite instructive: in essence it involves the problem of ordering rationals embedded in the unit interval on a hierarchical structure. From a physical point of view, the main problem is to identify a (number-theoretically) consistent hierarchy susceptible of experimental verification. We will now describe a few ways of organizing rationals along the unit interval: each has its own advantages as well as its drawbacks, when analyzed from both mathematical and physical perspective.


Figure 24.3: Critical circle map ( $k=1$ in (24.1)) with golden mean bare rotation number.

### 24.1.1 Hierarchical partitions of the rationals

Intuitively, the longer the cycle, the finer the tuning of the parameter $\Omega$ required to attain it; given finite time and resolution, we expect to be able to resolve cycles up to some maximal length $Q$. This is the physical motivation for partitioning mode lockings into sets of cycle length up to $Q$. In number theory such sets of rationals are called Farey series. They are denoted by $\mathcal{F}_{Q}$ and defined as follows. The Farey series of order $Q$ is the monotonically increasing sequence of all irreducible rationals between 0 and 1 whose denominators do not exceed $Q$. Thus $P_{i} / Q_{i}$ belongs to $\mathcal{F}_{Q}$ if $0<P_{i} \leq Q_{i} \leq Q$ and $\left(P_{i} \mid Q_{i}\right)=1$. For example

$$
\mathcal{F}_{5}=\left\{\frac{1}{5}, \frac{1}{4}, \frac{1}{3}, \frac{2}{5}, \frac{1}{2}, \frac{3}{5}, \frac{2}{3}, \frac{3}{4}, \frac{4}{5}, \frac{1}{1}\right\}
$$

A Farey series is characterized by the property that if $P_{i-1} / Q_{i-1}$ and $P_{i} / Q_{i}$ are consecutive terms of $\mathcal{F}_{Q}$, then

$$
P_{i} Q_{i-1}-P_{i-1} Q_{i}=1
$$

The number of terms in the Farey series $F_{Q}$ is given by

$$
\begin{equation*}
\Phi(Q)=\sum_{n=1}^{Q} \phi(Q)=\frac{3 Q^{2}}{\pi^{2}}+O(Q \ln Q) \tag{24.6}
\end{equation*}
$$

Here the Euler function $\phi(Q)$ is the number of integers not exceeding and relatively prime to $Q$. For example, $\phi(1)=1, \phi(2)=1, \phi(3)=$ $2, \ldots, \phi(12)=4, \phi(13)=12, \ldots$

From a number-theorist's point of view, the continued fraction partitioning of the unit interval is the most venerable organization of rationals, preferred already by Gauss. The continued fraction partitioning is obtained by ordering rationals corresponding to continued fractions of increasing length. If we turn this ordering into a way of covering the complementary set to mode-lockings in a circle map, then the first level is obtained by deleting $\Delta_{[1]}, \Delta_{[2]}, \cdots, \Delta_{\left[a_{1}\right]}, \cdots$ mode-lockings; their complement are the covering intervals $\ell_{1}, \ell_{2}, \ldots, \ell_{a_{1}}, \ldots$ which contain all windings, rational and irrational, whose continued fraction expansion starts with $\left[a_{1}, \ldots\right]$ and is of length at least 2 . The second level is obtained by deleting $\Delta_{[1,2]}, \Delta_{[1,3]}, \cdots, \Delta_{[2,2]}, \Delta_{[2,3]}, \cdots, \Delta_{[n, m]}, \cdots$ and so on.

The $n$th level continued fraction partition $\mathcal{S}_{n}=\left\{a_{1} a_{2} \cdots a_{n}\right\}$ is defined as the monotonically increasing sequence of all rationals $P_{i} / Q_{i}$ between 0 and 1 whose continued fraction expansion is of length n :

$$
\frac{P_{i}}{Q_{i}}=\left[a_{1}, a_{2}, \cdots, a_{n}\right]=\frac{1}{a_{1}+\frac{1}{a_{2}+\ldots \frac{1}{a_{n}}}}
$$

The object of interest, the set of the irrational winding numbers, is in this partitioning labeled by $\mathcal{S}_{\infty}=\left\{a_{1} a_{2} a_{3} \cdots\right\}, a_{k} \in Z^{+}$, that is, the set of winding numbers with infinite continued fraction expansions. The continued fraction labeling is particularly appealing in the present context because of the close connection of the Gauss shift to the renormalization transformation $R$, discussed below. The Gauss map

$$
\begin{array}{cc}
T(x)= & \frac{1}{x}-\left[\frac{1}{x}\right] \\
0, & x \neq 0  \tag{24.7}\\
x=0
\end{array}
$$

([ $\cdots$ ] denotes the integer part) acts as a shift on the continued fraction representation of numbers on the unit interval

$$
\begin{equation*}
x=\left[a_{1}, a_{2}, a_{3}, \ldots\right] \rightarrow T(x)=\left[a_{2}, a_{3}, \ldots\right] . \tag{24.8}
\end{equation*}
$$

into the "mother" interval $\ell_{a_{2} a_{3} \ldots}$.
However natural the continued fractions partitioning might seem to a number theorist, it is problematic in practice, as it requires measuring infinity of mode-lockings even at the first step of the partitioning. Thus numerical and experimental use of continued fraction partitioning requires
at least some understanding of the asymptotics of mode-lockings with large continued fraction entries.

The Farey tree partitioning is a systematic bisection of rationals: it is based on the observation that roughly halfways between any two large stability intervals (such as $1 / 2$ and $1 / 3$ ) in the devil's staircase of figure 24.2 there is the next largest stability interval (such as $2 / 5$ ). The winding number of this interval is given by the Farey mediant $\left(P+P^{\prime}\right) /\left(Q+Q^{\prime}\right)$ of the parent mode-lockings $P / Q$ and $P^{\prime} / Q^{\prime}$. This kind of cycle "gluing" is rather general and by no means restricted to circle maps; it can be attained whenever it is possible to arrange that the $Q$ th iterate deviation caused by shifting a parameter from the correct value for the $Q$-cycle is exactly compensated by the $Q^{\prime}$ th iterate deviation from closing the $Q^{\prime}$-cycle; in this way the two near cycles can be glued together into an exact cycle of length $Q+Q^{\prime}$. The Farey tree is obtained by starting with the ends of the unit interval written as $0 / 1$ and $1 / 1$, and then recursively bisecting intervals by means of Farey mediants.

We define the $n$th Farey tree level $T_{n}$ as the monotonically increasing sequence of those continued fractions $\left[a_{1}, a_{2}, \ldots, a_{k}\right]$ whose entries $a_{i} \geq$ $1, i=1,2, \ldots, k-1, \quad a_{k} \geq 2$, add up to $\sum_{i=1}^{k} a_{i}=n+2$. For example

$$
\begin{equation*}
T_{2}=\{[4],[2,2],[1,1,2],[1,3]\}=\left(\frac{1}{4}, \frac{1}{5}, \frac{3}{5}, \frac{3}{4}\right) . \tag{24.9}
\end{equation*}
$$

The number of terms in $T_{n}$ is $2^{n}$. Each rational in $T_{n-1}$ has two "daughters" in $T_{n}$, given by

$$
[\cdots, a-1,2]{ }^{[\cdots, a]} \quad[\cdots, a+1]
$$

Iteration of this rule places all rationals on a binary tree, labeling each by a unique binary label, figure 24.4.

The smallest and the largest denominator in $T_{n}$ are respectively given by

$$
\begin{equation*}
[n-2]=\frac{1}{n-2}, \quad[1,1, \ldots, 1,2]=\frac{F_{n+1}}{F_{n+2}} \propto \rho^{n} \tag{24.10}
\end{equation*}
$$

where the Fibonacci numbers $F_{n}$ are defined by $F_{n+1}=F_{n}+F_{n-1} ; \quad F_{0}=$ $0, F_{1}=1$, and $\rho$ is the golden mean ratio

$$
\begin{equation*}
\rho=\frac{1+\sqrt{5}}{2}=1.61803 \ldots \tag{24.11}
\end{equation*}
$$

Note the enormous spread in the cycle lengths on the same level of the Farey tree: $n \leq Q \leq \rho^{n}$. The cycles whose length grows only as a power of the


Figure 24.4: Farey tree: alternating binary ordered labeling of all Farey denominators on the $n$th Farey tree level.

Farey tree level will cause strong non-hyperbolic effects in the evaluation of various averages.

Having defined the partitioning schemes of interest here, we now briefly summarize the results of the circle-map renormalization theory.

### 24.2 Local theory: "Golden mean" renormalization

- ${ }^{3}$
The way to pinpoint a point on the border of order is to recursively adjust the parameters so that at the recurrence times $t=n_{1}, n_{2}, n_{3}, \cdots$ the trajectory passes through a region of contraction sufficiently strong to compensate for the accumulated expansion of the preceding $n_{i}$ steps, but not so strong as to force the trajectory into a stable attracting orbit. The renormalization operation $R$ implements this procedure by recursively magnifying the neighborhood of a point on the border in the dynamical space (by rescaling by a factor $\alpha$ ), in the parameter space (by shifting the parameter origin onto the border and rescaling by a factor $\delta$ ), and by replacing the initial map $f$ by the $n$th iterate $f^{n}$ restricted to the magnified neighborhood

$$
f_{p}(x) \rightarrow R f_{p}(x)=\alpha f_{p / \delta}^{n}(x / \alpha)
$$

There are by now many examples of such renormalizations in which the new function, framed in a smaller box, is a rescaling of the original function, that is, the fix-point function of the renormalization operator $R$. The best known is the period doubling renormalization, with the recurrence times $n_{i}=$ $2^{i}$. The simplest circle map example is the golden mean renormalization, with recurrence times $n_{i}=F_{i}$ given by the Fibonacci numbers (24.10). Intuitively, in this context a metric self-similarity arises because iterates of
critical maps are themselves critical, that is, they also have cubic inflection points with vanishing derivatives.

The renormalization operator appropriate to circle maps acts as a generalization of the Gauss shift (24.38); it maps a circle map (represented as a pair of functions $(g, f)$, of winding number $[a, b, c, \ldots]$ into a rescaled map of winding number $[b, c, \ldots]$ :

$$
\begin{equation*}
R_{a}\binom{g}{f}=\binom{\alpha g^{a-1} \circ f \circ \alpha^{-1}}{\alpha g^{a-1} \circ f \circ g \circ \alpha^{-1}}, \tag{24.12}
\end{equation*}
$$

Acting on a map with winding number $[a, a, a, \ldots], R_{a}$ returns a map with the same winding number $[a, a, \ldots]$, so the fixed point of $R_{a}$ has a quadratic irrational winding number $W=[a, a, a, \ldots]$. This fixed point has a single expanding eigenvalue $\delta_{a}$. Similarly, the renormalization transformation $R_{a_{p}} \ldots R_{a_{2}} R_{a_{1}} \equiv R_{a_{1} a_{2} \ldots a_{p}}$ has a fixed point of winding number $W_{p}=\left[a_{1}, a_{2}, \ldots, a_{n_{p}}, a_{1}, a_{2}, \ldots\right]$, with a single expanding eigenvalue $\delta_{p}$.

For short repeating blocks, $\delta$ can be estimated numerically by comparing successive continued fraction approximants to $W$. Consider the $P_{r} / Q_{r}$ rational approximation to a quadratic irrational winding number $W_{p}$ whose continued fraction expansion consists of $r$ repeats of a block $p$. Let $\Omega_{r}$ be the parameter for which the map (24.1) has a superstable cycle of rotation number $P_{r} / Q_{r}=[p, p, \ldots, p]$. The $\delta_{p}$ can then be estimated by extrapolating from

$$
\begin{equation*}
\Omega_{r}-\Omega_{r+1} \propto \delta_{p}^{-r} \tag{24.13}
\end{equation*}
$$

What this means is that the "devil's staircase" of figure 24.2 is self-similar under magnification by factor $\delta_{p}$ around any quadratic irrational $W_{p}$.

The fundamental result of the renormalization theory (and the reason why all this is so interesting) is that the ratios of successive $P_{r} / Q_{r}$ mode-locked intervals converge to universal limits. The simplest example of (24.13) is the sequence of Fibonacci number continued fraction approximants to the golden mean winding number $W=[1,1,1, \ldots]=(\sqrt{5}-1) / 2$.

When global problems are considered, it is useful to have at least and idea on extemal scaling laws for mode-lockings. This is achieved, in a first analysis, by fixing the cycle length $Q$ and describing the range of possible asymptotics.

For a given cycle length $Q$, it is found that the narrowest interval shrinks with a power law

$$
\begin{equation*}
\Delta_{1 / Q} \propto Q^{-3} \tag{24.14}
\end{equation*}
$$

For fixed $Q$ the widest interval is bounded by $P / Q=F_{n-1} / F_{n}$, the $n$th continued fraction approximant to the golden mean. The intuitive reason
is that the golden mean winding sits as far as possible from any short cycle mode-locking.

The golden mean interval shrinks with a universal exponent

$$
\begin{equation*}
\Delta_{P / Q} \propto Q^{-2 \mu_{1}} \tag{24.15}
\end{equation*}
$$

where $P=F_{n-1}, Q=F_{n}$ and $\mu_{1}$ is related to the universal Shenker number $\delta_{1}$ (24.13) and the golden mean (24.11) by

$$
\begin{equation*}
\mu_{1}=\frac{\ln \left|\delta_{1}\right|}{2 \ln \rho}=1.08218 \ldots \tag{24.16}
\end{equation*}
$$

The closeness of $\mu_{1}$ to 1 indicates that the golden mean approximant modelockings barely feel the fact that the map is critical (in the $\mathrm{k}=0$ limit this exponent is $\mu=1$ ).

To summarize: for critical maps the spectrum of exponents arising from the circle maps renormalization theory is bounded from above by the harmonic scaling, and from below by the geometric golden-mean scaling:

$$
\begin{equation*}
3 / 2>\mu_{m / n} \geq 1.08218 \cdots \tag{24.17}
\end{equation*}
$$

### 24.3 Global theory: Thermodynamic averaging

Consider the following average over mode-locking intervals (24.5):

$$
\begin{equation*}
\Omega(\tau)=\sum_{Q=1}^{\infty} \sum_{(P \mid Q)=1} \Delta_{P / Q}^{-\tau} \tag{24.18}
\end{equation*}
$$

The sum is over all irreducible rationals $P / Q, P<Q$, and $\Delta_{P / Q}$ is the width of the parameter interval for which the iterates of a critical circle map lock onto a cycle of length $Q$, with winding number $P / Q$.

The qualitative behavior of (24.18) is easy to pin down. For sufficiently negative $\tau$, the sum is convergent; in particular, for $\tau=-1, \Omega(-1)=1$, as for the critical circle maps the mode-lockings fill the entire $\Omega$ range [24.11]. However, as $\tau$ increases, the contributions of the narrow (large $Q$ ) modelocked intervals $\Delta_{P / Q}$ get blown up to $1 / \Delta_{P / Q}^{\tau}$, and at some critical value of $\tau$ the sum diverges. This occurs for $\tau<0$, as $\Omega(0)$ equals the number of all rationals and is clearly divergent.

The sum (24.18) is infinite, but in practice the experimental or numerical mode-locked intervals are available only for small finite $Q$. Hence it
is necessary to split up the sum into subsets $\mathcal{S}_{n}=\{i\}$ of rational winding numbers $P_{i} / Q_{i}$ on the "level" $n$, and present the set of mode-lockings hierarchically, with resolution increasing with the level:

$$
\begin{equation*}
\bar{Z}_{n}(\tau)=\sum_{i \in \mathcal{S}_{n}} \Delta_{i}^{-\tau} \tag{24.19}
\end{equation*}
$$

The original sum (24.18) can now be recovered as the $z=1$ value of a "generating" function $\Omega(z, \tau)=\sum_{n} z^{n} \bar{Z}_{n}(\tau)$. As $z$ is anyway a formal parameter, and $n$ is a rather arbitrary "level" in some ad hoc partitioning of rational numbers, we bravely introduce a still more general, $P / Q$ weighted generating function for (24.18):

$$
\begin{equation*}
\Omega(q, \tau)=\sum_{Q=1}^{\infty} \sum_{(P \mid Q)=1} e^{-q \nu_{P / Q}} Q^{2 \tau \mu_{P / Q}} . \tag{24.20}
\end{equation*}
$$

The sum (24.18) corresponds to $q=0$. Exponents $\nu_{P / Q}$ will reflect the importance we assign to the $P / Q$ mode-locking, that is, the measure used in the averaging over all mode-lockings. Three choices of of the $\nu_{P / Q}$ hierarchy that we consider here correspond respectively to the Farey series partitioning

$$
\begin{equation*}
\Omega(q, \tau)=\sum_{Q=1}^{\infty} \Phi(Q)^{-q} \sum_{(P \mid Q)=1} Q^{2 \tau \mu_{P / Q}} \tag{24.21}
\end{equation*}
$$

the continued fraction partitioning

$$
\begin{equation*}
\Omega(q, \tau)=\sum_{n=1}^{\infty} e^{-q n} \sum_{\left[a_{1}, \ldots, a_{n}\right]} Q^{2 \tau \mu_{\left[a_{1}, \ldots, a_{n}\right]}} \tag{24.22}
\end{equation*}
$$

and the Farey tree partitioning

$$
\begin{equation*}
\Omega(q, \tau)=\sum_{k=n}^{\infty} 2^{-q n} \sum_{i=1}^{2^{n}} Q_{i}^{2 \tau \mu_{i}}, \quad Q_{i} / P_{i} \in T_{n} \tag{24.23}
\end{equation*}
$$

We remark that we are investigating a set arising in the analysis of the parameter space of a dynamical system: there is no "natural measure" dictated by dynamics, and the choice of weights reflects only the choice of hierarchical presentation.

### 24.4 Hausdorff dimension of irrational windings

A finite cover of the set irrational windings at the " $n$th level of resolution" is obtained by deleting the parameter values corresponding to the modelockings in the subset $\mathcal{S}_{n}$; left behind is the set of complement covering intervals of widths

$$
\begin{equation*}
\ell_{i}=\Omega_{P_{r} / Q_{r}}^{\min }-\Omega_{P_{l} / Q_{l}}^{\max } . \tag{24.24}
\end{equation*}
$$

Here $\Omega_{P_{r} / Q_{r}}^{\min }\left(\Omega_{P_{l} / Q_{l}}^{\max }\right)$ are respectively the lower (upper) edges of the modelocking intervals $\Delta_{P_{r} / Q_{r}}\left(\Delta_{P_{l} / Q_{l}}\right)$ bounding $\ell_{i}$ and $i$ is a symbolic dynamics label, for example the entries of the continued fraction representation $P / Q=\left[a_{1}, a_{2}, \ldots, a_{n}\right]$ of one of the boundary mode-lockings, $i=a_{1} a_{2} \cdots a_{n}$. $\ell_{i}$ provide a finite cover for the irrational winding set, so one may consider the sum

$$
\begin{equation*}
Z_{n}(\tau)=\sum_{i \in \mathcal{S}_{n}} \ell_{i}^{-\tau} \tag{24.25}
\end{equation*}
$$

The value of $-\tau$ for which the $n \rightarrow \infty$ limit of the sum (24.25) is finite is the Hausdorff dimension $D_{H}$ of the irrational winding set. Strictly speaking, this is the Hausdorff dimension only if the choice of covering intervals $\ell_{i}$ is optimal; otherwise it provides an upper bound to $D_{H}$. As by construction the $\ell_{i}$ intervals cover the set of irrational winding with no slack, we expect that this limit yields the Hausdorff dimension. This is supported by all numerical evidence, but a proof that would satisfy mathematicians is lacking.

The physically relevant statement is that for critical circle maps $D_{H}=$ $0.870 \ldots$ is a (global) universal number.

### 24.4.1 The Hausdorff dimension in terms of cycles

Estimating the $n \rightarrow \infty$ limit of (24.25) from finite numbers of covering intervals $\ell_{i}$ is a rather unilluminating chore. Fortunately, there exist considerably more elegant ways of extracting $D_{H}$. We have noted that in the case of the "trivial" mode-locking problem (24.4), the covering intervals are generated by iterations of the Farey map (24.37) or the Gauss shift (24.38). The $n$th level sum (24.25) can be approximated by $\mathcal{L}_{\tau}^{n}$, where

$$
\mathcal{L}_{\tau}(y, x)=\delta\left(x-f^{-1}(y)\right)\left|f^{\prime}(y)\right|^{\tau}
$$

This amounts to approximating each cover width $\ell_{i}$ by $\left|d f^{n} / d x\right|$ evaluated on the $i$ th interval. We are thus led to the following determinant

$$
\operatorname{det}\left(1-z \mathcal{L}_{\tau}\right)=\exp \left(-\sum_{p} \sum_{r=1}^{\infty} \frac{z^{r n_{p}}}{r} \frac{\left|\Lambda_{p}^{r}\right|^{\tau}}{1-1 / \Lambda_{p}^{r}}\right)
$$

$$
\begin{equation*}
=\prod_{p} \prod_{k=0}^{\infty}\left(1-z^{n_{p}}\left|\Lambda_{p}\right|^{\tau} / \Lambda_{p}^{k}\right) \tag{24.26}
\end{equation*}
$$

The sum (24.25) is dominated by the leading eigenvalue of $\mathcal{L}_{\tau}$; the Hausdorff dimension condition $Z_{n}\left(-D_{H}\right)=O(1)$ means that $\tau=-D_{H}$ should be such that the leading eigenvalue is $z=1$. The leading eigenvalue is determined by the $k=0$ part of (24.26); putting all these pieces together, we obtain a pretty formula relating the Hausdorff dimension to the prime cycles of the map $f(x)$ :

$$
\begin{equation*}
0=\prod_{p}\left(1-1 /\left|\Lambda_{p}\right|^{D_{H}}\right) \tag{24.27}
\end{equation*}
$$

For the Gauss shift (24.38) the stabilities of periodic cycles are available analytically, as roots of quadratic equations: For example, the $x_{a}$ fixed points (quadratic irrationals with $x_{a}=[a, a, a \ldots]$ infinitely repeating continued fraction expansion) are given by

$$
\begin{equation*}
x_{a}=\frac{-a+\sqrt{a^{2}+4}}{2}, \quad \Lambda_{a}=-\left(\frac{a+\sqrt{a^{2}+4}}{2}\right)^{2} \tag{24.28}
\end{equation*}
$$

and the $x_{a b}=[a, b, a, b, a, b, \ldots] \quad 2$-cycles are given by

$$
\begin{align*}
& x_{a b}=\frac{-a b+\sqrt{(a b)^{2}+4 a b}}{2 b}  \tag{24.29}\\
& \Lambda_{a b}=\left(x_{a b} x_{b a}\right)^{-2}=\left(\frac{a b+2+\sqrt{a b(a b+4)}}{2}\right)^{2}
\end{align*}
$$

We happen to know beforehand that $D_{H}=1$ (the irrationals take the full measure on the unit interval, or, from another point of view the Gauss map is not a repeller), so is the infinite product (24.27) merely a very convoluted way to compute the number 1? Possibly so, but once the meaning of (24.27) has been grasped, the corresponding formula for the critical circle maps follows immediately:

$$
\begin{equation*}
0=\prod_{p}\left(1-1 /\left|\delta_{p}\right|^{D_{H}}\right) \tag{24.30}
\end{equation*}
$$

The importance of this formula relies on the fact that it expresses $D_{H}$ in terms of universal quantities, thus providing a nice connection from local universal exponents to global scaling quantities: actual computations using (24.30) are rather involved, as they require a heavy computational effort to extract Shenker's scaling $\delta_{p}$ for periodic continued fractions, and moreover dealing with an infinite alphabet requires control over tail summation if an accurate estimate is to be sought. In table 24.1 we give a small selection of computed Shenker's scalings.

| $p$ | $\delta_{p}$ |
| :---: | :---: |
|  | -2.833612 |
| [ $\left.\begin{array}{lllllll}2 & 2 & 2 & 2 & \ldots\end{array}\right]$ | -6.7992410 |
| $\left[\begin{array}{llllll}3 & 3 & 3 & 3 & \text {.. }\end{array}\right]$ | -13.760499 |
| [ $\left.\begin{array}{lllllll}4 & 4 & 4 & 4 & \ldots\end{array}\right]$ | -24.62160 |
| $\left[\begin{array}{llllll}5 & 5 & 5 & 5 & \ldots\end{array}\right]$ | -40.38625 |
| $\left[\begin{array}{llllll}6 & 6 & 6 & 6 & . .\end{array}\right]$ | -62.140 |
|  | 17.66549 |
| $\left[\begin{array}{llllll}1 & 3 & 1 & 3 & \ldots .\end{array}\right]$ | 31.62973 |
| $\left[\begin{array}{lllllll}1 & 4 & 1 & 4 & \ldots\end{array}\right]$ | 50.80988 |
| [15lllll | 76.01299 |
| [23 $233 \ldots$ | 91.29055 |

Table 24.1: Shenker's $\delta_{p}$ for a few periodic continued fractions, from ref. [24.1].

### 24.5 Thermodynamics of Farey tree: Farey model



We end this chapter by giving an example of a number theoretical model motivated by the mode-locking phenomenology. We will consider it by means of the thermodynamic formalism of chapter 20 , by looking at the free energy.

Consider the Farey tree partition sum (24.23): the narrowest modelocked interval (24.15) at the $n$th level of the Farey tree partition sum (24.23) is the golden mean interval

$$
\begin{equation*}
\Delta_{F_{n-1} / F_{n}} \propto\left|\delta_{1}\right|^{-n} \tag{24.31}
\end{equation*}
$$

It shrinks exponentially, and for $\tau$ positive and large it dominates $q(\tau)$ and bounds $d q(\tau) / d \tau$ :

$$
\begin{equation*}
q_{\max }^{\prime}=\frac{\ln \left|\delta_{1}\right|}{\ln 2}=1.502642 \ldots \tag{24.32}
\end{equation*}
$$

However, for $\tau$ large and negative, $q(\tau)$ is dominated by the interval (24.14) which shrinks only harmonically, and $q(\tau)$ approaches 0 as

$$
\begin{equation*}
\frac{q(\tau)}{\tau}=\frac{3 \ln n}{n \ln 2} \rightarrow 0 \tag{24.33}
\end{equation*}
$$

So for finite $\mathrm{n}, q_{n}(\tau)$ crosses the $\tau$ axis at $-\tau=D_{n}$, but in the $n \rightarrow \infty$ limit, the $q(\tau)$ function exhibits a phase transition; $q(\tau)=0$ for $\tau<-D_{H}$, but is a non-trivial function of $\tau$ for $-D_{H} \leq \tau$. This non-analyticity is rather severe - to get a clearer picture, we illustrate it by a few number-theoretic models (the critical circle maps case is qualitatively the same).

An approximation to the "trivial" Farey level thermodynamics is given by the "Farey model", in which the intervals $\ell_{P / Q}$ are replaced by $Q^{-2}$ :

$$
\begin{equation*}
Z_{n}(\tau)=\sum_{i=1}^{2^{n}} Q_{i}^{2 \tau} \tag{24.34}
\end{equation*}
$$

Here $Q_{i}$ is the denominator of the $i$ th Farey rational $P_{i} / Q_{i}$. For example (see figure 24.4),

$$
Z_{2}(1 / 2)=4+5+5+4 .
$$

By the annihilation property (24.38) of the Gauss shift on rationals, the $n$th Farey level sum $Z_{n}(-1)$ can be written as the integral

$$
Z_{n}(-1)=\int d x \delta\left(f^{n}(x)\right)=\sum 1 /\left|f_{a_{1} \ldots a_{k}}^{\prime}(0)\right|,
$$

and in general

$$
Z_{n}(\tau)=\int d x \mathcal{L}_{\tau}^{n}(0, x)
$$

with the sum restricted to the Farey level $a_{1}+\ldots+a_{k}=n+2$. It is easily checked that $f_{a_{1} \ldots a_{k}}^{\prime}(0)=(-1)^{k} Q_{\left[a_{1}, \ldots, a_{k}\right]}^{2}$, so the Farey model sum is a partition generated by the Gauss map preimages of $x=0$, that is, by rationals, rather than by the quadratic irrationals as in (24.26). The sums are generated by the same transfer operator, so the eigenvalue spectrum should be the same as for the periodic orbit expansion, but in this variant of the finite level sums we can can evaluate $q(\tau)$ exactly for $\tau=k / 2, k$ a nonnegative integer. First one observes that $Z_{n}(0)=2^{n}$. It is also easy to check that $Z_{n}(1 / 2)=\sum_{i} Q_{i}=2 \cdot 3^{n}$. More surprisingly, $Z_{n}(3 / 2)=$ $\sum_{i} Q^{3}=54 \cdot 7^{n-1}$. A few of these "sum rules" are listed in the table 24.2, they are consequence of the fact that the denominators on a given level are Farey sums of denominators on preceding levels.

A bound on $D_{H}$ can be obtained by approximating (24.34) by

$$
\begin{equation*}
Z_{n}(\tau)=n^{2 \tau}+2^{n} \rho^{2 n \tau} \tag{24.35}
\end{equation*}
$$

In this approximation we have replaced all $\ell_{P / Q}$, except the widest interval $\ell_{1 / n}$, by the narrowest interval $\ell_{F_{n-1} / F_{n}}$ (see (24.15)). The crossover from the harmonic dominated to the golden mean dominated behavior occurs at the $\tau$ value for which the two terms in (24.35) contribute equally:

$$
\begin{equation*}
D_{n}=\hat{D}+O\left(\frac{\ln n}{n}\right), \quad \hat{D}=\frac{\ln 2}{2 \ln \rho}=.72 \ldots \tag{24.36}
\end{equation*}
$$

For negative $\tau$ the sum (24.35) is the lower bound on the sum (24.25), so $\hat{D}$ is a lower bound on $D_{H}$.

From a general perspective the analysis of circle maps thermodynamics has revealed the fact that physically interesting dynamical systems often

| $\tau / 2$ | $Z_{n}(\tau / 2) / Z_{n-1}(\tau / 2)$ |
| :---: | :---: |
| 0 | 2 |
| 1 | 3 |
| 2 | $(5+\sqrt{17}) / 2$ |
| 3 | 7 |
| 4 | $(5+\sqrt{17}) / 2$ |
| 5 | $7+4 \sqrt{6}$ |
| 6 | $26.20249 \ldots$ |

Table 24.2: Partition function sum rules for the Farey model.
exhibit mixtures of hyperbolic and marginal stabilities. In such systems there are orbits that stay 'glued' arbitrarily close to stable regions for arbitrarily long times. This is a generic phenomenon for Hamiltonian systems, where elliptic islands of stability coexist with hyperbolic homoclinic webs. Thus the considerations of chapter 21 are important also in the analysis of renomarmalization at the onset of chaos.

## Commentary

Remark 24.1 The physics of circle maps. Mode-locking phenomenology is reviewed in ref. [24.5], a more theoretically oriented discussion is contained in ref. [24.3]. While representative of dissipative systems we may also consider circle mapsas a crude approximation to Hamiltonian local dynamics: a typical island of stability in a Hamiltonian 2-d map is an infinite sequence of concentric KAM tori and chaotic regions. In the crudest approximation, the radius can here be treated as an external parameter $\Omega$, and the angular motion can be modelled by a map periodic in the angular variable [24.8, 24.9]. By losing all of the "island-within-island" structure of real systems, circle map models skirt the problems of determining the symbolic dynamics for a realistic Hamiltonian system, but they do retain some of the essential features of such systems, such as the golden mean renormalization [17.4, 24.8] and non-hyperbolicity in form of sequences of cycles accumulating toward the borders of stability. In particular, in such systems there are orbits that stay "glued" arbitrarily close to stable regions for arbitrarily long times. As this is a generic phenomenon in physically interesting dynamical systems, such as the Hamiltonian systems with coexisting elliptic islands of stability and hyperbolic homoclinic webs, development of good computational techniques is here of utmost practical importance.

Remark 24.2 Critical mode-locking set The fact that mode-lockings completely fill the unit interval at the critical point has been proposed in refs. [24.3, 24.10]. The proof that the set of irrational windings is of zero Lebesgue measure in given in ref. [24.11].

Remark 24.3 Counting noise for Farey series. The number of rationals in the Farey series of order $Q$ is $\phi(Q)$, which is a highly irregular function of $Q$ : incrementing $Q$ by 1 increases $\Phi(Q)$ by anything from 2 to $Q$ terms. We refer to this fact as the "Euler noise".

The Euler noise poses a serious obstacle for numerical calculations with the Farey series partitionings; it blocks smooth extrapolations to $Q \rightarrow \infty$ limits from
finite $Q$ data. While this in practice renders inaccurate most Farey-sequence partitioned averages, the finite Q Hausdorff dimension estimates exhibit (for reasons that we do not understand) surprising numerical stability, and the Farey series partitioning actually yields the best numerical value of the Hausdorff dimension (24.25) of any methods used so far; for example the computation in ref. [24.12] for critical sine map (24.1), based on $240 \leq Q \leq 250$ Farey series partitions, yields $D_{H}=.87012 \pm .00001$. The quoted error refers to the variation of $D_{H}$ over this range of $Q$; as the computation is not asymptotic, such numerical stability can underestimate the actual error by a large factor.

Remark 24.4 Farey tree presentation function. The Farey tree rationals can be generated by backward iterates of $1 / 2$ by the Farey presentation function [24.13]:

$$
\begin{array}{lll}
f_{0}(x)=x /(1-x) & & 0 \leq x<1 / 2 \\
f_{1}(x)=(1-x) / x & & 1 / 2<x \leq 1 . \tag{24.37}
\end{array}
$$

The Gauss shift (24.7) corresponds to replacing the binary Farey presentation function branch $f_{0}$ in (24.37) by an infinity of branches

$$
\begin{align*}
f_{a}(x) & =f_{1} \circ f_{0}^{(a-1)}(x)=\frac{1}{x}-a, \quad \frac{1}{a-1}<x \leq \frac{1}{a}, \\
f_{a b \cdots c}(x) & =f_{c} \circ \circ \circ f_{b} \circ f_{a}(x) . \tag{24.38}
\end{align*}
$$

A rational $x=\left[a_{1}, a_{2}, \ldots, a_{k}\right]$ is annihilated by the $k$ th iterate of the Gauss shift, $f_{a_{1} a_{2} \cdots a_{k}}(x)=0$. The above maps look innocent enough, but note that what is being partitioned is not the dynamical space, but the parameter space. The flow described by (24.37) and by its non-trivial circle-map generalizations will turn out to be a renormalization group flow in the function space of dynamical systems, not an ordinary flow in the phase space of a particular dynamical system.

The Farey tree has a variety of interesting symmetries (such as "flipping heads and tails" relations obtained by reversing the order of the continued-fraction entries) with as yet unexploited implications for the renormalization theory: some of these are discussed in ref. [24.4].

An alternative labeling of Farey denominators has been introduced by Knauf [24.6] in context of number-theoretical modeling of ferromagnetic spin chains: it allows for a number of elegant manipulations in thermodynamic averages connected to the Farey tree hierarchy.

Remark 24.5 Circle map renormalization The idea underlying golden mean renormalization goes back to Shenker [24.9]. A renormalization group procedure was formulated in refs. [24.7, 24.14], where moreover the uniqueness of the relevant eigenvalue is claimed. This statement has been confirmed by a computerassisted proof [24.15], and in the following we will always assume it. There are a number of experimental evidences for local universality, see refs. [24.16, 24.17].

On the other side of the scaling tale, the power law scaling for harmonic fractions (discussed in refs. [24.2, 24.3, 24.4]) is derived by methods akin to those used in describing intermittency [24.21]: $1 / Q$ cycles accumulate toward the edge of $0 / 1$ mode-locked interval, and as the successive mode-locked intervals $1 / Q, 1 /(Q-$ 1) lie on a parabola, their differences are of order $Q^{-3}$.

Remark 24.6 Farey series and the Riemann hypothesis The Farey series thermodynamics is of a number theoretical interest, because the Farey series provide uniform coverings of the unit interval with rationals, and because they are closely related to the deepest problems in number theory, such as the Riemann hypothesis [24.22, 24.23] . The distribution of the Farey series rationals across the unit interval is surprisingly uniform - indeed, so uniform that in the pre-computer days it has motivated a compilation of an entire handbook of Farey series [24.24]. A quantitive measure of the non-uniformity of the distribution of Farey rationals is given by displacements of Farey rationals for $P_{i} / Q_{i} \in \mathcal{F}_{Q}$ from uniform spacing:

$$
\delta_{i}=\frac{i}{\Phi(Q)}-\frac{P_{i}}{Q_{i}}, \quad i=1,2, \cdots, \Phi(Q)
$$

The Riemann hypothesis states that the zeros of the Riemann zeta function lie on the $s=1 / 2+i \tau$ line in the complex $s$ plane, and would seem to have nothing to do with physicists' real mode-locking widths that we are interested in here. However, there is a real-line version of the Riemann hypothesis that lies very close to the mode-locking problem. According to the theorem of Franel and Landau [24.25, $24.22,24.23]$, the Riemann hypothesis is equivalent to the statement that

$$
\sum_{Q_{i} \leq Q}\left|\delta_{i}\right|=o\left(Q^{\frac{1}{2}+\epsilon}\right)
$$

for all $\epsilon$ as $Q \rightarrow \infty$. The mode-lockings $\Delta_{P / Q}$ contain the necessary information for constructing the partition of the unit interval into the $\ell_{i}$ covers, and therefore implicitly contain the $\delta_{i}$ information. The implications of this for the circle-map scaling theory have not been worked out, and is not known whether some conjecture about the thermodynamics of irrational windings is equivalent to (or harder than) the Riemann hypothesis, but the danger lurks.

Remark 24.7 Farey tree partitioning. The Farey tree partitioning was introduced in refs. [24.26, 24.27, 24.4] and its thermodynamics is discussed in detail in refs. [24.12, 24.13]. The Farey tree hierarchy of rationals is rather new, and, as far as we are aware, not previously studied by number theorists. It is appealing both from the experimental and from the the golden-mean renormalization point of view, but it has a serious drawback of lumping together mode-locking intervals of wildly different sizes on the same level of the Farey tree.

Remark 24.8 Local and global universality. Numerical evidences for global universal behavior have been presented in ref. [24.3]. The question was reexamined in ref. [24.12], where it was pointed out how a high-precision numerical estimate is in practice very hard to obtain. It is not at all clear whether this is the optimal global quantity to test but at least the Hausdorff dimension has the virtue of being independent of how one partitions mode-lockings and should thus be the same for the variety of thermodynamic averages in the literature.

The formula (24.30), linking local to global behavior, was proposed in ref. [24.1].
The derivation of (24.30) relies only on the following aspects of the "hyperbolicity conjecture" of refs. [24.4, 24.18, 24.19, 24.20]:

1. limits for Shenker $\delta$ 's exist and are universal. This should follow from the renormalization theory developed in refs. [24.7, 24.14, 24.15], though a general proof is still lacking.
2. $\delta_{p}$ grow exponentially with $n_{p}$, the length of the continued fraction block $p$.
3. $\delta_{p}$ for $p=a_{1} a_{2} \ldots n$ with a large continued fraction entry $n$ grows as a power of $n$. According to (24.14), $\lim _{n \rightarrow \infty} \delta_{p} \propto n^{3}$. In the calculation of ref. [24.1] the explicit values of the asymptotic exponents and prefactors were not used, only the assumption that the growth of $\delta_{p}$ with $n$ is not slower than a power of $n$.

Remark 24.9 Farey model. The Farey model (24.33) has been proposed in ref. [24.12]; though it might seem to have been pulled out of a hat, the Farey model is as sensible description of the distribution of rationals as the periodic orbit expansion (24.26).

Remark 24.10 Symbolic dynamics for Hamiltonian rotational orbits. The rotational codes of ref. [5.4] are closely related to those for maps with a natural angle variable, for example for circle maps [24.34, 24.36] and cat maps [24.37]. Ref. [5.4] also offers a systematic rule for obtaining the symbolic codes of "islands around islands" rotational orbits [24.39]. These correspond, for example, to orbits that rotate around orbits that rotate around the elliptic fixed point; thus they are defined by a sequence of rotation numbers.

A different method for constructing symbolic codes for "islands around islands" was given in refs. [24.42, 24.40]; however in these cases the entire set of orbits in an island was assigned the same sequence and the motivation was to study the transport implications for chaotic orbits outside the islands [24.39, 24.41].

## Résumé

The mode locking problem, and the quasiperiodic transition to chaos offer an opportunity to use cycle expansions on hierarchical structures in parameter space: this is not just an application of the conventional thermodynamic formalism, but offers a clue on how to extend universality theory from local scalings to global quantities.

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## Exercises

Exercise 24.1 Mode-locked intervals. Check that when $k \neq 0$ the interval $\Delta_{P / Q}$ have a non-zero width (look for instance at simple fractions, and consider $k$ small). Show that for small $k$ the width of $\Delta_{0 / 1}$ is an increasing function of $k$.

Exercise 24.2 Bounds on Hausdorff dimension. By making use of the bounds (24.17) show that the Hausdorff dimension for critical mode lockings may be bounded by

$$
2 / 3 \leq D_{H} \leq .9240 \ldots
$$

Exercise 24.3 Farey model sum rules. Verify the sum rules reported in table 24.2. An elegant way to get a number of sum rules for the Farey model is by taking into account an lexical ordering introduced by Contucci and Knauf, see ref. [24.28].

Exercise 24.4 Metric entropy of the Gauss shift. Check that the Lyapunov exponent of the Gauss map (24.7) is given by $\pi^{2} / 6 \ln 2$. This result has been claimed to be relevant in the discussion of "mixmaster" cosmologies, see ref. [24.30].

Exercise 24.5 Refined expansions. Show that the above estimates can be refined as follows:

$$
F(z, 2) \sim \zeta(2)+(1-z) \log (1-z)-(1-z)
$$

and

$$
F(z, s) \sim \zeta(s)+\Gamma(1-s)(1-z)^{s-1}-S(s)(1-z)
$$

for $s \in(1,2)$ ( $S(s)$ being expressed by a converging sum). You may use either more detailed estimate for $\zeta(s, a)$ (via Euler summation formula) or keep on subtracting leading contributions [24.31].

Exercise 24.6 Hitting condition. Prove (23.39). Hint: together with the real trajectory consider the line passing through the starting point, with polar angle $\theta_{m, n}$ : then draw the perpendiculars to the actual trajectory, passing through the center of the $(0,0)$ and $(m, n)$ disks.

Exercise $24.7 \quad j_{n}$ and $\alpha_{c r}$. Look at the integration region and how it scales by plotting it for increasing values of $n$.

Exercise 24.8 Estimates of the Riemann zeta function. Try to approximate numerically the Riemann zeta function for $s=2,4,6$ using different acceleration algorithms: check your results with refs. [24.32, 24.33].

Exercise 24.9 Farey tree and continued fractions I. Consider the Farey tree presentation function $f:[0,1] \mapsto[0,1]$, such that if $I=[0,1 / 2)$ and $J=[1 / 2,1]$, $\left.f\right|_{I}=x /(1-x)$ and $\left.f\right|_{J}=(1-x) / x$. Show that the corresponding induced map is the Gauss map $g(x)=1 / x-[1 / x]$.

Exercise 24.10 Farey tree and continued fraction II. (Lethal weapon II). Build the simplest piecewise linear approximation to the Farey tree presentation function (hint: substitute first the righmost, hyperbolic branch with a linear one): consider then the spectral determinant of the induced map $\hat{g}$, and calculate the first two eigenvalues besides the probability conservation one. Compare the results with the rigorous bound deduced in ref. [21.17].

## Chaos: Classical and Quantum

## Part II: Semiclassical Chaos



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## Chapter 25

## Prologue


#### Abstract

Anyone who uses words "quantum" and "chaos" in the same sentence should be hung by his thumbs on a tree in the park behind the Niels Bohr Institute. Joseph Ford


(G. Vattay, G. Tanner and P. Cvitanović)

You have read the first volume of this book. So far, so good - anyone can play a game of classical pinball, and a skilled neuroscientist can poke rat brains. We learned that information about chaotic dynamics can be obtained by calculating spectra of linear operators such as the evolution operator of sect. 10.2 or the associated partial differential equations such as the Liouville equation (9.35). The spectra of these operators can be expressed in terms of periodic orbits of the deterministic dynamics by means of trace formulas and cycle expansions.

But what happens quantum mechanically, that is, if we scatter waves rather than point-like pinballs? Can we turn the problem round and study linear PDE's in terms of the underlying deterministic dynamics? And, is there a link between structures in the spectrum or the eigenfunctions of a PDE and the dynamical properties of the underlying classical flow? The answer is yes, but $\ldots$ things are becoming somewhat more complicated when studying 2nd or higher order linear PDE's. We can find classical dynamics associated with a linear PDE, just take geometric optics as a familiar example. Propagation of light follows a second order wave equation but may in certain limits be well described in terms of geometric rays. A theory in terms of properties of the classical dynamics alone, referred to here as the semiclassical theory, will not be exact, in contrast to the classical periodic orbit formulas obtained so far. Waves exhibit new phenomena, such as interference, diffraction, and higher $\hbar$ corrections which will only be partially incorporated into the periodic orbit theory.

### 25.1 Quantum pinball

In what follows, we will restrict the discussion to the non-relativistic Schrödinger equation. The approach will be very much in the spirit of the early days of quantum mechanics, before its wave character has been fully uncovered by Schrödinger in the mid 1920's. Indeed, were physicists of the period as familiar with classical chaos as we are today, this theory could have been developed 80 years ago. It was the discrete nature of the hydrogen spectrum which inspired the Bohr - de Broglie picture of the old quantum theory: one places a wave instead of a particle on a Keplerian orbit around the hydrogen nucleus. The quantization condition is that only those orbits contribute for which this wave is stationary; from this followed the Balmer spectrum and the Bohr-Sommerfeld quantization which eventually led to the more sophisticated theory of Heisenberg, Schrödinger and others. Today we are very aware of the fact that elliptic orbits are an idiosyncracy of the Kepler problem, and that chaos is the rule; so can the Bohr quantization be generalized to chaotic systems?

The question was answered affirmatively by M. Gutzwiller, as late as 1971: a chaotic system can indeed be quantized by placing a wave on each of the infinity of unstable periodic orbits. Due to the instability of the orbits the wave does not stay localized but leaks into neighborhoods of other periodic orbits. Contributions of different periodic orbits interfere and the quantization condition can no longer be attributed to a single periodic orbit: A coherent summation over the infinity of periodic orbit contributions gives the desired spectrum.

The pleasant surprise is that the zeros of the dynamical zeta function (1.6.2) derived in the context of classical chaotic dynamics,

$$
1 / \zeta(z)=\prod_{p}\left(1-t_{p}\right)
$$

also yield excellent estimates of quantum resonances, with the quantum amplitude associated with a given cycle approximated semiclassically by the weight

$$
\begin{equation*}
t_{p}=\frac{1}{\left|\Lambda_{p}\right|^{\frac{1}{2}}} e^{\frac{i}{\hbar} S_{p}-i \pi m_{p} / 2} \tag{25.1}
\end{equation*}
$$

whose magnitude is the square root of the classical weight (15.10)

$$
t_{p}=\frac{1}{\left|\Lambda_{p}\right|} e^{\beta \cdot A_{p}-s T_{p}},
$$

and the phase is given by the Bohr-Sommerfeld action integral $S_{p}$, together with an additional topological phase $m_{p}$, the number of caustics
along the periodic trajectory, points where the naive semiclassical approximation fails.

In this approach, the quantal spectra of classically chaotic dynamical systems are determined from the zeros of dynamical zeta functions, defined by cycle expansions of infinite products of form

$$
\begin{equation*}
1 / \zeta=\prod_{p}\left(1-t_{p}\right)=1-\sum_{f} t_{f}-\sum_{k} c_{k} \tag{25.2}
\end{equation*}
$$

with weight $t_{p}$ associated to every prime (non-repeating) periodic orbit (or cycle) $p$.

The key observation is that the chaotic dynamics is often organized around a few fundamental cycles. These short cycles capture the skeletal topology of the motion in the sense that any long orbit can approximately be pieced together from the fundamental cycles. In chapter 18 it was shown that for this reason the cycle expansion (25.2) is a highly convergent expansion dominated by short cycles grouped into fundamental contributions, with longer cycles contributing rapidly decreasing curvature corrections. Computations with dynamical zeta functions are rather straightforward; typically one determines lengths and stabilities of a finite number of shortest periodic orbits, substitutes them into (25.2), and estimates the zeros of $1 / \zeta$ from such polynomial approximations.

From the vantage point of the dynamical systems theory, the trace formulas (both the exact Selberg and the semiclassical Gutzwiller trace formula) fit into a general framework of replacing phase space averages by sums over periodic orbits. For classical hyperbolic systems this is possible since the invariant density can be represented by sum over all periodic orbits, with weights related to their instability. The semiclassical periodic orbit sums differ from the classical ones only in phase factors and stability weights; such differences may be traced back to the fact that in quantum mechanics the amplitudes rather than the probabilities are added.

The type of dynamics has a strong influence on the convergence of cycle expansions and the properties of quantal spectra; this necessitates development of different approaches for different types of dynamical behavior such as, on one hand, the strongly hyperbolic and, on the other hand, the intermittent dynamics of chapters 18 and 21. For generic nonhyperbolic systems (which we shall not discuss here), with mixed phase space and marginally stable orbits, periodic orbit summations are hard to control, and it is still not clear that the periodic orbit sums should necessarily be the computational method of choice.

Where is all this taking us? The goal of this part of the book is to demonstrate that the cycle expansions, developed so far in classical settings, are also a powerful tool for evaluation of quantum resonances of classically chaotic systems.

Figure 25.1: A typical collinear helium trajectory in the $r_{1}-r_{2}$ plane; the trajectory enters along the $r_{1}$ axis and escapes to infinity along the $r_{2}$ axis.

$\mathrm{r}_{1}$

First we shall warm up playing our game of pinball, this time in a quantum version. Were the game of pinball a closed system, quantum mechanically one would determine its stationary eigenfunctions and eigenenergies. For open systems one seeks instead complex resonances, where the imaginary part of the eigenenergy describes the rate at which the quantum wave function leaks out of the central scattering region. This will turn out to work well, except who truly wants to know accurately the resonances of a quantum pinball?

### 25.2 Quantization of helium

Once we have derived the semiclassical weight associated with the periodic orbit $p$ (25.1), we will finally be in position to accomplish something altogether remarkable. We are now able to put together all ingredients that make the game of pinball unpredictable, and compute a "chaotic" part of the helium spectrum to shocking accuracy. From the classical dynamics point of view, helium is an example of Poincaré's dreaded and intractable 3-body problem. Undaunted, we forge ahead and consider the collinear helium, with zero total angular momentum, and the two electrons on the opposite sides of the nucleus.


We set the electron mass to 1 , the nucleus mass to $\infty$, the helium nucleus charge to 2 , the electron charges to -1 . The Hamiltonian is

$$
\begin{equation*}
H=\frac{1}{2} p_{1}^{2}+\frac{1}{2} p_{2}^{2}-\frac{2}{r_{1}}-\frac{2}{r_{2}}+\frac{1}{r_{1}+r_{2}} . \tag{25.3}
\end{equation*}
$$

Due to the energy conservation, only three of the phase space coordinates $\left(r_{1}, r_{2}, p_{1}, p_{2}\right)$ are independent. The dynamics can be visualized as a motion in the $\left(r_{1}, r_{2}\right), r_{i} \geq 0$ quadrant, figure 25.1 , or, better still, by a well chosen 2-dimensional Poincaré section.

The motion in the $\left(r_{1}, r_{2}\right)$ plane is topologically similar to the pinball motion in a 3 -disk system, except that the motion is not free, but in the

Coulomb potential. The classical collinear helium is also a repeller; almost all of the classical trajectories escape. Miraculously, the symbolic dynamics for the survivors turns out to be binary, just as in the 3 -disk game of pinball, so we know what cycles need to be computed for the cycle expansion (1.10). A set of shortest cycles up to a given symbol string length then yields an estimate of the helium spectrum. This simple calculation yields surprisingly accurate eigenvalues; even though the cycle expansion was based on the semiclassical approximation (25.1) which is expected to be good only in the classical large energy limit, the eigenenergies are good to $1 \%$ all the way down to the ground state.

Before we can get to this point, we first have to recapitulate some basic notions of quantum mechanics; after having defined the main quantum objects of interest, the quantum propagator and the Green's function, we will relate the quantum propagation to the classical flow of the underlying dynamical system. We will then proceed to construct semiclassical approximations to the quantum propagator and the Green's function. A rederivation of classical Hamiltonian dynamics starting from the Hamilton-Jacobi equation will be offered along the way. The derivation of the Gutzwiller trace formula and the semiclassical zeta function as a sum and as a product over periodic orbits will be given in chapter 30. In subsequent chapters we buttress our case by applying and extending the theory: a cycle expansion calculation of scattering resonances in a 3 -disk billiard in chapter 32 , the spectrum of helium in chapter 34, and the incorporation of diffraction effects in chapter 35.

## Guide to literature

A key prerequisite to developing any theory of "quantum chaos" is solid understanding of Hamiltonian mechanics. For that, Arnol'd monograph [1.27] is the essential reference. Ozorio de Almeida's monograph [7.11] offers a compact introduction to the aspects of Hamiltonian dynamics required for the quantization of integrable and nearly integrable systems, with emphasis on periodic orbits, normal forms, catastrophy theory and torus quantization. The book by Brack and Bhaduri [25.1] is an excellent introduction to the semiclassical methods. Gutzwiller's monograph [25.2] is an advanced introduction focusing on chaotic dynamics both in classical Hamiltonian settings and in the semiclassical quantization. This book is worth browsing through for its many insights and erudite comments on quantum and celestial mechanics even if one is not working on problems of quantum chaos. More suitable as a graduate course text is Reichl's exposition [25.3].

This book does not discuss the random matrix theory approach to chaos in quantal spectra; no randomness assumptions are made here, rather the goal is to milk the deterministic chaotic dynamics for its full worth. The book concentrates on the periodic orbit theory. For an introduction to "quantum chaos" that focuses on the random matrix theory the reader is referred to the excellent monograph by Haake [25.4], among others.

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## Chapter 26

## Quantum mechanics, briefly

We start with a review of standard quantum mechanical concepts prerequisite to the derivation of the semiclassical trace formula.

In coordinate representation the time evolution of a quantum mechanical wave function is governed by the Schrödinger equation

$$
\begin{equation*}
i \hbar \frac{\partial}{\partial t} \psi(q, t)=\hat{H}\left(q, \frac{\hbar}{i} \frac{\partial}{\partial q}\right) \psi(q, t), \tag{26.1}
\end{equation*}
$$

where the Hamilton operator $\hat{H}\left(q,-i \hbar \partial_{q}\right)$ is obtained from the classical Hamiltonian by substitution $p \rightarrow-i \hbar \partial_{q}$. Most of the Hamiltonians we shall consider here are of form

$$
\begin{equation*}
H(q, p)=T(p)+V(q), \quad T(p)=p^{2} / 2 m \tag{26.2}
\end{equation*}
$$

describing dynamics of a particle in a $D$-dimensional potential $V(q)$. For time independent Hamiltonians we are interested in finding stationary solutions of the Schrödinger equation of the form

$$
\begin{equation*}
\psi_{n}(q, t)=e^{-i E_{n} t / \hbar} \phi_{n}(q), \tag{26.3}
\end{equation*}
$$

where $E_{n}$ are the eigenenergies of the time-independent Schrödinger equation

$$
\begin{equation*}
\hat{H} \phi(q)=E \phi(q) . \tag{26.4}
\end{equation*}
$$

If the kinetic term can be separated out as in (26.2), the time-independent Schrödinger equation

$$
\begin{equation*}
-\frac{\hbar^{2}}{2 m} \partial^{2} \phi(q)+V(q) \phi(q)=E \phi(q) \tag{26.5}
\end{equation*}
$$

can be rewritten in terms of a local wavenumber

$$
\begin{equation*}
\left(\partial^{2}+k^{2}(q)\right) \phi=0, \quad \hbar^{2} k(q)=\sqrt{2 m(E-V(q))} . \tag{26.6}
\end{equation*}
$$

For bound systems the spectrum is discrete and the eigenfunctions form an orthonormal,

$$
\begin{equation*}
\int d q \phi_{n}(q) \phi_{m}^{*}(q)=\delta_{n m}, \tag{26.7}
\end{equation*}
$$

and complete,

$$
\begin{equation*}
\sum_{n} \phi_{n}(q) \phi_{n}^{*}\left(q^{\prime}\right)=\delta\left(q-q^{\prime}\right), \tag{26.8}
\end{equation*}
$$

set of functions in a Hilbert space. Here and throughout the text,

$$
\begin{equation*}
\int d q=\int d q_{1} d q_{2} \ldots d q_{D} \tag{26.9}
\end{equation*}
$$

For simplicity we will assume that the system is bound, although most of the results will be applicable to open systems, where one has complex resonances instead of real energies, and the spectrum has continuous components.

A given wave function can be expanded in the energy eigenbasis

$$
\begin{equation*}
\psi(q, t)=\sum_{n} c_{n} e^{-i E_{n} t / \hbar} \phi_{n}(q), \tag{26.10}
\end{equation*}
$$

where the expansion coefficient $c_{n}$ is given by the projection of the initial wave function $\psi(q, 0)$ onto the $n$th eigenstate

$$
\begin{equation*}
c_{n}=\int d q \phi_{n}^{*}(q) \psi(q, 0) \tag{26.11}
\end{equation*}
$$

By substituting (26.11) into (26.10), we can cast the evolution of a wave function into a multiplicative form

$$
\psi(q, t)=\int d q^{\prime} K\left(q, q^{\prime}, t\right) \psi\left(q^{\prime}, 0\right)
$$

with the kernel

$$
\begin{equation*}
K\left(q, q^{\prime}, t\right)=\sum_{n} \phi_{n}(q) e^{-i E_{n} t / \hbar} \phi_{n}^{*}\left(q^{\prime}\right) \tag{26.12}
\end{equation*}
$$

called the quantum evolution operator, or the propagator. Applied twice, first for time $t_{1}$ and then for time $t_{2}$, it propagates the initial wave function from $q^{\prime}$ to $q^{\prime \prime}$, and then from $q^{\prime \prime}$ to $q$

$$
\begin{equation*}
K\left(q, q^{\prime}, t_{1}+t_{2}\right)=\int d q^{\prime \prime} K\left(q, q^{\prime \prime}, t_{2}\right) K\left(q^{\prime \prime}, q^{\prime}, t_{1}\right) \tag{26.13}
\end{equation*}
$$

forward in time, hence the name "propagator". In non-relativistic quantum mechanics the range of $q^{\prime \prime}$ is infinite, meaning that the wave can propagate at any speed; in relativistic quantum mechanics this is rectified by restricting the propagation to the forward light cone.

Since the propagator is a linear combination of the eigenfunctions of the Schrödinger equation, it also satisfies the Schrödinger equation

$$
\begin{equation*}
i \hbar \frac{\partial}{\partial t} K\left(q, q^{\prime}, t\right)=\hat{H}\left(q, \frac{i}{\hbar} \frac{\partial}{\partial q}\right) K\left(q, q^{\prime}, t\right), \tag{26.14}
\end{equation*}
$$

and is thus a wave function defined for $t \geq 0$; from the completeness relation (26.8) we obtain the boundary condition at $t=0$ :

$$
\begin{equation*}
\lim _{t \rightarrow 0_{+}} K\left(q, q^{\prime}, t\right)=\delta\left(q-q^{\prime}\right) . \tag{26.15}
\end{equation*}
$$

The propagator thus represents the time evolution of a wave packet which starts out as a configuration space delta-function localized in the point $q^{\prime}$ at the initial time $t=0$.

For time independent Hamiltonians the time dependence of the wave functions is known as soon as the eigenenergies $E_{n}$ and eigenfunctions $\phi_{n}$ have been determined. With time dependence rendered "trivial", it makes sense to focus on the Green's function, the Laplace transformation of the propagator

$$
\begin{equation*}
G\left(q, q^{\prime}, E+i \epsilon\right)=\frac{1}{i \hbar} \int_{0}^{\infty} d t e^{\frac{i}{\hbar} E t-\frac{\epsilon}{\hbar} t} K\left(q, q^{\prime}, t\right)=\sum_{n} \frac{\phi_{n}(q) \phi_{n}^{*}\left(q^{\prime}\right)}{E-E_{n}+i \epsilon} . \tag{26.16}
\end{equation*}
$$

Here $\epsilon$ is a small positive number, ensuring the existence of the integral. The eigenenergies show up as poles in the Green's function with residues corresponding to the wave function amplitudes. If one is only interested in the spectrum, one may restrict the considerations to the (formal) trace of the Green's function,

$$
\begin{equation*}
\operatorname{tr} G\left(q, q^{\prime}, E\right)=\int d q G(q, q, E)=\sum_{n} \frac{1}{E-E_{n}} \tag{26.17}
\end{equation*}
$$



Figure 26.1: Schematic picture of a) the density of states $d(E)$, and $\mathbf{b}$ ) the spectral staircase function $N(E)$. The dashed lines denote the mean density of states $\bar{d}(E)$ and the average number of states $\bar{N}(E)$ discussed in more detail in sect. 30.1.1.
where $E$ is complex, with a positive imaginary part, and we have used the eigenfunction orthonormality (26.7). This trace is formal, since as it stands, the sum in (26.17) is often divergent. We shall return to this point in sects. 30.1.1 and 30.1.2.

A useful characterization of the set of eigenvalues is given in terms of the density of states, with a delta function peak at each eigenenergy, figure 26.1(a),

$$
\begin{equation*}
d(E)=\sum_{n} \delta\left(E-E_{n}\right) . \tag{26.18}
\end{equation*}
$$

Using the identity

$$
\begin{equation*}
\delta\left(E-E_{n}\right)=-\lim _{\epsilon \rightarrow+0} \frac{1}{\pi} \operatorname{Im} \frac{1}{E-E_{n}+i \epsilon} \tag{26.19}
\end{equation*}
$$

we can express the density of states in terms of the trace of the Green's function, that is

$$
\begin{equation*}
d(E)=\sum_{n} \delta\left(E-E_{n}\right)=-\lim _{\epsilon \rightarrow 0} \frac{1}{\pi} \operatorname{Im} \operatorname{tr} G\left(q, q^{\prime}, E+i \epsilon\right) \tag{26.20}
\end{equation*}
$$

As we shall see after "some" work, a semiclassical formula for right hand side of this relation will yield the quantum spectrum in terms of periodic orbits.

The density of states can be written as the derivative $d(E)=\mathrm{d} N(E) / \mathrm{d} E$ of the spectral staircase function

$$
\begin{equation*}
N(E)=\sum_{n} \Theta\left(E-E_{n}\right) \tag{26.21}
\end{equation*}
$$

which counts the number of eigenenergies below $E$, figure 26.1(b). Here $\Theta$ is the Heaviside function

$$
\begin{equation*}
\Theta(x)=1 \quad \text { if } x>0 ; \quad \Theta(x)=0 \quad \text { if } x<0 . \tag{26.22}
\end{equation*}
$$

The spectral staircase is a useful quantity in many contexts, both experimental and theoretical. This completes our lightning review of quantum mechanics.

## Exercises

Exercise 26.1 Dirac delta function, Lorentzian representation. Derive the representation (26.19)

$$
\delta\left(E-E_{n}\right)=-\lim _{\epsilon \rightarrow+0} \frac{1}{\pi} \operatorname{Im} \frac{1}{E-E_{n}+i \epsilon}
$$

of a delta function as imaginary part of $1 / x$.
(Hint: read up on principal parts, positive and negative frequency part of the delta function, the Cauchy theorem in a good quantum mechanics textbook).

Exercise 26.2 Green's function. Verify Green's function Laplace transform (26.16),

$$
\begin{aligned}
G\left(q, q^{\prime}, E+i \varepsilon\right) & =\frac{1}{i \hbar} \int_{0}^{\infty} d t e^{\frac{i}{\hbar} E t-\frac{\varepsilon}{\hbar} t} K\left(q, q^{\prime}, t\right) \\
& =\sum \frac{\phi_{n}(q) \phi_{n}^{*}\left(q^{\prime}\right)}{E-E_{n}+i \varepsilon}
\end{aligned}
$$

argue that positive $\epsilon$ is needed (hint: read a good quantum mechanics textbook).

## Chapter 27

## WKB quantization

The wave function for a particle of energy $E$ moving in a constant potential $V$ is

$$
\begin{equation*}
\psi=A e^{\frac{i}{\hbar} p q} \tag{27.1}
\end{equation*}
$$

with a constant amplitude $A$, and constant wavelength $\lambda=2 \pi / k, k=p / \hbar$, and $p= \pm \sqrt{2 m(E-V)}$ is the momentum. Here we generalize this solution to the case where the potential varies slowly over many wavelengths. This semiclassical (or WKB) approximate solution of the Schrödinger equation fails at classical turning points, configuration space points where the particle momentum vanishes. In such neighborhoods, where the semiclassical approximation fails, one needs to solve locally the exact quantum problem, in order to compute connection coefficients which patch up semiclassical segments into an approximate global wave function.

Two lessons follow. First, semiclassical methods can be very powerful - classical mechanics computations yield suprisingly accurate estimates of quantal spectra, without solving the Schrödinger equation. Second, semiclassical quantization does depend on a purely wave-mechanical phenomena, the coherent addition of phases accrued by all fixed energy phase-space trajectories that connect pairs of coordinate points, and the topological phase loss at every turning point, a topological property of the classical flow that plays no role in classical mechanics.

### 27.1 WKB ansatz

Consider a time-independent Schrödinger equation in 1 spatial dimension:

$$
\begin{equation*}
-\frac{\hbar^{2}}{2 m} \psi^{\prime \prime}(q)+V(q) \psi(q)=E \psi(q) \tag{27.2}
\end{equation*}
$$

Figure 27.1: A 1-dimensional potential, location of the two turning points at fixed energy $E$.

with potential $V(q)$ growing sufficiently fast as $q \rightarrow \pm \infty$ so that the classical particle motion is confined for any $E$. Define the local momentum $p(q)$ and the local wavenumber $k(q)$ by

$$
\begin{equation*}
p(q)= \pm \sqrt{2 m(E-V(q))}, \quad p(q)=\hbar k(q) . \tag{27.3}
\end{equation*}
$$

The variable wavenumber form of the Schrödinger equation

$$
\begin{equation*}
\psi^{\prime \prime}+k^{2}(q) \psi=0 \tag{27.4}
\end{equation*}
$$

sugests that the wave function be written as $\psi=A e^{\frac{2}{\hbar} S}, A$ and $S$ real functions of $q$. Substitution yields two equations, one for the real and other for the imaginary part:

$$
\begin{align*}
\left(S^{\prime}\right)^{2} & =p^{2}+\hbar^{2} \frac{A^{\prime \prime}}{A}  \tag{27.5}\\
S^{\prime \prime} A+2 S^{\prime} A^{\prime} & =\frac{1}{A} \frac{d}{d q}\left(S^{\prime} A^{2}\right)=0 \tag{27.6}
\end{align*}
$$

The Wentzel-Kramers-Brillouin (WKB) or semiclassical approximation consists of dropping the $\hbar^{2}$ term in (27.5). Recalling that $p=\hbar k$, this amounts to assuming that $k^{2} \gg \frac{A^{\prime \prime}}{A}$, which in turn implies that the phase of the wave function is changing much faster than its overall amplitude. So the WKB approximation can interpreted either as a short wavelength/high frequency approximation to a wave-mechanical problem, or as the semiclassical, $\hbar \ll 1$ approximation to quantum mechanics.

Setting $\hbar=0$ and integrating (27.5) we obtain the phase increment of a wave function initially at $q$, at energy $E$

$$
\begin{equation*}
S\left(q, q^{\prime}, E\right)=\int_{q^{\prime}}^{q} d q^{\prime \prime} p\left(q^{\prime \prime}\right) \tag{27.7}
\end{equation*}
$$

This integral over a particle trajectory of constant energy, called the action, will play a key role in all that follows. The integration of (27.6) is even easier

$$
\begin{equation*}
A(q)=\frac{C}{|p(q)|^{\frac{1}{2}}}, \quad C=\left|p\left(q^{\prime}\right)\right|^{\frac{1}{2}} \psi\left(q^{\prime}\right), \tag{27.8}
\end{equation*}
$$

where the integration constant $C$ is fixed by the value of the wave function at the initial point $q^{\prime}$. The $W K B$ (or semiclassical) ansatz wave function is given by

$$
\begin{equation*}
\psi_{s c}\left(q, q^{\prime}, E\right)=\frac{C}{|p(q)|^{\frac{1}{2}}} e^{\frac{i}{\hbar} S\left(q, q^{\prime}, E\right)} \tag{27.9}
\end{equation*}
$$

In what follows we shall supress dependence on the initial point and energy in such formulas, $\left(q, q^{\prime}, E\right) \rightarrow(q)$.

The WKB ansatz generalizes the free motion wave function (27.1), with the probability density $|A(q)|^{2}$ for finding a particle at $q$ now inversely proportional to the velocity at that point, and the phase $\frac{1}{\hbar} q p$ replaced by $\frac{1}{\hbar} \int d q p(q)$, the integrated action along the trajectory. This is fine, except at any turning point $q_{0}$, figure 27.1, where all energy is potential, and

$$
\begin{equation*}
p(q) \rightarrow 0 \quad \text { as } \quad q \rightarrow q_{0}, \tag{27.10}
\end{equation*}
$$

so that the assumption that $k^{2} \gg \frac{A^{\prime \prime}}{A}$ fails. What can one do in this case?
For the task at hand, a simple physical picture, due to Maslov, does the job. In the $q$ coordinate, the turning points are defined by the zero kinetic energy condition (see figure 27.1), and the motion appears singular. This is not so in the full phase space: the trajectory in a smooth confining 1-dimensional potential is always a smooth loop, with the "special" role of the turning points $q_{L}, q_{R}$ seen to be an artifact of a particular choice of the ( $q, p$ ) coordinate frame. Maslov's idea was to proceed from the initial point $\left(q^{\prime}, p^{\prime}\right)$ to a point $\left(q_{A}, p_{A}\right)$ preceeding the turning point in the $\psi(q)$ representation, then switch to the momentum representation

$$
\begin{equation*}
\widetilde{\psi}(p)=\frac{1}{\sqrt{2 \pi \hbar}} \int d q e^{-\frac{i}{\hbar} q p} \psi(q) \tag{27.11}
\end{equation*}
$$

continue from $\left(q_{A}, p_{A}\right)$ to $\left(q_{B}, p_{B}\right)$, switch back to the coordinate representation,

$$
\begin{equation*}
\psi(q)=\frac{1}{\sqrt{2 \pi \hbar}} \int d p e^{\frac{i}{\hbar} q p} \widetilde{\psi}(p) \tag{27.12}
\end{equation*}
$$

and so on.
The only rub is that one usually cannot evaluate these transforms exactly. But, as the WKB wave function (27.9) is approximate anyway, it suffices to estimate these transforms to leading order in $\hbar$ accuracy. This is accomplished by the method of stationary phase.

Figure 27.2: A 1-dof phase space trajectory of a particle moving in a bound potential.


### 27.2 Method of stationary phase

All "semiclassical" approximations are based on saddlepoint evaluations of integrals of the type

$$
\begin{equation*}
I=\int d x A(x) e^{i s \Phi(x)}, \quad x, \Phi(x) \in \mathbb{R} \tag{27.13}
\end{equation*}
$$

where $s$ is assumed to be a large, real parameter, and $\Phi(x)$ is a real-valued function. In our applications $s=1 / \hbar$ will always be assumed large.

For large $s$, the phase oscillates rapidly and "averages to zero" everywhere except at the extremal points $\Phi^{\prime}\left(x_{0}\right)=0$. The method of approximating an integral by its values at extremal points is called the method of stationary phase. Consider first the case of a 1-dimensional integral, and expand $\Phi\left(x_{0}+\delta x\right)$ around $x_{0}$ to second order in $\delta x$,

$$
\begin{equation*}
I=\int d x A(x) e^{i s\left(\Phi\left(x_{0}\right)+\frac{1}{2} \Phi^{\prime \prime}\left(x_{0}\right) \delta x^{2}+\ldots\right)} . \tag{27.14}
\end{equation*}
$$

Assume (for time being) that $\Phi^{\prime \prime}\left(x_{0}\right) \neq 0$, with either sign, $\operatorname{sgn}\left[\Phi^{\prime \prime}\right]=$ $\Phi^{\prime \prime} /\left|\Phi^{\prime \prime}\right|= \pm 1$. If in the neighborhood of $x_{0}$ the amplitude $A(x)$ varies slowly over many oscillations of the exponential function, we may retain the leading term in the Taylor expansion of the amplitude, and approximate the integral up to quadratic terms in the phase by

$$
\begin{equation*}
I \approx A\left(x_{0}\right) e^{i s \Phi\left(x_{0}\right)} \int d x e^{\frac{1}{2} i s \Phi^{\prime \prime}\left(x_{0}\right)\left(x-x_{0}\right)^{2}} \tag{27.15}
\end{equation*}
$$

Using the Fresnel integral formula

$$
\begin{equation*}
\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} d x e^{-\frac{x^{2}}{2 i a}}=\sqrt{i a}=|a|^{1 / 2} e^{i \frac{\pi}{4} \frac{a}{|a|}} \tag{27.16}
\end{equation*}
$$

we obtain

$$
\begin{equation*}
I \approx A\left(x_{0}\right)\left|\frac{2 \pi}{s \Phi^{\prime \prime}\left(x_{0}\right)}\right|^{1 / 2} e^{i s \Phi\left(x_{0}\right) \pm i \frac{\pi}{4}} \tag{27.17}
\end{equation*}
$$

where $\pm$ corresponds to the positive/negative sign of $s \Phi^{\prime \prime}\left(x_{0}\right)$.

### 27.3 WKB quantization

We can now evaluate the Fourier transforms (27.11), (27.12) to the same order in $\hbar$ as the WKB wave function using the stationary phase method,

$$
\begin{align*}
\widetilde{\psi}_{s c}(p) & =\frac{C}{\sqrt{2 \pi \hbar}} \int \frac{d q}{\mid p(q))^{\frac{1}{2}}} e^{\frac{i}{\hbar}(S(q)-q p)} \\
& \approx \frac{C}{\sqrt{2 \pi \hbar}} \frac{e^{\frac{i}{\hbar}\left(S\left(q^{*}\right)-q^{*} p\right)}}{\left|p\left(q^{*}\right)\right|^{\frac{1}{2}}} \int d q e^{\frac{i}{2 \hbar} S^{\prime \prime}\left(q^{*}\right)\left(q-q^{*}\right)^{2}}, \tag{27.18}
\end{align*}
$$

where $q^{*}$ is given implicitly by the stationary phase condition

$$
0=S^{\prime}\left(q^{*}\right)-p=p\left(q^{*}\right)-p
$$

and the sign of $S^{\prime \prime}\left(q^{*}\right)=p^{\prime}\left(q^{*}\right)$ determines the phase of the Fresnel integral (27.16)

$$
\begin{equation*}
\widetilde{\psi}_{s c}(p)=\frac{C}{\left|p\left(q^{*}\right) p^{\prime}\left(q^{*}\right)\right|^{\frac{1}{2}}} e^{\frac{i}{\hbar}\left[S\left(q^{*}\right)-q^{*} p\right]+\frac{i \pi}{4} \operatorname{sgn}\left[S^{\prime \prime}\left(q^{*}\right)\right]} . \tag{27.19}
\end{equation*}
$$

As we continue from $\left(q_{A}, p_{A}\right)$ to $\left(q_{B}, p_{B}\right)$, nothing problematic occurrs $p\left(q^{*}\right)$ is finite, and so is the acceleration $p^{\prime}\left(q^{*}\right)$. Otherwise, the trajectory would take infinitely long to get across. We recognize the exponent as the Legendre transform

$$
\tilde{S}(p)=S(q(p))-q(p) p
$$

which can be used to expresses everything in terms of the $p$ variable,

$$
\begin{equation*}
q^{*}=q(p), \quad \frac{d}{d q} q=1=\frac{d p}{d q} \frac{d q(p)}{d p}=q^{\prime}(p) p^{\prime}\left(q^{*}\right) . \tag{27.20}
\end{equation*}
$$

As the classical trajectory crosses $q_{L}$, the weight in (27.19),

$$
\begin{equation*}
\frac{d}{d q} p^{2}\left(q_{L}\right)=2 p\left(q_{L}\right) p^{\prime}\left(q_{L}\right)=-2 m V^{\prime}(q) \tag{27.21}
\end{equation*}
$$

is finite, and $S^{\prime \prime}\left(q^{*}\right)=p^{\prime}\left(q^{*}\right)<0$ for any point in the lower left quadrant, including $\left(q_{A}, p_{A}\right)$. Hence, the phase loss in (27.19) is $-\frac{\pi}{4}$. To go back from the $p$ to the $q$ representation, just turn figure $27.290^{\circ}$ anticlockwise. Everything is the same if you replace $(q, p) \rightarrow(-p, q)$; so, without much ado we get the semiclassical wave function at the point $\left(q_{B}, p_{B}\right)$,

$$
\begin{equation*}
\psi_{s c}(q)=\frac{e^{\frac{i}{\hbar}\left(\tilde{S}\left(p^{*}\right)+q p^{*}\right)-\frac{i \pi}{4}}}{\left|q^{*}\left(p^{*}\right)\right|^{\frac{1}{2}}} \widetilde{\psi}_{s c}\left(p^{*}\right)=\frac{C}{|p(q)|^{\frac{1}{2}}} e^{\frac{i}{\hbar} S(q)-\frac{i \pi}{2}} . \tag{27.22}
\end{equation*}
$$

Figure 27.3: $\quad S_{p}(E)$, the action of a periodic orbit $p$ at energy $E$, equals the area in the phase space traced out by the 1-dof trajectory.


The extra $\left|p^{\prime}\left(q^{*}\right)\right|^{1 / 2}$ weight in (27.19) is cancelled by the $\left|q^{\prime}\left(p^{*}\right)\right|^{1 / 2}$ term, by the Legendre relation (27.20).

The message is that going through a smooth potential turning point the WKB wave function phase slips by $-\frac{\pi}{2}$. This is equally true for the right and the left turning points, as can be seen by rotating figure 27.2 by $180^{\circ}$, and flipping coordinates $(q, p) \rightarrow(-q,-p)$. While a turning point is not an invariant concept (for a sufficiently short trajectory segment, it can be undone by a $45^{\circ}$ turn), for a complete period $(q, p)=\left(q^{\prime}, p^{\prime}\right)$ the total phase slip is always $-2 \cdot \pi / 2$, as a loop always has $m=2$ turning points.

The WKB quantization condition follows by demanding that the wave function computed after a complete period be single-valued. With the normalization (27.8), we obtain

$$
\psi\left(q^{\prime}\right)=\psi(q)=\left|\frac{p\left(q^{\prime}\right)}{p(q)}\right|^{\frac{1}{2}} e^{i\left(\frac{1}{\hbar} \oint p(q) d q-\pi\right)} \psi\left(q^{\prime}\right) .
$$

The prefactor is 1 by the periodic orbit condition $q=q^{\prime}$, so the phase must be a multiple of $2 \pi$,

$$
\begin{equation*}
\frac{1}{\hbar} \oint p(q) d q=2 \pi\left(n+\frac{m}{4}\right), \tag{27.23}
\end{equation*}
$$

where $m$ is the number of turning points along the trajectory - for this 1 -dof problem, $m=2$.

The action integral in (27.23) is the area (see figure 27.3) enclosed by the classical phase space loop of figure 27.2, and the quantization condition says that eigenenergies correspond to loops whose action is an integer multiple of the unit quantum of action, Planck's constant $\hbar$. The extra topological phase, which, although it had been discovered many times in centuries past, had to wait for its most recent quantum chaotic (re)birth until the 1970's. Despite its derivation in a noninvariant coordinate frame, the final result involves only canonically invariant classical quantities, the periodic orbit action $S$, and the topological index $m$.

Figure 27.4: Airy function $\operatorname{Ai}(q)$.


### 27.3.1 Harmonic oscillator quantization

Let us check the WKB quantization for one case (the only case?) whose quantum mechanics we fully understand: the harmonic oscillator

$$
E=\frac{1}{2 m}\left(p^{2}+(m \omega q)^{2}\right) .
$$

The loop in figure 27.2 is now a circle in the $(m \omega q, p)$ plane, the action is its area $S=2 \pi E / \omega$, and the spectrum in the WKB approximation

$$
\begin{equation*}
E_{n}=\hbar \omega(n+1 / 2) \tag{27.24}
\end{equation*}
$$

turns out to be the exact harmonic oscillator spectrum. The stationary phase condition (27.18) keeps $V(q)$ accurate to order $q^{2}$, which in this case is the whole answer (but we were simply lucky, really). For many 1-dof problems the WKB spectrum turns out to be very accurate all the way down to the ground state. Surprisingly accurate, if one interprets dropping the $\hbar^{2}$ term in (27.5) as a short wavelength approximation.

### 27.4 Beyond the quadratic saddle point

We showed, with a bit of Fresnel/Maslov voodoo, that in a smoothly varying potential the phase of the WKB wave function slips by a $\pi / 2$ for each turning point. This $\pi / 2$ came from a $\sqrt{i}$ in the Fresnel integral (27.16), one such factor for every time we switched representation from the configuration space to the momentum space, or back. Good, but what does this mean?

The stationary phase approximation (27.14) fails whenever $\Phi^{\prime \prime}(x)=0$, or, in our the WKB ansatz (27.18), whenever the momentum $p^{\prime}(q)=S^{\prime \prime}(q)$ vanishes. In that case we have to go beyond the quadratic approximation (27.15) to the first nonvanishing term in the Taylor expansion of the exponent. If $\Phi^{\prime \prime \prime}\left(x_{0}\right) \neq 0$, then

$$
\begin{equation*}
I \approx A\left(x_{0}\right) e^{i s \Phi\left(x_{0}\right)} \int_{-\infty}^{\infty} d x e^{i s \Phi^{\prime \prime \prime}\left(x_{0}\right) \frac{\left(x-x_{0}\right)^{3}}{6}} . \tag{27.25}
\end{equation*}
$$

Airy functions can be represented by integrals of the form

$$
\begin{equation*}
A i(x)=\frac{1}{2 \pi} \int_{-\infty}^{+\infty} d y e^{i\left(x y-\frac{y^{3}}{3}\right)} . \tag{27.26}
\end{equation*}
$$

Derivations of the WKB quantization condition given in standard quantum mechanics textbooks rely on expanding the potential close to the turning point

$$
V(q)=V\left(q_{0}\right)+\left(q-q_{0}\right) V^{\prime}\left(q_{0}\right)+\cdots,
$$

solving the Airy equation

$$
\begin{equation*}
\psi^{\prime \prime}=q \psi, \tag{27.27}
\end{equation*}
$$

and matching the oscillatory and the exponentially decaying "forbidden" region wave function pieces by means of the WKB connection formulas. That requires staring at Airy functions and learning about their asymptotics - a challenge that we will have to eventually overcome, in order to incorporate diffraction phenomena into semiclassical quantization.
2) what does the wave function look like?
3) generically useful when Gaussian approximations fail

The physical origin of the topological phase is illustrated by the shape of the Airy function, figure 27.4. For a potential with a finite slope $V^{\prime}(q)$ the wave function pentrates into the forbidden region, and accomodates a bit more of a stationary wavelength then what one would expect from the classical trajectory alone. For infinite walls (that is, billiards) a different argument applies: the wave function must vanish at the wall, and the phase slip due to a specular reflection is $-\pi$, rather than $-\pi / 2$.

## Commentary

Remark 27.1 Airy function. The stationary phase approximation is all that is needed for the semiclassical approximation, with the proviso that $D$ in (28.36) has no zero eigenvalues. The zero eigenvalue case would require going beyond the Gaussian saddle-point approximation, which typically leads to approximations of the integrals in terms of Airy functions [27.10].

Remark 27.2 Bohr-Sommerfeld quantization. Bohr-Sommerfeld quantization condition was the key result of the old quantum theory, in which the electron trajectories were purely classical. They were lucky - the symmetries of the Kepler problem work out in such a way that the total topological index $m=4$ amount effectively to numbering the energy levels starting with $n=1$. They were unlucky - because the hydrogen $m=4$ masked the topological index, they could never get the helium spectrum right - the semiclassical calculation had to wait for until 1980, when Leopold and Percival [A.6] added the topological indices.

## Résumé

The WKB ansatz wave function for 1-degree of freedom problems fails at the turning points of the classical trajectory. While in the $q$-representation the WKB ansatz a turning point is singular, along the $p$ direction the classical trajectory in the same neighborhood is smooth, as for any smooth bound potential the classical motion is topologically a circle around the origin in the ( $q, p$ ) space. The simplest way to deal with such singularities is as follows; follow the classical trajectory in $q$-space until the WKB approximation fails close to the turning point; then insert $\int d p|p\rangle\langle p|$ and follow the classical trajectory in the $p$-space until you encounter the next $p$-space turning point; go back to the $q$-space representation, an so on. Each matching involves a Fresnel integral, yielding an extra $e^{-i \pi / 4}$ phase shift, for a total of $e^{-i \pi}$ phase shift for a full period of a semiclassical particle moving in a soft potential. The condition that the wave-function be singlevalued then leads to the 1-dimensional WKB quantization, and its lucky cousin, the Bohr-Sommerfeld quantization.

Alternatively, one can linearize the potential around the turning point $a, V(q)=V(a)+(q-a) V^{\prime}(a)+\cdots$, and solve the quantum mechanical constant linear potential $V(q)=q F$ problem exactly, in terms of an Airy function. An approximate wave function is then patched together from an Airy function at each turning point, and the WKB ansatz wave-function segments inbetween via the WKB connection formulas. The single-valuedness condition again yields the 1-dimensional WKB quantization. This a bit more work than tracking the classical trajectory in the full phase space, but it gives us a better feeling for shapes of quantum eigenfunctions, and exemplifies the general strategy for dealing with other singularities, such as wedges, bifurcation points, creeping and tunneling: patch together the WKB segments by means of exact QM solutions to local approximations to singular points.

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## Exercises

## Exercise 27.1

WKB ansatz. Try to show that no other ansatz other than (28.1) gives a meaningful definition of the momentum in the $\hbar \rightarrow 0$ limit.

## Exercise 27.2 Fresnel integral. Derive the Fresnel integral

$$
\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} d x e^{-\frac{x^{2}}{2 i a}}=\sqrt{i a}=|a|^{1 / 2} e^{i \frac{\pi}{4} \frac{a}{|a|}}
$$

Exercise 27.3 Sterling formula for $n!$. Compute an approximate value of $n$ ! for large $n$ using the stationary phase approximation. Hint: $n!=\int_{0}^{\infty} d t t^{n} e^{-t}$.

## Exercise 27.4 Airy function for large arguments.

 contributions as stationary phase points may arise from extremal points where the first non-zero term in a Taylor expansion of the phase is of third or higher order. Such situations occur, for example, at bifurcation points or in diffraction effects, (such as waves near sharp corners, waves creeping around obstacles, etc.). In such calculations, one meets Airy functions integrals of the form$$
\begin{equation*}
A i(x)=\frac{1}{2 \pi} \int_{-\infty}^{+\infty} d y e^{i\left(x y-\frac{y^{3}}{3}\right)} \tag{27.28}
\end{equation*}
$$

Calculate the Airy function $A i(x)$ using the stationary phase approximation. What happens when considering the limit $x \rightarrow 0$. Estimate for which value of $x$ the stationary phase approximation breaks down.

## Chapter 28

## Semiclassical evolution


#### Abstract

William Rowan Hamilton was born in 1805. At three he could read English; by four he began to read Latin, Greek and Hebrew, by ten he read Sanskrit, Persian, Arabic, Chaldee, Syrian and sundry Indian dialects. At age seventeen he began to think about optics, and worked out his great principle of "Characteristic Function". Turnbull, Lives of Mathematicians


(G. Vattay, G. Tanner and P. Cvitanović)

Semiclassical approximations to quantum mechanics are valid in the regime where the de Broglie wavelength $\lambda \sim \hbar / p$ of a particle with momentum $p$ is much shorter than the length scales across which the potential of the system changes significantly. In the short wavelength approximation the particle is a point-like object bouncing off potential walls, the same way it does in the classical mechanics. The novelty of quantum mechanics is the interference of the point-like particle with other versions of itself traveling along different classical trajectories, a feat impossible in classical mechanics. The short wavelength - or semiclassical - formalism is developed by formally taking the limit $\hbar \rightarrow 0$ in quantum mechanics in such a way that quantum quantities go to their classical counterparts.

### 28.1 Hamilton-Jacobi theory

We saw in chapter 27 that for a 1-dof particle moving in a slowly varying potential, it makes sense to generalize the free particle wave function (27.1) to a wave function

$$
\begin{equation*}
\psi(q, t)=A(q, t) e^{i R(q, t) / \hbar} \tag{28.1}
\end{equation*}
$$

with slowly varying (real) amplitude $A(q, t)$ and rapidly varying (real) phase $R(q, t)$. its phase and magnitude. The time evolution of the phase and
the magnitude of $\psi$ follows from the Schrödinger equation (26.1)

$$
\begin{equation*}
\left(i \hbar \frac{\partial}{\partial t}+\frac{\hbar^{2}}{2 m} \frac{\partial^{2}}{\partial q^{2}}-V(q)\right) \psi(q, t)=0 . \tag{28.2}
\end{equation*}
$$

Assume $A \neq 0$, and separate out the real and the imaginary parts. We get two equations: The real part governs the time evolution of the phase

$$
\begin{equation*}
\frac{\partial R}{\partial t}+\frac{1}{2 m}\left(\frac{\partial R}{\partial q}\right)^{2}+V(q)-\frac{\hbar^{2}}{2 m} \frac{1}{A} \frac{\partial^{2}}{\partial q^{2}} A=0 \tag{28.3}
\end{equation*}
$$

and the imaginary part the time evolution of the amplitude

$$
\begin{equation*}
\frac{\partial A}{\partial t}+\frac{1}{m} \sum_{i=1}^{D} \frac{\partial A}{\partial q_{i}} \frac{\partial R}{\partial q_{i}}+\frac{1}{2 m} A \frac{\partial^{2} R}{\partial q^{2}}=0 . \tag{28.4}
\end{equation*}
$$

In this way a linear PDE for a complex wave function is converted into a set of coupled non-linear PDE's for real-valued functions $R$ and $A$. The coupling term in (28.3) is, however, of order $\hbar^{2}$ and thus small in the semiclassical limit $\hbar \rightarrow 0$.

Now we generalize the Wentzel-Kramers-Brillouin (WKB) ansatz for 1-dof dynamics to the Van Vleck ansatz in arbitrary dimension: we assume the magnitude $A(q, t)$ varies slowly compared to the phase $R(q, t) / \hbar$, so we drop the $\hbar$-dependent term. In this approximation the phase $R(q, t)$ and the corresponding "momentum field" $\frac{\partial R}{\partial q}(q, t)$ can be determined from the amplitude independent equation

$$
\begin{equation*}
\frac{\partial R}{\partial t}+H\left(q, \frac{\partial R}{\partial q}\right)=0 \tag{28.5}
\end{equation*}
$$

In classical mechanics this equation is known as the Hamilton-Jacobi equation. We will refer to this step (as well as all leading order in $\hbar$ approximations to follow) as the semiclassical approximation to wave mechanics, and from now on work only within this approximation.

### 28.1.1 Hamilton's equations

We now solve the nonlinear partial differential equation (28.5) in a way the 17 year old Hamilton might have solved it. The main step is the step leading from the nonlinear PDE (28.9) to Hamilton's ODEs (28.10). If you already understand the Hamilton-Jacobi theory, you can safely skip this section.


Figure 28.1: (a) A phase $R(q, t)$ plotted as a function of the position $q$ for two infinitesimally close times. (b) The phase $R(q, t)$ transported by a swarm of "particles"; The Hamilton's equations (28.10) construct $R(q, t)$ by transporting $q_{0} \rightarrow q(t)$ and the slope of $R\left(q_{0}, t_{0}\right)$, that is $p_{0} \rightarrow p(t)$.

The wave equation (26.1) describes how the wave function $\psi$ evolves with time, and if you think of $\psi$ as an (infinite dimensional) vector, position $q$ plays a role of an index. In one spatial dimension the phase $R$ plotted as a function of the position $q$ for two different times looks something like figure 28.1(a): The phase $R\left(q, t_{0}\right)$ deforms smoothly with time into the phase $R(q, t)$ at time $t$. Hamilton's idea was to let a swarm of particles transport $R$ and its slope $\partial R / \partial q$ at $q$ at initial time $t=t_{0}$ to a corresponding $R(q, t)$ and its slope at time $t$, figure 28.1(b). For notational convenience, define

$$
\begin{equation*}
p_{i}=p_{i}(q, t):=\frac{\partial R}{\partial q_{i}}, \quad i=1,2, \ldots, D . \tag{28.6}
\end{equation*}
$$

We saw earlier that (28.3) reduces in the semiclassical approximation to the Hamilton-Jacobi equation (28.5). To make life simple, we shall assume throughout this chapter that the Hamilton's function $H(q, p)$ does not depend explicitly on time $t$, that is, the energy is conserved.

To start with, we also assume that the function $R(q, t)$ is smooth and well defined for every $q$ at the initial time $t$. This is true for sufficiently short times; as we will see later, $R$ develops folds and becomes multi-valued as $t$ progresses. Consider now the variation of the function $R(q, t)$ with respect to independent infinitesimal variations of the time and space coordinates $d t$ and $d q$, figure 28.1(a)

$$
\begin{equation*}
d R=\frac{\partial R}{\partial t} d t+\frac{\partial R}{\partial q} d q \tag{28.7}
\end{equation*}
$$

Dividing through by $d t$ and substituting (28.5) we obtain the total derivative of $R(q, t)$ with respect to time along the as yet arbitrary direction $\dot{q}$, that is,

$$
\begin{equation*}
\frac{d R}{d t}(q, \dot{q}, t)=-H(q, p)+\dot{q} \cdot p . \tag{28.8}
\end{equation*}
$$

Note that the "momentum" $p=\partial R / \partial q$ is a well defined function of $q$ and $t$. In order to integrate $R(q, t)$ with the help of (28.8) we also need to know how $p=\partial R / \partial q$ changes along $\dot{q}$. Varying $p$ with respect to independent infinitesimal variations $d t$ and $d q$ and substituting the Hamilton-Jacobi equation (28.5) yields

$$
d \frac{\partial R}{\partial q}=\frac{\partial^{2} R}{\partial q \partial t} d t+\frac{\partial^{2} R}{\partial q^{2}} d q=-\left(\frac{\partial H}{\partial q}+\frac{\partial H}{\partial p} \frac{\partial p}{\partial q}\right) d t+\frac{\partial p}{\partial q} d q
$$

Note that $H(q, p)$ depends on $q$ also through $p(q, t)=\partial R / \partial q$, hence the $\frac{\partial H}{\partial p}$ term in the above equation. Dividing again through by $d t$ we get the time derivative of $\partial R / \partial q$, that is,

$$
\begin{equation*}
\dot{p}(q, \dot{q}, t)+\frac{\partial H}{\partial q}=\left(\dot{q}-\frac{\partial H}{\partial p}\right) \frac{\partial p}{\partial q} . \tag{28.9}
\end{equation*}
$$

Time variation of $p$ depends not only on the yet unknown $\dot{q}$, but also on the second derivatives of $R$ with respect to $q$ with yet unknown time dependence. However, if we choose $\dot{q}$ (which was arbitrary, so far) such that the right hand side of the above equation vanishes, we can calculate the function $R(q, t)$ along a specific trajectory $(q(t), p(t))$ given by integrating the ordinary differential equations

$$
\begin{equation*}
\dot{q}=\frac{\partial H(q, p)}{\partial p}, \quad \dot{p}=-\frac{\partial H(q, p)}{\partial q} \tag{28.10}
\end{equation*}
$$

with initial conditions

$$
\begin{equation*}
q\left(t_{0}\right)=q^{\prime}, \quad p\left(t_{0}\right)=p^{\prime}=\frac{\partial R}{\partial q}\left(q^{\prime}, t_{0}\right) . \tag{28.11}
\end{equation*}
$$

We recognize (28.10) as Hamilton's equations of motion of classical mechanics. The miracle happens in the step leading from (28.5) to (28.9) - if you missed it, you have missed the point. Hamilton derived his equations contemplating optics - it took him three more years to realize that all of Newtonian dynamics can be profitably recast in this form.
$\dot{q}$ is no longer an independent function, and the phase $R(q, t)$ can now be computed by integrating equation (28.8) along the trajectory $(q(t), p(t))$

$$
\begin{align*}
R(q, t) & =R\left(q^{\prime}, t_{0}\right)+R\left(q, t ; q^{\prime}, t_{0}\right) \\
R\left(q, t ; q^{\prime}, t_{0}\right) & =\int_{t_{0}}^{t} d \tau[\dot{q}(\tau) \cdot p(\tau)-H(q(\tau), p(\tau))] \tag{28.12}
\end{align*}
$$

with the initial conditions (28.11). In this way the Hamilton-Jacobi partial differential equation (28.3) is solved by integrating a set of ordinary
differential equations, Hamilton's equations. In order to determine $R(q, t)$ for arbitrary $q$ and $t$ we have to find a $q^{\prime}$ such that the trajectory starting in $\left(q^{\prime}, p^{\prime}=\partial_{q} R\left(q^{\prime}, t_{0}\right)\right)$ reaches $q$ in time $t$ and then compute $R$ along this trajectory, see figure 28.1(b). The integrand of (28.12) is known as the Lagrangian,

$$
\begin{equation*}
L(q, \dot{q}, t)=\dot{q} \cdot p-H(q, p, t) . \tag{28.13}
\end{equation*}
$$

A variational principle lurks here, but we shall not make much fuss about it as yet.

Throughout this chapter we assume that the energy is conserved, and that the only time dependence of $H(q, p)$ is through $(q(\tau), p(\tau))$, so the value of $R\left(q, t ; q^{\prime}, t_{0}\right)$ does not depend on $t_{0}$, but only on the elapsed time $t-t_{0}$. To simplify notation we will set $t_{0}=0$ and write

$$
R\left(q, q^{\prime}, t\right)=R\left(q, t ; q^{\prime}, 0\right) .
$$

The initial momentum of the particle must coincide with the initial momentum of the trajectory connecting $q^{\prime}$ and $q$ :

$$
\begin{equation*}
p^{\prime}=\frac{\partial}{\partial q^{\prime}} R\left(q^{\prime}, 0\right)=-\frac{\partial}{\partial q^{\prime}} R\left(q, q^{\prime}, t\right) . \tag{28.14}
\end{equation*}
$$

The function $R\left(q, q^{\prime}, t\right)$ is known as Hamilton's principal function.
To summarize: Hamilton's achievement was to trade in the HamiltonJacobi partial differential equation (28.5) describing the evolution of a wave front for a finite number of ordinary differential equations of motion, with the initial phase $R(q, 0)$ incremented by the integral (28.12) evaluated along the phase space trajectory $(q(\tau), p(\tau))$.

### 28.1.2 Action

Before proceeding, we note in passing a few facts about Hamiltonian dynamics that will be needed for the construction of semiclassical Green's functions. If the energy is conserved, the $\int H(q, p) d \tau$ integral in (28.12) is simply Et. The first term, or the action

$$
\begin{equation*}
S\left(q, q^{\prime}, E\right)=\int_{0}^{t} d \tau \dot{q}(\tau) \cdot p(\tau)=\int_{q^{\prime}}^{q} d q \cdot p \tag{28.15}
\end{equation*}
$$

is integrated along a trajectory from $q^{\prime}$ to $q$ with a fixed energy $E$. By (28.12) the action is a Legendre transform of Hamilton's principal function

$$
\begin{equation*}
S\left(q, q^{\prime}, E\right)=R\left(q, q^{\prime}, t\right)+E t . \tag{28.16}
\end{equation*}
$$

The time of flight $t$ along the trajectory connecting $q^{\prime} \rightarrow q$ with fixed energy $E$ is given by

$$
\begin{equation*}
\frac{\partial}{\partial E} S\left(q, q^{\prime}, E\right)=t \tag{28.17}
\end{equation*}
$$

The way to think about the formula (28.16) for action is that the time of flight is a function of the energy, $t=t\left(q, q^{\prime}, E\right)$. The left hand side is explicitly a function of $E$; the right hand side is an implicit function of $E$ through energy dependence of the flight time $t$.

Going in the opposite direction, the energy of a trajectory $E=E\left(q, q^{\prime}, t\right)$ connecting $q^{\prime} \rightarrow q$ with a given time of flight $t$ is given by the derivative of Hamilton's principal function

$$
\begin{equation*}
\frac{\partial}{\partial t} R\left(q, q^{\prime}, t\right)=-E \tag{28.18}
\end{equation*}
$$

and the second variations of $R$ and $S$ are related in the standard way of Legendre transforms:

$$
\begin{equation*}
\frac{\partial^{2}}{\partial t^{2}} R\left(q, q^{\prime}, t\right) \frac{\partial^{2}}{\partial E^{2}} S\left(q, q^{\prime}, E\right)=-1 \tag{28.19}
\end{equation*}
$$

A geometric visualization of what the phase evolution looks like is very helpful in understanding the origin of topological indices to be introduced in what follows. Given an initial phase $R\left(q, t_{0}\right)$, the gradient $\partial_{q} R$ defines a $D$-dimensional Lagrangian manifold $\left(q, p=\partial_{q} R(q)\right)$ in the full $2 d$ dimensional phase space $(q, p)$. The defining property of this manifold is that any contractible loop $\gamma$ in it has zero action,

$$
0=\oint_{\gamma} d q \cdot p
$$

a fact that follows from the definition of $p$ as a gradient, and the Stokes theorem. Hamilton's equations of motion preserve this property and map a Lagrangian manifold into a Lagrangian manifold at a later time. $t$

Returning back to the main line of our argument: so far we have determined the wave function phase $R(q, t)$. Next we show that the velocity field given by the Hamilton's equations together with the continuity equation determines the amplitude of the wave function.

### 28.1.3 Density evolution

To obtain the full solution of the Schrödinger equation (26.1), we also have to integrate (28.4).

$$
\rho(q, t):=A^{2}=\psi^{*} \psi
$$

plays the role of a density. To the leding order in $\hbar$, the gradient of $R$ may be interpreted as the semiclassical momentum density

$$
\psi(q, t)^{*}\left(-i \hbar \frac{\partial}{\partial q}\right) \psi(q, t)=-i \hbar A \frac{\partial A}{\partial q}+\rho \frac{\partial R}{\partial q} .
$$

Evaluated along the trajectory $(q(t), p(t))$, the amplitude equation (28.4) is equivalent to the continuity equation (9.34) after multiplying (28.4) by $2 A$, that is

$$
\begin{equation*}
\frac{\partial \rho}{\partial t}+\frac{\partial}{\partial q_{i}}\left(\rho v_{i}\right)=0 . \tag{28.20}
\end{equation*}
$$

Here, $v_{i}=\dot{q}_{i}=p_{i} / m$ denotes a velocity field, which is in turn determined by the gradient of $R(q, t)$, or the Lagrangian manifold $\left(q(t), p(t)=\partial_{q} R(q, t)\right)$,

$$
v=\frac{1}{m} \frac{\partial}{\partial q} R(q, t) .
$$

As we already know how to solve the Hamilton-Jacobi equation (28.5), we can also solve for the density evolution as follows:

The density $\rho(q)$ can be visualized as the density of a configuration space flow $q(t)$ of a swarm of hypothetical particles; the trajectories $q(t)$ are solutions of Hamilton's equations with initial conditions given by $(q(0)=$ $\left.q^{\prime}, p(0)=p^{\prime}=\partial_{q} R\left(q^{\prime}, 0\right)\right)$.

If we take a small configuration space volume $d^{D} q$ around some point $q$ at time $t$, then the number of particles in it is $\rho(q, t) d^{D} d q$. They started initially in a small volume $d^{D} q^{\prime}$ around the point $q^{\prime}$ of the configuration space. For the moment, we assume that there is only one solution, the case of several paths will be considered below. The number of particles at time $t$ in the volume is the same as the number of particles in the initial volume at $t=0$,

$$
\rho(q(t), t) d^{D} q=\rho\left(q^{\prime}, 0\right) d^{D} q^{\prime}
$$

see figure 28.2. The ratio of the initial and the final volumes can be expressed as

$$
\begin{equation*}
\rho(q(t), t)=\left|\operatorname{det} \frac{\partial q^{\prime}}{\partial q}\right| \rho\left(q^{\prime}, 0\right) . \tag{28.21}
\end{equation*}
$$

As we know how to compute trajectories $(q(t), p(t))$, we know how to compute this Jacobian and, by (28.21), the density $\rho(q(t), t)$ at time $t$.

Figure 28.2: Density evolution of an initial surface $\left(q^{\prime}, p^{\prime}=\partial_{q} R\left(q^{\prime}, 0\right)\right.$ into $(q(t), p(t))$ surface time $t$ later, sketched in 1 dimension. While the number of trajectories and the phase space Liouville volume are conserved, the density of trajectories projected on the $q$ coordinate varies; trajectories which started in $d q^{\prime}$ at time zero end up in the interval $d q$.


### 28.1.4 Semiclassical wave function

Now we have all ingredients to write down the semiclassical wave function at time $t$. Consider first the case when our initial wave function can be written in terms of single-valued functions $A\left(q^{\prime}, 0\right)$ and $R\left(q^{\prime}, 0\right)$. For sufficiently short times, $R(q, t)$ will remain a single-valued function of $q$, and every $d^{D} q$ configuration space volume element keeps its orientation. The evolved wave function is in the semiclassical approximation then given by

$$
\begin{aligned}
\psi_{s c}(q, t) & =A(q, t) e^{i R(q, t) / \hbar}=\sqrt{\operatorname{det} \frac{\partial q^{\prime}}{\partial q}} A\left(q^{\prime}, 0\right) e^{i\left(R\left(q^{\prime}, 0\right)+R\left(q, q^{\prime}, t\right)\right) / \hbar} \\
& =\sqrt{\operatorname{det} \frac{\partial q^{\prime}}{\partial q}} e^{i R\left(q, q^{\prime}, t\right) / \hbar} \psi\left(q^{\prime}, 0\right)
\end{aligned}
$$

As the time progresses the Lagrangian manifold $\partial_{q} R(q, t)$ can develop folds, so for longer times the value of the phase $R(q, t)$ is not necessarily unique; in general more than one trajectory will connect points $q$ and $q^{\prime}$ with different phases $R\left(q, q^{\prime}, t\right)$ accumulated along these paths, see figure 28.3.

We thus expect in general a collection of different trajectories from $q^{\prime}$ to $q$ which we will index by $j$, with different phase increments $R_{j}\left(q, q^{\prime}, t\right)$. The hypothetical particles of the density flow at a given configuration space point can move with different momenta $p=\partial_{q} R_{j}(q, t)$. This is not an ambiguity, since in the full $(q, p)$ phase space each particle follows its own trajectory with a unique momentum.

Whenever the Lagrangian manifold develops a fold, the density of the phase space trajectories in the fold projected on the configuration coordinates diverges. As illustrated in figure 28.3, when the Lagrangian manifold develops a fold at $q=q_{1}$; the volume element $d q_{1}$ in the neighborhood of the folding point is proportional to $\sqrt{d q^{\prime}}$ instead of $d q^{\prime}$. The Jacobian $\partial q^{\prime} / \partial q$ diverges like $1 / \sqrt{q_{1}-q(t)}$ when computed along the trajectory going trough the folding point at $q_{1}$. After the folding the orientation of the interval $d q^{\prime}$ has changed when being mapped into $d q_{2}$; in addition the function $R$, as well as its derivative which defines the Lagrangian manifold, becomes multi-valued. Distinct trajectories starting from different initial

Figure 28.3: Folding of the Lagrangian surface $\left(q, \partial_{q} R(q, t)\right)$.

points $q^{\prime}$ can now reach the same final point $q_{2}$. (That is, the point $q^{\prime}$ may have more than one pre-image.) The projection of a simple fold, or of an envelope of a family of phase space trajectories, is called a caustic; this expression comes from the Greek word for "capable of burning", evoking the luminous patterns that one observes swirling across the bottom of a swimming pool.

The folding also changes the orientation of the pieces of the Lagrangian manifold ( $q, \partial_{q} R(q, t)$ ) with respect to the initial manifold, so the eigenvalues of the Jacobian determinant change sign at each fold crossing. We can keep track of the signs by writing the Jacobian determinant as

$$
\left.\operatorname{det} \frac{\partial q^{\prime}}{\partial q}\right|_{j}=e^{-i \pi m_{j}\left(q, q^{\prime}, t\right)}\left|\operatorname{det} \frac{\partial q^{\prime}}{\partial q}\right|_{j},
$$

where $m_{j}\left(q, q^{\prime}, t\right)$ counts the number of sign changes of the Jacobian determinant on the way from $q^{\prime}$ to $q$ along the trajectory indexed with $j$, see figure 28.3. We shall refer to the integer $m_{j}\left(q, q^{\prime}, t\right)$ as the topological of the trajectory. So in general the semiclassical approximation to the wave function is thus a sum over possible trajectories that start at any inital $q^{\prime}$ and end in $q$ in time $t$

$$
\begin{equation*}
\psi_{s c}(q, t)=\int d q^{\prime} \sum_{j}\left|\operatorname{det} \frac{\partial q^{\prime}}{\partial q}\right|_{j}^{1 / 2} e^{i R_{j}\left(q, q^{\prime}, t\right) / \hbar-i \pi m_{j}\left(q, q^{\prime}, t\right) / 2} \psi\left(q_{j}^{\prime}, 0\right), \tag{28.22}
\end{equation*}
$$

each contribution weighted by corresponding density, phase increment and the topological index.

That the correct topological index is obtained by simply counting the number of eigenvalue sign changes and taking the square root is not obvious - the careful argument requires that quantum wave functions evaluated across the folds remain single valued.

### 28.2 Semiclassical propagator

We saw in chapter 26 that the evolution of an initial wave function $\psi(q, 0)$ is completely determined by the propagator (26.12). As $K\left(q, q^{\prime}, t\right)$ itself sat-
isfies the Schrödinger equation (26.14), we can treat it as a wave function parameterized by the configuration point $q^{\prime}$. In order to obtain a semiclassical approximation to the propagator we follow now the ideas developed in the last section. There is, however, one small complication: the initial condition (26.15) demands that the propagator at $t=0$ is a $\delta$-function at $q=q^{\prime}$, that is, the amplitude is infinite at $q^{\prime}$ and the phase is not well defined. Our hypothetical cloud of particles is thus initially localized at $q=q^{\prime}$ with any initial velocity. This is in contrast to the situation in the previous section where we assumed that the particles at a given point $q$ have well defined velocity (or a discrete set of velocities) given by $\dot{q}=\partial_{p} H(q, p)$. We will now derive at a semiclassical expression for $K\left(q, q^{\prime}, t\right)$ by considering the propagator for short times first, and extrapolating from there to arbitrary times $t$.

### 28.2.1 Short time propagator

For infinitesimally short times $\delta t$ away from the singular point $t=0$ we assume that it is again possible to write the propagator in terms of a well defined phase and amplitude, that is

$$
K\left(q, q^{\prime}, \delta t\right)=A\left(q, q^{\prime}, \delta t\right) e^{\frac{i}{\hbar} R\left(q, q^{\prime}, \delta t\right)} .
$$

As all particles start at $q=q^{\prime}, R\left(q, q^{\prime}, \delta t\right)$ will be of the form (28.12), that is

$$
\begin{equation*}
R\left(q, q^{\prime}, \delta t\right)=p \dot{q} \delta t-H(q, p) \delta t, \tag{28.23}
\end{equation*}
$$

with $\dot{q} \approx\left(q-q^{\prime}\right) / \delta t$. For Hamiltonians of the form (26.2) we have $\dot{q}=p / m$, which leads to

$$
R\left(q, q^{\prime}, \delta t\right)=\frac{m\left(q-q^{\prime}\right)^{2}}{2 \delta t}-V(q) \delta t
$$

Here $V$ can be evaluated any place along the trajectory from $q$ to $q^{\prime}$, for example at the midway point $V\left(\left(q+q^{\prime}\right) / 2\right)$. Inserting this into our ansatz for the propagator we obtain

$$
\begin{equation*}
K_{s c}\left(q, q^{\prime}, \delta t\right) \approx A\left(q, q^{\prime}, \delta t\right) e^{\frac{i}{\hbar}\left(\frac{m}{2 \delta t}\left(q-q^{\prime}\right)^{2}-V(q) \delta t\right)} . \tag{28.24}
\end{equation*}
$$

For infinitesimal times we can neglect the term $V(q) \delta t$, so $K_{s c}\left(q, q^{\prime}, \delta t\right)$ is a $d$-dimensional Gaussian with width $\sigma^{2}=i \hbar \delta t / m$. This Gaussian is a finite width approximation to the Dirac delta function

$$
\begin{equation*}
\delta(z)=\lim _{\sigma \rightarrow 0} \frac{1}{\sqrt{2 \pi \sigma^{2}}} e^{-z^{2} / 2 \sigma^{2}} \tag{28.25}
\end{equation*}
$$

if $A=(m / 2 \pi i \hbar \delta t)^{D / 2}$, with $A\left(q, q^{\prime}, \delta t\right)$ fixed by the Dirac delta function normalization condition. The correctly normalized propagator for infinitesimal times $\delta t$ is therefore

$$
\begin{equation*}
K_{s c}\left(q, q^{\prime}, \delta t\right) \approx\left(\frac{m}{2 \pi i \hbar \delta t}\right)^{D / 2} e^{\left.\frac{i}{\hbar} \frac{m\left(q-q^{\prime}\right)^{2}}{2 \delta t}-V(q) \delta t\right)} \tag{28.26}
\end{equation*}
$$

The short time dynamics of the Lagrangian manifold $\left(q, \partial_{q} R\right)$ which corresponds to the quantum propagator can now be deduced from (28.23); one obtains

$$
\frac{\partial R}{\partial q}=p \approx \frac{m}{\delta t}\left(q-q^{\prime}\right)
$$

that is, is the particles start for short times on a Lagrangian manifold which is a plane in phase space, see figure 28.4. Note, that for $\delta t \rightarrow 0$, this plane is given by the condition $q=q^{\prime}$, that is, particles start on a plane parallel to the momentum axis. As we have already noted, all particles start at $q=q^{\prime}$ but with different velocities for $t=0$. The inital surface $\left(q^{\prime}, p^{\prime}=\partial_{q} R\left(q^{\prime}, 0\right)\right)$ is mapped into the surface $(q(t), p(t))$ some time $t$ later. The slope of the Lagrangian plane for a short finite time is given as

$$
\frac{\partial p_{i}}{\partial q_{j}}=-\frac{\partial^{2} R}{\partial q_{j} \partial q_{i}^{\prime}}=-\frac{\partial p_{i}^{\prime}}{\partial q_{j}}=\frac{m}{\delta t} \delta_{i j}
$$

The prefactor $(m / \delta t)^{D / 2}$ in (28.26) can therefore be interpreted as the determinant of the Jacobian of the transformation from final position coordinates $q$ to initial momentum coordinates $p^{\prime}$, that is

$$
\begin{equation*}
K_{s c}\left(q, q^{\prime}, \delta t\right)=\frac{1}{(2 \pi i \hbar)^{D / 2}}\left(\operatorname{det} \frac{\partial p^{\prime}}{\partial q}\right)^{1 / 2} e^{i R\left(q, q^{\prime}, \delta t\right) / \hbar} \tag{28.27}
\end{equation*}
$$

where

$$
\begin{equation*}
\left.\frac{\partial p_{i}^{\prime}}{\partial q_{j}}\right|_{t, q^{\prime}}=\frac{\partial^{2} R\left(q, q^{\prime}, \delta t\right)}{\partial q_{j} \partial q_{i}^{\prime}} \tag{28.28}
\end{equation*}
$$

The subscript $\left.\cdots\right|_{t, q^{\prime}}$ indicates that the partial derivatives are to be evaluated with $t, q^{\prime}$ fixed.

The propagator in (28.27) has been obtained for short times. It is, however, already more or less in its final form. We only have to evolve our short time approximation of the propagator according to (28.22)

$$
K_{s c}\left(q^{\prime \prime}, q^{\prime}, t^{\prime}+\delta t\right)=\sum_{j}\left|\operatorname{det} \frac{\partial q}{\partial q^{\prime \prime}}\right|_{j}^{1 / 2} e^{i R_{j}\left(q^{\prime \prime}, q, t^{\prime}\right) / \hbar-i \pi m_{j}\left(q^{\prime \prime}, q, t^{\prime}\right) / 2} K\left(q, q_{j}^{\prime}, \delta t\right)
$$

Figure 28.4: Evolution of the semiclassical propagator. The configuration which corresponds to the initial conditions of the propagator is a Lagrangian manifold $q=q^{\prime}$, that is, a plane parallel to the $p$ axis. The hypothetical particles are thus initially all placed at $q^{\prime}$ but take on all possible momenta $p^{\prime}$. The Jacobian matrix $C(28.29)$ relates an initial volume element in momentum space $d p^{\prime}$ to a final configuration space volume $d q$.

and we included here already the possibility that the phase becomes multivalued, that is, that there is more than one path from $q^{\prime}$ to $q^{\prime \prime}$. The topological index $m_{j}=m_{j}\left(q^{\prime \prime}, q^{\prime}, t\right)$ is the number of singularities in the Jacobian along the trajectory $j$ from $q^{\prime}$ to $q^{\prime \prime}$. We can write $K_{s c}\left(q^{\prime \prime}, q^{\prime}, t^{\prime}+\delta t\right)$ in closed form using the fact that $R\left(q^{\prime \prime}, q, t^{\prime}\right)+R\left(q, q^{\prime}, \delta t\right)=R\left(q^{\prime \prime}, q^{\prime}, t^{\prime}+\delta t\right)$ and the multiplicativity of Jacobian determinants, that is

$$
\begin{equation*}
\left.\left.\operatorname{det} \frac{\partial q}{\partial q^{\prime \prime}}\right|_{t} \operatorname{det} \frac{\partial p^{\prime}}{\partial q}\right|_{q^{\prime}, \delta t}=\left.\operatorname{det} \frac{\partial p^{\prime}}{\partial q^{\prime \prime}}\right|_{q^{\prime}, t^{\prime}+\delta t} \tag{28.29}
\end{equation*}
$$

The final form of the semiclassical or Van Vleck propagator, is thus

$$
\begin{equation*}
K_{s c}\left(q, q^{\prime}, t\right)=\sum_{j} \frac{1}{(2 \pi i \hbar)^{D / 2}}\left|\operatorname{det} \frac{\partial p^{\prime}}{\partial q}\right|^{1 / 2} e^{i R_{j}\left(q, q^{\prime}, t\right) / \hbar-i m_{j} \pi / 2} \tag{28.30}
\end{equation*}
$$

This Van Vleck propagator is the essential ingredient of the semiclassical quantization to follow.

The apparent simplicity of the semiclassical propagator is deceptive. The wave function is not evolved simply by multiplying by a complex number of magnitude $\sqrt{\operatorname{det} \partial p^{\prime} / \partial q}$ and phase $R\left(q, q^{\prime}, t\right)$; the more difficult task in general is to find the trajectories connecting $q^{\prime}$ and $q$ in a given time $t$.

In addition, we have to treat the approximate propagator (28.30) with some care. Unlike the full quantum propagator, which satisfies the group property (26.13) exactly, the semiclassical propagator performs this only approximately, that is

$$
\begin{equation*}
K_{s c}\left(q, q^{\prime}, t_{1}+t_{2}\right) \approx \int d q^{\prime \prime} K_{s c}\left(q, q^{\prime \prime}, t_{2}\right) K_{s c}\left(q^{\prime \prime}, q^{\prime}, t_{1}\right) . \tag{28.31}
\end{equation*}
$$

The connection can be made explicit by the stationary phase approximation, sect. 27.2. Approximating the integral in (28.31) by integrating only over regions near points $q^{\prime \prime}$ at which the phase is stationary, leads to the stationary phase condition

$$
\begin{equation*}
\frac{\partial R\left(q, q^{\prime \prime}, t_{2}\right)}{\partial q_{i}^{\prime \prime}}+\frac{\partial R\left(q^{\prime \prime}, q^{\prime}, t_{1}\right)}{\partial q_{i}^{\prime \prime}}=0 . \tag{28.32}
\end{equation*}
$$

Classical trajectories contribute whenever the final momentum for a path from $q^{\prime}$ to $q^{\prime \prime}$ and the initial momentum for a path from $q^{\prime \prime}$ to $q$ coincide. Unlike the classical evolution of sect. 10.2, the semiclassical evolution is not an evolution by linear operator multiplication, but evolution supplemented by a stationary phase condition $p_{\text {out }}=p_{\text {in }}$ that matches up the classical momenta at each evolution step.

### 28.2.2 Free particle propagator

To develop some intuition about the above formalism, consider the case of a free particle. For a free particle the potential energy vanishes, the kinetic energy is $\frac{m}{2} \dot{q}^{2}$, and the Hamilton's principal function (28.12) is

$$
\begin{equation*}
R\left(q, q^{\prime}, t\right)=\frac{m\left(q-q^{\prime}\right)^{2}}{2 t} . \tag{28.33}
\end{equation*}
$$

The weight det $\frac{\partial p^{\prime}}{\partial q}$ from (28.28) can be evaluated explicitly, and the Van Vleck propagator is

$$
\begin{equation*}
K_{s c}\left(q, q^{\prime}, t\right)=\left(\frac{m}{2 \pi i \hbar t}\right)^{D / 2} e^{i m\left(q-q^{\prime}\right)^{2} / 2 \hbar t} \tag{28.34}
\end{equation*}
$$

identical to the short time propagator (28.26), with $V(q)=0$. This case is rather exceptional: for a free particle the semiclassical propagator turns out to be the exact quantum propagator $K\left(q, q^{\prime}, t\right)$, as can be checked by substitution in the Schrödinger equation (28.2). The Feynman path integral formalism uses this fact to construct an exact quantum propagator by integrating the free particle propagator (with $V(q)$ treated as constant for short times) along all possible (not necessarily classical) paths from $q^{\prime}$ to $q$.

### 28.3 Semiclassical Green's function

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So far we have derived semiclassical formulas for the time evolution of wave functions, that is, we obtained approximate solutions to the time dependent Schrödinger equation (26.1). Even though we assumed in the calculation a time independent Hamiltonian of the special form (26.2), the derivation would lead to the same final result (28.30) were one to consider more complicated or explicitly time dependent Hamiltonians. The propagator is thus important when we are interested in finite time quantum mechanical effects. For time independent Hamiltonians, the time dependence of the propagator as well as of wave functions is, however, essentially given in terms of the energy eigen-spectrum of the system, as in (26.10). It is therefore advantageous to switch from a time representation to an energy representation, that is from the propagator (26.12) to the energy
dependent Green's function (26.16). A semiclassical approximation of the Green's function $G_{s c}\left(q, q^{\prime}, E\right)$ is given by the Laplace transform (26.16) of the Van Vleck propagator $K_{s c}\left(q, q^{\prime}, t\right)$ :

$$
\begin{equation*}
G_{s c}\left(q, q^{\prime}, E\right)=\frac{1}{i \hbar} \int_{0}^{\infty} d t e^{i E t / \hbar} K_{s c}\left(q, q^{\prime}, t\right) . \tag{28.35}
\end{equation*}
$$

The expression as it stands is not very useful; in order to evaluate the integral, at least to the leading order in $\hbar$, we need to turn to the method of stationary phase again.

### 28.3.1 Stationary phase in higher dimensions

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Generalizing the method of sect. 27.2 to $d$ dimensions, consider stationary phase points fulfilling

$$
\left.\frac{d}{d x_{i}} \Phi(x)\right|_{x=x_{0}}=0 \quad \forall i=1, \ldots d
$$

An expansion of the phase up to second order involves now the symmetric matrix of second derivatives of $\Phi(x)$, that is

$$
D_{i j}\left(x_{0}\right)=\left.\frac{\partial^{2}}{\partial x_{i} \partial x_{j}} \Phi(x)\right|_{x=x_{0}}
$$

After choosing a suitable coordinate system which diagonalizes $D$, we can approximate the $d$-dimensional integral by $d$ one-dimensional Fresnel integrals; the stationary phase estimate of (27.13) is then

$$
\begin{equation*}
I \approx \sum_{x_{0}}(2 \pi i / s)^{d / 2}\left|\operatorname{det} D\left(x_{0}\right)\right|^{-1 / 2} A\left(x_{0}\right) e^{i s \Phi\left(x_{0}\right)-\frac{i \pi}{2} m\left(x_{0}\right)}, \tag{28.36}
\end{equation*}
$$

where the sum runs over all stationary phase points $x_{0}$ of $\Phi(x)$ and $m\left(x_{0}\right)$ counts the number of negative eigenvalues of $D\left(x_{0}\right)$.

The stationary phase approximation is all that is needed for the semiclassical approximation, with the proviso that $D$ in (28.36) has no zero eigenvalues.

### 28.3.2 Long trajectories

When evaluating the integral (28.35) approximately we have to distinguish between two types of contributions: those coming from stationary points of the phase and those coming from infinitesimally short times. The first type
of contributions can be obtained by the stationary phase approximation and will be treated in this section. The latter originate from the singular behavior of the propagator for $t \rightarrow 0$ where the assumption that the amplitude changes slowly compared to the phase is not valid. The short time contributions therefore have to be treated separately, which we will do in sect. 28.3.3.

The stationary phase points $t^{*}$ of the integrand in (28.35) are given by the condition

$$
\begin{equation*}
\frac{\partial}{\partial t} R\left(q, q^{\prime}, t^{*}\right)+E=0 \tag{28.37}
\end{equation*}
$$

We recognize this condition as the solution of (28.18), the time $t^{*}=t^{*}\left(q, q^{\prime}, E\right)$ in which a particle of energy $E$ starting out in $q^{\prime}$ reaches $q$. Taking into account the second derivative of the phase evaluated at the stationary phase point,

$$
R\left(q, q^{\prime}, t\right)+E t=R\left(q, q^{\prime}, t^{*}\right)+E t^{*}+\frac{1}{2}\left(t-t^{*}\right)^{2} \frac{\partial^{2}}{\partial t^{2}} R\left(q, q^{\prime}, t^{*}\right)+\cdots
$$

the stationary phase approximation of the integral corresponding to a classical trajectory $j$ in the Van Vleck propagator sum (28.30) yields

$$
\begin{equation*}
G_{j}\left(q, q^{\prime}, E\right)=\frac{1}{i \hbar(2 i \pi \hbar)^{(D-1) / 2}}\left|\operatorname{det} C_{j}\left(\frac{\partial^{2} R_{j}}{\partial t^{2}}\right)^{-1}\right|^{1 / 2} e^{\frac{i}{\hbar} S_{j}-\frac{i \pi}{2} m_{j}}, \tag{28.38}
\end{equation*}
$$

where $m_{j}=m_{j}\left(q, q^{\prime}, E\right)$ now includes a possible additional phase arising from the time stationary phase integration (27.16), and $C_{j}=C_{j}\left(q, q^{\prime}, t^{*}\right)$, $R_{j}=R_{j}\left(q, q^{\prime}, t^{*}\right)$ are evaluated at the transit time $t^{*}$. We re-express the phase in terms of the energy dependent action (28.16)

$$
\begin{equation*}
S\left(q, q^{\prime}, E\right)=R\left(q, q^{\prime}, t^{*}\right)+E t^{*}, \quad \text { with } \quad t^{*}=t^{*}\left(q, q^{\prime}, E\right), \tag{28.39}
\end{equation*}
$$

the Legendre transform of Hamilton's principal function. Note that the partial derivative of the action (28.39) with respect to $q_{i}$

$$
\frac{\partial S\left(q, q^{\prime}, E\right)}{\partial q_{i}}=\frac{\partial R\left(q, q^{\prime}, t^{*}\right)}{\partial q_{i}}+\left(\frac{\partial R\left(q, q^{\prime}, t\right)}{\partial t^{*}}+E\right) \frac{\partial t}{\partial q_{i}} .
$$

is equal to

$$
\begin{equation*}
\frac{\partial S\left(q, q^{\prime}, E\right)}{\partial q_{i}}=\frac{\partial R\left(q, q^{\prime}, t^{*}\right)}{\partial q_{i}} \tag{28.40}
\end{equation*}
$$

due to the stationary phase condition (28.37), so the definition of momentum as a partial derivative with respect to $q$ remains unaltered by the Legendre transform from time to energy domain.

Next we will simplify the amplitude term in (28.38) and rewrite it as an explicit function of the energy. Consider the $[(D+1) \times(D+1)]$ matrix

$$
D\left(q, q^{\prime}, E\right)=\left(\begin{array}{cc}
\frac{\partial^{2} S}{\partial q^{\prime} \partial q} & \frac{\partial^{2} S}{\partial q^{\prime} \partial E}  \tag{28.41}\\
\frac{\partial^{S} S}{\partial q \partial E} & \frac{\partial^{2} S}{\partial E^{2}}
\end{array}\right)=\left(\begin{array}{cc}
-\frac{\partial p^{\prime}}{\partial q} & -\frac{\partial p^{\prime}}{\partial E} \\
\frac{\partial t}{\partial q} & \frac{\partial t}{\partial E}
\end{array}\right),
$$

where $S=S\left(q, q^{\prime}, E\right)$ and we used (28.14-28.17) here to obtain the left hand side of (28.41). The minus signs follow from observing from the definition of (28.15) that $S\left(q, q^{\prime}, E\right)=-S\left(q^{\prime}, q, E\right)$. Note that $D$ is nothing but the Jacobian matrix of the coordinate transformation $(q, E) \rightarrow\left(p^{\prime}, t\right)$ for fixed $q^{\prime}$. We can therefore use the multiplication rules of determinants of Jacobians, which are just ratios of volume elements, to obtain

$$
\begin{aligned}
\operatorname{det} D & =(-1)^{D+1}\left(\operatorname{det} \frac{\partial\left(p^{\prime}, t\right)}{\partial(q, E)}\right)_{q^{\prime}}=(-1)^{D+1}\left(\operatorname{det} \frac{\partial\left(p^{\prime}, t\right)}{\partial(q, t)} \frac{\partial(q, t)}{\partial(q, E)}\right)_{q^{\prime}} \\
& =(-1)^{D+1}\left(\operatorname{det} \frac{\partial p^{\prime}}{\partial q}\right)_{t, q^{\prime}}\left(\operatorname{det} \frac{\partial t}{\partial E}\right)_{q^{\prime}, q}=\operatorname{det} C\left(\frac{\partial^{2} R}{\partial t^{2}}\right)^{-1} .
\end{aligned}
$$

We use here the notation (det. $)_{q^{\prime}, t}$ for a Jacobian determinant with partial derivatives evaluated at $t, q^{\prime}$ fixed, and likewise for other subscripts. Using the relation (28.19) which relates the term $\frac{\partial t}{\partial E}$ to $\partial_{t}^{2} R$ we can write the determinant of $D$ as a product of the Van Vleck determinant (28.28) and the amplitude factor arising from the stationary phase approximation. The amplitude in (28.38) can thus be interpreted as the determinant of a Jacobian of a coordinate transformation which includes time and energy as independent coordinates. This causes the increase in the dimensionality of the matrix $D$ relative to the Van Vleck determinant (28.28).

We can now write down the semiclassical approximation of the contribution of the $j$ th trajectory to the Green's function (28.38) in explicitly energy dependent form:

$$
\begin{equation*}
G_{j}\left(q, q^{\prime}, E\right)=\frac{1}{i \hbar(2 i \pi \hbar)^{(D-1) / 2}}\left|\operatorname{det} D_{j}\right|^{1 / 2} e^{\frac{i}{\hbar} S_{j}-\frac{i \pi}{2} m_{j}} . \tag{28.42}
\end{equation*}
$$

However, this is still not the most convenient form of the Green's function.
The trajectory contributing to $G_{j}\left(q, q^{\prime}, E\right)$ is constrained to a given energy $E$, and will therefore be on a phase space manifold of constant energy, that is $H(q, p)=E$. Writing this condition as a partial differential equation for $S\left(q, q^{\prime}, E\right)$, that is

$$
H\left(q, \frac{\partial S}{\partial q}\right)=E
$$

one obtains

$$
\begin{align*}
& \frac{\partial}{\partial q_{i}^{\prime}} H(q, p)=0=\frac{\partial H}{\partial p_{j}} \frac{\partial p_{j}}{\partial q_{i}^{\prime}}=\dot{q}_{j} \frac{\partial^{2} S}{\partial q_{j} \partial q_{i}^{\prime}} \\
& \frac{\partial}{\partial q_{i}} H\left(q^{\prime}, p^{\prime}\right) \quad=0=\frac{\partial^{2} S}{\partial q_{i} \partial q_{j}^{\prime}} \dot{q}_{j}^{\prime}, \tag{28.43}
\end{align*}
$$

that is the sub-matrix $\partial^{2} S / \partial q_{i} \partial q_{j}^{\prime}$ has (left- and right-) eigenvectors corresponding to an eigenvalue 0 . Rotate the local coordinate system at the either end of the trajectory

$$
\left(q_{1}, q_{2}, q_{3}, \cdots, q_{d}\right) \rightarrow\left(q_{\|}, q_{\perp 1}, q_{\perp 2}, \cdots, q_{\perp(D-1)}\right)
$$

so that one axis points along the trajectory and all others are perpendicular to it

$$
\left(\dot{q_{1}}, \dot{q_{2}}, \dot{q_{3}}, \cdots, \dot{q}_{d}\right) \rightarrow(\dot{q}, 0,0, \cdots, 0) .
$$

With such local coordinate systems at both ends, with the longitudinal coordinate axis $q_{\|}$pointing along the velocity vector of magnitude $\dot{q}$, the matrix of variations of $S\left(q, q^{\prime}, E\right)$ has a column and a row of zeros as (28.43) takes form

$$
\dot{q} \frac{\partial^{2} S}{\partial q_{\|} \partial q_{i}^{\prime}}=\frac{\partial^{2} S}{\partial q_{i} \partial q_{\|}^{\prime}} \dot{q}^{\prime}=0 .
$$

The initial and final velocities are non-vanishing except for points $|\dot{q}|=0$. These are the turning points (where all energy is potential), and we assume that neither $q$ nor $q^{\prime}$ is a turning point (in our application - periodic orbits - we can always chose $q=q^{\prime}$ not a turning point). In the local coordinate system with one axis along the trajectory and all other perpendicular to it the determinant of (28.41) is of the form

$$
\operatorname{det} D\left(q, q^{\prime}, E\right)=(-1)^{D+1}\left(\begin{array}{cccc} 
& 0 & 0 & \frac{\partial^{2} S}{\partial E \partial q_{\|}^{\prime}}  \tag{28.44}\\
\operatorname{det} & 0 & \frac{\partial^{2} S}{\partial q_{\perp} \partial q_{\perp}^{\prime}} & * \\
& \frac{\partial^{2} S}{\partial q_{\|} \partial E} & * & *
\end{array}\right)
$$

The corner entries can be evaluated using (28.17)

$$
\frac{\partial^{2} S}{\partial q_{\|} \partial E}=\frac{\partial}{\partial q_{\|}} t=\frac{1}{\dot{q}}, \quad \frac{\partial^{2} S}{\partial E \partial q_{\|}^{\prime}}=\frac{1}{\dot{q}^{\prime}} .
$$

As the $q_{\|}$axis points along the velocity direction, velocities $\dot{q}, \dot{q}^{\prime}$ are by construction almost always positive non-vanishing numbers. In this way
the determinant of the $[(D+1) \times(D+1)]$ dimensional matrix $D\left(q, q^{\prime}, E\right)$ can be reduced to the determinant of a $[(D-1) \times(D-1)]$ dimensional transverse matrix $D_{\perp}\left(q, q^{\prime}, E\right)$

$$
\begin{align*}
\operatorname{det} D\left(q, q^{\prime}, E\right) & =\frac{1}{\dot{q} \dot{q}^{\prime}} \operatorname{det} D_{\perp}\left(q, q^{\prime}, E\right) \\
D_{\perp}\left(q, q^{\prime}, E\right)_{i k} & =-\frac{\partial^{2} S\left(q, q^{\prime}, E\right)}{\partial q_{\perp i} \partial q_{\perp k}^{\prime}} \tag{28.45}
\end{align*}
$$

Putting everything together we obtain the $j$ th trajectory contribution to

$$
\begin{equation*}
G_{j}\left(q, q^{\prime}, E\right)=\frac{1}{i \hbar(2 \pi i \hbar)^{(D-1) / 2}} \frac{1}{\left|\dot{q} \dot{q}^{\prime}\right|^{1 / 2}}\left|\operatorname{det} D_{\perp}^{j}\right|^{1 / 2} e^{\frac{i}{\hbar} S_{j}-\frac{i \pi}{2} m_{j}}, \tag{28.46}
\end{equation*}
$$

where the topological index $m_{j}=m_{j}\left(q, q^{\prime}, E\right)$ now counts the number of changes of sign of det $D_{\perp}^{j}$ along the trajectory $j$ which connects $q^{\prime}$ to $q$ at energy $E$.

The endpoint velocities $\dot{q}, \dot{q}^{\prime}$ also depend on $\left(q, q^{\prime}, E\right)$ and the trajectory $j$.

### 28.3.3 Short trajectories

The stationary phase method cannot be used when $t^{*}$ is small, both because we cannot extend the integration in (27.16) to $-\infty$, and because the amplitude of $K\left(q, q^{\prime}, t\right)$ is divergent. In this case we have to evaluate the integral involving the short time form of the exact quantum mechanical propagator (28.26)

$$
\begin{equation*}
G_{0}\left(q, q^{\prime}, E\right)=\frac{1}{i \hbar} \int_{0}^{\infty} d t\left(\frac{m}{2 \pi i \hbar t}\right)^{D / 2} e^{\frac{i}{\hbar}\left(\frac{m\left(q-q^{\prime}\right)^{2}}{2 t}-V(q) t+E t\right)} . \tag{28.47}
\end{equation*}
$$

By introducing a dimensionless variable $\tau=t \sqrt{2 m(E-V(q))} / m\left|q-q^{\prime}\right|$, the integral can be rewritten as

$$
G_{0}\left(q, q^{\prime}, E\right)=\frac{m}{i \hbar^{2}(2 \pi i)^{D / 2}}\left(\frac{\sqrt{2 m(E-V)}}{\hbar\left|q-q^{\prime}\right|}\right)^{\frac{D}{2}-1} \int_{0}^{\infty} \frac{d \tau}{\tau^{D / 2}} e^{\frac{i}{2 \hbar} S_{0}\left(q, q^{\prime}, E\right)(\tau+1 / \tau)},
$$

where $S_{0}\left(q, q^{\prime}, E\right)=\sqrt{2 m(E-V)}\left|q-q^{\prime}\right|$ is the short distance form of the action. Using the integral representation of the Hankel function of first kind

$$
H_{\nu}^{+}(z)=-\frac{i}{\pi} e^{-i \nu \pi / 2} \int_{0}^{\infty} e^{\frac{1}{2} i z(\tau+1 / \tau)} \tau^{-\nu-1} d \tau
$$

we can write the short distance form of the Green's function as

$$
\begin{equation*}
G_{0}\left(q, q^{\prime}, E\right) \approx-\frac{i m}{2 \hbar^{2}}\left(\frac{\sqrt{2 m(E-V)}}{2 \pi \hbar\left|q-q^{\prime}\right|}\right)^{\frac{D-2}{2}} H_{\frac{D-2}{+}}^{+}\left(S_{0}\left(q, q^{\prime}, E\right) / \hbar\right) \tag{28.48}
\end{equation*}
$$

Hankel functions are stabdard, and their the short wavelength asymptotics is described in standard reference books. The short distance Green's function approximation is valid when $S_{0}\left(q, q^{\prime}, E\right) \leq \hbar$.

## Commentary

Remark 28.1 Limit $\hbar \rightarrow 0$. The semiclassical limit " $\hbar \rightarrow 0$ " discussed in sect. 28 is a shorthand notation for the limit in which typical quantities like the actions $R$ or $S$ in semiclassical expressions for the propagator or the Green's function become large compared to $\hbar$. In the world that we live in the quantity $\hbar$ is a fixed physical constant whose value [28.8] is $1.054571596(82) 10^{-34} \mathrm{Js}$.

Remark 28.2 Madelung's fluid dynamics. Already Schrödinger [28.3] noted that

$$
\rho=\rho(q, t):=A^{2}=\psi^{*} \psi
$$

plays the role of a density, and that the gradient of $R$ may be interpreted as a local semiclassical momentum, as the momentum density is

$$
\psi(q, t)^{*}\left(-i \hbar \frac{\partial}{\partial q}\right) \psi(q, t)=-i \hbar A \frac{\partial A}{\partial q}+\rho \frac{\partial R}{\partial q} .
$$

A very different interpretation of (28.3-28.4) has been given by Madelung [28.2], and then built upon by Bohm [28.6] and others [28.3, 28.7]. Keeping the $\hbar$ dependent term in (28.3), the ordinary differential equations driving the flow (28.10) have to be altered; if the Hamiltonian can be written as kinetic plus potential term $V(q)$ as in (26.2), the $\hbar^{2}$ term modifies the $p$ equation of motion as

$$
\begin{equation*}
\dot{p}_{i}=-\frac{\partial}{\partial q_{i}}(V(q)+Q(q, t)), \tag{28.49}
\end{equation*}
$$

where, for the example at hand,

$$
\begin{equation*}
Q(q, t)=-\frac{\hbar^{2}}{2 m} \frac{1}{\sqrt{\rho}} \frac{\partial^{2}}{\partial q^{2}} \sqrt{\rho} \tag{28.50}
\end{equation*}
$$

interpreted by Bohm [28.6] as the "quantum potential". Madelung observed that Hamilton's equation for the momentum (28.49) can be rewritten as

$$
\begin{equation*}
\frac{\partial v_{i}}{\partial t}+\left(v \cdot \frac{\partial}{\partial q}\right) v_{i}=-\frac{1}{m} \frac{\partial V}{\partial q_{i}}-\frac{1}{m \rho} \frac{\partial}{\partial q_{j}} \sigma_{i j}, \tag{28.51}
\end{equation*}
$$

where $\sigma_{i j}=\frac{\hbar^{2} \rho}{4 m} \frac{\partial^{2} \ln \rho}{\partial q_{i} \partial q_{j}}$ is the "pressure" stress tensor, $v_{i}=p_{i} / m$, and $\rho=A^{2}$ as defined [28.3] in sect. 28.1.3. We recall that the Eulerian $\frac{\partial}{\partial t}+\frac{\partial q_{i}}{\partial t} \frac{\partial}{\partial q_{i}}$ is the ordinary derivative of Lagrangian mechanics, that is $\frac{d}{d t}$. For comparison, the Euler equation for classical hydrodynamics is

$$
\frac{\partial v_{i}}{\partial t}+\left(v \cdot \frac{\partial}{\partial q}\right) v_{i}=-\frac{1}{m} \frac{\partial V}{\partial q_{i}}-\frac{1}{m \rho} \frac{\partial}{\partial q_{j}}\left(p \delta_{i j}\right)
$$

where $p \delta_{i j}$ is the pressure tensor.
The classical dynamics corresponding to quantum evolution is thus that of an "hypothetical fluid" experiencing $\hbar$ and $\rho$ dependent stresses. The "hydrodynamic" interpretation of quantum mechanics has, however, not been very fruitful in practice.

Remark 28.3 Path integrals. The semiclassical propagator (28.30) can also be derived from Feynman's path integral formalism. Dirac was the first to discover that in the short-time limit the quantum propagator (28.34) is exact. Feynman noted in 1946 that one can construct the exact propagator of the quantum Schrödinger equation by formally summing over all possible (and emphatically not classical) paths from $q^{\prime}$ to $q$.

Gutzwiller started from the path integral to rederive Van Vleck's semiclassical expression for the propagator; Van Vleck's original derivation is very much in the spirit of what has presented in this chapter. He did, however, not consider the possibility of the formation of caustics or folds of Lagrangian manifolds and thus did not include the topological phases in his semiclassical expression for the propagator. Some 40 years later Gutzwiller [30.4] added the topological indices when deriving the semiclassical propagator from Feynman's path integral by stationary phase conditions.

## Résumé

The aim of the semiclassical or short-wavelength methods is to approximate a solution of the Schrödinger equation with a semiclassical wave function

$$
\psi_{s c}(q, t)=\sum_{j} A_{j}(q, t) e^{i R_{j}(q, t) / \hbar}
$$

accurate to the leading order in $\hbar$. Here the sum is over all classical trajectories that connect the initial point $q^{\prime}$ to the final point $q$ in time $t$. "Semi-" refers to $\hbar$, the quantum unit of phase in the exponent. The quantum mechanics enters only through this atomic scale, in units of which the variation of the phase across the classical potential is assumed to be large. "-classical" refers to the rest - both the amplitudes $A_{j}(q, t)$ and the phases $R_{j}(q, t)$ - which are determined by the classical Hamilton-Jacobi equations.

In the semiclassical approximation the quantum time evolution operator is given by the semiclassical propagator

$$
K_{s c}\left(q, q^{\prime}, t\right)=\frac{1}{(2 \pi i \hbar)^{D / 2}} \sum_{j}\left|\operatorname{det} \frac{\partial p^{\prime}}{\partial q}\right|_{j}^{1 / 2} e^{\frac{i}{\hbar} R_{j}-\frac{i \pi}{2} m_{j}}
$$

where the topological index $m_{j}\left(q, q^{\prime}, t\right)$ counts the number of the direction reversal along the $j$ th classical trajectory that connects $q^{\prime} \rightarrow q$ in time $t$. Until very recently it was not possible to resolve quantum evolution on quantum time scales (such as one revolution of electron around a nucleus) - physical measurements are almost always done at time scales asymptotically large compared to the intrinsic quantum time scale. Formally this information is extracted by means of a Laplace transform of the propagator which yields the energy dependent semiclassical Green's function

$$
\begin{align*}
G_{s c}\left(q, q^{\prime}, E\right) & =G_{0}\left(q, q^{\prime}, E\right)+\sum_{j} G_{j}\left(q, q^{\prime}, E\right) \\
G_{j}\left(q, q^{\prime}, E\right) & =\frac{1}{i \hbar(2 \pi i \hbar)^{\frac{(D-1)}{2}}}\left|\frac{1}{\dot{q} \dot{q}^{\prime}} \operatorname{det} \frac{\partial p_{\perp}^{\prime}}{\partial q_{\perp}}\right|_{j}^{1 / 2} e^{\frac{i}{\hbar} S_{j}-\frac{i \pi}{2} m_{j}} \tag{28.52}
\end{align*}
$$

where $G_{0}\left(q, q^{\prime}, E\right)$ is the contribution of short trajectories with $S_{0}\left(q, q^{\prime}, E\right) \leq$ $\hbar$, while the sum is over the contributions of long trajectories (28.46) going from $q^{\prime}$ to $q$ with fixed energy $E$, with $S_{j}\left(q, q^{\prime}, E\right) \gg \hbar$.

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## Exercises

Exercise 28.1 Who ordered $\sqrt{\pi}$ ? Derive the Gaussian integral

$$
\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} d x e^{-\frac{x^{2}}{2 a}}=\sqrt{a}, \quad a>0 .
$$

assuming only that you know to integrate the exponential function $e^{-x}$. Hint, hint: $x^{2}$ is a radius-squared of something. $\pi$ is related to the area or circumference of something.

Exercise 28.2 Dirac delta function, Gaussian representation. Consider the Gaussian distribution function

$$
\delta_{\sigma}(z)=\frac{1}{\sqrt{2 \pi \sigma^{2}}} e^{-z^{2} / 2 \sigma^{2}}
$$

Show that in $\sigma \rightarrow 0$ limit this is the Dirac delta function

$$
\int_{\mathcal{M}} d x \delta(x)=1 \text { if } 0 \in \mathcal{M}, \text { zero otherwise. }
$$

Exercise 28.3 D-dimensional Gaussian integrals. Show that the Gaussian integral in $D$-dimensions is given by

$$
\begin{equation*}
\frac{1}{(2 \pi)^{d / 2}} \int d^{d} \phi e^{-\frac{1}{2} \phi^{T} \cdot M^{-1} \cdot \phi+\phi \cdot J}=|\operatorname{det} M|^{\frac{1}{2}} e^{\frac{1}{2} J^{T} \cdot M \cdot J}, \tag{28.53}
\end{equation*}
$$

where $M$ is a real positive definite $[d \times d]$ matrix, that is a matrix with strictly positive eigenvalues. $x, J$ are $D$-dimensional vectors, and $x^{T}$ is the transpose of $x$.

Exercise 28.4 Stationary phase approximation in higher dimensions. All semiclassical approximations are based on saddlepoint evaluations of integrals of type

$$
\begin{equation*}
I=\int d^{D} x A(x) e^{i \Phi(x) / \hbar} \tag{28.54}
\end{equation*}
$$

for small values of $\hbar$. Obtain the stationary phase estimate

$$
I \approx \sum_{n} A\left(x_{n}\right) e^{i \Phi\left(x_{n}\right) / \hbar} \frac{(2 \pi i \hbar)^{D / 2}}{\sqrt{\operatorname{det} \mathbf{D}^{2} \Phi\left(x_{n}\right)}},
$$

where $\mathbf{D}^{2} \Phi\left(x_{n}\right)$ denotes the second derivative matrix.

Exercise 28.5 Schrödinger equation in the Madelung form. Verify the decomposition of Schrödinger equation into real and imaginary parts, eqs. (28.3) and (28.4).

## Exercise 28.6 Transport equations.

Write the wave-function in the asymptotic form

$$
\psi(q, t)=e^{\frac{i}{\hbar} R(x, t)+\frac{i}{\hbar} \varepsilon t} \sum_{n \geq 0}(i \hbar)^{n} A_{n}(x, t) .
$$

Derive the transport equations for the $A_{n}$ by substituting this into the Schrödinger equation and then collecting terms by orders of $\hbar$. Notice that equation for $\dot{A}_{n}$ only requires knowledge of $A_{n-1}$ and $R$.

Exercise 28.7 Easy examples of the Hamilton's principal function. Calculate $R\left(q, q^{\prime}, t\right)$ for
a) a D-dimensional free particle
b) a 3-dimensional particle in constant magnetic field
c) a 1-dimensional harmonic oscillator.
(Continuation: exercise 28.15.)

Exercise 28.8 1-dimensional harmonic oscillator. Take a 1-dimensional harmonic oscillator $U(q)=\frac{1}{2} k q^{2}$. Take a WKB wave function of form $A(q, t)=$ $a(t)$ and $R(q, t)=r(t)+b(t) q+c(t) q^{2}$, where $r(t), a(t), b(t)$ and $c(t)$ are time dependent coefficients. Derive ordinary differential equations by using (28.3) and (28.4) and solve them. (Continuation: exercise 28.11.)

Exercise 28.9 1-dimensional linear potential. Take a 1-dimensional linear potential $U(q)=-F q$. Take a WKB wave function of form $A(q, t)=$ $a(t)$ and $R(q, t)=r(t)+b(t) q+c(t) q^{2}$, where $r(t), a(t), b(t)$ and $c(t)$ are time dependent coefficients. Derive and solve the ordinary differential equations from (28.3) and (28.4).

Exercise $28.10 \quad D$-dimensional quadratic potentials. Generalize the above method to general $D$-dimensional quadratic potentials.

Exercise $28.11 \quad$ Time evolution of $R$. (Continuation of exercise 28.8). Calculate the time evolution of $R(q, 0)=a+b q+c q^{2}$ for a 1-dimensional harmonic oscillator using (28.12) and (28.14).

Exercise 28.12 D-dimensional free particle propagator. Verify the results in sect. 28.2.2; show explicitely that (28.34), the semiclassical Van Vleck propagator in $D$ dimensions, solves the Schrödinger's equation.

## Exercise 28.13 Propagator, charged particle in constant magnetic field.

 Calculate the semiclassical propagator for a charged particle in constant magnetic field in 3 dimensions. Verify that the semiclassical expression coincides with the exact solution.Exercise 28.14 1-dimensional harmonic oscillator propagator. Calculate the semiclassical propagator for a 1-dimensional harmonic oscillator and verify that it is identical to the exact quantum propagator.

Exercise 28.15 Free particle action. Calculate the energy dependent action for a free particle, a charged particle in a constant magnetic field and for the harmonic oscillator.

Exercise 28.16 Zero length orbits.
Derive the classical trace (14.1) rigorously and either add the $t \rightarrow 0_{+}$zero length contribution to the trace formula, or show that it vanishes. Send us a reprint of Phys. Rev. Lett. with the correct derivation.

Exercise 28.17 Free particle semiclassical Green's functions. Calculate the semiclassical Green's functions for the systems of exercise 28.15.

## Chapter 29

## Noise

He who establishes his argument by noise and command shows that his reason is weak. M. de Montaigne
(G. Vattay and P. Cvitanović)


This chapter (which reader can safely skip on the first reading) is about noise, how it affects classical dynamics, and the ways it mimicks quantum dynamics.

fast track:
chapter 30, p. 515
Why - in a study of deterministic and quantum chaos - start discussing noise? First, in physical settings any dynamics takes place against a noisy background, and whatever prediction we might have, we have to check its robustness to noise. Second, as we show in this chapter, to the leading order in noise strength the semiclassical Hamilton-Jacobi formalism carries over to weakly stochastic flows in totto. As classical noisy dynamics is more intuitive than quantum dynamics, this exercise helps demistify some of the formal machinery of semiclassical quantization. Surprisingly, symplectic structure emerges here not as a deep principle of mechanics, but an artifact of the leading approximation to quantum/noisy dynamics, not respected by higher order corrections. Third, the variational principle derived here will be refashioned into a powerful tool for determining periodic orbits in chapter 31.

We start by deriving the continuity equation for purely deterministic, noiseless flow, and then incorporate noise in stages: diffusion equation, Langevin equation, Fokker-Planck equation, Hamilton-Jacobi formulation, stochastic path integrals.

### 29.1 Deterministic transport

(E.A. Spiegel and P. Cvitanović)

Fluid dynamics is about physical flows of media with continuous densities. On the other hand, the flows in state spaces of dynamical systems frequently require more abstract tools. To sharpen our intuition about those, it is helpful to outline the more tangible fluid dynamical vision.

Consider first the simplest property of a fluid flow called material invariant. A material invariant $I(x)$ is a property attached to each point $x$ that is preserved by the flow, $I(x)=I\left(f^{t}(x)\right)$; for example, at this point a green particle (more formally: a passive scalar) is embedded into the fluid. As $I(x)$ is invariant, its total time derivative vanishes, $\dot{I}(x)=0$. Written in terms of partial derivatives this is the conservation equation for the material invariant

$$
\begin{equation*}
\partial_{t} I+v \cdot \partial I=0 . \tag{29.1}
\end{equation*}
$$

Let the density of representative points be $\rho(x, t)$. The manner in which the flow redistributes $I(x)$ is governed by a partial differential equation whose form is relatively simple because the representative points are neither created nor destroyed. This conservation property is expressed in the integral statement

$$
\partial_{t} \int_{V} d x \rho I=-\int_{\partial V} d \sigma \hat{n}_{i} v_{i} \rho I
$$

where $V$ is an arbitrary volume in the state space $\mathcal{M}, \partial V$ is its surface, $\hat{n}$ is its outward normal, and repeated indices are summed over throughout. The divergence theorem turns the surface integral into a volume integral,

$$
\int_{V}\left[\partial_{t}(\rho I)+\partial_{i}\left(v_{i} \rho I\right)\right] d x=0
$$

where $\partial_{i}$ is the partial derivative operator with respect to $x_{i}$. Since the integration is over an arbitrary volume, we conclude that

$$
\begin{equation*}
\partial_{t}(\rho I)+\partial_{i}\left(\rho I v_{i}\right)=0 . \tag{29.2}
\end{equation*}
$$

The choice $I \equiv 1$ yields the continuity equation for the density:

$$
\begin{equation*}
\partial_{t} \rho+\partial_{i}\left(\rho v_{i}\right)=0 . \tag{29.3}
\end{equation*}
$$

We have used here the language of fluid mechanics to ease the visualization, but, as we already saw in the discussion of infinitesimal action of the Perron-Frobenius operator (9.24), continuity equation applies to any deterministic state space flow.

### 29.2 Brownian difussion

Consider tracer molecules, let us say green molecules, embedded in a denser gas of transparent molecules. Assume that the density of tracer molecules $\rho$ compared to the background gas density is low, so we can neglect greengreen collisions. Each green molecule, jostled by frequent collisions with the background gas, executes its own Brownian motion. The molecules are neither created nor destroyed, so their number within an arbitrary volume $V$ changes with time only by the current density $j_{i}$ flow through its surface $\partial V$ (with $\hat{n}$ its outward normal):

$$
\begin{equation*}
\partial_{t} \int_{V} d x \rho=-\int_{\partial V} d \sigma \hat{n}_{i} j_{i} . \tag{29.4}
\end{equation*}
$$

The divergence theorem turns this into the conservation law for tracer density:

$$
\begin{equation*}
\partial_{t} \rho+\partial_{i} j_{i}=0 \tag{29.5}
\end{equation*}
$$

The tracer density $\rho$ is defined as the average density of a "material particle", averaged over a subvolume large enough to contain many green (and still many more background) molecules, but small compared to the macroscopic observational scales. What is $j$ ? If the density is constant, on the average as many molecules leave the material particle volume as they enter it, so a reasonable phenomenological assumption is that the average current density (not the individual particle current density $\rho v_{i}$ in (29.3)) is driven by the density gradient

$$
\begin{equation*}
j_{i}=-D \frac{\partial \rho}{\partial x_{i}} \tag{29.6}
\end{equation*}
$$

This is the Fick law, with the diffusion constant $D$ a phenomenological parameter. For simplicity here we assume that $D$ is a scalar; in general $D \rightarrow D_{i j}(x, t)$ is a space- and time-dependent tensor. Substituting this $j$ into (29.5) yields the diffusion equation

$$
\begin{equation*}
\frac{\partial}{\partial t} \rho(x, t)=D \frac{\partial^{2}}{\partial x^{2}} \rho(x, t) \tag{29.7}
\end{equation*}
$$

This linear equation has an exact solution in terms of an initial Dirac delta density distribution, $\rho(x, 0)=\delta\left(x-x_{0}\right)$,

$$
\begin{equation*}
\rho(x, t)=\frac{1}{(4 \pi D t)^{d / 2}} e^{-\frac{\left(x-x_{0}\right)^{2}}{4 D t}}=\frac{1}{(4 \pi D t)^{d / 2}} e^{-\frac{\dot{x}^{2}}{4 D} t}, \tag{29.8}
\end{equation*}
$$

in the spirit of the quantum free particle propagation of sect. 28.2.2. The average distance covered in time $t$ obeys the Einstein diffusion formula

$$
\begin{equation*}
\left\langle\left(x-x_{0}\right)^{2}\right\rangle_{t}=\int d x \rho(x, t)\left(x-x_{0}\right)^{2}=2 D t \tag{29.9}
\end{equation*}
$$

### 29.3 Weak noise

The connection between path integration and Brownian motion is so close that they are nearly indistingushable. Unfortunately though, like a body and its mirror image, the sum over paths for Brownian motion is a theory having substance, while its path integral image exists mainly in the eye of the beholder. L. S. Schulman

So far we have considered tracer molecule dynamics which is purely Brownian, with no deterministic "drift". Consider next a deterministic flow $\dot{x}=v(x)$ perturbed by a stochastic term $\xi(t)$,

$$
\begin{equation*}
\dot{x}=v(x)+\xi(t) \tag{29.10}
\end{equation*}
$$

Assume that $\xi(t)$ 's fluctuate around $[\dot{x}-v(x)]$ with a Gaussian probability density

$$
\begin{equation*}
p(\xi, \delta t)=\left(\frac{\delta t}{4 \pi D}\right)^{d / 2} e^{-\frac{\xi^{2}}{4 D} \delta t} \tag{29.11}
\end{equation*}
$$

and are uncorrelated in time (white noise)

$$
\begin{equation*}
\left\langle\xi(t) \xi\left(t^{\prime}\right)\right\rangle=2 D \delta\left(t-t^{\prime}\right) \tag{29.12}
\end{equation*}
$$

The normalization factors in (29.8) and (29.11) differ, as $p(\xi, \delta t)$ is a probability density for velocity $\xi$, and $\rho(x, t)$ is a probability density for position $x$. The material particle now drifts along the trajectory $x(t)$, so the velocity diffusion follows (29.8) for infinitesimal time $\delta t$ only. As $D \rightarrow 0$, the distribution tends to the (noiseless, deterministic) Dirac delta function.

An example is the Langevin equation for a Brownian particle, in which one replaces the Newton's equation for force by two counter-balancing forces: random accelerations $\xi(t)$ which tend to smear out a particle trajectory, and a damping term which drives the velocity to zero.

The phenomenological Fick law current (29.6) is now a sum of two components, the material particle center-of-mass deterministic drift $v(x)$ and the weak noise term

$$
\begin{equation*}
j_{i}=\rho v_{i}-D \frac{\partial \rho}{\partial x_{i}}, \tag{29.13}
\end{equation*}
$$

Substituting this $j$ into (29.5) yields the Fokker-Planck equation

$$
\begin{equation*}
\partial_{t} \rho+\partial_{i}\left(\rho v_{i}\right)=D \partial^{2} \rho \tag{29.14}
\end{equation*}
$$

The left hand side, $d \rho / d t=\partial_{t} \rho+\partial \cdot(\rho v)$, is deterministic, with the continuity equation (29.3) recovered in the weak noise limit $D \rightarrow 0$. The right hand side describes the diffusive transport in or out of the material particle volume. If the density is lower than in the immediate neighborhood, the local curvature is positive, $\partial^{2} \rho>0$, and the density grows. Conversely, for negative curvature diffusion lowers the local density, thus smoothing the variability of $\rho$. Where is the density going globally?

If the system is bound, the probability density vanishes sufficiently fast outside the central region, $\rho(x, t) \rightarrow 0$ as $|x| \rightarrow \infty$, and the total probability is conserved

$$
\int d x \rho(x, t)=1
$$

Any initial density $\rho(x, 0)$ is smoothed by diffusion and with time tends to the invariant density

$$
\begin{equation*}
\rho_{0}(x)=\lim _{t \rightarrow \infty} \rho(x, t) \tag{29.15}
\end{equation*}
$$

an eigenfunction $\rho(x, t)=e^{s t} \rho_{0}(x)$ of the time-independent Fokker-Planck equation

$$
\begin{equation*}
\left(\partial_{i} v_{i}-D \partial^{2}+s_{\alpha}\right) \rho_{\alpha}=0 \tag{29.16}
\end{equation*}
$$

with vanishing eigenvalue $s_{0}=0$. Provided the noiseless classical flow is hyperbolic, in the vanishing noise limit the leading eigenfunction of the Fokker-Planck equation tends to natural measure (9.16) of the corresponding deterministic flow, the leading eigenvector of the Perron-Frobenius operator.

If the system is open, there is a continuous outflow of probability from the region under study, the leading eigenvalue is contracting, $s_{0}<0$, and the density of the system tends to zero. In this case the leading eigenvalue $s_{0}$ of the time-independent Fokker-Planck equation (29.16) can be interpreted by saying that a finite density can be maintained by pumping back probability into the system at a constant rate $\gamma=-s_{0}$. The value of $\gamma$ for which any initial probability density converges to a finite equilibrium density is called the escape rate. In the noiseless limit this coincides with the deterministic escape rate (10.15).

We have introduced noise phenomenologically, and used the weak noise assumption in retaining only the first derivative of $\rho$ in formulating the Fick law (29.6) and including noise additively in (29.13). A full theory of stochastic ODEs is much subtler, but this will do for our purposes.

### 29.4 Weak noise approximation

In the spirit of the WKB approximation, we shall now study the evolution of the probability distribution by rewriting it as

$$
\begin{equation*}
\rho(x, t)=e^{-R(x, t) / D} . \tag{29.17}
\end{equation*}
$$

The time evolution of $R$ is given by

$$
\partial_{t} R+v \partial R+(\partial R)^{2}=D \partial v+D \partial^{2} R .
$$

Consider now the weak noise limit and drop the terms proportional to $D$. The remaining equation

$$
\partial_{t} R+H(x, \partial R)=0
$$

is the Hamilton-Jacobi equation (28.5). The function $R$ can be interpreted as the Hamilton's principal function, corresponding to the Hamiltonian

$$
H(x, p)=p v(x)+p^{2},
$$

with the Hamilton's equations of motion

$$
\begin{align*}
\dot{x} & =\partial_{p} H=v+2 p \\
\dot{p} & =-\partial_{x} H=-A p, \tag{29.18}
\end{align*}
$$

where $A$ is the matrix of variations (4.3)

$$
A_{i j}(x)=\frac{\partial v_{i}(x)}{\partial x_{j}} .
$$

The noise Lagrangian (28.13) is then

$$
\begin{equation*}
L(x, \dot{x})=\dot{x} \cdot p-H=\frac{1}{4}[\dot{x}-v(x)]^{2} . \tag{29.19}
\end{equation*}
$$

We have come the full circle - the Lagrangian is the exponent of our assumed Gaussian distribution (29.11) for noise $\xi^{2}=[\dot{x}-v(x)]^{2}$. What is the meaning of this Hamiltonian, Lagrangian? Consider two points $x_{0}$ and $x$. Which noisy path is the most probable path that connects them in time $t$ ? The probability of a given path $\mathcal{P}$ is given by the probability of the noise sequence $\xi(t)$ which generates the path. This probability is proportional to the product of the noise probability functions (29.11) along the path, and
the total probabbility for reaching $x$ from $x_{0}$ in time $t$ is given by the sum over all paths, or the stochastic path integral (Wiener integral)

$$
\begin{align*}
P\left(x, x_{0}, t\right) & \sim \sum_{\mathcal{P}} \prod_{j} p\left(\xi\left(\tau_{j}\right), \delta \tau_{j}\right)=\int \prod_{j} d \xi_{j}\left(\frac{\delta \tau_{j}}{4 \pi D}\right)^{d / 2} e^{-\frac{\xi\left(\tau_{j}\right)^{2}}{4 D} \delta \tau_{i}} \\
& \rightarrow \frac{1}{Z} \sum_{\mathcal{P}} \exp \left(-\frac{1}{4 D} \int_{0}^{t} d \tau \xi^{2}(\tau)\right), \tag{29.20}
\end{align*}
$$

where $\delta \tau_{i}=\tau_{i}-\tau_{i}$, and the normalization constant is

$$
\frac{1}{Z}=\lim \prod_{i}\left(\frac{\delta \tau_{i}}{4 \pi D}\right)^{d / 2}
$$

The most probable path is the one maximizing the integral inside the exponential. If we express the noise (29.10) as

$$
\xi(t)=\dot{x}(t)-v(x(t)),
$$

the probability is maximized by the variational principle

$$
\min \int_{0}^{t} d \tau[\dot{x}(\tau)-v(x(\tau))]^{2}=\min \int_{0}^{t} L(x(\tau), \dot{x}(\tau)) d \tau
$$

By the standard arguments, for a given $x, x^{\prime}$ and $t$ the the probability is maximized by a solution of Hamilton's equations (29.18) that connects the two points $x_{0} \rightarrow x^{\prime}$ in time $t$.

## Commentary

Remark 29.1 Literature. The theory of stochastic processes is a vast subject, spanning over centuries and over disciplines ranging from pure mathematics to impure finance. We enjoyed reading van Kampen classic [29.1], especially his railings against those who blunder carelessly into nonlinear landscapes (with this chapter we join the list of van Kampen's sinners). A more specialized monograph like Risken's [29.2] will do just as well. The "Langevin equation" introduces noise and damping only into the acceleration of Newton's equations; here we are considering more general stochastic differential equations in the weak noise limit. OnsagerMachlup seminal paper [29.18] was the first to introduce a variational method - the "principle of least dissipation" - based on the Lagrangian of form (29.19). This paper deals only with a finite set of linearly damped thermodynamic variables. Here the setting is much more general: we study fluctuations over a state space varying velocity field $v(x)$. Schulman's monograph [29.11] contains a very readable summary of Kac's [29.12] exposition of Wiener's integral over stchastic paths.

## Résumé

When a deterministic trajectory is smeared out under the influence of Gaussian noise of strength $D$, the deterministic dynamics is recovered in the weak noise limit $D \rightarrow 0$. The effect of the noise can be taken into account by adding noise corrections to the classical trace formula.

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## Chapter 30

## Semiclassical quantization

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We derive here the Gutzwiller trace formula and the semiclassical zeta function, the central results of the semiclassical quantization of classically chaotic systems. In chapter 32 we will rederive these formulas for the case of scattering in open systems. Quintessential wave mechanics effects such as creeping, diffraction and tunneling will be taken up in chapter 35 .

### 30.1 Trace formula

Our next task is to evaluate the Green's function trace (26.17) in the semiclassical approximation. The trace

$$
\operatorname{tr} G_{s c}(E)=\int d^{D} q G_{s c}(q, q, E)=\operatorname{tr} G_{0}(E)+\sum_{j} \int d^{D} q G_{j}(q, q, E)
$$

receives contributions from "long" classical trajectories labeled by $j$ which start and end in $q$ after finite time, and the "zero length" trajectories whose lengths approach zero as $q^{\prime} \rightarrow q$.

First we work out the contributions coming from the finite time returning classical orbits, that is, trajectories that originate and end at a given configuration point $q$. As we are identifying $q$ with $q^{\prime}$, taking of a trace involves (still another!) stationary phase condition in the $q^{\prime} \rightarrow q$ limit,

$$
\left.\frac{\partial S_{j}\left(q, q^{\prime}, E\right)}{\partial q_{i}}\right|_{q^{\prime}=q}+\left.\frac{\partial S_{j}\left(q, q^{\prime}, E\right)}{\partial q_{i}^{\prime}}\right|_{q^{\prime}=q}=0
$$

meaning that the initial and final momenta (28.40) of contributing trajectories should coincide

$$
\begin{equation*}
p_{i}(q, q, E)-p_{i}^{\prime}(q, q, E)=0, \quad q \in j \text { th periodic orbit }, \tag{30.1}
\end{equation*}
$$

Figure 30.1: A returning trajectory in the configuration space. The orbit is periodic in the full phase space only if the initial and the final momenta of a returning trajectory coincide as well.


Figure 30.2: A romanticized sketch of $S_{p}(E)=S(q, q, E)=\oint p(q, E) d q$ landscape orbit. Unstable periodic orbits traverse isolated ridges and saddles of the mountainous landscape of the action $S\left(q_{\|}, q_{\perp}, E\right)$. Along a periodic orbit $S_{p}(E)$ is constant; in the transverse directions it generically changes quadratically.

so the trace receives contributions only from those long classical trajectories which are periodic in the full phase space.

For a periodic orbit the natural coordinate system is the intrinsic one, with $q_{\|}$axis pointing in the $\dot{q}$ direction along the orbit, and $q_{\perp}$, the rest of the coordinates transverse to $\dot{q}$. The $j$ th periodic orbit contribution to the trace of the semiclassical Green's function in the intrinsic coordinates is

$$
\operatorname{tr} G_{j}(E)=\frac{1}{i \hbar(2 \pi \hbar)^{(d-1) / 2}} \oint_{j} \frac{d q_{\|}}{\dot{q}} \int_{j} d^{d-1} q_{\perp}\left|\operatorname{det} D_{\perp}^{j}\right|^{1 / 2} e^{\frac{i}{\hbar} S_{j}-\frac{i \pi}{2} m_{j}}
$$

where the integration in $q_{\|}$goes from 0 to $L_{j}$, the geometric length of small tube around the orbit in the configuration space. As always, in the stationary phase approximation we worry only about the fast variations in the phase $S_{j}\left(q_{\|}, q_{\perp}, E\right)$, and assume that the density varies smoothly and is well approximated by its value $D_{\perp}^{j}\left(q_{\|}, 0, E\right)$ on the classical trajectory, $q_{\perp}=$ 0 . The topological index $m_{j}\left(q_{\|}, q_{\perp}, E\right)$ is an integer which does not depend on the initial point $q_{\|}$and not change in the infinitesimal neighborhood of an isolated periodic orbit, so we set $m_{j}(E)=m_{j}\left(q_{\|}, q_{\perp}, E\right)$.

The transverse integration is again carried out by the stationary phase method, with the phase stationary on the periodic orbit, $q_{\perp}=0$. The result of the transverse integration can depend only on the parallel coordinate

$$
\operatorname{tr} G_{j}(E)=\frac{1}{i \hbar} \oint \frac{d q_{\|}}{\dot{q}}\left|\frac{\operatorname{det} D_{\perp j}\left(q_{\|}, 0, E\right)}{\operatorname{det} D_{\perp j}^{\prime}\left(q_{\|}, 0, E\right)}\right|^{1 / 2} e^{\frac{i}{\hbar} S_{j}-\frac{i \pi}{2} m_{j}}
$$

where the new determinant in the denominator, $\operatorname{det} D_{\perp j}^{\prime}=$

$$
\operatorname{det}\left(\frac{\partial^{2} S\left(q, q^{\prime}, E\right)}{\partial q_{\perp i} \partial q_{\perp j}}+\frac{\partial^{2} S\left(q, q^{\prime}, E\right)}{\partial q_{\perp i}^{\prime} \partial q_{\perp j}}+\frac{\partial^{2} S\left(q, q^{\prime}, E\right)}{\partial q_{\perp i} \partial q_{\perp j}^{\prime}}+\frac{\partial^{2} S\left(q, q^{\prime}, E\right)}{\partial q_{\perp i}^{\prime} \partial q_{\perp j}^{\prime}}\right)
$$

is the determinant of the second derivative matrix coming from the stationary phase integral in transverse directions. Mercifully, this integral also removes most of the $2 \pi \hbar$ prefactors in (??).

The ratio $\operatorname{det} D_{\perp j} / \operatorname{det} D_{\perp j}^{\prime}$ is here to enforce the periodic boundary condition for the semiclassical Green's function evaluated on a periodic orbit. It can be given a meaning in terms of the monodromy matrix of the periodic orbit by following observations

$$
\begin{aligned}
\operatorname{det} D_{\perp} & =\left\|\frac{\partial p_{\perp}^{\prime}}{\partial q_{\perp}}\right\|=\left\|\frac{\partial\left(q_{\perp}^{\prime}, p_{\perp}^{\prime}\right)}{\partial\left(q_{\perp}, q_{\perp}^{\prime}\right)}\right\| \\
\operatorname{det} D_{\perp}^{\prime} & =\left\|\frac{\partial p_{\perp}}{\partial q_{\perp}}-\frac{\partial p_{\perp}^{\prime}}{\partial q_{\perp}}+\frac{\partial p_{\perp}}{\partial q_{\perp}^{\prime}}-\frac{\partial p_{\perp}^{\prime}}{\partial q_{\perp}^{\prime}}\right\|=\left\|\frac{\partial\left(p_{\perp}-p_{\perp}^{\prime}, q_{\perp}-q_{\perp}^{\prime}\right)}{\partial\left(q_{\perp}, q_{\perp}^{\prime}\right)}\right\| .
\end{aligned}
$$

Defining the $2(D-1)$-dimensional transverse vector $x_{\perp}=\left(q_{\perp}, p_{\perp}\right)$ in the full phase space we can express the ratio

$$
\begin{align*}
\frac{\operatorname{det} D_{\perp}^{\prime}}{\operatorname{det} D_{\perp}} & =\left\|\frac{\partial\left(p_{\perp}-p_{\perp}^{\prime}, q_{\perp}-q_{\perp}^{\prime}\right)}{\partial\left(q_{\perp}^{\prime}, p_{\perp}^{\prime}\right)}\right\|=\left\|\frac{\partial\left(x_{\perp}-x_{\perp}^{\prime}\right)}{\partial x_{\perp}^{\prime}}\right\| \\
& =\operatorname{det}(\mathbf{J}-\mathbf{1}), \tag{30.2}
\end{align*}
$$

in terms of the monodromy matrix $\mathbf{J}$ for a surface of section transverse to the orbit within the constant energy $E=H(q, p)$ shell.

The classical periodic orbit action $S_{j}(E)=\oint p\left(q_{\|}, E\right) d q_{\|}$is an integral around a loop defined by the periodic orbit, and does not depend on the starting point $q_{\|}$along the orbit, see figure 30.2. The eigenvalues of the monodromy matrix are also independent of where $\mathbf{J}_{j}$ is evaluated along the orbit, so $\operatorname{det}\left(1-\mathbf{J}_{j}\right)$ can also be taken out of the the $q_{\|}$integral

$$
\operatorname{tr} G_{j}(E)=\frac{1}{i \hbar} \sum_{j} \frac{1}{\left|\operatorname{det}\left(1-\mathbf{J}_{j}\right)\right|^{1 / 2}} e^{r\left(\frac{i}{\hbar} S_{j}-\frac{i \pi}{2} m_{j}\right)} \oint \frac{d q_{\|}}{\dot{q}_{\|}} .
$$

Here we have assumed that $\mathbf{J}_{j}$ has no marginal eigenvalues. The determinant of the monodromy matrix, the action $S_{p}(E)=\oint p\left(q_{\|}, E\right) d q_{\|}$and the topological index are all classical invariants of the periodic orbit. The integral in the parallel direction we now do exactly.

First we take into account the fact that any repeat of a periodic orbit is also a periodic orbit. The action and the topological index are additive along the trajectory, so for $r$ th repeat they simply get multiplied by $r$. The monodromy matrix of the $r$ th repeat of a prime cycle $p$ is (by the chain rule for derivatives) $\mathbf{J}_{p}^{r}$, where $\mathbf{J}_{p}$ is the prime cycle monodromy matrix. Let us denote the time period of the prime cycle $p$, the single, shortest traversal of a periodic orbit by $T_{p}$. The remaining integral can be carried out by change of variables $d t=d q_{\|} / \dot{q}(t)$

$$
\int_{0}^{L_{p}} \frac{d q_{\|}}{\dot{q}(t)}=\int_{0}^{T_{p}} d t=T_{p}
$$

Note that the spatial integral corresponds to a single traversal. If you do not see why this is so, rethink the derivation of the classical trace formula (14.19) - that derivation takes only three pages of text. Regrettably, in the quantum case we do not know of an honest derivation that takes less than 30 pages. The final result, the Gutzwiller trace formula

$$
\begin{equation*}
\operatorname{tr} G_{s c}(E)=\operatorname{tr} G_{0}(E)+\frac{1}{i \hbar} \sum_{p} T_{p} \sum_{r=1}^{\infty} \frac{1}{\left|\operatorname{det}\left(1-\mathbf{J}_{p}^{r}\right)\right|^{1 / 2}} e^{r\left(\frac{i}{\hbar} S_{p}-\frac{i \pi}{2} m_{p}\right)} \tag{30.3}
\end{equation*}
$$

an expression for the trace of the semiclassical Green's function in terms of periodic orbits, is beautiful in its simplicity and elegance.

The topological index $m_{p}(E)$ counts the number of changes of sign of the matrix of second derivatives evaluated along the prime periodic orbit $p$. By now we have gone through so many stationary phase approximations that you have surely lost track of what the total $m_{p}(E)$ actually is. The rule is this: The topological index of a closed curve in a $2 D$ phase space is the sum of the number of times the partial derivatives $\frac{\partial p_{i}}{\partial q_{i}}$ for each dual pair $\left(q_{i}, p_{i}\right), i=1,2, \ldots, D$ (no sum on $i$ ) change their signs as one goes once around the curve.

### 30.1.1 Average density of states

We still have to evaluate $\operatorname{tr} G_{0}(E)$, the contribution coming from the infinitesimal trajectories. The real part of $\operatorname{tr} G_{0}(E)$ is infinite in the $q^{\prime} \rightarrow q$ limit, so it makes no sense to write it down explicitly here. However, the imaginary part is finite, and plays an important role in the density of states formula, which we derive next.

The semiclassical contribution to the density of states (26.17) is given by the imaginary part of the Gutzwiller trace formula (30.3) multiplied with $-1 / \pi$. The contribution coming from the zero length trajectories is the imaginary part of (28.48) for $q^{\prime} \rightarrow q$ integrated over the configuration space

$$
d_{0}(E)=-\frac{1}{\pi} \int d^{D} q \operatorname{Im} G_{0}(q, q, E),
$$

The resulting formula has a pretty interpretation; it estimates the number of quantum states that can be accomodated up to the energy $E$ by counting the available quantum cells in the phase space. This number is given by the Weyl rule, as the ratio of the phase space volume bounded by energy $E$ divided by $h^{D}$, the volume of a quantum cell,

$$
\begin{equation*}
N_{s c}(E)=\frac{1}{h^{D}} \int d^{D} p d^{D} q \Theta(E-H(q, p)) . \tag{30.4}
\end{equation*}
$$

where $\Theta(x)$ is the Heaviside function (26.22). $N_{s c}(E)$ is an estimate of the spectral staircase (26.21), so its derivative yields the average density of states

$$
\begin{equation*}
d_{0}(E)=\frac{d}{d E} N_{s c}(E)=\frac{1}{h^{D}} \int d^{D} p d^{D} q \delta(E-H(q, p)) \tag{30.5}
\end{equation*}
$$

precisely the semiclassical result (30.6). For Hamiltonians of type $p^{2} / 2 m+$ $V(q)$, the energy shell volume in (30.5) is a sphere of radius $\sqrt{2 m(E-V(q))}$. The surface of a $d$-dimensional sphere of radius $r$ is $\pi^{d / 2} r^{d-1} / \Gamma(d / 2)$, so the average density of states is given by

$$
\begin{equation*}
d_{0}(E)=\frac{2 m}{\hbar^{D} 2^{d} \pi^{D 2} \Gamma(D / 2)} \int_{V(q)<E} d^{D} q[2 m(E-V(q))]^{D / 2-1}, \tag{30.6}
\end{equation*}
$$

and

$$
\begin{equation*}
N_{s c}(E)=\frac{1}{h^{D}} \frac{\pi^{D / 2}}{\Gamma(1+D / 2)} \int_{V(q)<E} d^{D} q[2 m(E-V(q))]^{D / 2} . \tag{30.7}
\end{equation*}
$$

Physically this means that at a fixed energy the phase space can support $N_{s c}(E)$ distinct eigenfunctions; anything finer than the quantum cell $h^{D}$ cannot be resolved, so the quantum phase space is effectively finite dimensional. The average density of states is of a particularly simple form in one spatial dimension

$$
\begin{equation*}
d_{0}(E)=\frac{T(E)}{2 \pi \hbar}, \tag{30.8}
\end{equation*}
$$

where $T(E)$ is the period of the periodic orbit of fixed energy $E$. In two spatial dimensions the average density of states is

$$
\begin{equation*}
d_{0}(E)=\frac{m \mathcal{A}(E)}{2 \pi \hbar^{2}}, \tag{30.9}
\end{equation*}
$$

where $\mathcal{A}(E)$ is the classically allowed area of configuration space for which $V(q)<E$.

The semiclassical density of states is a sum of the average density of states and the oscillation of the density of states around the average, $d_{s c}(E)=d_{0}(E)+d_{o s c}(E)$, where

$$
\begin{equation*}
d_{o s c}(E)=\frac{1}{\pi \hbar} \sum_{p} T_{p} \sum_{r=1}^{\infty} \frac{\cos \left(r S_{p}(E) / \hbar-r m_{p} \pi / 2\right)}{\left|\operatorname{det}\left(1-\mathbf{J}_{p}^{r}\right)\right|^{1 / 2}} \tag{30.10}
\end{equation*}
$$

follows from the trace formula (30.3).

### 30.1.2 Regularization of the trace

The real part of the $q^{\prime} \rightarrow q$ zero length Green's function (28.48) is ultraviolet divergent in dimensions $d>1$, and so is its formal trace (26.17). The short distance behavior of the real part of the Green's function can be extracted from the real part of (28.48) by using the Bessel function expansion for $\operatorname{small} z$

$$
Y_{\nu}(z) \approx\left\{\begin{array}{cl}
-\frac{1}{\pi} \Gamma(\nu)\left(\frac{z}{2}\right)^{-\nu} & \text { for } \nu \neq 0 \\
\frac{2}{\pi}(\ln (z / 2)+\gamma) & \text { for } \nu=0
\end{array},\right.
$$

where $\gamma=0.577 \ldots$ is the Euler constant. The real part of the Green's function for short distance is dominated by the singular part

$$
G_{\text {sing }}\left(\left|q-q^{\prime}\right|, E\right)=\left\{\begin{array}{cl}
-\frac{m}{2 \hbar^{2} \pi^{\frac{d}{2}}} \Gamma((d-2) / 2) \frac{1}{\left|q-q^{\prime}\right|^{d-2}} & \text { for } d \neq 2 \\
\frac{m}{2 \pi \hbar^{2}}\left(\ln \left(2 m(E-V)\left|q-q^{\prime}\right| / 2 \hbar\right)+\gamma\right) & \text { for } d=2
\end{array} .\right.
$$

The regularized Green's function

$$
G_{\text {reg }}\left(q, q^{\prime}, E\right)=G\left(q, q^{\prime}, E\right)-G_{\text {sing }}\left(\left|q-q^{\prime}\right|, E\right)
$$

is obtained by subtracting the $q^{\prime} \rightarrow q$ ultraviolet divergence. For the regularized Green's function the Gutzwiller trace formula is

$$
\begin{equation*}
\operatorname{tr} G_{r e g}(E)=-i \pi d_{0}(E)+\frac{1}{i \hbar} \sum_{p} T_{p} \sum_{r=1}^{\infty} \frac{e^{r\left(\frac{i}{\hbar} S_{p}(E)-\frac{i \pi}{2} m_{p}(E)\right)}}{\left|\operatorname{det}\left(1-\mathbf{J}_{p}^{r}\right)\right|^{1 / 2}} . \tag{30.11}
\end{equation*}
$$

Now you stand where Gutzwiller stood in 1990. You hold the trace formula in your hands. You have no clue how good is the $\hbar \rightarrow 0$ approximation, how to take care of the sum over an infinity of periodic orbits, and whether the formula converges at all.

### 30.2 Semiclassical spectral determinant

The problem with trace formulas is that they diverge where we need them, at the individual energy eigenvalues. What to do? Much of the quantum chaos literature responds to the challenge of wrestling the trace formulas by replacing the delta functions in the density of states (26.18) by Gaussians. But there is no need to do this - we can compute the eigenenergies without any further ado by remembering that the smart way to determine the eigenvalues of linear operators is by determining zeros of their spectral determinants.

Figure 30.3: A sketch of how spectral determinants convert poles into zeros: The trace shows $1 /\left(E-E_{n}\right)$ type singularities at the eigenenergies while the spectral determinant goes smoothly through zeroes.


A sensible way to compute energy levels is to construct the spectral determinant whose zeroes yield the eigenenergies, $\operatorname{det}(\hat{H}-E)_{s c}=0$. A first guess might be that the spectral determinant is the Hadamard product of form

$$
\operatorname{det}(\hat{H}-E)=\prod_{n}\left(E-E_{n}\right)
$$

but this product is not well defined, since for fixed $E$ we multiply larger and larger numbers $\left(E-E_{n}\right)$. This problem is dealt with by regularization, discussed below in appendix 30.1.2. Here we offer an impressionistic sketch of regularization.

The logarithmic derivative of $\operatorname{det}(\hat{H}-E)$ is the (formal) trace of the Green's function

$$
-\frac{d}{d E} \ln \operatorname{det}(\hat{H}-E)=\sum_{n} \frac{1}{E-E_{n}}=\operatorname{tr} G(E) .
$$

This quantity, not surprisingly, is divergent again. The relation, however, opens a way to derive a convergent version of $\operatorname{det}(\hat{H}-E)_{s c}$, by replacing the trace with the regularized trace

$$
-\frac{d}{d E} \ln \operatorname{det}(\hat{H}-E)_{s c}=\operatorname{tr} G_{r e g}(E)
$$

The regularized trace still has $1 /\left(E-E_{n}\right)$ poles at the semiclassical eigenenergies, poles which can be generated only if $\operatorname{det}(\hat{H}-E)_{s c}$ has a zero at $E=E_{n}$, see figure 30.3. By integrating and exponentiating we obtain

$$
\operatorname{det}(\hat{H}-E)_{s c}=\exp \left(-\int^{E} d E^{\prime} \operatorname{tr} G_{r e g}\left(E^{\prime}\right)\right)
$$

Now we can use (30.11) and integrate the terms coming from periodic orbits, using the relation (28.17) between the action and the period of a periodic orbit, $d S_{p}(E)=T_{p}(E) d E$, and the relation (26.21) between the density of states and the spectral staircase, $d N_{s c}(E)=d_{0}(E) d E$. We obtain the semiclassical zeta function

$$
\begin{equation*}
\operatorname{det}(\hat{H}-E)_{s c}=e^{i \pi N_{s c}(E)} \exp \left(-\sum_{p} \sum_{r=1}^{\infty} \frac{1}{r} \frac{e^{i r\left(S_{p} / \hbar-m_{p} \pi / 2\right)}}{\left.\operatorname{det}\left(1-\mathbf{J}_{p}^{r}\right)\right|^{1 / 2}}\right) . \tag{30.12}
\end{equation*}
$$

We already know from the study of classical evolution operator spectra of chapter 15 that this can be evaluated by means of cycle expansions. The beauty of this formula is that everything on the right side - the cycle action $S_{p}$, the topological index $m_{p}$ and monodromy matrix $\mathbf{J}_{p}$ determinant - is intrinsic, coordinate-choice independent property of the cycle $p$.

### 30.3 One-dof systems

It has been a long trek, a stationary phase upon stationary phase. Let us check whether the result makes sense even in the simplest case, for quantum mechanics in one spatial dimension.

In one dimension the average density of states follows from the one-dof form of the oscillating density (30.10) and of the average density (30.8)

$$
\begin{equation*}
d(E)=\frac{T_{p}(E)}{2 \pi \hbar}+\sum_{r} \frac{T_{p}(E)}{\pi \hbar} \cos \left(r S_{p}(E) / \hbar-r m_{p}(E) \pi / 2\right) \tag{30.13}
\end{equation*}
$$

The classical particle oscillates in a single potential well with period $T_{p}(E)$. There is no monodromy matrix to evaluate, as in one dimension there is only the parallel coordinate, and no transverse directions. The $r$ repetition sum in (30.13) can be rewritten by using the Fourier series expansion of a delta spike train

$$
\sum_{n=-\infty}^{\infty} \delta(x-n)=\sum_{k=-\infty}^{\infty} e^{i 2 \pi k x}=1+\sum_{k=1}^{\infty} 2 \cos (2 \pi k x)
$$

We obtain

$$
\begin{equation*}
d(E)=\frac{T_{p}(E)}{2 \pi \hbar} \sum_{n} \delta\left(S_{p}(E) / 2 \pi \hbar-m_{p}(E) / 4-n\right) \tag{30.14}
\end{equation*}
$$

This expression can be simplified by using the relation (28.17) between $T_{p}$ and $S_{p}$, and the identity (9.7) $\delta\left(x-x^{*}\right)=\left|f^{\prime}(x)\right| \delta(f(x))$, where $x^{*}$ is the only zero of the function $f\left(x^{*}\right)=0$ in the interval under consideration. We obtain

$$
d(E)=\sum_{n} \delta\left(E-E_{n}\right)
$$

where the energies $E_{n}$ are the zeroes of the arguments of delta functions in (30.14)

$$
S_{p}\left(E_{n}\right) / 2 \pi \hbar=n-m_{p}(E) / 4
$$

where $m_{p}(E)=m_{p}=2$ for smoth potential at both turning points, and $m_{p}(E)=m_{p}=4$ for two billiard (infinite potential) walls. These are precisely the Bohr-Sommerfeld quantized energies $E_{n}$, defined by the condition

$$
\begin{equation*}
\oint p\left(q, E_{n}\right) d q=h\left(n-\frac{m_{p}}{4}\right) . \tag{30.15}
\end{equation*}
$$

In this way the trace formula recovers the well known 1-dof quantization rule. In one dimension, the average of states can be expressed from the quantization condition. At $E=E_{n}$ the exact number of states is $n$, while the average number of states is $n-1 / 2$ since the staircase function $N(E)$ has a unit jump in this point

$$
\begin{equation*}
N_{s c}(E)=n-1 / 2=S_{p}(E) / 2 \pi \hbar-m_{p}(E) / 4-1 / 2 . \tag{30.16}
\end{equation*}
$$

The 1-dof spectral determinant follows from (30.12) by dropping the monodromy matrix part and using (30.16)

$$
\begin{equation*}
\operatorname{det}(\hat{H}-E)_{s c}=\exp \left(-\frac{i}{2 \hbar} S_{p}+\frac{i \pi}{2} m_{p}\right) \exp \left(-\sum_{r} \frac{1}{r} e^{\frac{i}{\hbar} r S_{p}-\frac{i \pi}{2} r m_{p}}\right) . \tag{30.17}
\end{equation*}
$$

Summation yields a logarithm by $\sum_{r} t^{r} / r=-\ln (1-t)$ and we get

$$
\begin{aligned}
\operatorname{det}(\hat{H}-E)_{s c} & =e^{-\frac{i}{2 \hbar} S_{p}+\frac{i m_{p}}{4}+\frac{i \pi}{2}}\left(1-e^{\frac{i}{\hbar} S_{p}-i \frac{m_{p}}{2}}\right) \\
& =2 \sin \left(S_{p}(E) / \hbar-m_{p}(E) / 4\right) .
\end{aligned}
$$

So in one dimension, where there is only one periodic orbit for a given energy E, nothing is gained by going from the trace formula to the spectral determinant. The spectral determinant is a real function for real energies, and its zeros are again the Bohr-Sommerfeld quantized eigenenergies (30.15).

### 30.4 Two-dof systems

For flows in two configuration dimensions the monodromy matrix $\mathbf{J}_{p}$ has two eigenvalues $\Lambda_{p}$ and $1 / \Lambda_{p}$, as explained in sect. 5.2. Isolated periodic orbits can be elliptic or hyperbolic. Here we discuss only the hyperbolic case, when the eigenvalues are real and their absolute value is not equal to one. The determinant appearing in the trace formulas can be written in terms of the expanding eigenvalue as

$$
\left|\operatorname{det}\left(1-\mathbf{J}_{p}^{r}\right)\right|^{1 / 2}=\left|\Lambda_{p}^{r}\right|^{1 / 2}\left(1-1 / \Lambda_{p}^{r}\right)
$$

and its inverse can be expanded as a geometric series

$$
\frac{1}{\left|\operatorname{det}\left(1-\mathbf{J}_{p}^{r}\right)\right|^{1 / 2}}=\sum_{k=0}^{\infty} \frac{1}{\left|\Lambda_{p}^{r}\right|^{1 / 2} \Lambda_{p}^{k r}} .
$$

With the 2-dof expression for the average density of states (30.9) the spectral determinant becomes

$$
\begin{align*}
\operatorname{det}(\hat{H}-E)_{s c} & =e^{i \frac{m \mathcal{A} E}{2 \hbar^{2}}} \exp \left(-\sum_{p} \sum_{r=1}^{\infty} \sum_{k=0}^{\infty} \frac{e^{i r\left(S_{p} / \hbar-m_{p} \pi / 2\right)}}{r\left|\Lambda_{p}^{r}\right|^{1 / 2} \Lambda_{p}^{k r}}\right) \\
& =e^{i \frac{m \mathcal{A} E}{2 \hbar^{2}}} \prod_{p} \prod_{k=0}^{\infty}\left(1-\frac{e^{\frac{i}{\hbar} S_{p}-\frac{i \pi}{2} m_{p}}}{\left|\Lambda_{p}\right|^{1 / 2} \Lambda_{p}^{k}}\right) . \tag{30.18}
\end{align*}
$$

## Commentary

Remark 30.1 Zeta functions. For "zeta function" nomenclature, see remark 15.4 on page 259 .

## Résumé

Spectral determinants and dynamical zeta functions arise both in classical and quantum mechanics because in both the dynamical evolution can be described by the action of linear evolution operators on infinite-dimensional vector spaces. In quantum mechanics the periodic orbit theory arose from studies of semi-conductors, and the unstable periodic orbits have been measured in experiments on the very paradigm of Bohr's atom, the hydrogen atom, this time in strong external fields.

In practice, most "quantum chaos" calculations take the stationary phase approximation to quantum mechanics (the Gutzwiller trace formula, possibly improved by including tunneling periodic trajectories, diffraction corrections, etc.) as the point of departure. Once the stationary phase approximation is made, what follows is classical in the sense that all quantities used in periodic orbit calculations - actions, stabilities, geometrical phases - are classical quantities. The problem is then to understand and control the convergence of classical periodic orbit formulas.

While various periodic orbit formulas are formally equivalent, practice shows that some are preferable to others. Three classes of periodic orbit formulas are frequently used:

Trace formulas. The trace of the semiclassical Green's function

$$
\operatorname{tr} G_{s c}(E)=\int d q G_{s c}(q, q, E)
$$

is given by a sum over the periodic orbits (30.11). While easiest to derive, in calculations the trace formulas are inconvenient for anything other than the leading eigenvalue estimates, as they tend to be divergent in the region of physical interest. In classical dynamics trace formulas hide under a variety of appelations such as the $f-\alpha$ or multifractal formalism; in quantum mechanics they are known as the Gutzwiller trace formulas.

Zeros of Ruelle or dynamical zeta functions

$$
1 / \zeta(s)=\prod_{p}\left(1-t_{p}\right), \quad t_{p}=\frac{1}{\left|\Lambda_{p}\right|^{1 / 2}} e^{\frac{i}{\hbar} S_{p}-i \pi m_{p} / 2}
$$

yield, in combination with cycle expansions, the semiclassical estimates of quantum resonances. For hyperbolic systems the dynamical zeta functions have good convergence and are a useful tool for determination of classical and quantum mechanical averages.

Spectral determinants, Selberg-type zeta functions, Fredholm determinants, functional determinants are the natural objects for spectral calculations, with convergence better than for dynamical zeta functions, but with less transparent cycle expansions. The 2 -dof semiclassical spectral determinant (30.18)

$$
\operatorname{det}(\hat{H}-E)_{s c}=e^{i \pi N_{s c}(E)} \prod_{p} \prod_{k=0}^{\infty}\left(1-\frac{e^{i S_{p} / \hbar-i \pi m_{p} / 2}}{\left|\Lambda_{p}\right|^{1 / 2} \Lambda_{p}^{k}}\right)
$$

is a typical example. Most periodic orbit calculations are based on cycle expansions of such determinants.

As we have assumed repeatedly during the derivation of the trace formula that the periodic orbits are isolated, and do not form families (as is the case for integrable systems or in KAM tori of systems with mixed phase space), the formulas discussed so far are valid only for hyperbolic and elliptic periodic orbits.

For deterministic dynamical flows and number theory, spectral determinants and zeta functions are exact. The quantum-mechanical ones, derived by the Gutzwiller approach, are at best only the stationary phase approximations to the exact quantum spectral determinants, and for quantum mechanics an important conceptual problem arises already at the level of derivation of the semiclassical formulas; how accurate are they, and can the periodic orbit theory be systematically improved?

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## Exercises

## Exercise 30.1 Monodromy matrix from second variations of the action.

 Show that$$
\begin{equation*}
D_{\perp j} / D_{\perp j}^{\prime}=(\mathbf{1}-\mathbf{J}) \tag{30.19}
\end{equation*}
$$

Exercise 30.2 Jacobi gymnastics. Prove that the ratio of determinants in (N.37) can be expressed as

$$
\frac{\operatorname{det} D_{\perp j}^{\prime}\left(q_{\|}, 0, E\right)}{\operatorname{det} D_{\perp j}\left(q_{\|}, 0, E\right)}=\operatorname{det}\left(\begin{array}{cc}
I-\mathbf{J}_{q q} & -\mathbf{J}_{q p}  \tag{30.20}\\
-\mathbf{J}_{p q} & I-\mathbf{J}_{p p}
\end{array}\right)=\operatorname{det}\left(1-\mathbf{J}_{j}\right),
$$

where $\mathbf{J}_{j}$ is the monodromy matrix of the periodic orbit.

Exercise 30.3 Volume of $d$-dimensional sphere. Show that the volume of a $d$-dimensional sphere of radius $r$ equals $\pi^{d / 2} r^{d} / \Gamma(1+d / 2)$. Show that $\Gamma(1+d / 2)=$ $\Gamma(d / 2) d / 2$.

Exercise 30.4 Average density of states in 1 dimension. Show that in one dimension the average density of states is given by (30.8)

$$
\bar{d}(E)=\frac{T(E)}{2 \pi \hbar},
$$

where $T(E)$ is the time period of the 1-dimensional motion and show that

$$
\begin{equation*}
\bar{N}(E)=\frac{S(E)}{2 \pi \hbar} \tag{30.21}
\end{equation*}
$$

where $S(E)=\oint p(q, E) d q$ is the action of the orbit.

Exercise 30.5 Average density of states in 2 dimensions. Show that in 2 dimensions the average density of states is given by (30.9)

$$
\bar{d}(E)=\frac{m \mathcal{A}(E)}{2 \pi \hbar^{2}},
$$

where $\mathcal{A}(E)$ is the classically allowed area of configuration space for which $U(q)<E$.

## Chapter 31

## Relaxation for cyclists

Cycles, that is, solutions of the periodic orbit condition (17.1)

$$
\begin{equation*}
f^{t+T}(x)=f^{t}(x), \quad T>0 \tag{31.1}
\end{equation*}
$$

are prerequisite to chapters 14 and 15 evaluation of spectra of classical evolution operators, and, as we shall see in chapter 30, the semiclassical approximations to quantum evolution operators. Chapter 17 offered an introductory, hands-on guide to extraction of periodic orbits by means of the Newton-Raphson method. Here we take a very different tack, drawing inspiration from variational principles of classical mechanics, and path integrals of quantum mechanics.

In sect. 17.2.1 we converted orbits unstable forward in time into orbits stable backwards in time. Indeed, all methods for finding unstable cycles are based on the idea of constructing a new dynamical system such that (i) the position of the cycle is the same for the original system and the transformed one, (ii) the unstable cycle in the original system is a stable cycle of the transformed system.

The Newton-Raphson method for determining a fixed point $x_{*}$ for a map $x^{\prime}=f(x)$ is an example. The method replaces iteration of $f(x)$ by iteration of the Newton-Raphson map (17.5)

$$
\begin{equation*}
x_{i}^{\prime}=g_{i}(x)=x_{i}-\left(\frac{1}{\mathbf{J}(x)-\mathbf{1}}\right)_{i j}(f(x)-x)_{j} . \tag{31.2}
\end{equation*}
$$

A fixed point $x_{*}$ for a map $f(x)$ is also a fixed point of $g(x)$, indeed a superstable fixed point since $\partial g_{i}\left(x_{*}\right) / \partial x_{j}=0$. This makes the convergence to the fixed point super-exponential.

We also learned in chapter 17 that methods that start with initial guesses for a number of points along a cycle are considerably more robust and safer than searches based on direct solution of the fixed-point condition
(31.1). The relaxation (or variational) methods that we shall now describe take this multipoint approach to its logical extreme, and start by a guess of not a few points along a periodic orbit, but a guess of the entire orbit.

The idea is to make an informed rough guess of what the desired periodic orbit looks like globally, and then use variational methods to drive the initial guess toward the exact solution. Sacrificing computer memory for robustness of the method, we replace a guess that a point is on the periodic orbit by a guess of the entire orbit. And, sacrificing speed for safety, in sect. 31.1 we replace the Newton-Raphson iteration by a fictitious time flow that minimizes a cost function computed as deviation of the approximate flow from the true flow along a loop approximation to a periodic orbit.

If you have some insight into the topology of the flow and its symbolic dynamics, or have already found a set of short cycles, you might be able to construct an initial approximation to a longer cycle $p$ as a sequence of $N$ points $\left(\tilde{x}_{1}^{(0)}, \tilde{x}_{2}^{(0)}, \cdots, \tilde{x}_{N}^{(0)}\right)$ with the periodic boundary condition $\tilde{x}_{N+1}=\tilde{x}_{1}$. Suppose you have an iterative method for improving your guess; after $k$ iterations the cost function

$$
\begin{equation*}
F^{2}\left(\tilde{x}^{(k)}\right)=\sum_{i}^{N}\left(\tilde{x}_{i+1}^{(k)}-f\left(\tilde{x}_{i}^{(k)}\right)\right)^{2} \tag{31.3}
\end{equation*}
$$

or some other more cleverly constructed function (for classical mechanics - action) is a measure of the deviation of the $k$ th approximate cycle from the true cycle. This observation motivates variational approaches to determining cycles.

We give here three examples of such methods, two for maps, and one for billiards. In sect. 31.1 we start out by converting a problem of finding an unstable fixed point of a map into a problem of constructing a differential flow for which the desired fixed point is an attracting equilibrium point. Solving differential equations can be time intensive, so in sect. 31.2 we replace such flows by discrete iterations. In sect. 31.3 we show that for $2 D$-dimensional billiard flows variation of $D$ coordinates (where $D$ is the number of Hamiltonian degrees of freedom) suffices to determine cycles in the full $2 D$-dimensional phase space.

### 31.1 Fictitious time relaxation

(O. Biham, C. Chandre and P. Cvitanović)

The relaxation (or gradient) algorithm for finding cycles is based on the observation that a trajectory of a map such as the Hénon map (3.12),

$$
\begin{align*}
x_{i+1} & =1-a x_{i}^{2}+b y_{i} \\
y_{i+1} & =x_{i}, \tag{31.4}
\end{align*}
$$

Figure 31.1: "Potential" $V_{i}(x)$ (31.7) for a typical point along an inital guess trajectory. For $\sigma_{i}=+1$ the flow is toward the local maximum of $V_{i}(x)$, and for $\sigma_{i}=-1$ toward the local minimum. A large deviation of $x_{i}$ 's is needed to destabilize a trajectory passing through such local extremum of $V_{i}(x)$, hence the basin of attraction is expected to be large.

is a stationary solution of the relaxation dynamics defined by the flow

$$
\begin{equation*}
\frac{d x_{i}}{d \tau}=v_{i}, \quad i=1, \ldots, n \tag{31.5}
\end{equation*}
$$

for any vector field $v_{i}=v_{i}(x)$ which vanishes on the trajectory. Here $\tau$ is a "fictitious time" variable, unrelated to the dynamical time (in this example, the discrete time of map iteration). As the simplest example, take $v_{i}$ to be the deviation of an approximate trajectory from the exact 2-step recurrence form of the Hénon map (3.13)

$$
\begin{equation*}
v_{i}=x_{i+1}-1+a x_{i}^{2}-b x_{i-1} \tag{31.6}
\end{equation*}
$$

For fixed $x_{i-1}, x_{i+1}$ there are two values of $x_{i}$ satisfying $v_{i}=0$. These solutions are the two extremal points of a local "potential" function (no sum on $i$ )

$$
\begin{equation*}
v_{i}=\frac{\partial}{\partial x_{i}} V_{i}(x), \quad V_{i}(x)=x_{i}\left(x_{i+1}-b x_{i-1}-1\right)+\frac{a}{3} x_{i}^{3} \tag{31.7}
\end{equation*}
$$

Assuming that the two extremal points are real, one is a local minimum of $V_{i}(x)$ and the other is a local maximum. Now here is the idea; replace (31.5) by

$$
\begin{equation*}
\frac{d x_{i}}{d \tau}=\sigma_{i} v_{i}, \quad i=1, \ldots, n \tag{31.8}
\end{equation*}
$$

where $\sigma_{i}= \pm 1$.
The modified flow will be in the direction of the extremal point given by the local maximum of $V_{i}(x)$ if $\sigma_{i}=+1$ is chosen, or in the direction of the one corresponding to the local minimum if we take $\sigma_{i}=-1$. This is not quite what happens in solving (31.8) - all $x_{i}$ and $V_{i}(x)$ change at each integration step - but this is the observation that motivates the method. The differential equations (31.8) then drive an approximate initial guess toward the exact trajectory. A sketch of the landscape in which $x_{i}$ converges towards the proper fixed point is given in figure 31.1. As the "potential" function (31.7) is not bounded for a large $\left|x_{i}\right|$, the flow diverges for initial

Figure 31.2: The repeller for the Hénon map at $a=1.8, b=0.3$.

guesses which are too distant from the true trajectory. However, the basin of attraction of initial guesses that converge to a given cycle is nevertheless very large, with the spread in acceptable initial guesses for figure 31.1 of order 1, in contrast to the exponential precision required of initial guesses by the Newton-Raphson method.

Example 31.1 Hénon map cycles. Our aim in this calculation is to find all periodic orbits of period $n$ for the Hénon map (31.4), in principle at most $2^{n}$ orbits. We start by choosing an initial guess trajectory $\left(x_{1}, x_{2}, \cdots, x_{n}\right)$ and impose the periodic boundary condition $x_{n+1}=x_{1}$. The simplest and a rather crude choice of the initial condition in the Hénon map example is $x_{i}=0$ for all $i$. In order to find a given orbit one sets $\sigma_{i}=-1$ for all iterates $i$ which are local minima of $V_{i}(x)$, and $\sigma_{i}=1$ for iterates which are local maxima. In practice one runs through a complete list of prime cycles, such as the table 11.1. The real issue for all searches for periodic orbits, this one included, is how large is the basin of attraction of the desired periodic orbit? There is no easy answer to this question, but empirically it turns out that for the Hénon map such initial guess almost always converges to the desired trajectory as long as the initial $|x|$ is not too large compared to $1 / \sqrt{a}$. Figure 31.1 gives some indication of a typical basin of attraction of the method (see also figure 31.3).

The calculation is carried out by solving the set of $n$ ordinary differential equations (31.8) using a simple Runge-Kutta method with a relatively large step size ( $h=0.1$ ) until $|v|$ becomes smaller than a given value $\varepsilon$ (in a typical calculation $\varepsilon \sim 10^{-7}$ ). Empirically, in the case that an orbit corresponding to the desired itinerary does not exist, the initial guess escapes to infinity since the "potential" $V_{i}(x)$ grows without bound.

Applied to the Hénon map at the Hénon's parameters choice $a=1.4, b=0.3$, the method has yielded all periodic orbits to periods as long as $n=28$, as well as selected orbits up to period $n=1000$. All prime cycles up to period 10 for the Hénon map, $a=1.4$ and $b=0.3$, are listed in table 31.1. The number of unstable periodic orbits for periods $n \leq 28$ is given in table 31.2. Comparing this with the list of all possible 2-symbol alphabet prime cycles, table 11.1, we see that the pruning is quite extensive, with the number of cycle points of period $n$ growing as $e^{0.4645 \cdot n}=(1.592)^{n}$ rather than as $2^{n}$.

As another example we plot all unstable periodic points up to period $n=14$ for $a=1.8, b=0.3$ in figure 31.2. Comparing this repelling set with the strange attractor for the Hénon's parameters figure 3.4, we note the existence of gaps in the set, cut out by the preimages of the escaping regions.

In practice, the relaxation flow (31.8) finds (almost) all periodic orbits which exist and indicates which ones do not. For the Hénon map the method enables us to calculate almost all unstable cycles of essentially any desired length and accuracy.

| $n$ | $p$ | $\left(y_{p}, x_{p}\right)$ | $\lambda_{p}$ |
| :---: | :---: | :---: | :---: |
| 1 | 0 | (-1.13135447, -1.13135447) | 1.18167262 |
|  | 1 | (0.63135447, 0.63135447) | 0.65427061 |
| 2 | 01 | (0.97580005, , 0.47580005) | 0.55098676 |
| 4 | 0111 | (-0.70676677, 0.63819399) | 0.53908457 |
| 6 | 010111 | (-0.41515894, 1.07011813) | 0.55610982 |
|  | 011111 | (-0.80421990, 0.44190995) | 0.55245341 |
| 7 | 0011101 | (-1.04667757, -0.17877958) | 0.40998559 |
|  | 0011111 | (-1.08728604, -0.28539206) | 0.46539757 |
|  | 0101111 | (-0.34267842, 1.14123046) | 0.41283650 |
|  | 0111111 | (-0.88050537, 0.26827759) | 0.51090634 |
| 8 | 00011101 | (-1.25487963, -0.82745422) | 0.43876727 |
|  | 00011111 | (-1.25872451, , 0.83714168) | 0.43942101 |
|  | 00111101 | (-1.14931330, -0.48368863) | 0.47834615 |
|  | 00111111 | (-1.14078564, -0.44837319) | 0.49353764 |
|  | 01010111 | (-0.52309999, 0.93830866) | 0.54805453 |
|  | 01011111 | (-0.38817041, 1.09945313) | 0.55972495 |
|  | 01111111 | (-0.83680827, 0.36978609) | 0.56236493 |
| 9 | 000111101 | (-1.27793296, -0.90626780) | 0.38732115 |
|  | 000111111 | (-1.27771933, -0.90378859) | 0.39621864 |
|  | 001111101 | (-1.10392601, -0.34524675) | 0.51112950 |
|  | 001111111 | (-1.11352304, -0.36427104) | 0.51757012 |
|  | 010111111 | (-0.36894919, 1.11803210) | 0.54264571 |
|  | 011111111 | (-0.85789748, 0.32147653) | 0.56016658 |
| 10 | 0001111101 | (-1.26640530, -0.86684837) | 0.47738235 |
|  | 0001111111 | (-1.26782752, -0.86878943) | 0.47745508 |
|  | 0011111101 | (-1.12796804, -0.41787432) | 0.52544529 |
|  | 0011111111 | (-1.12760083, -0.40742737) | 0.53063973 |
|  | 0101010111 | (-0.48815908, 0.98458725) | 0.54989554 |
|  | 0101011111 | (-0.53496022, 0.92336925) | 0.54960607 |
|  | 0101110111 | (-0.42726915, 1.05695851) | 0.54836764 |
|  | 0101111111 | (-0.37947780, 1.10801373) | 0.56915950 |
|  | 0111011111 | (-0.69555680, 0.66088560) | 0.54443884 |
|  | 0111111111 | (-0.84660200, 0.34750875) | 0.57591048 |
| 13 | 1110011101000 | (-1.2085766485, -0.6729999948) | 0.19882434 |
|  | 1110011101001 | (-1.0598110494, -0.2056310390) | 0.21072511 |

Table 31.1: All prime cycles up to period 10 for the Hénon map, $a=1.4$ and $b=0.3$. The columns list the period $n_{p}$, the itinerary (defined in remark 31.4), a cycle point ( $y_{p}, x_{p}$ ), and the cycle Lyapunov exponent $\lambda_{p}=\ln \left|\Lambda_{p}\right| / n_{p}$. While most of the cycles have $\lambda_{p} \approx 0.5$, several significantly do not. The $\overline{0}$ cycle point is very unstable, isolated and transient fixed point, with no other cycles returning close to it. At period 13 one finds a pair of cycles with exceptionally low Lyapunov exponents. The cycles are close for most of the trajectory, differing only in the one symbol corresponding to two cycle points straddle the (partition) fold of the attractor. As the system is not hyperbolic, there is no known lower bound on cycle Lyapunov exponents, and the Hénon's strange "attractor" might some day turn out to be nothing but a transient on the way to a periodic attractor of some long period.

| $n$ | $M_{n}$ | $N_{n}$ | $n$ | $M_{n}$ | $N_{n}$ | $n$ | $M_{n}$ | $N_{n}$ |
| ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: |
| 11 | 14 | 156 | 17 | 166 | 2824 | 23 | 1930 | 44392 |
| 12 | 19 | 248 | 18 | 233 | 4264 | 24 | 2902 | 69952 |
| 13 | 32 | 418 | 19 | 364 | 6918 | 25 | 4498 | 112452 |
| 14 | 44 | 648 | 20 | 535 | 10808 | 26 | 6806 | 177376 |
| 15 | 72 | 1082 | 21 | 834 | 17544 | 27 | 10518 | 284042 |
| 16 | 102 | 1696 | 22 | 1225 | 27108 | 28 | 16031 | 449520 |
|  |  |  |  |  |  |  |  |  |

Table 31.2: The number of unstable periodic orbits of the Hénon map for $a=1.4$, $b=0.3$, of all periods $n \leq 28 . M_{n}$ is the number of prime cycles of length $n$, and $N_{n}$ is the total number of periodic points of period $n$ (including repeats of shorter prime cycles).

Figure 31.3: Typical trajectories of the vector field (31.9) for the stabilization of a hyperbolic fixed point of the Ikeda map (31.11) located at $(x, y) \approx(0.53275,0.24689)$. The circle indicates the position of the fixed point. Note that the basin of attraction of this fixed point is large, larger than the entire Ikeda attractor.


The idea of the relaxation algorithm illustrated by the above Hénon map example is that instead of searching for an unstable periodic orbit of a map, one searches for a stable attractor of a vector field. More generally, consider a $d$-dimensional map $x^{\prime}=f(x)$ with a hyperbolic fixed point $x_{*}$. Any fixed point $x_{*}$ is by construction an equilibrium point of the fictitious time flow

$$
\begin{equation*}
\frac{d x}{d \tau}=f(x)-x \tag{31.9}
\end{equation*}
$$

If all eigenvalues of the Jacobian matrix $J\left(x_{*}\right)=D f\left(x_{*}\right)$ have real parts smaller than unity, then $x_{*}$ is a stable equilibrium point of the flow.

If some of the eigenvalues have real parts larger than unity, then one needs to modify the vector field so that the corresponding directions of the flow are turned into stable directions in a neighborhood of the fixed point. In the spirit of (31.8), modify the flow by

$$
\begin{equation*}
\frac{d x}{d \tau}=\mathbf{C}(f(x)-x), \tag{31.10}
\end{equation*}
$$

where $\mathbf{C}$ is a $[d \times d]$ invertible matrix. The aim is to turn $x_{*}$ into a stable equilibrium point of the flow by an appropriate choice of $\mathbf{C}$. It can be shown that a set of permutation / reflection matrices with one and only one nonvanishing entry $\pm 1$ per row or column (for $d$-dimensional systems, there are $d!2^{d}$ such matrices) suffices to stabilize any fixed point. In practice, one chooses a particular matrix $\mathbf{C}$, and the flow is integrated. For each choice of $\mathbf{C}$, one or more hyperbolic fixed points of the map may turn into stable equilibria of the flow.

Example 31.2 Ikeda map. We illustrate the method with the determination of the periodic orbits of the Ikeda map:

$$
\begin{align*}
& x^{\prime}=1+a(x \cos w-y \sin w) \\
& y^{\prime}=a(x \sin w+y \cos w)  \tag{31.11}\\
& \text { where } w=b-\frac{c}{1+x^{2}+y^{2}}
\end{align*}
$$

with $a=0.9, b=0.4, c=6$. The fixed point $x_{*}$ is located at $(x, y) \approx(0.53275,0.24689)$, with eigenvalues of the Jacobian matrix $\left(\Lambda_{1}, \Lambda_{2}\right) \approx(-2.3897,-0.3389)$, so the flow is


Figure 31.4: Typical trajectories of the vector field (31.10) for a hyperbolic fixed point $(x, y) \approx(-0.13529,-0.37559)$ of $f^{3}$, where $f$ is the Ikeda map (31.11). The circle indicates the position of the fixed point. For the vector field corresponding to (a) $\mathbf{C}=\mathbf{1}, x_{*}$ is a hyperbolic equilibrium point of the flow, while for $(b) \mathbf{C}=\left(\begin{array}{cc}1 & 0 \\ 0 & -1\end{array}\right)$, $x_{*}$ is an attracting equilibrium point.
already stabilized with $\mathbf{C}=1$. Figure 31.3 depicts the flow of the vector field around the fixed point $x_{*}$.

In order to determine $x_{*}$, one needs to integrate the vector field (31.9) forward in time (the convergence is exponential in time), using a fourth order Runge-Kutta or any other integration routine.

In contrast, determination of the 3 -cycles of the Ikeda map requires nontrivial C matrices, different from the identity. Consider for example the hyperbolic fixed point $(x, y) \approx(-0.13529,-0.37559)$ of the third iterate $f^{3}$ of the Ikeda map. The flow of the vector field for $\mathbf{C}=1$, Figure 31.4(a), indicates a hyperbolic equilibrium point, while for $\mathbf{C}=\left(\begin{array}{cc}1 & 0 \\ 0 & -1\end{array}\right)$ the flow of the vector field, figure 31.4(b) indicates that $x_{*}$ is an attracting equilibrium point, reached at exponential speed by integration forward in time.

The generalization from searches for fixed points to searches for cycles is straightforward. In order to determine a prime cycle $x=\left(x_{1}, x_{2}, \ldots, x_{n}\right)$ of a $d$-dimensional map $x^{\prime}=f(x)$, we modify the multipoint shooting method of sect. 17.4.1, and consider the $n d$-dimensional vector field

$$
\begin{equation*}
\frac{d x}{d \tau}=\mathbf{C}(f(x)-x), \tag{31.12}
\end{equation*}
$$

where $f(x)=\left(f\left(x_{n}\right), f\left(x_{1}\right), f\left(x_{2}\right), \ldots, f\left(x_{n-1}\right)\right)$, and $\mathbf{C}$ is an invertible $[n d \times n d]$ matrix. For the Hénon map, it is sufficient to consider a set of $2^{n}$ diagonal matrices with eigenvalues $\pm 1$. Risking a bit of confusion, we denote by $x, f(x)$ both the $d$-dimensional vectors in (31.10), and $n d$ dimensional vectors in (31.12), as the structure of the equations is the same.

### 31.2 Discrete iteration relaxation method

(C. Chandre, F.K. Diakonos and P. Schmelcher)

The problem with the Newton-Raphson iteration (31.2) is that it requires very precise initial guesses. For example, the $n$th iterate of a unimodal
map has as many as $2^{n}$ periodic points crammed into the unit interval, so determination of all cycles of length $n$ requires that the initial guess for each one of them has to be accurate to roughly $2^{-n}$. This is not much of a problem for 1-dimensional maps, but making a good initial guess for where a cycle might lie in a $d$-dimensional phase space can be a challenge.

Emboldened by the success of the cyclist relaxation trick (31.8) of manually turning instability into stability by a sign change, we now $(i)$ abandon the Newton-Raphson method altogether, (ii) abandon the continuous fictitious time flow (31.9) with its time-consuming integration, replacing it by a map $g$ with a larger basin of attraction (not restricted to a linear neighborhood of the fixed point). The idea is to construct a very simple map $g$, a linear transformation of the original $f$, for which the fixed point is stable. We replace the Jacobian matrix prefactor in (31.2) (whose inversion can be time-consuming) by a constant matrix prefactor

$$
\begin{equation*}
x^{\prime}=g(x)=x+\Delta \tau \mathbf{C}(f(x)-x), \tag{31.13}
\end{equation*}
$$

where $\Delta \tau$ is a positive real number, and $\mathbf{C}$ is a $[d \times d]$ permutation and reflection matrix with one and only one non-vanishing entry $\pm 1$ per row or column. A fixed point of $f$ is also a fixed point of $g$. Since $\mathbf{C}$ is invertible, the inverse is also true.

This construction is motivated by the observation that for small $\Delta \tau \rightarrow$ $d \tau$ the map (31.13) is the Euler method for integrating the modified flow (31.10), with the integration step $\Delta \tau$.

The argument why a suitable choice of matrix $\mathbf{C}$ can lead to the stabilization of an unstable periodic orbit is similar to the one used to motivate the construction of the modified vector field in sect. 31.1. Indeed, the flow (31.8) is the simplest example of this method, with the infinitesimal fictitious time increment $\Delta \tau \rightarrow d \tau$, the infinitesimal coordinate correction $\left(x-x^{\prime}\right) \rightarrow d x_{i}$, and the $[n \times n]$ diagonal matrix $\mathbf{C} \rightarrow \sigma_{i}= \pm 1$.

For a given fixed point of $f(x)$ we again chose a $\mathbf{C}$ such that the flow in the expanding directions of $\mathbf{J}\left(x_{*}\right)$ is turned into a contracting flow. The aim is to stabilize $x_{*}$ by a suitable choice of $\mathbf{C}$. In the case where the map has multiple fixed points, the set of fixed points is obtained by changing the matrix $\mathbf{C}$ (in general different for each unstable fixed point) and varying initial conditions for the map $g$. For example, for 2-dimensional dissipative maps it can be shown that the 3 matrices

$$
\mathbf{C} \in\left\{\left(\begin{array}{ll}
1 & 0 \\
0 & 1
\end{array}\right),\left(\begin{array}{cc}
-1 & 0 \\
0 & 1
\end{array}\right),\left(\begin{array}{cc}
1 & 0 \\
0 & -1
\end{array}\right)\right\}
$$

suffice to stabilize all kinds of possible hyperbolic fixed points.
If $\Delta \tau$ is chosen sufficiently small, the magnitude of the eigenvalues of the fixed point $x_{*}$ in the transformed system are smaller than one, and
one has a stable fixed point. However, $\Delta \tau$ should not be chosen too small: Since the convergence is geometrical with a ratio $1-\alpha \Delta \tau$ (where the value of constant $\alpha$ depends on the stability of the fixed point in the original system), small $\Delta \tau$ can slow down the speed of convergence. The critical value of $\Delta \tau$, which just suffices to make the fixed point stable, can be read off from the quadratic equations relating the stability coefficients of the original system and those of the transformed system. In practice, one can find the optimal $\Delta \tau$ by iterating the dynamical system stabilized with a given $\mathbf{C}$ and $\Delta \tau$. In general, all starting points converge on the attractor provided $\Delta \tau$ is small enough. If this is not the case, the trajectory either diverges (if $\Delta \tau$ is far too large) or it oscillates in a small section of the phase space (if $\Delta \tau$ is close to its stabilizing value).

The search for the fixed points is now straightforward: A starting point chosen in the global neighborhood of the fixed point iterated with the transformed dynamical system $g$ converges to the fixed point due to its stability. Numerical investigations show that the domain of attraction of a stabilized fixed point is a rather extended connected area, by no means confined to a linear neighborhood. At times the basin of attraction encompasses the complete phase space of the attractor, so one can be sure to be within the attracting basin of a fixed point regardless of where on the on the attractor on picks the initial condition.

The step size $|g(x)-x|$ decreases exponentially when the trajectory approaches the fixed point. To get the coordinates of the fixed points with a high precision, one therefore needs a large number of iterations for the trajectory which is already in the linear neighborhood of the fixed point. To speed up the convergence of the final part of the approach to a fixed point we recommend a combination of the above approach with the NewtonRaphson method (31.2).

The fixed points of the $n$th iterate $f^{n}$ are cycle points of a cycle of period $n$. If we consider the map

$$
\begin{equation*}
x^{\prime}=g(x)=x+\Delta \tau \mathbf{C}\left(f^{n}(x)-x\right), \tag{31.14}
\end{equation*}
$$

the iterates of $g$ converge to a fixed point provided that $\Delta \tau$ is sufficiently small and $\mathbf{C}$ is a $[d \times d]$ constant matrix chosen such that it stabilizes the flow. As $n$ grows, $\Delta \tau$ has to be chosen smaller and smaller. In the case of the Ikeda map example 31.2 the method works well for $n \leq 20$. As in (31.12), the multipoint shooting method is the method of preference for determining longer cycles. Consider $x=\left(x_{1}, x_{2}, \ldots, x_{n}\right)$ and the $n d$-dimensional map

$$
x^{\prime}=f(x)=\left(f\left(x_{n}\right), f\left(x_{1}\right), \ldots, f\left(x_{n-1}\right)\right) .
$$

Determining cycles with period $n$ for the $d$-dimensional $f$ is equivalent to determining fixed points of the multipoint $d n$-dimensional $f$. The idea is to construct a matrix $\mathbf{C}$ such that the fixed point of $f$ becomes stable for the map:

$$
x^{\prime}=x+\Delta \tau \mathbf{C}(f(x)-x),
$$

where $\mathbf{C}$ is now a $[n d \times n d]$ permutation/reflection matrix with only one non-zero matrix element $\pm 1$ per row or column. For any given matrix $\mathbf{C}$, a certain fraction of the cycles becomes stable and can be found by iterating the transformed map which is now a $n d$ dimensional map.

From a practical point of view, the main advantage of this method compared to the Newton-Raphson method is twofold: $(i)$ the stability matrix of the flow need not be computed, so there is no large matrix to invert, simplifying considerably the implementation, and (ii) empirical basins of attractions for individual $\mathbf{C}$ are much larger than for the Newton-Raphson method. The price is a reduction in the speed of convergence.

### 31.3 Least action method

The methods of sects. 31.1 and 31.2 are somewhat $a d$ hoc, as for general flows and iterated maps there is no fundamental principle to guide us in chosing the cost function, such as (31.3), to vary.

For Hamiltonian dynamics, we are on much firmer ground; Maupertuis least action principle. You yawn your way through it in every mechanics course - Maupertuis believed that the principle provided a proof of the existence of God - but as we shall now see, it is a very hands-on numerical method for finding cycles.

Indeed, the simplest and numerically most robust method for determining cycles of planar billiards is given by the principle of least action, or equivalently, by extremizing the length of an approximate orbit that visits a given sequence of disks. In contrast to the multipoint shooting method of sect. 17.4.1 which requires variation of $2 N$ phase-space points, extremization of a cycle length requires variation of only $N$ bounce positions $s_{i}$.

The problem is to find the extremum values of cycle length $L(s)$ where $s=\left(s_{1}, \ldots, s_{N}\right)$, that is find the roots of $\partial_{i} L(s)=0$. Expand to first order

$$
\partial_{i} L\left(s_{0}+\delta s\right)=\partial_{i} L\left(s_{0}\right)+\sum_{j} \partial_{i} \partial_{j} L\left(s_{0}\right) \delta s_{j}+\ldots
$$

and use $\mathbf{J}_{i j}\left(s_{0}\right)=\partial_{i} \partial_{j} L\left(s_{0}\right)$ in the $N$-dimensional Newton-Raphson iteration scheme of sect. 17.2.2

$$
\begin{equation*}
s_{i} \mapsto s_{i}-\sum_{j}\left(\frac{1}{\mathbf{J}(s)}\right)_{i j} \partial_{j} L(s) \tag{31.15}
\end{equation*}
$$

The extremization is achieved by recursive implementation of the above algorithm, with proviso that if the dynamics is pruned, one also has to

| $p$ | $\Lambda_{p}$ | $T_{p}$ |
| :--- | ---: | ---: |
| 0 | 9.898979485566 | 4.000000000000 |
| 1 | $-1.177145519638 \times 10^{1}$ | 4.267949192431 |
| 01 | $-1.240948019921 \times 10^{2}$ | 8.316529485168 |
| 001 | $-1.240542557041 \times 10^{3}$ | 12.321746616182 |
| 011 | $1.449545074956 \times 10^{3}$ | 12.580807741032 |
| 0001 | $-1.229570686196 \times 10^{4}$ | 16.322276474382 |
| 0011 | $1.445997591902 \times 10^{4}$ | 16.585242906081 |
| 0111 | $-1.707901900894 \times 10^{4}$ | 16.849071859224 |
| 00001 | $-1.217338387051 \times 10^{5}$ | 20.322330025739 |
| 00011 | $1.432820951544 \times 10^{5}$ | 20.585689671758 |
| 00101 | $1.539257907420 \times 10^{5}$ | 20.638238386018 |
| 00111 | $-1.704107155425 \times 10^{5}$ | 20.853571517227 |
| 01011 | $-1.799019479426 \times 10^{5}$ | 20.897369388186 |
| 01111 | $2.010247347433 \times 10^{5}$ | 21.116994322373 |
| 000001 | $-1.205062923819 \times 10^{6}$ | 24.322335435738 |
| 000011 | $1.418521622814 \times 10^{6}$ | 24.585734788507 |
| 000101 | $1.525597448217 \times 10^{6}$ | 24.638760250323 |
| 000111 | $-1.688624934257 \times 10^{6}$ | 24.854025100071 |
| 001011 | $-1.796354939785 \times 10^{6}$ | 24.902167001066 |
| 001101 | $-1.796354939785 \times 10^{6}$ | 24.902167001066 |
| 001111 | $2.005733106218 \times 10^{6}$ | 25.121488488111 |
| 010111 | $2.119615015369 \times 10^{6}$ | 25.165628236279 |
| 011111 | $-2.366378254801 \times 10^{6}$ | 25.384945785676 |

Table 31.3: All prime cycles up to 6 bounces for the 3-disk fundamental domain, center-to-center separation $R=6$, disk radius $a=1$. The columns list the cycle itinerary, its expanding eigenvalue $\Lambda_{p}$, and the length of the orbit (if the velocity=1 this is the same as its period or the action). Note that the two 6 cycles $\overline{001011}$ and $\overline{001101}$ are degenerate due to the time reversal symmetry, but are not related by any discrete spatial symmetry. (Computed by P.E. Rosenqvist.)
check that the final extremal length orbit does not penetrate a billiard wall.

As an example, the short periods and stabilities of 3-disk cycles computed this way are listed table 31.3.

## Commentary

Remark 31.1 Piece-wise linear maps. The Lozi map (3.14) is linear, and 100,000's of cycles can be easily computed by [2x2] matrix multiplication and inversion.

Remark 31.2 Relaxation method. The relaxation (or gradient) algorithm is one of the methods for solving extremal problems [31.13]. The method described above was introduced by Biham and Wenzel [31.1], who have also generalized it (in the case of the Hénon map) to determination of all $2^{n}$ cycles of period $n$, real or complex [31.2]. The applicability and reliability of the method is discussed in detail by Grassberger, Kantz and Moening [31.5], who give examples of the ways in which the method fails: (a) it might reach a limit cycle rather than a equilibrium saddlepoint (that can be remedied by the complex Biham-Wenzel algorithm [31.2]) (b) different symbol sequences can converge to the same cycle (that is, more refined initial conditions might be needed). Furthermore, Hansen (ref. [31.7] and chapter
4. of ref. [1.3]) has pointed out that the method cannot find certain cycles for specific values of the Hénon map parameters.

In practice, the relaxation method for determining periodic orbits of maps appears to be effective almost always, but not always. It is much slower than the multipoint shooting method of sect. 17.4.1, but also much quicker to program, as it does not require evaluation of stability matrices and their inversion. If the complete set of cycles is required, the method has to be supplemented by other methods.

Remark 31.3 Hybrid Newton-Raphson/relaxation methods. The method discussed in sect. 31.2 was introduced by Schmelcher et al [31.9]. The method was extended to flows by means of the Poincaré surface of section technique in ref. [31.10]. It is also possible to combine the Newton-Raphson method and (31.13) in the construction of a transformed map [31.14]. In this approach, each step of the iteration scheme is a linear superposition of a step of the stability transformed system and a step of the Newton-Raphson algorithm. Far from the linear neighborhood the weight is dominantly on the globally acting stability transformation algorithm. Close to the fixed point, the steps of the iteration are dominated by the NewtonRaphson procedure.

Remark 31.4 Relation to the Smale horseshoe symbolic dynamics. For a complete horseshoe Hénon repeller ( $a$ sufficiently large), such as the one given in figure 31.2 , the signs $\sigma_{i} \in\{1,-1\}$ are in a 1-to-1 correspondence with the Smale horsheshoe symbolic dynamics $s_{i} \in\{0,1\}$ :

$$
s_{i}=\left\{\begin{array}{lll}
0 & \text { if } \sigma_{i}=-1, & x_{i}<0  \tag{31.16}\\
1 & \text { if } \sigma_{i}=+1, & x_{i}>0
\end{array} .\right.
$$

For arbitrary parameter values with a finite subshift symbolic dynamics or with arbitrarily complicated pruning, the relation of sign sequences $\left\{\sigma_{1}, \sigma_{2}, \cdots, \sigma_{n}\right\}$ to the itineraries $\left\{s_{1}, s_{2}, \cdots, s_{n}\right\}$ can be much subtler; this is discussed in ref. [31.5].

Remark 31.5 Ikeda map. Ikeda map (31.11) was introduced in ref. [31.12] is a model which exhibits complex dynamics observed in nonlinear optical ring cavities.

Remark 31.6 Relaxation for continuous time flows. For a $d$-dimensional flow $\dot{x}=v(x)$, the method described above can be extended by considering a Poincaré surface of section. The Poincaré section yields a map $f$ with dimension $d-1$, and the above discrete iterative maps procedures can be carried out. A method that keeps the trial orbit continuous throughout the calculation is the Newton descent, a variational method for finding periodic orbits of continuous time flows, is described in refs. [31.15, 31.16].

Remark 31.7 Stability ordering. The parameter $\Delta \tau$ in (31.13) is a key quantity here. It is related to the stability of the desired cycle in the transformed system: The more unstable a fixed point is, the smaller $\Delta \tau$ has to be to stabilize
it. With increasing cycle periods, the unstable eigenvalue of the stability matrix increases and therefore $\Delta \tau$ has to be reduced to achieve stabilization of all fixed points. In many cases the least unstable cycles of a given period $n$ are of physically most important [31.11]. In this context $\Delta \tau$ operates as a stability filter. It allows the selective stabilization of only those cycles which posses Lyapunov exponents smaller than a cut-off value. If one starts the search for cycles within a given period $n$ with a value $\Delta \tau \approx O\left(10^{-1}\right)$, and gradually lowers $\Delta \tau$ one obtains the sequence of all unstable orbits of order $n$ sorted with increasing values of their Lyapunov exponents. For the specific choice of $\mathbf{C}$ the relation between $\Delta \tau$ and the stability coefficients of the fixed points of the original system is strictly monotonous. Transformed dynamical systems with other C's do not obey such a strict behavior but show a rough ordering of the sequence of stability eigenvalues of the fixed points stabilized in the course of decreasing values for $\Delta \tau$. As explained in sect. 18.4, stability ordered cycles are needed to order cycle expansions of dynamical quantities of chaotic systems for which a symbolic dynamics is not known. For such systems, an ordering of cycles with respect to their stability has been proposed [18.13, 18.14, 18.12], and shown to yield good results in practical applications.

Remark 31.8 Action extremization method. The action extremization (sect. 31.3) as a numerical method for finding cycles has been introduced independently by many people. We have learned it from G. Russberg, and from M. Sieber's and F. Steiner's hyperbola billiard computations [31.17, 31.18]. The convergence rate is really impressive, for the Sinai billiard some 5000 cycles are computed within CPU seconds with rather bad initial guesses.

Variational methods are the key ingredient of the Aubry-Mather theory of areapreserving twist maps (known in the condensed matter literature as the FrenkelKontorova models of 1-dimensional crystals), discrete-time Hamiltonian dynamical systems particularly suited to explorations of the K.A.M. theorem. Proofs of the Aubry-Mather theorem [31.20] on existence of quasi-periodic solutions are variational. It was quickly realized that the variational methods can also yield reliable, high precision computations of long periodic orbits of twist map models in 2 or more dimensions, needed for K.A.M. renormalization studies [31.19].

A fictitious time gradient flow similar to the one discussed here in sect. 31.1 was introduced by Anegent [31.21] for twist maps, and used by Gole [31.22] in his proof of the Aubry-Mather theorem. Mathematical bounds on the regions of stability of K.A.M. tori are notoriously restrictive compared to the numerical indications, and de la Llave, Falcolini and Tompaidis [31.23, 31.24] have found the gradient flow formulation advantageous both in studies of the analyticity domains of the K.A.M. stability, as well as proving the Aubry-Mather theorem for extended systems (for a pedagogical introduction, see the lattice dynamics section of ref. [31.25]).

All of the twist-maps work is based on extremizing the discrete dynamics version of the action $S$ (in this context sometimes called a "generating function"). However, in their investigations in the complex plane, Falcolini and de la Llave [31.23] do find it useful to minimize instead $S \bar{S}$, analogous to our cost function (31.3).

## Résumé

Unlike the Newton-Raphson method, variational methods are very robust. As each step around a cycle is short, they do not suffer from exponential instabilities, and with rather coarse initial guesses one can determine cycles of arbitrary length.

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## Exercises

Exercise 31.1 Evaluation of cycles by minimization*. Given a symbol sequence, you can construct a guess trajectory by taking a point on the boundary of each disk in the sequence, and connecting them by straight lines. If this were a rubber band wrapped through 3 rings, it would shrink into the physical trajectory, which minimizes the action (in this case, the length) of the trajectory.

Write a program to find the periodic orbits for your billiard simulator. Use the least action principle to extremize the length of the periodic orbit, and reproduce the periods and stabilities of 3-disk cycles, table 31.3. After that check the accuracy of the computed orbits by iterating them forward with your simulator. What is $\left|f^{T_{p}}(x)-x\right|$ ?

## Exercise 31.2 Tracking cycles adiabatically*. Once a cycle has been

 found, orbits for different system parameters values may be obtained by varying slowly (adiabatically) the parameters, and using the old orbit points as starting guesses in the Newton method. Try this method out on the 3-disk system. It works well for $R: a$ sufficiently large. For smaller values, some orbits change rather quickly and require very small step sizes. In addition, for ratios below $R: a=2.04821419 \ldots$ families of cycles are pruned, that is some of the minimal length trajectories are blocked by intervening disks.Exercise 31.3 Find cycles of the Hénon map. Apply the method of sect. 31.1 to the Hénon map at the Hénon's parameters choice $a=1.4, b=0.3$, and compute all prime cycles for at least $n \leq 6$. Estimate the topological entropy, either from the definition (13.1), or as the zero of a truncated topological zeta function (13.21). Do your cycles agree with the cycles listed in table 31.1?

## Chapter 32

## Quantum scattering

Scattering is easier than gathering. Irish proverb
(A. Wirzba, P. Cvitanović and N. Whelan)

So far the trace formulas have been derived assuming that the system under consideration is bound. As we shall now see, we are in luck - the semiclassics of bound systems is all we need to understand the semiclassics for open, scattering systems as well. We start by a brief review of the quantum theory of elastic scattering of a point particle from a (repulsive) potential, and then develop the connection to the standard Gutzwiller theory for bound systems. We do this in two steps - first, a heuristic derivation which helps us understand in what sense density of states is "density", and then we sketch a general derivation of the central result of the spectral theory of quantum scattering, the Krein-Friedel-Lloyd formula. The end result is that we establish a connection between the scattering resonances (both positions and widths) of an open quantum system and the poles of the trace of the Green function, which we learned to analyze in earlier chapters.

### 32.1 Density of states

For a scattering problem the density of states (26.18) appear ill defined since formulas such as (30.6) involve integration over infinite spatial extent. What we will now show is that a quantity that makes sense physically is the difference of two densities - the first with the scatterer present and the second with the scatterer absent.

In nonrelativistic dynamics the relative motion can be separated from the center-of-mass motion. Therefore the elastic scattering of two particles can be treated as the scattering of one particle from a static potential $V(q)$. We will study the scattering of a point-particle of (reduced) mass $m$ by a short-range potential $V(q)$, excluding inter alia the Coulomb potential. (The Coulomb potential decays slowly as a function of $q$ so that various
asymptotic approximations which apply to general potentials fail for it.) Although we can choose the spatial coordinate frame freely, it is advisable to place its origin somewhere near the geometrical center of the potential. The scattering problem is solved, if a scattering solution to the time-independent Schrödinger equation (26.5)

$$
\begin{equation*}
\left(-\frac{\hbar^{2}}{2 m} \frac{\partial^{2}}{\partial q^{2}}+V(q)\right) \phi_{\vec{k}}(q)=E \phi_{\vec{k}}(q) \tag{32.1}
\end{equation*}
$$

can be constructed. Here $E$ is the energy, $\vec{p}=\hbar \vec{k}$ the initial momentum of the particle, and $\vec{k}$ the corresponding wave vector.

When the argument $r=|q|$ of the wave function is large compared to the typical size $a$ of the scattering region, the Schrödinger equation effectively becomes a free particle equation because of the short-range nature of the potential. In the asymptotic domain $r \gg a$, the solution $\phi_{\vec{k}}(q)$ of (32.1) can be written as superposition of ingoing and outgoing solutions of the free particle Schrödinger equation for fixed angular momentum:

$$
\phi(q)=A \phi^{(-)}(q)+B \phi^{(+)}(q), \quad(+ \text { boundary conditions }),
$$

where in 1-dimensional problems $\phi^{(-)}(q), \phi^{(+)}(q)$ are the "left", "right" moving plane waves, and in higher-dimensional scattering problems the "incoming", "outgoing" radial waves, with the constant matrices $A, B$ fixed by the boundary conditions. What are the boundary conditions? The scatterer can modify only the outgoing waves (see figure 32.1 ), since the incoming ones, by definition, have yet to encounter the scattering region. This defines the quantum mechanical scattering matrix, or the $S$ matrix

$$
\begin{equation*}
\phi_{m}(r)=\phi_{m}^{(-)}(r)+S_{m m^{\prime}} \phi_{m^{\prime}}^{(+)}(r) . \tag{32.2}
\end{equation*}
$$

All scattering effects are incorporated in the deviation of $\mathbf{S}$ from the unit matrix, the transition matrix $\mathbf{T}$

$$
\begin{equation*}
\mathbf{S}=\mathbf{1}-i \mathbf{T} . \tag{32.3}
\end{equation*}
$$

For concreteness, we have specialized to two dimensions, although the final formula is true for arbitrary dimensions. The indices $m$ and $m^{\prime}$ are the angular momenta quantum numbers for the incoming and outgoing state of the scattering wave function, labeling the $S$-matrix elements $S_{m m^{\prime}}$. More generally, given a set of quantum numbers $\beta, \gamma$, the $S$ matrix is a collection $S_{\beta \gamma}$ of transition amplitudes $\beta \rightarrow \gamma$ normalized such that $\left|S_{\beta \gamma}\right|^{2}$ is the probability of the $\beta \rightarrow \gamma$ transition. The total probability that the ingoing state $\beta$ ends up in some outgoing state must add up to unity

$$
\begin{equation*}
\sum_{\gamma}\left|S_{\beta \gamma}\right|^{2}=1, \tag{32.4}
\end{equation*}
$$



Figure 32.1: (a) Incoming spherical waves running into an obstacle. (b) Superposition of outgoing spherical waves scattered from an obstacle.
so the $S$ matrix is unitary: $\mathbf{S}^{\dagger} \mathbf{S}=\mathbf{S} \mathbf{S}^{\dagger}=\mathbf{1}$.
We have already encountered a solution to the 2-dimensional problem; free particle propagation Green's function (28.48) is a radial solution, given in terms of the Hankel function

$$
G_{0}(r, 0, E)=-\frac{i m}{2 \hbar^{2}} H_{0}^{(+)}(k r),
$$

where we have used $S_{0}(r, 0, E) / \hbar=k r$ for the action. The $m$ th angular momentum eigenfunction is proportional to $\phi_{m}^{( \pm)}(q) \propto H_{m}^{( \pm)}(k r)$, and given a potential $V(q)$ we can in principle compute the infinity of matrix elements $S_{m m^{\prime}}$. We will not need much information about $H_{m}^{(t)}(k r)$, other than that for large $r$ its asymptotic form is

$$
H^{ \pm} \propto e^{ \pm i k r}
$$

In general, the potential $V(q)$ is not radially symmetric and (32.1) has to be solved numerically, by explicit integration, or by diagonalizing a large matrix in a specific basis. To simplify things a bit, we assume for the time being that a radially symmetric scatterer is centered at the origin; the final formula will be true for arbitrary asymmetric potentials. Then the solutions of the Schrödinger equation (26.5) are separable, $\phi_{m}(q)=\phi(r) e^{i m \theta}, r=|q|$, the scattering matrix cannot mix different angular momentum eigenstates, and $S$ is diagonal in the radial basis (32.2) with matrix elements given by

$$
\begin{equation*}
S_{m}(k)=e^{2 i \delta_{m}(k)} . \tag{32.5}
\end{equation*}
$$

The matrix is unitary so in a diagonal basis all entries are pure phases. This means that an incoming state of the form $H_{m}^{(-)}(k r) e^{i m \theta}$ gets scattered into an outgoing state of the form $S_{m}(k) H_{m}^{(+)}(k r) e^{i m \theta}$, where $H_{m}^{(\mp)}(z)$ are incoming and outgoing Hankel functions respectively. We now embed the scatterer in a infinite cylindrical well of radius $R$, and will later take $R \rightarrow \infty$.

Figure 32.2: The "difference" of two bounded reference systems, one with and one without the scattering system.


Angular momentum is still conserved so that each eigenstate of this (now bound) problem corresponds to some value of $m$. For large $r \gg a$ each eigenstate is of the asymptotically free form

$$
\begin{align*}
\phi_{m}(r) & \approx e^{i m \theta}\left(S_{m}(k) H_{m}^{(+)}(k r)+H_{m}^{(-)}(k r)\right) \\
& \approx \cdots \cos \left(k r+\delta_{m}(k)-\chi_{m}\right) \tag{32.6}
\end{align*}
$$

where $\cdots$ is a common prefactor, and $\chi_{m}=m \pi / 2+\pi / 4$ is an annoying phase factor from the asymptotic expansion of the Hankel functions that will play no role in what follows.

The state (32.6) must satisfy the external boundary condition that it vanish at $r=R$. This implies the quantization condition

$$
k_{n} R+\delta_{m}\left(k_{n}\right)-\chi_{m}=\pi(n+12) .
$$

We now ask for the difference in the eigenvalues of two consecutive states of fixed $m$. Since $R$ is large, the density of states is high, and the phase $\delta_{m}(k)$ does not change much over such a small interval. Therefore, to leading order we can include the effect of the change of the phase on state $n+1$ by Taylor expanding. is

$$
k_{n+1} R+\delta_{m}\left(k_{n}\right)+\left(k_{n+1}-k_{n}\right) \delta_{m}^{\prime}\left(k_{n}\right)-\chi_{m} \approx \pi+\pi(n+12)
$$

Taking the difference of the two equations we obtain $\Delta k \approx \pi\left(R+\delta_{m}^{\prime}(k)\right)^{-1}$. This is the eigenvalue spacing which we now interpret as the inverse of the density of states within $m$ angular momentum sbuspace

$$
d_{m}(k) \approx \frac{1}{\pi}\left(R+\delta_{m}^{\prime}(k)\right)
$$

The $R$ term is essentially the $1-d$ Weyl term (30.8), appropriate to $1-d$ radial quantization. For large $R$, the dominant behavior is given by the size of the circular enclosure with a correction in terms of the derivative of the scattering phase shift, approximation accurate to order $1 / R$. However, not all is well: the area under consideration tends to infinity. We regularize this by subtracting from the result from the free particle density of states $d_{0}(k)$, for the same size container, but this time without any scatterer, figure 32.2. We also sum over all $m$ values so that

$$
\begin{align*}
d(k)-d_{0}(k)=\frac{1}{\pi} \sum_{m} \delta_{m}^{\prime}(k) & =\frac{1}{2 \pi i} \sum_{m} \frac{d}{d k} \log S_{m} \\
& =\frac{1}{2 \pi i} \operatorname{Tr}\left(S^{\dagger} \frac{d S}{d k}\right) . \tag{32.7}
\end{align*}
$$

The first line follows from the definition of the phase shifts (32.5) while the second line follows from the unitarity of $S$ so that $S^{-1}=S^{\dagger}$. We can now take the limit $R \rightarrow \infty$ since the $R$ dependence has been cancelled away.

This is essentially what we want to prove since for the left hand side we already have the semiclassical theory for the trace of the difference of Green's functions,

$$
\begin{equation*}
d(k)-d_{0}(k)=-\frac{1}{2 \pi k} \operatorname{Im}\left(\operatorname{tr}\left(G(k)-G_{0}(k)\right) .\right. \tag{32.8}
\end{equation*}
$$

There are a number of generalizations. This can be done in any number of dimensions. It is also more common to do this as a function of energy and not wave number $k$. However, as the asymptotic dynamics is free wave dynamics labeled by the wavenumber $k$, we have adapted $k$ as the natural variable in the above discussion.

Finally, we state without proof that the relation (32.7) applies even when there is no circular symmetry. The proof is more difficult since one cannot appeal to the phase shifts $\delta_{m}$ but must work directly with a nondiagonal $S$ matrix.

### 32.2 Quantum mechanical scattering matrix

The results of the previous section indicate that there is a connection between the scattering matrix and the trace of the quantum Green's function (more formally between the difference of the Green's function with and without the scattering center.) We now show how this connection can be derived in a more rigorous manner. We will also work in terms of the energy $E$ rather than the wavenumber $k$, since this is the more usual exposition. Suppose particles interact via forces of sufficiently short range, so that in the remote past they were in a free particle state labeled $\beta$, and in the distant future they will likewise be free, in a state labeled $\gamma$. In the Heisenberg picture the $S$-matrix is defined as $\mathbf{S}=\Omega_{-} \Omega_{+}^{\dagger}$ in terms of the Møller operators

$$
\begin{equation*}
\Omega_{ \pm}=\lim _{t \rightarrow \pm \infty} e^{i H t / \hbar} e^{-i H_{0} t / \hbar} \tag{32.9}
\end{equation*}
$$

where $H$ is the full Hamiltonian, whereas $H_{0}$ is the free Hamiltonian. In the interaction picture the $S$-matrix is given by

$$
\mathbf{S}=\Omega_{+}^{\dagger} \Omega_{-}=\lim _{t \rightarrow \infty} e^{i H_{0} t / \hbar} e^{-2 i H t / \hbar} e^{i H_{0} t / \hbar}
$$

$$
\begin{equation*}
=T \exp \left(-i \int_{-\infty}^{+\infty} d t H^{\prime}(t)\right) \tag{32.10}
\end{equation*}
$$

where $H^{\prime}=V=H-H_{0}$ is the interaction Hamiltonian and $T$ is the timeordering operator. In stationary scattering theory the $S$ matrix has the following spectral representation

$$
\begin{aligned}
S & =\int_{0}^{\infty} d E S(E) \delta\left(H_{0}-E\right) \\
S(E) & =Q_{+}(E) Q_{-}^{-1}(E), \quad Q_{ \pm}(E)=\mathbf{1}+\left(H_{0}-E \pm i \epsilon\right)^{-1}(32.11)
\end{aligned}
$$

such that

$$
\begin{equation*}
\operatorname{Tr}\left[S^{\dagger}(E) \frac{d}{d E} S(E)\right]=\operatorname{Tr}\left[\frac{1}{H_{0}-E-i \epsilon}-\frac{1}{H-E-i \epsilon}-(\epsilon \leftrightarrow-\epsilon)\right] \tag{32.12}
\end{equation*}
$$

The manipulations leading to (32.12) are justified if the operators $Q_{ \pm}(E)$ can be linked to trace-class operators.

We can now use this result to derive the Krein-Lloyd formula which is the central result of this chapter. The Krein-Lloyd formula provides the connection between the trace of the Green's function and the poles of the scattering matrix, implicit in all of the trace formulas for open quantum systems which will be presented in the subsequent chapters.

### 32.3 Krein-Friedel-Lloyd formula

The link between quantum mechanics and semiclassics for scattering problems is provided by the semiclassical limit of the Krein-Friedel-Lloyd sum for the spectral density which we now derive. This derivation builds on the results of the last section and extends the discussion of the opening section.

In chapter 28 we linked the spectral density (see (26.18)) of a bounded system

$$
\begin{equation*}
d(E) \equiv \sum_{n} \delta\left(E_{n}-E\right) \tag{32.13}
\end{equation*}
$$

via the identity

$$
\begin{align*}
\delta\left(E_{n}-E\right) & =-\lim _{\epsilon \rightarrow 0} \frac{1}{\pi} \operatorname{Im} \frac{1}{E-E_{n}+i \epsilon} \\
& =-\lim _{\epsilon \rightarrow 0} \frac{1}{\pi} \operatorname{Im}\left\langle E_{n}\right| \frac{1}{E-H+i \epsilon}\left|E_{n}\right\rangle \\
& =\frac{1}{2 \pi i} \lim _{\epsilon \rightarrow 0}\left\langle E_{n}\right| \frac{1}{E-H-i \epsilon}-\frac{1}{E-H+i \epsilon}\left|E_{n}\right\rangle 3 ? \tag{32.14}
\end{align*}
$$

to the trace of the Green's function (30.1.1). Furthermore, in the semiclassical approximation, the trace of the Green's function is given by the Gutzwiller trace formula (30.11) in terms of a smooth Weyl term and an oscillating contribution of periodic orbits.

Therefore, the task of constructing the semiclassics of a scattering system is completed, if we can find a connection between the spectral density $d(E)$ and the scattering matrix $S$. We will see that (32.12) provides the clue. Note that the right hand side of (32.12) has nearly the structure of (32.14) when the latter is inserted into (32.13). The principal difference between these two types of equations is that the $S$ matrix refers to outgoing scattering wave functions which are not normalizable and which have a continuous spectrum, whereas the spectral density $d(E)$ refers to a bound system with normalizable wave functions with a discrete spectrum. Furthermore, the bound system is characterized by a hermitian operator, the Hamiltonian $H$, whereas the scattering system is characterized by a unitary operator, the $S$-matrix. How can we reconcile these completely different classes of wave functions, operators and spectra? The trick is to put our scattering system into a finite box as in the opening section. We choose a spherical conatiner with radius $R$ and with its center at the center of our finite scattering system. Our scattering potential $V(\vec{r})$ will be unaltered within the box, whereas at the box walls we will choose an infinitely high potential, with the Dirichlet boundary conditions at the outside of the box:

$$
\begin{equation*}
\left.\phi(\vec{r})\right|_{r=R}=0 . \tag{32.15}
\end{equation*}
$$

In this way, for any finite value of the radius $R$ of the box, we have mapped our scattering system into a bound system with a spectral density $d(E ; R)$ over discrete eigenenergies $E_{n}(R)$. It is therefore important that our scattering potential was chosen to be short-ranged to start with. (Which explains why the Coulomb potential requires special care.) The hope is that in the limit $R \rightarrow \infty$ we will recover the scattering system. But some care is required in implementing this. The smooth Weyl term $\bar{d}(E ; R)$ belonging to our box with the enclosed potential $V$ diverges for a spherical two-dimensional box of radius $R$ quadratically, as $\pi R^{2} /(4 \pi)$ or as $R^{3}$ in the three-dimensional case. This problem can easily be cured if the spectral density of an empty reference box of the same size (radius $R$ ) is subtracted (see figure 32.2). Then all the divergences linked to the increasing radius $R$ in the limit $R \rightarrow \infty$ drop out of the difference. Furthermore, in the limit $R \rightarrow \infty$ the energy-eigenfunctions of the box are only normalizable as a delta distribution, similarly to a plane wave. So we seem to recover a continous spectrum. Still the problem remains that the wave functions do not discriminate between incoming and outgoing waves, whereas this symmetry, namely the hermiticity, is broken in the scattering problem. The last problem can be tackled if we replace the spectral density over discrete delta distributions by a smoothed spectral density with a small finite imaginary part $\eta$ in the energy $E$ :

$$
\begin{equation*}
d(E+i \eta ; R) \equiv \frac{1}{i 2 \pi} \sum_{n}\left\{\frac{1}{E-E_{n}(R)-i \eta}-\frac{1}{E-E_{n}(R)+i \eta}\right\} . \tag{32.16}
\end{equation*}
$$

Note that $d(E+i \eta ; R) \neq d(E-i \eta ; R)=-d(E+i \eta ; R)$. By the introduction of the positive finite imaginary part $\eta$ the time-dependent behavior of the wave function has effectively been altered from an oscillating one to a decaying one and the hermiticity of the Hamiltonian is removed. Finally the limit $\eta \rightarrow 0$ can be carried out, respecting the order of the limiting procedures. First the limit $R \rightarrow \infty$ has to be performed for a finite value of $\eta$, only then the limit $\eta \rightarrow 0$ is allowed. In practice, one can try to work with a finite value of $R$, but then it will turn out (see below) that the scattering system is only recovered if $R \sqrt{\eta} \gg 1$.

Let us summarize the relation between the smoothed spectral densities $d(E+i \eta ; R)$ of the boxed potential and $d^{(0)}(E+i \eta ; R)$ of the empty reference system and the $S$ matrix of the corresponding scattering system:

$$
\begin{aligned}
\lim _{\eta \rightarrow+0} \lim _{R \rightarrow \infty}\left(d(E+i \eta ; R)-d^{(0)}(E+i \eta ; R)\right) & =\frac{1}{2 \pi i} \operatorname{Tr}\left[S^{\dagger}(E) \frac{d}{d E} S(E)\right] \\
=\frac{1}{2 \pi i} \operatorname{Tr} \frac{d}{d E} \ln S(E) & =\frac{1}{2 \pi i} \frac{d}{d E} \ln \operatorname{det} S(\text { E我2.17) }
\end{aligned}
$$

This is the Krein-Friedel-Lloyd formula. It replaces the scattering problem by the difference of two bounded reference billiards of the same radius $R$ which finally will be taken to infinity. The first billiard contains the scattering region or potentials, whereas the other does not (see figure 32.2). Here $d(E+i \eta ; R)$ and $d^{(0)}(E+i \eta ; R)$ are the smoothed spectral densities in the presence or in the absence of the scatterers, respectively. In the semiclassical approximation, they are replaced by a Weyl term (30.10) and an oscillating sum over periodic orbits. As in (30.2), the trace formula (32.17) can be integrated to give a relation between the smoothed staircase functions and the determinant of the $S$-matrix:

$$
\lim _{\eta \rightarrow+0} \lim _{R \rightarrow \infty}\left(N(E+i \eta ; R)-N^{(0)}(E+i \eta ; R)\right)=\frac{1}{2 \pi i} \ln \operatorname{det} S\left(\left(B^{2}\right) .18\right)
$$

Furthermore, in both versions of the Krein-Friedel-Lloyd formulas the energy argument $E+i \eta$ can be replaced by the wavenumber argument $k+i \eta^{\prime}$. These expressions only make sense for wavenumbers on or above the real $k$-axis. In particular, if $k$ is chosen to be real, $\eta^{\prime}$ must be greater than zero. Otherwise, the exact left hand sides (32.18) and (32.17) would give discontinuous staircase or even delta function sums, respectively, whereas the right hand sides are continuous to start with, since they can be expressed by continuous phase shifts. Thus the order of the two limits in (32.18) and (32.17) is essential.

The necessity of the $+i \eta$ prescription can also be understood by purely phenomenological considerations in the semiclassical approximation: Without the $i \eta$ term there is no reason why one should be able to neglect spurious periodic orbits which are there solely because of the introduction of the confining boundary. The subtraction of the second (empty) reference system removes those spurious periodic orbits which never encounter the
scattering region - in addition to the removal of the divergent Weyl term contributions in the limit $R \rightarrow \infty$. The periodic orbits that encounter both the scattering region and the external wall would still survive the first limit $R \rightarrow \infty$, if they were not exponentially suppressed by the $+i \eta$ term because of their

$$
e^{i L(R) \sqrt{2 m(E+i \eta)}}=e^{i L(R) k} e^{-L(R) \eta^{\prime}}
$$

behavior. As the length $L(R)$ of a spurious periodic orbit grows linearly with the radius $R$. The bound $R \eta^{\prime} \gg 1$ is an essential precondition on the suppression of the unwanted spurious contributions of the container if the Krein-Friedel-Lloyd formulas (32.17) and (32.18) are evaluated at a finite value of $R$.

Finally, the semiclassical approximation can also help us in the interpretation of the Weyl term contributions for scattering problems. In scattering problems the Weyl term appears with a negative sign. The reason is the subtraction of the empty container from the container with the potential. If the potential is a dispersing billiard system (or a finite collection of dispersing billiards), we expect an excluded volume (or the sum of excluded volumes) relative to the empty container. In other words, the Weyl term contribution of the empty container is larger than of the filled one and therefore a negative net contribution is left over. Secondly, if the scattering potential is a collection of a finite number of non-overlapping scattering regions, the Krein-Friedel-Lloyd formulas show that the corresponding Weyl contributions are completely independent of the position of the single scatterers, as long as these do not overlap.

### 32.4 Wigner time delay

The term $\frac{d}{d E} \ln \operatorname{det} S$ in the density formula (32.17) is dimensionally time. This suggests another, physically important interpretation of such formulas for scattering systems, the Wigner delay, defined as

$$
\begin{align*}
d(k) & =\frac{d}{d k} \operatorname{Argdet}(\mathbf{S}(k)) \\
& =-i \frac{d}{d k} \log \operatorname{det}(\mathbf{S}(k) \\
& =-i \operatorname{tr}\left(\mathbf{S}^{\dagger}(k) \frac{d \mathbf{S}}{d k}(k)\right) \tag{32.19}
\end{align*}
$$

and can be shown to equal the total delay of a wave packet in a scattering system. We now review this fact.

A related quantity is the total scattering phase shift $\Theta(k)$ defined as

$$
\operatorname{det} \mathbf{S}(k)=e^{+i \Theta(k)},
$$

so that $d(k)=\frac{d}{d k} \Theta(k)$.
The time delay may be both positive and negative, reflecting attractive respectively repulsive features of the scattering system. To elucidate the connection between the scattering determinant and the time delay we study a plane wave:

The phase of a wave packet will have the form:

$$
\phi=\vec{k} \cdot \vec{x}-\omega t+\Theta
$$

Here the term in the parenthesis refers to the phase shift that will occur if scattering is present. The center of the wave packet will be determined by the principle of stationary phase:

$$
0=d \phi=d \vec{k} \cdot \vec{x}-d \omega t+d \Theta
$$

Hence the packet is located at

$$
\vec{x}=\frac{\partial \omega}{\partial \vec{k}} t-\frac{\partial \Theta}{\partial \vec{k}}
$$

The first term is just the group velocity times the given time $t$. Thus the the packet is retarded by a length given by the derivative of the phase shift with respect to the wave vector $\vec{k}$. The arrival of the wave packet at the position $\vec{x}$ will therefore be delayed. This time delay can similarly be found as

$$
\tau(\omega)=\frac{\partial \Theta(\omega)}{\partial \omega}
$$

To show this we introduce the slowness of the phase $\vec{s}=\vec{k} / \omega$ for which $\vec{s} \cdot \vec{v}_{g}=1$, where $\vec{v}_{g}$ is the group velocity to get

$$
d \vec{k} \cdot \vec{x}=\vec{s} \cdot \vec{x} d \omega=\frac{x}{v_{g}} d \omega
$$

since we may assume $\vec{x}$ is parallel to the group velocity (consistent with the above). Hence the arrival time becomes

$$
t=\frac{x}{v_{g}}+\frac{\partial \Theta(\omega)}{\partial \omega}
$$

If the scattering matrix is not diagonal, one interprets

$$
\Delta t_{i j}=\operatorname{Re}\left(-i S_{i j}^{-1} \frac{\partial S_{i j}}{\partial \omega}\right)=\operatorname{Re}\left(\frac{\partial \Theta_{i j}}{\partial \omega}\right)
$$

as the delay in the $j$ th scattering channel after an injection in the $i$ th. The probability for appearing in channel $j$ goes as $\left|S_{i j}\right|^{2}$ and therefore the average delay for the incoming states in channel $i$ is

$$
\begin{aligned}
\left\langle\Delta t_{i}\right\rangle & =\sum_{j}\left|S_{i j}\right|^{2} \Delta t_{i j}=\operatorname{Re}\left(-i \sum_{j} S_{i j}^{*} \frac{\partial S_{i j}}{\partial \omega}\right)=\operatorname{Re}\left(-i \mathbf{S}^{\dagger} \cdot \frac{\partial \mathbf{S}}{\partial \omega}\right)_{i i} \\
& =-i\left(\mathbf{S}^{\dagger} \cdot \frac{\partial \mathbf{S}}{\partial \omega}\right)_{i i}
\end{aligned}
$$

where we have used the derivative, $\partial / \partial \omega$, of the unitarity relation $\mathbf{S} \cdot \mathbf{S}^{\dagger}=\mathbf{1}$ valid for real frequencies. This discussion can in particular be made for wave packets related to partial waves and superpositions of these like an incoming plane wave corresponding to free motion. The total Wigner delay therefore corresponds to the sum over all channel delays (32.19).

## Commentary

Remark 32.1 Krein-Friedel-Lloyd formula. The third volume of Thirring [32.1], sections 3.6.14 (Levison Theorem) and 3.6.15 (the proof), or P. Scherer's thesis [32.15] (appendix) discusses the Levison Theorem.

It helps to start with a toy example or simplified example instead of the general theorem, namely for the radially symmetric potential in a symmetric cavity. Have a look at the book of K. Huang, chapter 10 (on the "second virial coefficient"), or Beth and Uhlenbeck [32.5], or Friedel [32.7]. These results for the correction to the density of states are particular cases of the Krein formula [32.3]. The Krein-Friedel-Lloyd formula (32.17) was derived in refs. [32.3, 32.7, 32.8, 32.9], see also refs. [32.11, $32.14,32.15,32.17,32.18]$. The original papers are by Krein and Birman [32.3, 32.4] but beware, they are mathematicans.

Also, have a look at pages 15-18 of Wirzba's talk on the Casimir effect [32.16]. Page 16 discusses the Beth-Uhlenbeck formula [32.5], the predecessor of the more general Krein formula for spherical cases.

Remark 32.2 Weyl term for empty container. For a discussion of why the Weyl term contribution of the empty container is larger than of the filled one and therefore a negative net contribution is left over, see ref. [32.15].

Remark 32.3 Wigner time delay. Wigner time delay and the Wigner-Smith time delay matrix, are powerful concepts for a statistical description of scattering. The diagonal elements $Q_{a a}$ of the lifetime matrix $\mathbf{Q}=-i \mathbf{S}^{-1} \partial \mathbf{S} / \partial \omega$, where $\mathbf{S}$ is the $[2 N \times 2 N]$ scattering matrix, are interpreted in terms of the time spent in the scattering region by a wave packet incident in one channel. As shown by Smith [32.26], they are the sum over all ouput channels (both in reflection and transmission) of $\Delta t_{a b}=\operatorname{Re}\left[\left(-i / S_{a b}\right)\left(\partial S_{a b} / \partial \omega\right)\right]$ weighted by the probability of emerging from that channel. The sum of the $Q_{a a}$ over all $2 N$ channels is the Wigner time delay $\tau_{W}=\sum_{a} Q_{a a}$, which is the trace of the lifetime matrix and is proportional to the density of states.

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## Exercises

Exercise 32.1 Spurious orbits under the Krein-Friedel-Lloyd contruction. Draw examples for the three types of period orbits under the Krein-Friedel-Lloyd construction: (a) the genuine periodic orbits of the scattering region, (b) spurious periodic orbits which can be removed by the subtraction of the reference system, (c) spurious periodic orbits which cannot be removed by this subtraction. What is the role of the double limit $\eta \rightarrow 0$, container size $b \rightarrow \infty$ ?

Exercise 32.2 The one-disk scattering wave function. Derive the one-disk scattering wave function.
(Andreas Wirzba)

Exercise 32.3 Quantum two-disk scattering. Compute the quasiclassical spectral determinant

$$
Z(\varepsilon)=\prod_{p, j, l}\left(1-\frac{t_{p}}{\Lambda_{p}^{j+2 l}}\right)^{j+1}
$$

for the two disk problem. Use the geometry


The full quantum mechanical version of this problem can be solved by finding the zeros in $k$ for the determinant of the matrix

$$
M_{m, n}=\delta_{m, n}+\frac{(-1)^{n}}{2} \frac{J_{m}(k a)}{H_{n}^{(1)}(k a)}\left(H_{m-n}^{(1)}(k R)+(-1)^{n} H_{m+n}^{(1)}(k R)\right),
$$

where $J_{n}$ is the nth Bessel function and $H_{n}^{(1)}$ is the Hankel function of the first kind. Find the zeros of the determinant closest to the origin by solving $\operatorname{det} M(k)=0$. (Hints: notice the structure $M=I+A$ to approximate the determinant; or read Chaos 2, 79 (1992))

Exercise 32.4 Pinball topological index. Upgrade your pinball simulator so that it computes the topological index for each orbit it finds.

## Chapter 33

## Chaotic multiscattering

(A. Wirzba and P. Cvitanović)

We will discuss here the semiclassics of scattering in open systems with a finite number of non-overlapping finite scattering regions. Why is this interesting at all? The semiclassics of scattering systems has five advantages compared to the bound-state problems such as the helium quantization discussed in chapter 34.

- For bound-state problems the semiclassical approximation does not respect quantum-mechanical unitarity, and the semi-classical eigenenergies are not real. In contrast, the semiclassical scattering matrix that we construct is manifestly unitary.
- The Weyl-term contributions decouple from the multi-scattering system.
- The close relation to the classical escape processes discussed in chapter 1.
- For scattering systems the derivation of cycle expansions is more direct and controlled than in the bound-state case. It will be seen that the semiclassical cycle expansion is the saddle-point approximation to the cumulant expansion of the determinant of the exact quantummechanical multi-scattering matrix.
- The region of convergence of the semiclassical spectral function is larger than is the case for the bound-state case.

We start by a brief review of the elastic scattering of a point particle from finite collection of non-overlapping scattering regions in terms of the standard textbook scattering theory, and then develop the semiclassical scattering trace formulas and spectral determinants for scattering off $N$ disks in a plane.

### 33.1 Quantum mechanical scattering matrix

We now specialize to the elastic scattering of a point particle from finite collection of $N$ non-overlapping reflecting disks in a 2-dimensional plane. As the point particle moves freely between the static scatterers, the time independent Schrödinger equation outside the scattering regions is the Helmholtz equation:

$$
\begin{equation*}
\left(\vec{\nabla}_{r}^{2}+\vec{k}^{2}\right) \psi(\vec{r})=0, \quad \vec{r} \text { outside the scattering regions. } \tag{33.1}
\end{equation*}
$$

Here $\psi(\vec{r})$ is the wave function of the point particle at spatial position $\vec{r}$ and $E=\hbar^{2} \vec{k}^{2} / 2 m$ is its energy written in terms of its mass $m$ and the wave vector $\vec{k}$ of the incident wave. For refelcting wall billiards the scattering problem is a boundary value problem with Dirichlet boundary conditions:

$$
\begin{equation*}
\psi(\vec{r})=0, \quad \vec{r} \text { on the perimeter of a scatterer. } \tag{33.2}
\end{equation*}
$$

As usual for scattering problems, we expand the wave function $\psi(\vec{r})$ in the (2-dimensional) angular momentum eigenfunctions basis

$$
\begin{equation*}
\psi(\vec{r})=\sum_{m=-\infty}^{\infty} \psi_{m}^{k}(\vec{r}) e^{-i m \Phi_{k}}, \tag{33.3}
\end{equation*}
$$

where $k$ and $\Phi_{k}$ are the length and angle of the wave vector, respectively. A plane wave in two dimensions expaned in the angular momentum basis is

$$
\begin{equation*}
e^{i \vec{k} \cdot \vec{r}}=e^{i k r \cos \left(\Phi_{r}-\Phi_{k}\right)}=\sum_{m=-\infty}^{\infty} J_{m}(k r) e^{i m\left(\Phi_{r}-\Phi_{k}\right)} \tag{33.4}
\end{equation*}
$$

where $r$ and $\Phi_{r}$ denote the distance and angle of the spatial vector $\vec{r}$ as measured in the global 2-dimensional coordinate system.

The $m$ th angular component $J_{m}(k r) e^{i m \Phi_{r}}$ of a plane wave is split into a superposition of incoming and outgoing 2-dimensional spherical waves by decomposing the ordinary Bessel function $J_{m}(z)$ into the sum

$$
\begin{equation*}
J_{m}(z)=\frac{1}{2}\left(H_{m}^{(1)}(z)+H_{m}^{(2)}(z)\right) \tag{33.5}
\end{equation*}
$$

of the Hankel functions $H_{m}^{(1)}(z)$ and $H_{m}^{(2)}(z)$ of the first and second kind. For $|z| \gg 1$ the Hankel functions behave asymptotically as:

$$
\begin{align*}
& H_{m}^{(2)}(z) \sim \sqrt{\frac{2}{\pi z}} e^{-i\left(z-\frac{\pi}{2} m-\frac{\pi}{4}\right)} \quad \text { incoming, } \\
& H_{m}^{(1)}(z) \sim \sqrt{\frac{2}{\pi z}} e^{+i\left(z-\frac{\pi}{2} m-\frac{\pi}{4}\right)} \quad \text { outgoing. } \tag{33.6}
\end{align*}
$$

Thus for $r \rightarrow \infty$ and $k$ fixed, the $m$ th angular component $J_{m}(k r) e^{i m \Phi_{r}}$ of the plane wave can be written as superposition of incoming and outgoing 2-dimensional spherical waves:

$$
\begin{equation*}
J_{m}(k r) e^{i m \Phi_{r}} \sim \frac{1}{\sqrt{2 \pi k r}}\left[e^{-i\left(k r-\frac{\pi}{2} m-\frac{\pi}{4}\right)}+e^{i\left(k r-\frac{\pi}{2} m-\frac{\pi}{4}\right)}\right] e^{i m \Phi_{r}} . \tag{33.7}
\end{equation*}
$$

In terms of the asymptotic (angular momentum) components $\psi_{m}^{k}$ of the wave function $\psi(\vec{r})$, the scattering matrix (32.3) is defined as

$$
\begin{equation*}
\psi_{m}^{k} \sim \frac{1}{\sqrt{2 \pi k r}} \sum_{m^{\prime}=-\infty}^{\infty}\left[\delta_{m m^{\prime}} e^{-i\left(k r-\frac{\pi}{2} m^{\prime}-\frac{\pi}{4}\right)}+S_{m m^{\prime}} e^{i\left(k r-\frac{\pi}{2} m^{\prime}-\frac{\pi}{4}\right)}\right] e^{i m^{\prime} \Phi_{r}} \tag{33.8}
\end{equation*}
$$

The matrix element $S_{m m^{\prime}}$ describes the scattering of an incoming wave with angular momentum $m$ into an outgoing wave with angular momentum $m^{\prime}$. If there are no scatterers, then $\mathbf{S}=\mathbf{1}$ and the asymptotic expression of the plane wave $e^{i \vec{k} \cdot \vec{r}}$ in two dimensions is recovered from $\psi(\vec{r})$.

### 33.1.1 1-disk scattering matrix

In general, $\mathbf{S}$ is nondiagonal and nonseparable. An exception is the 1-disk scatterer. If the origin of the coordinate system is placed at the center of the disk, by (33.5) the $m$ th angular component of the time-independent scattering wave function is a superposition of incoming and outgoing 2dimensional spherical waves

$$
\begin{aligned}
\psi_{m}^{k} & =\frac{1}{2}\left(H_{m}^{(2)}(k r)+S_{m m} H_{m}^{(1)}(k r)\right) e^{i m \Phi_{r}} \\
& =\left(J_{m}(k r)-\frac{i}{2} T_{m m} H_{m}^{(1)}(k r)\right) e^{i m \Phi_{r}} .
\end{aligned}
$$

The vanishing (33.2) of the wave function on the disk perimeter

$$
0=J_{m}(k a)-\frac{i}{2} T_{m m} H_{m}^{(1)}(k a)
$$

yields the 1-disk scattering matrix in analytic form:

$$
\begin{equation*}
S_{m m^{\prime}}^{s}(k)=\left(1-\frac{2 J_{m}\left(k a_{s}\right)}{H_{m}^{(1)}\left(k a_{s}\right)}\right) \delta_{m m^{\prime}}=-\frac{H_{m}^{(2)}\left(k a_{s}\right)}{H_{m}^{(1)}\left(k a_{s}\right)} \delta_{m m^{\prime}} \tag{33.9}
\end{equation*}
$$

where $a=a_{s}$ is radius of the disk and the suffix ${ }^{s}$ indicates that we are dealing with a disk whose label is $s$. We shall derive a semiclassical approximation to this 1-disk S-matrix in sect. 33.3.

### 33.1.2 Multi-scattering matrix

Consider next a scattering region consisting of $N$ non-overlapping disks labeled $s \in\{1,2, \cdots, N\}$, following the notational conventions of sect. 11.6. The strategy is to construct the full T-matrix (32.3) from the exact 1 disk scattering matrix (33.9) by a succession of coordinate rotations and translations such that at each step the coordinate system is centered at the origin of a disk. Then the T-matrix can be split into a product over three kinds of matrices,

$$
\begin{aligned}
& S_{m m^{\prime}}(k)=\delta_{m m^{\prime}}-i T_{m m^{\prime}}(k) \\
& T_{m m^{\prime}}(k)=\sum_{s, s^{\prime}=1}^{N} C_{m l_{s}}^{s}(k)\left(\mathbf{M}^{-1}(k)\right)_{l_{s} l_{s^{\prime}}}^{s s^{\prime}} D_{l_{s^{\prime}} m^{\prime}}^{s^{\prime}}(k) .
\end{aligned}
$$

The outgoing spherical wave scattered by the disk $s$ is obtained by shifting the (in general arbitrary) origin the distance $R_{s}$ to the center of the disk $s$, and measuring the angle $\Phi_{s}$ with respect to direction $\mathbf{k}$ of the outgoing spherical wave. As in (33.9) the matrix $\mathbf{C}^{s}$ takes form

$$
\begin{equation*}
C_{m l_{s}}^{s}=\frac{2 i}{\pi a_{s}} \frac{J_{m-l_{s}}\left(k R_{s}\right)}{H_{l_{s}}^{(1)}\left(k a_{s}\right)} e^{i m \Phi_{s}} . \tag{33.10}
\end{equation*}
$$

If we now describe the ingoing spherical wave in the disk $s^{\prime}$ coordinate frame by the matrix $\mathbf{D}^{s^{\prime}}$

$$
\begin{equation*}
D_{l_{s^{\prime}}^{\prime}}^{s^{\prime}}=-\pi a_{s^{\prime}} J_{m^{\prime}-l_{s^{\prime}}}\left(k R_{s^{\prime}}\right) J_{l_{s^{\prime}}}\left(k a_{s^{\prime}}\right) e^{-i m^{\prime} \Phi_{s^{\prime}}} \tag{33.11}
\end{equation*}
$$

and apply the Bessel function addition theorem

$$
J_{m}(y+z)=\sum_{\ell=-\infty}^{\infty} J_{m-\ell}(y) J_{\ell}(z)
$$

we obtain the correct T-matrix (33.9) for the single disk $s=s^{\prime}$ scattering. The Bessel function sum is a statement of the completness of the spherical wave basis; as we shift the origin from the disk $s$ to the disk $s^{\prime}$ by distance $R_{s^{\prime}}$, we have to reexpand all basis functions in the new coordinate frame.

The labels $m$ and $m^{\prime}$ refer to the angular momentum quantum numbers of the ingoing and outgoing waves in the global coordinate system, and $l_{s}, l_{s^{\prime}}$ refer to the (angular momentum) basis fixed at the $s$ th and $s$ 'th scatterer, respectively. Thus, $\mathbf{C}^{s}$ and $\mathbf{D}^{s^{\prime}}$ depend on the origin and orientation of the global coordinate system of the 2-dimensional plane as well as on the internal coordinates of the scatterers. As they can be made separable in the scatterer label $s$, they describe the single scatterer aspects of what, in general, is a multi-scattering problem.

Figure 33.1: Global and local coordinates for a general 3 -disk problem.


The matrix $\mathbf{M}$ is called the multi-scattering matrix. If the scattering problem consists only of one scatterer, $\mathbf{M}$ is simply the unit matrix $M_{l_{s} l_{s^{\prime}}}^{s s^{\prime}}=$ $\delta^{s s^{\prime}} \delta_{l_{s} l_{s} l^{\prime}}$. For scattering from more than one scatterer we separate out a "single traversal" matrix A which transports the scattered wave from a scattering region $\mathcal{M}_{s}$ to the scattering region $\mathcal{M}_{s^{\prime}}$,

$$
\begin{equation*}
M_{l_{s} l_{s^{\prime}}}^{s s^{\prime}}=\delta^{s s^{\prime}} \delta_{l_{s} l_{s^{\prime}}}-A_{l_{s} l_{s^{\prime}}}^{s s^{\prime}} . \tag{33.12}
\end{equation*}
$$

The matrix $\mathbf{A}^{s s^{\prime}}$ reads:

$$
\begin{equation*}
A_{l_{s} l_{s^{\prime}}}^{s s^{\prime}}=-\left(1-\delta^{s s^{\prime}}\right) \frac{a_{s}}{a_{s^{\prime}}} \frac{J_{l_{s}}\left(k a_{s}\right)}{H_{l_{s^{\prime}}}^{(1)}\left(k a_{s^{\prime}}\right)} H_{l_{s}-l_{s^{\prime}}}^{(1)}\left(k R_{s s^{\prime}}\right) e^{i\left(l_{s} \alpha_{s^{\prime}}-l_{s^{\prime}}\left(\alpha_{s s^{\prime}}-\pi\right)\right)} . \tag{33.13}
\end{equation*}
$$

Here, $a_{s}$ is the radius of the $s$ th disk. $R_{s}$ and $\Phi_{s}$ are the distance and angle, respectively, of the ray from the origin in the 2-dimensional plane to the center of disk $s$ as measured in the global coordinate system. Furthermore, $R_{s s^{\prime}}=R_{s^{\prime} s}$ is the separation between the centers of the $s$ th and $s^{\prime}$ th disk and $\alpha_{s^{\prime} s}$ of the ray from the center of disk $s$ to the center of disk $s^{\prime}$ as measured in the local (body-fixed) coordinate system of disk $s$ (see figure 33.1).

Expanded as a geometrical series about the unit matrix 1, the inverse matrix $\mathbf{M}^{-1}$ generates a multi-scattering series in powers of the singletraversal matrix $\mathbf{A}$. All genuine multi-scattering dynamics is contained in the matrix $\mathbf{A}$; by construction $\mathbf{A}$ vanishes for a single-scatterer system.

### 33.2 N -scatterer spectral determinant

In the following we limit ourselves to a study of the spectral properties of the $\mathbf{S}$-matrix: resonances, time delays and phase shifts. The resonances are given by the poles of the $\mathbf{S}$-matrix in the lower complex wave number $(k)$ plane; more precisely, by the poles of the $\mathbf{S}$ on the second Riemann sheet of the complex energy plane. As the $\mathbf{S}$-matrix is unitary, it is also natural to focus on its total phase shift $\eta(k)$ defined by $\operatorname{det} \mathbf{S}=\exp (i \eta(k))$. The
time-delay is proportional to the derivative of the phase shift with respect to the wave number $k$.

As we are only interested in spectral properties of the scattering problem, it suffices to study the determinant, $\operatorname{det} \mathbf{S}$ which is basis and coordinatesystem independent, whereas the $\mathbf{S}$-matrix itself depends on the global coordinate system and on the choice of basis for the point particle wave function.

As the $\mathbf{S}$-matrix is, in general, an infinite dimensional matrix, it is not clear whether the corresponding determinant exists at all. If $\mathbf{T}$-matrix is trace-class, the determinant does exist. What does this mean?

### 33.2.1 Trace-class operators

An operator (an infinite-dimensional matrix) is called trace-class if and only if, for any choice of orthonormal basis, the sum of the diagonal matrix elements converges absolutely; it is called "Hilbert-Schmidt", if the sum of the absolute squared diagonal matrix elements converges. Once an operator is diagnosed as trace-class, we are allowed to manipulate it as we manipulate finite-dimensional matrices. We review the theory of trace-class operators in appendix K; here we will assume that the T-matrix (32.3) is trace-class, and draw the conlusions.

If $\mathbf{A}$ is trace-class, the determinant $\operatorname{det}(\mathbf{1}-z \mathbf{A})$, as defined by the cumulant expansion, exists and is an entire function of $z$. Furthermore, the determinant is invariant under any unitary transformation.

The cumulant expansion is the analytical continuation (as Taylor expansion in the book-keeping variable $z$ ) of the determinant

$$
\operatorname{det}(\mathbf{1}-z \mathbf{A})=\exp [\operatorname{tr} \ln (\mathbf{1}-z \mathbf{A})]=\exp \left(-\sum_{n=1}^{\infty} \frac{z^{n}}{z^{n}} \operatorname{tr}\left(\mathbf{A}^{n}\right)\right) .
$$

That means

$$
\begin{equation*}
\operatorname{det}(\mathbf{1}-z \mathbf{A}):=\sum_{m=0}^{\infty} z^{m} Q_{m}(\mathbf{A}) \tag{33.14}
\end{equation*}
$$

where the cumulants $Q_{m}(\mathbf{A})$ satisfy the Plemelj-Smithies recursion formula (K.26), a generalization of Newton's formula to determinants of infinitedimensional matrices,

$$
\begin{align*}
Q_{0}(\mathbf{A}) & =1 \\
Q_{m}(\mathbf{A}) & =-\frac{1}{m} \sum_{j=1}^{m} Q_{m-j}(\mathbf{A}) \operatorname{tr}\left(\mathbf{A}^{j}\right) \quad \text { for } m \geq 1 \tag{33.15}
\end{align*}
$$

in terms of cumulants of order $n<m$ and traces of order $n \leq m$. Because of the trace-class property of $\mathbf{A}$, all cumulants and traces exist separately.

For the general case of $N<\infty$ non-overlapping scatterers, T-matrix can be shown to be trace-class, so the determinant of the $\mathbf{S}$-matrix is well defined. What does trace-class propoerty mean for the corresponding matrices $\mathbf{C}^{s}, \mathbf{D}^{s}$ and $\mathbf{A}^{s s^{\prime}}$ ? Manipulating the operators as though theye were finite matrices, we we can perform the following transformations:

$$
\begin{align*}
\operatorname{det} \mathbf{S} & =\operatorname{det}\left(\mathbf{1}-i \mathbf{C M}^{-1} \mathbf{D}\right)=\operatorname{Det}\left(\mathbf{1}-i \mathbf{M}^{-1} \mathbf{D C}\right) \\
& =\operatorname{Det}\left(\mathbf{M}^{-1}(\mathbf{M}-i \mathbf{D C})\right) \\
& =\frac{\operatorname{Det}(\mathbf{M}-i \mathbf{D C})}{\operatorname{Det}(\mathbf{M})} \ldots \tag{33.16}
\end{align*}
$$

In the first equality (33.16) the determinant is taken over small $\ell$ (the angular momentum with respect to the global system). In the remainder of (33.16) the determinant is evaluated over the multiple indices $L_{s}=\left(s, l_{s}\right)$. In order to signal this difference we use the following notation: det... and $\operatorname{tr} \ldots$ refer to the $|\ell\rangle$ space, Det $\ldots$ and $\operatorname{Tr} \ldots$ refer to the multiple index space. The matrices in the multiple index space are expanded in the complete basis $\left\{\left|L_{s}\right\rangle\right\}=\left\{\left|s, \ell_{s}\right\rangle\right\}$ which refers for fixed index $s$ to the origin of the $s$ th scatterer and not any longer to the origin of the 2-dimensional plane.

Let us explicitly extract the product of the determinants of the subsystems from the determinant of the total system (33.16):

$$
\begin{align*}
\operatorname{det} \mathbf{S} & =\frac{\operatorname{Det}(\mathbf{M}-i \mathbf{D C})}{\operatorname{Det}(\mathbf{M})} \\
& =\frac{\operatorname{Det}(\mathbf{M}-i \mathbf{D C})}{\operatorname{Det} \mathbf{M}} \frac{\prod_{s=1}^{N} \operatorname{det} \mathbf{S}^{s}}{\prod_{s=1}^{N} \operatorname{det} \mathbf{S}^{s}} \\
& =\left(\prod_{s=1}^{N} \operatorname{det} \mathbf{S}^{s}\right) \frac{\operatorname{Det}(\mathbf{M}-i \mathbf{D C}) / \prod_{s=1}^{N} \operatorname{det} \mathbf{S}^{s}}{\operatorname{Det} \mathbf{M}} . \tag{33.17}
\end{align*}
$$

The final step in the reformulation of the determinant of the $\mathbf{S}$-matrix of the $N$-scatterer problem follows from the unitarity of the $\mathbf{S}$-matrix. The unitarity of $\mathbf{S}^{\dagger}\left(k^{*}\right)$ implies for the determinant

$$
\begin{equation*}
\operatorname{det}\left(\mathbf{S}\left(k^{*}\right)^{\dagger}\right)=1 / \operatorname{det} \mathbf{S}(k), \tag{33.18}
\end{equation*}
$$

where this manipulation is allowed because the $\mathbf{T}$-matrix is trace-class. The unitarity condition should apply for the $\mathbf{S}$-matrix of the total system, $\mathbf{S}$, as for the each of the single subsystems, $\mathbf{S}^{s}, s=1, \cdots, N$. In terms of the result of (33.17), this implies

$$
\frac{\operatorname{Det}(\mathbf{M}(k)-i \mathbf{D}(k) \mathbf{C}(k))}{\prod_{s=1}^{N} \operatorname{det} \mathbf{S}^{s}}=\operatorname{Det}\left(\mathbf{M}\left(k^{*}\right)^{\dagger}\right)
$$

since all determinants in (33.17) exist separately and since the determinants $\operatorname{det} \mathbf{S}^{s}$ respect unitarity by themselves. Thus, we finally have

$$
\begin{equation*}
\operatorname{det} \mathbf{S}(k)=\left\{\prod_{s=1}^{N}\left(\operatorname{det} \mathbf{S}^{s}(k)\right)\right\} \frac{\operatorname{Det} \mathbf{M}\left(k^{*}\right)^{\dagger}}{\operatorname{Det} \mathbf{M}(k)}, \tag{33.19}
\end{equation*}
$$

where all determinants exist separately.
In summary: We assumed a scattering system of a finite number of non-overlapping scatterers which can be of different shape and size, but are all of finite extent. We assumed the trace-class character of the T-matrix belonging to the total system and of the single-traversal matrix $\mathbf{A}$ and finally unitarity of the $\mathbf{S}$-matrices of the complete and all subsystems.

What can one say about the point-particle scattering from a finite number of scatterers of arbitrary shape and size? As long as each of $N<\infty$ single scatterers has a finite spatial extent, i.e., can be covered by a finite disk, the total system has a finite spatial extent as well. Therefore, it too can be put insided a circular domain of finite radius $b$, e.g., inside a single disk. If the impact parameter of the point particle measured with respect to the origin of this disk is larger than the disk size (actually larger than $(e / 2) \times b)$, then the $\mathbf{T}$ matrix elements of the $N$-scatterer problem become very small. If the wave number $k$ is kept fixed, the modulus of the diagonal matrix elements, $\left|T_{m m}\right|$ with the angular momentum $m>(e / 2) k b$, is bounded by the corresponding quantity of the covering disk.

### 33.2.2 Quantum cycle expansions

In formula (33.19) the genuine multi-scattering terms are separated from the single-scattering ones. We focus on the multi-scattering terms, that is, on the ratio of the determinants of the multi-scattering matrix $\mathbf{M}=$ $\mathbf{1} \mathbf{-} \mathbf{A}$ in (33.19), since they are the origin of the periodic orbit sums in the semiclassical reduction. The resonances of the multi-scattering system are given by the zeros of $\operatorname{Det} \mathbf{M}(k)$ in the lower complex wave number plane.

In order to set up the problem for the semiclassical reduction, we express the determinant of the multi-scattering matrix in terms of the traces of the powers of the matrix $\mathbf{A}$, by means of the cumulant expansion (33.14). Because of the finite number $N \geq 2$ of scatterers $\operatorname{tr}\left(\mathbf{A}^{n}\right)$ receives contributions corresponding to all periodic itineraries $s_{1} s_{2} s_{3} \cdots s_{n-1} s_{n}$ of total symbol length $n$ with an alphabet of $N$ symbols

$$
\operatorname{tr} \mathbf{A}^{s_{1} s_{2}} \mathbf{A}^{s_{2} s_{3}} \cdots \mathbf{A}^{s_{n-1} s_{n}} \mathbf{A}_{l_{s_{1}}=-\infty}^{+\infty} \sum_{l_{s_{2}}=-\infty}^{+\infty} \cdots \sum_{l_{s_{n}}=-\infty}^{+\infty} A_{l_{s_{1} s_{2}}^{s_{n}}}^{s_{1} s_{2}} A_{l_{s_{2}} l_{s_{3}}}^{s_{2} s_{3}} \cdots A_{l_{s_{n-1}} l_{s_{n}}}^{s_{n-1} s_{n}} A_{l_{s_{n}} s_{s_{1}}}^{s_{1} s_{1}}\left({ }_{3}\right.
$$

with $s_{i} \in\{1,2, \ldots, N\}$. Remember our notation that the trace $\operatorname{tr}(\cdots)$ refers only to the $|l\rangle$ space. By construction A describes only scatterer-toscatterer transitions, so the symbolic dynamics has to respect the no-selfreflection pruning rule: for admissible itineraries the successive symbols have to be different. This rule is implemented by the factor $1-\delta^{s s^{\prime}}$ in (33.13).

The trace $\operatorname{tr} \mathbf{A}^{n}$ is the sum of all itineraries of length $n$,

$$
\begin{equation*}
\operatorname{tr} \mathbf{A}^{n}=\sum_{\left\{s_{1} s_{2} \cdots s_{n}\right\}} \operatorname{tr} \mathbf{A}^{s_{1} s_{2}} \mathbf{A}^{s_{2} s_{3}} \cdots \mathbf{A}^{s_{n-1} s_{n}} \mathbf{A}^{s_{n} s_{1}} \tag{33.21}
\end{equation*}
$$

We will show for the $N$-disk problem that these periodic itineraries correspond in the semiclassical limit, $k a_{s_{i}} \gg 1$, to geometrical periodic orbits with the same symbolic dynamics. For periodic orbits with creeping sections the symbolic alphabet has to be extended, see sect. 33.3.1. Furthermore, depending on the geometry, there might be nontrivial pruning rules based on the so called ghost orbits, see sect. 33.4.1.

### 33.2.3 Symmetry reductions

The determinants over the multi-scattering matrices run over the multiple index $L$ of the multiple index space. This is the proper form for the symmetry reduction (in the multiple index space), e.g., if the scatterer configuration is characterized by a discrete symmetry group $G$, we have

$$
\operatorname{Det} \mathbf{M}=\prod_{\alpha}\left(\operatorname{det} \mathbf{M}_{\mathrm{D}_{\alpha}}(k)\right)^{d_{\alpha}},
$$

where the index $\alpha$ runs over all conjugate classes of the symmetry group $G$ and $\mathrm{D}_{\alpha}$ is the $\alpha$ th representation of dimension $d_{\alpha}$. The symmetry reduction on the exact quantum mechanical level is the same as for the classical evolution operators spectral determinant factorization (22.17) of sect. 22.4.2.

### 33.3 Semiclassical reduction for 1-disk scattering

We start by focusing on the single-scatterer problem. In order to be concrete, we will consider the semiclassical reduction of the scattering of a single disk in plane.

Instead of calculating the semiclassical approximation to the determinant of the one-disk system scattering matrix (33.9), we do so for

$$
\begin{equation*}
\mathbf{d}(k) \equiv \frac{1}{2 \pi i} \frac{d}{d k} \ln \operatorname{det} \mathbf{S}^{1}(k a), \tag{33.22}
\end{equation*}
$$

the so called time delay. Recall that the corresponding $\mathbf{T}^{1}$-matrix is trace-class. Thus, the following operations are justified:

$$
\begin{align*}
\mathbf{d}(k) & =\frac{1}{2 \pi i} \frac{d}{d k} \operatorname{tr}\left(\ln \operatorname{det} \mathbf{S}^{1}(k a)\right)=\frac{1}{2 \pi i} \operatorname{tr}\left(\frac{H_{m}^{(1)}(k a)}{H_{m}^{(2)}(k a)} \frac{d}{d k} \frac{H_{m}^{(2)}(k a)}{H_{m}^{(1)}(k a)}\right) \\
& =\frac{a}{2 \pi i} \operatorname{tr}\left(\frac{H_{m}^{(2)}{ }^{\prime}(k a)}{H_{m}^{(2)}(k a)}-\frac{H_{m}^{(1)^{\prime}}(k a)}{H_{m}^{(1)}(k a)}\right) . \tag{33.23}
\end{align*}
$$

Here the prime denotes the derivative with respect to the argument of the Hankel functions. Let us introduce the abbreviation

$$
\begin{equation*}
\chi_{\nu}=\frac{H_{\nu}^{(2)^{\prime}}(k a)}{H_{\nu}^{(2)}(k a)}-\frac{H_{\nu}^{(1)^{\prime}}(k a)}{H_{\nu}^{(1)}(k a)} . \tag{33.24}
\end{equation*}
$$

We apply the Watson contour method to (33.23)

$$
\begin{equation*}
\mathbf{d}(k)=\frac{a_{j}}{2 \pi i} \sum_{m=-\infty}^{+\infty} \chi_{m}=\frac{a_{j}}{2 \pi i} \frac{1}{2 i} \oint_{C} \mathrm{~d} \nu \frac{e^{-i \nu \pi}}{\sin (\nu \pi)} \chi_{\nu} . \tag{33.25}
\end{equation*}
$$

Here the contour $C$ encircles in a counter-clockwise manner a small semiinfinite strip $D$ which completely covers the real $\nu$-axis but which only has a small finite extend into the positive and negative imaginary $\nu$ direction. The contour $C$ is then split up in the path above and below the real $\nu$-axis such that

$$
\mathbf{d}(k)=\frac{a}{2 \pi i}\left\{-\frac{1}{2 i} \int_{-\infty+i \epsilon}^{+\infty+i \epsilon} \mathrm{~d} \nu \frac{e^{-i \nu \pi}}{\sin (\nu \pi)} \chi_{\nu}+\frac{1}{2 i} \int_{-\infty-i \epsilon}^{+\infty-i \epsilon} \mathrm{~d} \nu \frac{e^{-i \nu \pi}}{\sin (\nu \pi)} \chi_{\nu}\right\} .
$$

Then, we perform the substitution $\nu \rightarrow-\nu$ in the second integral so as to get

$$
\begin{align*}
\mathbf{d}(k) & =\frac{a}{2 \pi i}\left\{-\frac{1}{2 i} \int_{-\infty+i \epsilon}^{+\infty+i \epsilon} \mathrm{~d} \nu \frac{e^{-i \nu \pi}}{\sin (\nu \pi)} \chi_{\nu}-\frac{1}{2 i} \int_{-\infty+i \epsilon}^{+\infty+i \epsilon} \mathrm{~d} \nu \frac{e^{+i \nu \pi}}{\sin (\nu \pi)} \chi_{-\nu}\right\} \\
& =\frac{a}{2 \pi i}\left\{2 \int_{-\infty+i \epsilon}^{+\infty+i \epsilon} \mathrm{~d} \nu \frac{e^{2 i \nu \pi}}{1-e^{2 i \nu \pi}} \chi_{\nu}+\int_{-\infty}^{+\infty} \mathrm{d} \nu \chi_{\nu}\right\}, \tag{33.26}
\end{align*}
$$

where we used the fact that $\chi_{-\nu}=\chi_{\nu}$. The contour in the last integral could be deformed to pass over the real $\nu$-axis since its integrand has no Watson denominator any longer.

We will now approximate the last expression semiclassically, that is, under the assumption $k a \gg 1$. As the two contributions in the last line of (33.26) differ by the presence or absence of the Watson denominator, they will have to be handled semiclassically in different ways: the first will
be closed in the upper complex plane and evaluated at the poles of $\chi_{\nu}$, the second integral will be evaluated on the real $\nu$-axis under the Debye approximation for Hankel functions.

We will now work out the first term. The poles of $\chi_{\nu}$ in the upper complex plane are given by the zeros of $H_{\nu}^{(1)}(k a)$ which will be denoted by $\nu_{\ell}(k a)$ and by the zeros of $H_{\nu}^{(2)}(k a)$ which we will denote by $-\bar{\nu}_{\ell}(k a)$, $\ell=1,2,3, \cdots$. In the Airy approximation to the Hankel functions they are given by

$$
\begin{align*}
\nu_{\ell}(k a) & =k a+i \alpha_{\ell}(k a),  \tag{33.27}\\
-\bar{\nu}_{\ell}(k a) & =-k a+i\left(\alpha_{\ell}\left(k^{*} a\right)\right)^{*}=-\left(\nu_{\ell}\left(k^{*} a\right)\right)^{*}, \tag{33.28}
\end{align*}
$$

with

$$
\begin{align*}
i \alpha_{\ell}(k a) & =e^{i \frac{\pi}{3}}\left(\frac{k a}{6}\right)^{1 / 3} q_{\ell}-e^{-i \frac{\pi}{3}}\left(\frac{6}{k a}\right)^{1 / 3} \frac{q_{\ell}^{2}}{180}-\frac{1}{70 k a}\left(1-\frac{q_{\ell}^{3}}{30}\right) \\
& +e^{i \frac{\pi}{3}}\left(\frac{6}{k a}\right)^{\frac{5}{3}} \frac{1}{3150}\left(\frac{29 q_{\ell}}{6^{2}}-\frac{281 q_{\ell}^{4}}{180 \cdot 6^{3}}\right)+\cdots . \tag{33.29}
\end{align*}
$$

Here $q_{\ell}$ labels the zeros of the Airy integral

$$
A(q) \equiv \int_{0}^{\infty} \mathrm{d} \tau \cos \left(q \tau-\tau^{3}\right)=3^{-1 / 3} \pi \operatorname{Ai}\left(-3^{-1 / 3} q\right)
$$

with $\operatorname{Ai}(z)$ being the standard Airy function; approximately, $q_{\ell} \approx 6^{1 / 3}[3 \pi(\ell-$ $1 / 4)]^{2 / 3} / 2$. In order to keep the notation simple, we will abbreviate $\nu_{\ell} \equiv$ $\nu_{\ell}(k a)$ and $\bar{\nu}_{\ell} \equiv \bar{\nu}_{\ell}(k a)$. Thus the first term of (33.26) becomes finally

$$
\frac{a}{2 \pi i}\left\{2 \int_{-\infty+i \epsilon}^{+\infty+i \epsilon} \mathrm{~d} \nu \frac{e^{2 i \nu \pi}}{1-e^{2 i \nu \pi}} \chi_{\nu}\right\}=2 a \sum_{\ell=1}^{\infty}\left(\frac{e^{2 i \nu_{\ell} \pi}}{1-e^{2 i \nu_{\ell} \pi}}+\frac{e^{-2 i \bar{\nu}_{\ell} \pi}}{1-e^{-2 i \bar{\nu}_{\ell} \pi}}\right) .
$$

In the second term of (33.26) we will insert the Debye approximations for the Hankel functions:

$$
\begin{array}{ll}
H_{\nu}^{(1 / 2)}(x) \sim \sqrt{\frac{2}{\pi \sqrt{x^{2}-\nu^{2}}}} \exp \left( \pm i \sqrt{x^{2}-\nu^{2}} \mp i \nu \arccos \frac{\nu}{x} \mp i \frac{\pi}{4}\right) & \text { for }|x|>\nu \\
H_{\nu}^{(1 / 2)}(x) \sim \mp i \sqrt{\frac{2}{\pi \sqrt{\nu^{2}-x^{2}}}} \exp \left(-\sqrt{\nu^{2}-x^{2}}+\nu \operatorname{ArcCosh} \frac{\nu}{x}\right) & \text { for }|x|<\nu . \tag{33.30}
\end{array}
$$

Note that for $\nu>k a$ the contributions in $\chi_{\nu}$ cancel. Thus the second integral of (33.26) becomes

$$
\begin{align*}
\frac{a}{2 \pi i} \int_{-\infty}^{+\infty} \mathrm{d} \nu \chi_{\nu} & =\frac{a}{2 \pi i} \int_{-k a}^{+k a} \mathrm{~d} \nu \frac{(-2 i)}{a} \frac{d}{d k}\left(\sqrt{k^{2} a^{2}-\nu^{2}}-\nu \arccos \frac{\nu}{k a}\right)+\cdots \\
& =-\frac{1}{k \pi} \int_{-k a}^{k a} \mathrm{~d} \nu \sqrt{k^{2} a^{2}-\nu^{2}}+\cdots=-\frac{a^{2}}{2} k+\cdots, \tag{33.31}
\end{align*}
$$

where $\cdots$ takes care of the polynomial corrections in the Debye approximation and the boundary correction terms in the $\nu$ integration.

In summary, our semiclassical approximation to $\mathbf{d}(k)$ reads

$$
\mathbf{d}(k)=2 a \sum_{\ell=1}^{\infty}\left(\frac{e^{2 i \nu_{\ell} \pi}}{1-e^{2 i \nu_{\ell} \pi}}+\frac{e^{-2 i \bar{\nu}_{\ell} \pi}}{1-e^{-2 i \bar{\nu}_{\ell} \pi}}\right)-\frac{a^{2}}{2} k+\cdots .
$$

Using the definition of the time delay (33.22), we get the following expression for $\operatorname{det} \mathbf{S}^{1}(k a)$ :

$$
\begin{align*}
\ln & \operatorname{det} \mathbf{S}^{1}(k a)-\lim _{k_{0} \rightarrow 0} \ln \operatorname{det} \mathbf{S}^{1}\left(k_{0} a\right)  \tag{33.32}\\
& =\int_{0}^{k} \mathrm{~d} \tilde{k}\left(-i 2 \pi \frac{a^{2} \tilde{k}}{2}+2(i 2 \pi a) \sum_{\ell=1}^{\infty}\left(\frac{e^{i 2 \pi \nu_{\ell}(\tilde{k} a)}}{1-e^{i 2 \pi \nu_{\ell}(\tilde{k} a)}}+\frac{e^{-i 2 \pi \bar{\nu}_{\ell}(\tilde{k} a)}}{1-e^{-i 2 \pi \bar{\nu}_{\ell}(\tilde{k} a)}}\right)\right)+\cdots \\
& \sim-2 \pi i N(k)+2 \sum_{\ell=1}^{\infty} \int_{0}^{k} \mathrm{~d} \tilde{k} \frac{d}{d \tilde{k}}\left\{-\ln \left(1-e^{i 2 \pi \nu_{\ell}(\tilde{k} a)}\right)+\ln \left(1-e^{-i 2 \pi \bar{\nu}_{\ell}(\tilde{k} a)}\right)\right\}+\cdots,
\end{align*}
$$

where in the last expression it has been used that semiclassically $\frac{d}{d k} \nu_{\ell}(k a) \sim$ $\frac{d}{d k} \bar{\nu}_{\ell}(k a) \sim a$ and that the Weyl term for a single disk of radius $a$ goes like $N(k)=\pi a^{2} k^{2} /(4 \pi)+\cdots$ (the next terms come from the boundary terms in the $\nu$-integration in (33.31)). Note that for the lower limit, $k_{0} \rightarrow 0$, we have two simplifications: First,

$$
\begin{aligned}
\lim _{k_{0} \rightarrow 0} S_{m m^{\prime}}^{1}\left(k_{0} a\right)= & \lim _{k_{0} \rightarrow 0} \frac{-H_{m}^{(2)}\left(k_{0} a\right)}{H_{m}^{(1)}\left(k_{0} a\right)} \delta_{m m^{\prime}}=1 \times \delta_{m m^{\prime}} \quad \forall m, m^{\prime} \\
& \rightarrow \quad \lim _{k_{0} \rightarrow 0} \operatorname{det} \mathbf{S}^{1}\left(k_{0} a\right)=1 .
\end{aligned}
$$

Secondly, for $k_{0} \rightarrow 0$, the two terms in the curly bracket of (33.32) cancel.

### 33.3.1 1-disk spectrum interpreted; pure creeping

Hence, we finally obtain the semiclassical result for the determinant of $\mathbf{S}^{1}(k a)$

$$
\begin{equation*}
\operatorname{det} \mathbf{S}^{1}(k a) \sim e^{-i 2 \pi N(k)} \frac{\prod_{\ell=1}^{\infty}\left(1-e^{-2 i \pi \bar{\nu}_{\ell}(k a)}\right)^{2}}{\prod_{\ell=1}^{\infty}\left(1-e^{2 i \pi \nu_{\ell}(k a)}\right)^{2}}, \tag{33.33}
\end{equation*}
$$

with

$$
\begin{aligned}
\nu_{\ell}(k a) & =k a+i \alpha_{\ell}(k a)=k a+e^{+i \pi / 3}(k a / 6)^{1 / 3} q_{\ell}+\cdots \\
\bar{\nu}_{\ell}(k a) & =k a-i\left(\alpha_{\ell}\left(k^{*} a\right)\right)^{*}=k a+e^{-i \pi / 3}(k a / 6)^{1 / 3} q_{\ell}+\cdots \\
& =\left(\nu_{\ell}\left(k^{*} a\right)\right)^{*}
\end{aligned}
$$



Figure 33.2: Left- and right-handed diffractive creeping paths of increasing mode number $\ell$ for a single disk.
and $N(k a)=\left(\pi a^{2} k^{2}\right) / 4 \pi+\cdots$ the leading term in the Weyl approximation for the staircase function of the wavenumber eigenvalues in the disk interior. From the point of view of the scattering particle, the interior domains of the disks are excluded relatively to the free evolution without scattering obstacles. Therefore the negative sign in front of the Weyl term. For the same reason, the subleading boundary term has here a Neumann structure, although the disks have Dirichlet boundary conditions.

Lets us abbreviate the r.h.s. of (33.33) for a specified disk $s$ as

$$
\begin{equation*}
\operatorname{det} \mathbf{S}^{s}\left(k a_{s}\right) \sim\left(e^{-i \pi N\left(k a_{s}\right)}\right)^{2} \frac{\widetilde{Z}_{1-\operatorname{disk}(\mathrm{l})}\left(k^{*} a_{s}\right)^{*}}{\widetilde{Z}_{1-\operatorname{disk}(\mathrm{l})}\left(k a_{s}\right)} \frac{\widetilde{Z}_{1-\operatorname{disk}(\mathrm{r})}\left(k^{*} a_{s}\right)^{*}}{\widetilde{Z}_{1-\operatorname{disk}(\mathrm{r})}\left(k a_{s}\right)} \tag{33.34}
\end{equation*}
$$

where $\widetilde{Z}_{1-\operatorname{disk}(\mathrm{l})}\left(k a_{s}\right)$ and $\widetilde{Z}_{1-\operatorname{disk}(\mathrm{r})}\left(k a_{s}\right)$ are the diffractional zeta functions (here and in the following we will label semiclassical zeta functions with diffractive corrections by a tilde) for creeping orbits around the $s$ th disk in the left-handed sense and the right-handed sense, respectively (see figure 33.2). The two orientations of the creeping orbits are the reason for the exponents two in (33.33). Equation (33.33) describes the semiclassical approximation to the incoherent part (= the curly bracket on the r.h.s.) of the exact expression (33.19) for the case that the scatterers are disks.

In the following we will discuss the semiclassical resonances in the 1disk scattering problem with Dirichlet boundary conditions, i.e. the socalled shape resonances. The quantummechanical resonances are the poles of the $S$-matrix in the complex $k$-plane. As the 1 -disk scattering problem is separable, the $S$-matrix is already diagonalized in the angular momentum eigenbasis and takes the simple form (33.9). The exact quantummechanical poles of the scattering matrix are therefore given by the the zeros, $k_{n m}^{\text {res }}$, of the Hankel functions $H_{m}^{(1)}(k a)$ in the lower complex $k$ plane which can be labeled by two indices, $m$ and $n$, where $m$ denotes the angular quantum number of the Hankel function and $n$ is a radial quantum number. As the Hankel functions have to vanish at specific $k$ values, one cannot use the usual Debye approximation as semiclassical approximation for the Hankel function, since this approximation only works in case the Hankel function is dominated by only one saddle. However, for the vanishing of the Hankel function, one has to have the interplay of two saddles, thus an Airy approximation is needed as in the case of the creeping poles discussed above.

The Airy approximation of the Hankel function $H_{\nu}^{(1)}(k a)$ of complex-valued index $\nu$ reads

$$
H_{\nu}^{(1)}(k a) \sim \frac{2}{\pi} e^{-i \frac{\pi}{3}}\left(\frac{6}{k a}\right)^{1 / 3} A\left(q^{(1)}\right),
$$

with

$$
q^{(1)}=e^{-i \frac{\pi}{3}}\left(\frac{6}{k a}\right)^{1 / 3}(\nu-k a)+\mathcal{O}\left((k a)^{-1}\right)
$$

Hence the zeros $\nu_{\ell}$ of the Hankel function in the complex $\nu$ plane follow from the zeros $q_{\ell}$ of the Airy integral $A(q)$ (see (33.3). Thus if we set $\nu_{\ell}=m$ (with $m$ integer), we have the following semiclassical condition on $k^{\text {res }}$

$$
\begin{align*}
m & \sim k^{\mathrm{res}} a+i \alpha_{\ell}\left(k^{\mathrm{res}} a\right) \\
& =e^{i \frac{\pi}{3}}\left(\frac{k^{\mathrm{res}} a}{6}\right)^{1 / 3} q_{\ell}-e^{-i \frac{\pi}{3}}\left(\frac{6}{k^{\mathrm{res}} a}\right)^{1 / 3} \frac{q_{\ell}^{2}}{180}-\frac{1}{70 k^{\mathrm{res}} a}\left(1-\frac{q_{\ell}^{3}}{30}\right) \\
& +e^{i \frac{\pi}{3}}\left(\frac{6}{k^{\mathrm{res}} a}\right)^{\frac{5}{3}} \frac{1}{3150}\left(\frac{29 q_{\ell}}{6^{2}}-\frac{281 q_{\ell}^{4}}{180 \cdot 6^{3}}\right)+\cdots, \tag{33.35}
\end{align*}
$$

$$
\begin{equation*}
\text { with } l=1,2,3, \cdots \text {. } \tag{33.36}
\end{equation*}
$$

For a given index $l$ this is equivalent with

$$
0 \sim 1-e^{\left(i k^{\text {res }}-\alpha_{\ell}\right) 2 \pi a},
$$

the de-Broglie condition on the wave function that encircles the disk. Thus the semiclassical resonances of the 1-disk problem are given by the zeros of the following product

$$
\prod_{l=1}^{\infty}\left(1-e^{\left(i k-\alpha_{\ell}\right) 2 \pi a}\right)
$$

which is of course nothing else than $\widetilde{Z}_{1 \text {-disk }}(k)$, the semiclassical diffraction zeta function of the 1 -disk scattering problem, see (33.34). Note that this expression includes just the pure creeping contribution and no genuine geometrical parts. Because of

$$
H_{-m}^{(1)}(k a)=(-1)^{m} H_{m}^{(1)}(k a),
$$

the zeros are doubly degenerate if $m \neq 0$, corresponding to right- and left handed creeping turns. The case $m=0$ is unphysical, since all zeros of the Hankel function $H_{0}^{(1)}(k a)$ have negative real value.


Figure 33.3: The shape resonances of the 1-disk system in the complex $k$ plane in units of the disk radius $a$. The boxes label the exact quantum mechanical resonances (given by the zeros of $H_{m}^{(1)}(k a)$ for $m=0,1,2$ ), the crosses label the diffractional semiclassical resonances (given by the zeros of the creeping formula in the Airy approximation (33.36) up to the order $\mathcal{O}\left([k a]^{1 / 3}\right)$ ).


Figure 33.4: Same as in figure 33.3. However, the subleading terms in the Airy approximation (33.36) are taken into account up to the order $\mathcal{O}\left([k a]^{-1 / 3}\right)$ (upper panel) and up to $\operatorname{order} \mathcal{O}\left([k a]^{-1}\right)$ (lower panel).

From figure 33.3 one finds that the creeping terms in the Airy order $\mathcal{O}\left([k a]^{1 / 3}\right)$, which are used in the Keller construction, systematically underestimate the magnitude of the imaginary parts of the exact data. However, the creeping data become better for increasing $\operatorname{Re} k$ and decreasing $|\operatorname{Im} k|$, as they should as semiclassical approximations.

In the upper panel of figure 33.4 one sees the change, when the next order in the Airy approximation, (33.36), is taken into account. The approximation is nearly perfect, especially for the leading row of resonances. The second Airy approximation using eq.(33.36) up to order $\mathcal{O}\left([k a]^{-1}\right)$ is perfect up to the drawing scale of figure 33.4 (lower panel).

### 33.4 From quantum cycle to semiclassical cycle

The procedure for the semiclassical approximation of a general periodic itinerary (33.20) of length $n$ is somewhat laborious, and we will only sketch the procedure here. It follows, in fact, rather closely the methods developed for the semiclassical reduction of the determinant of the 1-disk system.

Figure 33.5: A 4-disk problem with three specular reflections, one ghost tunneling, and distinct creeping segments from which all associated creeping paths can be constructed.


The quantum cycle

$$
\operatorname{tr} \mathbf{A}^{s_{1} s_{2}} \cdots \mathbf{A}^{s_{m} s_{1}}=\sum_{l_{s_{1}}=-\infty}^{\infty} \cdots \sum_{l_{s_{m}}=-\infty}^{\infty} A_{l_{s_{1}} l_{s_{2}}}^{s_{1} s_{2}} \cdots A_{l_{s_{m}} s_{s_{1}}}^{s_{m} s_{1}}
$$

still has the structure of a "multi-trace" with respect to angular momentum.
Each of the sums $\sum_{l_{s_{i}}=-\infty}^{\infty}$ - as in the 1-disk case - will be replaced by a Watson contour resummation in terms of complex angular momentum $\nu_{s_{i}}$. Then the paths below the real $\nu_{s_{i}}$-axes will be transformed to paths above these axes. Thus the the integrals split into expressions with and without an explicit Watson $\sin \left(\nu_{s_{i}} \pi\right)$ denominator.

1. In the $\sin \left(\nu_{s_{i}} \pi\right)$-independent integrals we replace all Hankel and Bessel functions by Debye approximations. Then we evaluate the expression in the saddle point approximation: either left or right specular reflection at disk $s_{i}$ or ghost tunneling through disk $s_{i}$ result.
2. For the $\sin \left(\nu_{s_{i}} \pi\right)$-dependent integrals, we close the contour in the upper $\nu_{s_{i}}$ plane and evaluate the integral at the residua $H_{\nu_{s_{i}}}^{(1)}\left(k a_{s_{i}}\right)=0$. Then we use the Airy approximation for $J_{\nu_{s_{i}}}\left(k a_{s_{i}}\right)$ and $H_{\nu_{s_{i}}}^{(1)}\left(k a_{s_{i}}\right)$ : left and right creeping paths around disk $s_{i}$ result.

In the above we have assumed that no grazing geometrical paths appear. If they do show up, the analysis has to be extended to the case of coninciding saddles between the geometrical paths with $\pi / 2$ angle reflection from the disk surface and paths with direct ghost tunneling through the disk.

There are three possibilities of "semiclassical" contact of the point particle with the disk $s_{i}$ :

1. either geometrical which in turn splits into three alternatives
(a) specular reflection to the right,
(b) specular reflection to the left,
(c) or 'ghost tunneling' where the latter induce the nontrivial pruning rules (as discussed above)
2. or right-handed creeping turns

## 3. or left-handed creeping turns,

see figure 33.5. The specular reflection to the right is linked to left-handed creeping paths with at least one knot. The specular reflection to the left matches a right-handed creeping paths with at least one knot, whereas the shortest left- and right-handed creeping paths in the ghost tunneling case are topologically trivial. In fact, the topology of the creeping paths encodes the choice between the three alternatives for the geometrical contact with the disk. This is the case for the simple reason that creeping sections have to be positive definite in length: the creeping amplitude has to decrease during the creeping process, as tangential rays are constantly emitted. In mathematical terms, it means that the creeping angle has to be positive. Thus, the positivity of the two creeping angles for the shortest left and right turn uniquely specifies the topology of the creeping sections which in turn specifies which of the three alternatives, either specular reflection to the right or to the left or straight "ghost" tunneling through disk $j$, is realized for the semiclassical geometrical path. Hence, the existence of a unique saddlepoint is guaranteed.

In order to be concrete, we will restrict ourselves in the following to the scattering from $N<\infty$ non-overlapping disks fixed in the 2-dimensional plane. The semiclassical approximation of the periodic itinerary

$$
\operatorname{tr} \mathbf{A}^{s_{1} s_{2}} \mathbf{A}^{s_{2} s_{3}} \ldots \mathbf{A}^{s_{n-1} s_{n}} \mathbf{A}^{s_{n} s_{1}}
$$

becomes a standard periodic orbit labeled by the symbol sequence $s_{1} s_{2} \cdots s_{n}$. Depending on the geometry, the individual legs $s_{i-1} \rightarrow s_{i} \rightarrow s_{i+1}$ result either from a standard specular reflection at disk $s_{i}$ or from a ghost path passing straight through disk $s_{i}$. If furthermore creeping contributions are taken into account, the symbolic dynamics has to be generalized from singleletter symbols $\left\{s_{i}\right\}$ to triple-letter symbols $\left\{s_{i}, \sigma_{i} \times \ell_{i}\right\}$ with $\ell_{i} \geq 1$ integer valued and $\sigma_{i}=0, \pm 1{ }^{1}$ By definition, the value $\sigma_{i}=0$ represents the non-creeping case, such that $\left\{s_{i}, 0 \times \ell_{i}\right\}=\left\{s_{i}, 0\right\}=\left\{s_{i}\right\}$ reduces to the old single-letter symbol. The magnitude of a nonzero $\ell_{i}$ corresponds to creeping sections of mode number $\left|\ell_{i}\right|$, whereas the sign $\sigma_{i}= \pm 1$ signals whether the creeping path turns around the disk $s_{i}$ in the positive or negative sense. Additional full creeping turns around a disk $s_{i}$ can be summed up as a geometrical series; therefore they do not lead to the introduction of a further symbol.

### 33.4.1 Ghost contributions

An itinerary with a semiclassical ghost section at, say, disk $s_{i}$ can be shown to have the same weight as the corresponding itinerary without the $s_{i}$ th symbol. Thus, semiclassically, they cancel each other in the $\operatorname{tr} \ln (\mathbf{1}-\mathbf{A})$ expansion, where they are multiplied by the permutation factor $n / r$ with

[^3]

Figure 33.6: (a) The ghost itinerary $(1, \underline{2}, 3,4)$. (b) The parent itinerary $(1,3,4)$.
the integer $r$ counting the repeats. For example, let $(1, \underline{2}, 3,4)$ be a nonrepeated periodic itinerary with a ghost section at disk 2 steming from the 4 th-order trace $\operatorname{tr} A^{4}$. By convention, an underlined disk index signals a ghost passage (as in figure 33.6a), with corresponding semiclassical ghost
 geometrical contribution to $\operatorname{tr} \ln (\mathbf{1}-\mathbf{A})$ cancels exactly against the one of its "parent" itinerary ( $1,3,4$ ) (see figure 33.6b) resulting from the 3rd-order trace:

$$
\begin{aligned}
- & \frac{1}{4}\left(4 \underline{\mathbf{A}}^{1, \underline{2}} \underline{\mathbf{A}}^{\underline{2}, 3} \mathbf{A}^{3,4} \mathbf{A}^{4,1}\right)-\frac{1}{3}\left(3 \mathbf{A}^{1,3} \mathbf{A}^{3,4} \mathbf{A}^{4,1}\right) \\
& =(+1-1) \mathbf{A}^{1,3} \mathbf{A}^{3,4} \mathbf{A}^{4,1}=0
\end{aligned}
$$

The prefactors $-1 / 3$ and $-1 / 4$ are due to the expansion of the logarithm, the factors 3 and 4 inside the brackets result from the cyclic permutation of the periodic itineraries, and the cancellation stems from the rule

$$
\begin{equation*}
\cdots \underline{\mathbf{A}^{i}}, \underline{i+1} \underline{\mathbf{A}} \underline{i+1, i+2} \cdots=\cdots\left(-\mathbf{A}^{i, i+2}\right) \cdots . \tag{33.37}
\end{equation*}
$$

The reader might study more complicated examples and convince herself that the rule (33.37).is sufficient to cancel any primary or repeated periodic orbit with one or more ghost sections completely out of the expansion of $\operatorname{tr} \ln (\mathbf{1}-\mathbf{A})$ and therefore also out of the cumulant expansion in the semiclassical limit: Any periodic orbit of length $m$ with $n(<m)$ ghost sections is cancelled by the sum of all 'parent' periodic orbits of length $m-i$ (with $1 \leq i \leq n$ and $i$ ghost sections removed) weighted by their cyclic permutation factor and by the prefactor resulting from the trace-log expansion. This is the way in which the nontrivial pruning for the $N$-disk billiards can be derived from the exact quantum mechanical expressions in the semiclassical limit. Note that there must exist at least one index $i$ in any given periodic itinerary which corresponds to a non-ghost section, since otherwise the itinerary in the semiclassical limit could only be straight and therefore nonperiodic. Furthermore, the series in the ghost cancelation has to stop at the 2 nd-order trace, $\operatorname{tr} \mathbf{A}^{2}$, as $\operatorname{tr} \mathbf{A}$ itself vanishes identically in the full domain which is considered here.

### 33.5 Heisenberg uncertainty

Where is the boundary $k a \approx 2^{m-1} \bar{L} / a$ coming from?
This boundary follows from a combination of the uncertainty principle with ray optics and the non-vanishing value for the topological entropy of the 3 -disk repeller. When the wave number $k$ is fixed, quantum mechanics can only resolve the classical repelling set up to the critical topological order $n$.The quantum wave packet which explores the repelling set has to disentangle $2^{n}$ different sections of size $d \sim a / 2^{n}$ on the "visible" part of the disk surface (which is of order $a$ ) between any two successive disk collisions. Successive collisions are separated spatially by the mean flight length $\bar{L}$, and the flux spreads with a factor $\bar{L} / a$. In other words, the uncertainty principle bounds the maximal sensible truncation in the cycle expansion order by the highest quantum resolution attainable for a given wavenumber $k$.

## Commentary

Remark 33.1 Sources. This chapter is based in its entirety on ref. [K.1]; the reader is referred to the full exposition for the proofs and discussion of details omitted here. sect. 33.3 is based on appendix E of ref. [K.1]. We follow Franz [32.19] in applying the Watson contour method [32.20] to (33.23). The Airy and Debye approximations to the Hankel functions are given in ref. [32.21], the Airy expansion of the 1 -disk zeros can be found in ref. [32.22].For details see refs. [32.19, $32.22,32.23$, K.1]. That the interior domains of the disks are excluded relatively to the free evolution without scattering obstacles was noted in refs. [32.24, 32.15].

The procedure for the semiclassical approximation of a general periodic itinerary (33.20) of length $n$ can be found in ref. [K.1] for the case of the $N$-disk systems. The reader interested in the details of the semiclassical reduction is advised to consult this reference.

The ghost orbits were introduced in refs. [32.12, 32.24].

Remark 33.2 Krein-Friedel-Lloyd formula. In the literature (see, e.g., refs. [32.14, $32.15]$ based on ref. [32.11] or ref. [32.1]) the transition from the quantum mechanics to the semiclassics of scattering problems has been performed via the semiclassical limit of the left hand sides of the Krein-Friedel-Lloyd sum for the (integrated) spectral density [K.5, K.6, 32.8, 32.9]. See also ref. [32.13] for a modern discussion of the Krein-Friedel-Lloyd formula and refs. [32.1, 32.17] for the connection of (32.17) to the the Wigner time delay.

The order of the two limits in (32.18) and (32.17) is essential, see e.g. Balian and Bloch [32.11] who stress that smoothed level densities should be inserted into the Friedel sums.

The necessity of the $+i \epsilon$ in the semiclassical calculation can be understood by purely phenomenological considerations: Without the $i \epsilon$ term there is no reason why one should be able to neglect spurious periodic orbits which solely are there
because of the introduction of the confining boundary. The subtraction of the second (empty) reference system helps just in the removal of those spurious periodic orbits which never encounter the scattering region. The ones that do would still
 explicitly been shown that the $\mathbf{T}$-matrix as well as the $\mathbf{C}^{s}, \mathbf{D}^{s}$ and $\mathbf{A}^{s s^{\prime}}$-matrices of the scattering problem from $N<\infty$ non-overlapping finite disks are all traceclass. The corresponding properties for the single-disk systems is particulary easy to prove.

## Chapter 34

## Helium atom

"But," Bohr protested, "nobody will believe me unless I can explain every atom and every molecule." Rutherford was quick to reply, "Bohr, you explain hydrogen and you explain helium and everybody will believe the rest."
John Archibald Wheeler (1986)
(G. Tanner)

So far much has been said about 1-dimensional maps, game of pinball and other curious but rather idealized dynamical systems. If you have become impatient and started wondering what good are the methods learned so far in solving real physical problems, we have good news for you. We will show in this chapter that the concepts of symbolic dynamics, unstable periodic orbits, and cycle expansions are essential tools to understand and calculate classical and quantum mechanical properties of nothing less than the helium, a dreaded three-body Coulomb problem.

This sounds almost like one step too much at a time; we all know how rich and complicated the dynamics of the three-body problem is - can we really jump from three static disks directly to three charged particles moving under the influence of their mutually attracting or repelling forces? It turns out, we can, but we have to do it with care. The full problem is indeed not accessible in all its detail, but we are able to analyze a somewhat simpler subsystem - collinear helium. This system plays an important role in the classical dynamics of the full three-body problem and its quantum spectrum.

The main work in reducing the quantum mechanics of helium to a semiclassical treatment of collinear helium lies in understanding why we are allowed to do so. We will not worry about this too much in the beginning; after all, 80 years and many failed attempts separate Heisenberg, Bohr and others in the 1920ties from the insights we have today on the role chaos plays for helium and its quantum spectrum. We have introduced collinear helium and learned how to integrate its trajectories in sect. 7.3. Here we

Figure 34.1: Coordinates for the helium three body problem in the plane.

will find periodic orbits and determine the relevant eigenvalues of the Jacobian matrix in sect. 34.1. We will explain in sect. 34.5 why a quantization of the collinear dynamics in helium will enable us to find parts of the full helium spectrum; we then set up the semiclassical spectral determinant and evaluate its cycle expansion. A full quantum justification of this treatment of helium is briefly discussed in sect. 34.5.1.

### 34.1 Classical dynamics of collinear helium

Recapitulating briefly what we learned in sect. 7.3: the collinear helium system consists of two electrons of mass $m_{e}$ and charge $-e$ moving on a line with respect to a fixed positively charged nucleus of charge $+2 e$, as in figure 34.2.

The Hamiltonian can be brought to a non-dimensionalized form

$$
\begin{equation*}
H=\frac{p_{1}^{2}}{2}+\frac{p_{2}^{2}}{2}-\frac{2}{r_{1}}-\frac{2}{r_{2}}+\frac{1}{r_{1}+r_{2}}=-1 \tag{34.1}
\end{equation*}
$$

The case of negative energies chosen here is the most interesting one for us. It exhibits chaos, unstable periodic orbits and is responsible for the bound states and resonances of the quantum problem treated in sect. 34.5.

There is another classical quantity important for a semiclassical treatment of quantum mechanics, and which will also feature prominently in the discussion in the next section; this is the classical action (28.15) which scales with energy as

$$
\begin{equation*}
S(E)=\oint \mathrm{d} \mathbf{q}(E) \cdot \mathbf{p}(E)=\frac{e^{2} m_{e}^{1 / 2}}{(-E)^{1 / 2}} S, \tag{34.2}
\end{equation*}
$$

with $S$ being the action obtained from (34.1) for $E=-1$, and coordinates $\mathbf{q}=\left(r_{1}, r_{2}\right), \mathbf{p}=\left(p_{1}, p_{2}\right)$. For the Hamiltonian (34.1), the period of a cycle and its action are related by $(28.17), T_{p}=\frac{1}{2} S_{p}$.


Figure 34.3: (a) A typical trajectory in the $r_{1}-r_{2}$ plane; the trajectory enters here along the $r_{1}$ axis and escapes to infinity along the $r_{2}$ axis; (b) Poincaré map ( $r_{2}=0$ ) for collinear helium. Strong chaos prevails for small $r_{1}$ near the nucleus.

After a Kustaanheimo-Stiefel transformation

$$
\begin{equation*}
r_{1}=Q_{1}^{2}, \quad r_{2}=Q_{2}^{2}, \quad p_{1}=\frac{P_{1}}{2 Q_{1}}, \quad p_{2}=\frac{P_{2}}{2 Q_{2}} \tag{34.3}
\end{equation*}
$$

and reparametrization of time by $d \tau=d t / r_{1} r_{2}$, the equations of motion take form (7.18)

$$
\begin{array}{ll}
\dot{P}_{1}=2 Q_{1}\left[2-\frac{P_{2}^{2}}{8}-Q_{2}^{2}\left(1+\frac{Q_{2}^{2}}{R_{12}^{4}}\right)\right] ; & \dot{Q}_{1}=\frac{1}{4} P_{1} Q_{2}^{2}  \tag{34.4}\\
\dot{P}_{2}=2 Q_{2}\left[2-\frac{P_{1}^{2}}{8}-Q_{1}^{2}\left(1+\frac{Q_{1}^{2}}{R_{12}^{4}}\right)\right] ; & \dot{Q}_{2}=\frac{1}{4} P_{2} Q_{1}^{2} .
\end{array}
$$

Individual electron-nucleus collisions at $r_{1}=Q_{1}^{2}=0$ or $r_{2}=Q_{2}^{2}=0$ no longer pose a problem to a numerical integration routine. The equations (7.18) are singular only at the triple collision $R_{12}=0$, that is, when both electrons hit the nucleus at the same time.

The new coordinates and the Hamiltonian (7.17) are very useful when calculating trajectories for collinear helium; they are, however, less intuitive as a visualization of the three-body dynamics. We will therefore refer to the old coordinates $r_{1}, r_{2}$ when discussing the dynamics and the periodic orbits.

### 34.2 Chaos, symbolic dynamics and periodic orbits

Let us have a closer look at the dynamics in collinear helium. The electrons are attracted by the nucleus. During an electron-nucleus collision momentum is transferred between the inner and outer electron. The inner electron has a maximal screening effect on the charge of the nucleus, diminishing the

Figure 34.4: The cycle 011 in the fundamental domain $r_{1} \geq r_{2}$ (full line) and in the full domain (dashed line).

attractive force on the outer electron. This electron - electron interaction is negligible if the outer electron is far from the nucleus at a collision and the overall dynamics is regular like in the 1-dimensional Kepler problem.

Things change drastically if both electrons approach the nucleus nearly simultaneously. The momentum transfer between the electrons depends now sensitively on how the particles approach the origin. Intuitively, these nearly missed triple collisions render the dynamics chaotic. A typical trajectory is plotted in figure 34.3(a) where we used $r_{1}$ and $r_{2}$ as the relevant axis. The dynamics can also be visualized in a Poincaré surface of section, see figure 34.3(b). We plot here the coordinate and momentum of the outer electron whenever the inner particle hits the nucleus, that is, $r_{1}$ or $r_{2}=$ 0 . As the unstructured gray region of the Poincaré section for small $r_{1}$ illustrates, the dynamics is chaotic whenever the outer electron is close to the origin during a collision. Conversely, regular motions dominate whenever the outer electron is far from the nucleus. As one of the electrons escapes for almost any starting condition, the system is unbounded: one electron (say electron 1) can escape, with an arbitrary amount of kinetic energy taken by the fugative. The remaining electron is trapped in a Kepler ellipse with total energy in the range $[-1,-\infty]$. There is no energy barrier which would separate the bound from the unbound regions of the phase space. From general kinematic arguments one deduces that the outer electron will not return when $p_{1}>0, r_{2} \leq 2$ at $p_{2}=0$, the turning point of the inner electron. Only if the two electrons approach the nucleus almost symmetrically along the line $r_{1}=r_{2}$, and pass close to the triple collision can the momentum transfer between the electrons be large enough to kick one of the particles out completely. In other words, the electron escape originates from the near triple collisions.

The collinear helium dynamics has some important properties which we now list.

### 34.2.1 Reflection symmetry

The Hamiltonian (7.9) is invariant with respect to electron-electron exchange; this symmetry corresponds to the mirror symmetry of the potential along the line $r_{1}=r_{2}$, figure 34.4. As a consequence, we can restrict
ourselves to the dynamics in the fundamental domain $r_{1} \geq r_{2}$ and treat a crossing of the diagonal $r_{1}=r_{2}$ as a hard wall reflection. The dynamics in the full domain can then be reconstructed by unfolding the trajectory through back-reflections. As explained in chapter 22, the dynamics in the fundamental domain is the key to the factorization of spectral determinants, to be implemented here in (34.15). Note also the similarity between the fundamental domain of the collinear potential figure 34.4, and the fundamental domain figure $11.6(\mathrm{~b})$ in the 3 -disk system, a simpler problem with the same binary symbolic dynamics.


### 34.2.2 Symbolic dynamics

We have already made the claim that the triple collisions render the collinear helium fully chaotic. We have no proof of the assertion, but the analysis of the symbolic dynamics lends further credence to the claim.

The potential in (34.1) forms a ridge along the line $r_{1}=r_{2}$. One can show that a trajectory passing the ridge must go through at least one twobody collision $r_{1}=0$ or $r_{2}=0$ before coming back to the diagonal $r_{1}=r_{2}$. This suggests a binary symbolic dynamics corresponding to the dynamics in the fundamental domain $r_{1} \geq r_{2}$; the symbolic dynamics is linked to the Poincaré map $r_{2}=0$ and the symbols 0 and 1 are defined as

0 : if the trajectory is not reflected from the line $r_{1}=r_{2}$ between two collisions with the nucleus $r_{2}=0$;

1: if a trajectory is reflected from the line $r_{1}=r_{2}$ between two collisions with the nucleus $r_{2}=0$.

Empirically, the symbolic dynamics is complete for a Poincaré map in the fundamental domain, that is, there exists a one-to-one correspondence between binary symbol sequences and collinear trajectories in the fundamental domain, with exception of the $\overline{0}$ cycle.

### 34.2.3 Periodic orbits

The existence of a binary symbolic dynamics makes it easy to count the number of periodic orbits in the fundamental domain, as in sect. 13.5.2. However, mere existence of these cycles does not suffice to calculate semiclassical spectral determinants. We need to determine their phase space trajectories and calculate their periods, topological indices and stabilities. A restriction of the periodic orbit search to a suitable Poincaré surface of section, e.g. $r_{2}=0$ or $r_{1}=r_{2}$, leaves us in general with a 2-dimensional
search. Methods to find periodic orbits in multi-dimensional spaces have been described in chapter 17. They depend sensitively on good starting guesses. A systematic search for all orbits can be achieved only after combining multi-dimensional Newton methods with interpolation algorithms based on the binary symbolic dynamics phase space partitioning. All cycles up to symbol length 16 (some 8000 primitive cycles) have been computed by such methods, with some examples shown in figure 34.5. All numerical evidence indicates that the dynamics of collinear helium is hyperbolic, and that all periodic orbits are unstable.

Note that the fixed point $\overline{0}$ cycle is not in this list. The $\overline{0}$ cycle would correspond to the situation where the outer electron sits at rest infinitely far from the nucleus while the inner electron bounces back and forth into the nucleus. The orbit is the limiting case of an electron escaping to infinity with zero kinetic energy. The orbit is in the regular (that is, separable) limit of the dynamics and is thus marginally stable. The existence of this orbit is also related to intermittent behavior generating the quasi-regular dynamics for large $r_{1}$ that we have already noted in figure 34.3(b).

Search algorithm for an arbitrary periodic orbit is quite cumbersome to program. There is, however, a class of periodic orbits, orbits with symmetries, which can be easily found by a one-parameter search. The only symmetry left for the dynamics in the fundamental domain is time reversal symmetry; a time reversal symmetric periodic orbit is an orbit whose trajectory in phase space is mapped onto itself when changing $\left(p_{1}, p_{2}\right) \rightarrow\left(-p_{1},-p_{2}\right)$, by reversing the direction of the momentum of the orbit. Such an orbit must be a "libration" or self-retracing cycle, an orbit that runs back and forth along the same path in the $\left(r_{1}, r_{2}\right)$ plane. The cycles $\overline{1}, \overline{01}$ and $\overline{001}$ in figure 34.5 are examples of self-retracing cycles. Luckily, the shortest cycles that we desire most ardently have this symmetry.

Why is this observation helpful? A self-retracing cycle must start perpendicular to the boundary of the fundamental domain, that is, on either of the axis $r_{2}=0$ or $r_{1}=r_{2}$, or on the potential boundary $-\frac{2}{r_{1}}-\frac{2}{r_{2}}+\frac{1^{\prime}}{r_{1}+r_{2}}=$ -1 . By shooting off trajectories perpendicular to the boundaries and monitoring the orbits returning to the boundary with the right symbol length we will find time reversal symmetric cycles by varying the starting point on the boundary as the only parameter. But how can we tell whether a given cycle is self-retracing or not? All the relevant information is contained in the itineraries; a cycle is self-retracing if its itinerary is invariant under time reversal symmetry (that is, read backwards) and a suitable number of cyclic permutations. All binary strings up to length 5 fulfill this condition. The symbolic dynamics contains even more information; we can tell at which boundary the total reflection occurs. One finds that an orbit starts out perpendicular

- to the diagonal $r_{1}=r_{2}$ if the itinerary is time reversal invariant and has an odd number of 1's; an example is the cycle $\overline{001}$ in figure 34.5;
- to the axis $r_{2}=0$ if the itinerary is time reversal invariant and has an even number of symbols; an example is the cycle $\overline{0011}$ in figure 34.5;


Figure 34.5: Some of the shortest cycles in collinear helium. The classical collinear electron motion is bounded by the potential barrier $-1=-2 / r_{1}-2 / r_{2}+1 /\left(r_{1}+r_{2}\right)$ and the condition $r_{i} \geq 0$. The orbits are shown in the full $r_{1}-r_{2}$ domain, the itineraries refers to the dynamics in the $r_{1} \geq r_{2}$ fundamental domain. The last figure, the 14cycle 00101100110111 , is an example of a typical cycle with no symmetry.

- to the potential boundary if the itinerary is time reversal invariant and has an odd number of symbols; an example is the cycle $\overline{011}$ in figure 34.5.

All cycles up to symbol length 5 are time reversal invariant, the first two non-time reversal symmetric cycles are cycles $\overline{001011}$ and $\overline{001101}$ in figure 34.5. Their determination would require a two-parameter search. The two cycles are mapped onto each other by time reversal symmetry, that is, they have the same trace in the $r_{1}-r_{2}$ plane, but they trace out distinct cycles in the full phase space.

We are ready to integrate trajectories for classical collinear helium with the help of the equations of motions (7.18) and to find all cycles up to length for the matrix elements of the Jacobian matrix along a trajectory in order to calculate stability indices. We will provide the main equations in the next section, with the details of the derivation relegated to the appendix C.2.

### 34.3 Local coordinates, Jacobian matrix

In this section, we will derive the equations of motion for the Jacobian matrix along a collinear helium trajectory. The Jacobian matrix is 4dimensional; the two trivial eigenvectors corresponding to the conservation of energy and displacements along a trajectory can, however, be projected out by suitable orthogonal coordinates transformations, see appendix C. We will give the transformation to local coordinates explicitly, here for the regularized coordinates (7.16), and state the resulting equations of motion for the reduced [ $2 \times 2$ ] Jacobian matrix.

The vector locally parallel to the trajectory is pointing in the direction of the phase space velocity (5.7)

$$
v_{m}=\dot{x}_{m}(t)=\omega_{m n} \frac{\partial H}{\partial x_{n}}=\left(H_{P_{1}}, H_{P_{2}},-H_{Q_{1}},-H_{Q_{2}}\right)^{T},
$$

with $H_{Q_{i}}=\frac{\partial H}{\partial Q_{i}}$, and $H_{P_{i}}=\frac{\partial H}{\partial P_{i}}, i=1,2$. The vector perpendicular to a trajectory $x(t)=\left(Q_{1}(t), Q_{2}(t), P_{1}(t), P_{2}(t)\right)$ and to the energy manifold is given by the gradient of the Hamiltonian (7.17)

$$
\gamma=\nabla H=\left(H_{Q_{1}}, H_{Q_{2}}, H_{P_{1}}, H_{P_{2}}\right)^{T} .
$$

By symmetry $v_{m} \gamma_{m}=\omega_{m n} \frac{\partial H}{\partial x_{n}} \frac{\partial H}{\partial x_{m}}=0$, so the two vectors are orthogonal.
Next, we consider the orthogonal matrix

$$
\begin{equation*}
\mathbf{O}=\left(\gamma_{1}, \gamma_{2}, \gamma / R, v\right) \tag{34.5}
\end{equation*}
$$

$$
=\left(\begin{array}{cccc}
-H_{P_{2}} / R & H_{Q_{2}} & H_{Q_{1}} / R & H_{P_{1}} \\
H_{P_{1}} / R & -H_{Q_{1}} & H_{Q_{2}} / R & H_{P_{2}} \\
-H_{Q_{2}} / R & -H_{P_{2}} & H_{P_{1}} / R & -H_{Q_{1}} \\
H_{Q_{1}} / R & H_{P_{1}} & H_{P_{2}} / R & -H_{Q_{2}}
\end{array}\right)
$$

with $R=|\nabla H|^{2}=\left(H_{Q_{1}}^{2}+H_{Q_{2}}^{2}+H_{P_{1}}^{2}+H_{P_{2}}^{2}\right)$, which provides a transformation to local phase space coordinates centered on the trajectory $x(t)$ along the two vectors $(\gamma, v)$. The vectors $\gamma_{1,2}$ are phase space vectors perpendicular to the trajectory and to the energy manifold in the 4-dimensional phase space of collinear helium. The Jacobian matrix (4.5) rotated to the local coordinate system by $\mathbf{O}$ then has the form

$$
\mathbf{m}=\left(\begin{array}{cccc}
m_{11} & m_{12} & * & 0 \\
m_{21} & m_{22} & * & 0 \\
0 & 0 & 1 & 0 \\
* & * & * & 1
\end{array}\right), \mathbf{J}=\mathbf{O}^{\mathbf{T}} \mathbf{m O}
$$

The linearized motion perpendicular to the trajectory on the energy manifold is described by the $[2 \times 2]$ matrix $\mathbf{m}$; the 'trivial' directions correspond to unit eigenvalues on the diagonal in the 3rd and 4th column and row.

The equations of motion for the reduced Jacobian matrix $\mathbf{m}$ are given by

$$
\begin{equation*}
\dot{\mathbf{m}}=\mathbf{l}(t) \mathbf{m}(t), \tag{34.6}
\end{equation*}
$$

with $\mathbf{m}(0)=\mathbf{1}$. The matrix $\mathbf{l}$ depends on the trajectory in phase space and has the form

$$
\mathbf{l}=\left(\begin{array}{cccc}
l_{11} & l_{12} & * & 0 \\
l_{21} & l_{22} & * & 0 \\
0 & 0 & 0 & 0 \\
* & * & * & 0
\end{array}\right),
$$

where the relevant matrix elements $l_{i j}$ are given by

$$
\begin{aligned}
l_{11}= & \frac{1}{R}\left[2 H_{Q_{1} Q_{2}}\left(H_{Q_{2}} H_{P_{1}}+H_{Q_{1}} H_{P_{2}}\right)\right. \\
& \left.+\left(H_{Q_{1}} H_{P_{1}}-H_{Q_{2}} H_{P_{2}}\right)\left(H_{Q_{1} Q_{1}}-H_{Q_{2} Q_{2}}-H_{P_{1} P_{1}}+H_{P_{2} P_{2}}\right)\right] \\
l_{12}= & -2 H_{Q_{1} Q_{2}}\left(H_{Q_{1}} H_{Q_{2}}-H_{P_{1}} H_{P_{2}}\right) \\
& +\left(H_{Q_{1}}^{2}+H_{P_{2}}^{2}\right)\left(H_{Q_{2} Q_{2}}+H_{P_{1} P_{1}}\right)+\left(H_{Q_{2}}^{2}+H_{P_{1}}^{2}\right)\left(H_{Q_{1} Q_{1}}+H_{P_{2} P_{2}}\right) \\
l_{21}= & \frac{1}{R^{2}}\left[2\left(H_{Q_{1} P_{2}}+H_{Q_{2} P_{1}}\right)\left(H_{Q_{2}} H_{P_{1}}+H_{Q_{1}} H_{P_{8}}\right)\right. \\
& \left.-\left(H_{P_{1}}^{2}+H_{P_{2}}^{2}\right)\left(H_{Q_{1} Q_{1}}+H_{Q_{2} Q_{2}}\right)-\left(H_{Q_{1}}^{2}+H_{Q_{2}}^{2}\right)\left(H_{P_{1} P_{1}}+H_{P_{2} P_{2}}\right)\right] \\
l_{22}= & -l_{11} .
\end{aligned}
$$

Here $H_{Q_{i} Q_{j}}, H_{P_{i} P_{j}}, i, j=1,2$ are the second partial derivatives of $H$ with respect to the coordinates $Q_{i}, P_{i}$, evaluated at the phase space coordinate of the classical trajectory.

| p | $S_{p} / 2 \pi$ | $\ln \left\|\Lambda_{p}\right\|$ | $\sigma_{p}$ | $m_{p}$ |
| ---: | ---: | :---: | :---: | :---: |
| 1 | 1.82900 | 0.6012 | 0.5393 | 2 |
| 01 | 3.61825 | 1.8622 | 1.0918 | 4 |
| 001 | 5.32615 | 3.4287 | 1.6402 | 6 |
| 011 | 5.39451 | 1.8603 | 1.6117 | 6 |
| 0001 | 6.96677 | 4.4378 | 2.1710 | 8 |
| 0011 | 7.04134 | 2.3417 | 2.1327 | 8 |
| 0111 | 7.25849 | 3.1124 | 2.1705 | 8 |
| 00001 | 8.56618 | 5.1100 | 2.6919 | 10 |
| 00011 | 8.64306 | 2.7207 | 2.6478 | 10 |
| 00101 | 8.93700 | 5.1562 | 2.7291 | 10 |
| 00111 | 8.94619 | 4.5932 | 2.7173 | 10 |
| 01011 | 9.02689 | 4.1765 | 2.7140 | 10 |
| 01111 | 9.07179 | 3.3424 | 2.6989 | 10 |
| 000001 | 10.13872 | 5.6047 | 3.2073 | 12 |
| 000011 | 10.21673 | 3.0323 | 3.1594 | 12 |
| 000101 | 10.57067 | 6.1393 | 3.2591 | 12 |
| 000111 | 10.57628 | 5.6766 | 3.2495 | 12 |
| 001011 | 10.70698 | 5.3251 | 3.2519 | 12 |
| 001101 | 10.70698 | 5.3251 | 3.2519 | 12 |
| 001111 | 10.74303 | 4.3317 | 3.2332 | 12 |
| 010111 | 10.87855 | 5.0002 | 3.2626 | 12 |
| 011111 | 10.91015 | 4.2408 | 3.2467 | 12 |

Table 34.1: Action $S_{p}$ (in units of $2 \pi$ ), Lyapunov exponent $\left|\Lambda_{p}\right| / T_{p}$ for the motion in the collinear plane, winding number $\sigma_{p}$ for the motion perpendicular to the collinear plane, and the topological index $m_{p}$ for all fundamental domain cycles up to topological length 6.

### 34.4 Getting ready

Now everything is in place: the regularized equations of motion can be implemented in a Runge-Kutta or any other integration scheme to calculate trajectories. We have a symbolic dynamics and know how many cycles there are and how to find them (at least up to symbol length 5). We know how to compute the Jacobian matrix whose eigenvalues enter the semiclassical spectral determinant (30.12). By (28.17) the action $S_{p}$ is proportional to the period of the orbit, $S_{p}=2 T_{p}$.

There is, however, still a slight complication. Collinear helium is an invariant 4 -dimensional subspace of the full helium phase space. If we restrict the dynamics to angular momentum equal zero, we are left with 6 phase space coordinates. That is not a problem when computing periodic orbits, they are oblivious to the other dimensions. However, the Jacobian matrix does pick up extra contributions. When we calculate the Jacobian matrix for the full problem, we must also allow for displacements out of the collinear plane, so the full Jacobian matrix for dynamics for $L=0$ angular momentum is 6 dimensional. Fortunately, the linearized dynamics in and off the collinear helium subspace decouple, and the Jacobian matrix can be written in terms of two distinct [ $2 \times 2$ ] matrices, with trivial eigendirections providing the remaining two dimensions. The submatrix related to displacements off the linear configuration characterizes the linearized dynamics in the additional degree of freedom, the $\Theta$-coordinate in figure 34.1. It turns out that the linearized dynamics in the $\Theta$ coordinate is stable, corresponding to a bending type motion of the two electrons. We will need the stability exponents for all degrees of freedom in evaluating the semiclassical spectral determinant in sect. 34.5.

The numerical values of the actions, stability exponents, stability angles, and topological indices for the shortest cycles are listed in table 34.3. These numbers, needed for the semiclassical quantization implemented in the next section, an also be helpful in checking your own calculations.

### 34.5 Semiclassical quantization of collinear helium

Before we get down to a serious calculation of the helium quantum energy levels let us have a brief look at the overall structure of the spectrum. This will give us a preliminary feel for which parts of the helium spectrum are accessible with the help of our collinear model - and which are not. In order to keep the discussion as simple as possible and to concentrate on the semiclassical aspects of our calculations we offer here only a rough overview. For a guide to more detailed accounts see remark 34.4.

### 34.5.1 Structure of helium spectrum

We start by recalling Bohr's formula for the spectrum of hydrogen like one-electron atoms. The eigenenergies form a Rydberg series

$$
\begin{equation*}
E_{N}=-\frac{e^{4} m_{e}}{\hbar^{2}} \frac{Z^{2}}{2 N^{2}}, \tag{34.8}
\end{equation*}
$$

where $Z e$ is the charge of the nucleus and $m_{e}$ is the mass of the electron. Through the rest of this chapter we adopt the atomic units $e=m_{e}=\hbar=1$.

The simplest model for the helium spectrum is obtained by treating the two electrons as independent particles moving in the potential of the nucleus neglecting the electron-electron interaction. Both electrons are then bound in hydrogen like states; the inner electron will see a charge $Z=2$, screening at the same time the nucleus, the outer electron will move in a Coulomb potential with effective charge $Z-1=1$. In this way obtain a first estimate for the total energy

$$
\begin{equation*}
E_{N, n}=-\frac{2}{N^{2}}-\frac{1}{2 n^{2}} \quad \text { with } \quad n>N . \tag{34.9}
\end{equation*}
$$

This double Rydberg formula contains already most of the information we need to understand the basic structure of the spectrum. The (correct) ionizations thresholds $E_{N}=-\frac{2}{N^{2}}$ are obtained in the limit $n \rightarrow \infty$, yielding the ground and excited states of the helium ion $\mathrm{He}^{+}$. We will therefore refer to $N$ as the principal quantum number. We also see that all states $E_{N, n}$ with $N \geq 2$ lie above the first ionization threshold for $N=1$. As soon as we switch on electron-electron interaction these states are no longer bound states; they turn into resonant states which decay into a bound state of the helium ion and a free outer electron. This might not come as a big surprise if we have the classical analysis of the previous section in mind: we already found that one of the classical electrons will almost always escape after some finite time. More remarkable is the fact that the first, $N=1$ series consists of true bound states for all $n$, an effect which can only be understood by quantum arguments.

The hydrogen-like quantum energies (34.8) are highly degenerate; states with different angular momentum but the same principal quantum number $N$ share the same energy. We recall from basic quantum mechanics of hydrogen atom that the possible angular momenta for a given $N$ span $l=0,1 \ldots N-1$. How does that affect the helium case? Total angular momentum $L$ for the helium three-body problem is conserved. The collinear helium is a subspace of the classical phase space for $L=0$; we thus expect that we can only quantize helium states corresponding to the total angular momentum zero, a subspectrum of the full helium spectrum. Going back to our crude estimate (34.9) we may now attribute angular momenta to the two independent electrons, $l_{1}$ and $l_{2}$ say. In order to obtain total angular momentum $L=0$ we need $l_{1}=l_{2}=l$ and $l_{z 1}=-l_{z 2}$, that is, there


Figure 34.6: The exact quantum helium spectrum for $L=0$. The energy levels denoted by bars have been obtained from full 3-dimensional quantum calculations [34.3].
are $N$ different states corresponding to $L=0$ for fixed quantum numbers $N, n$. That means that we expect $N$ different Rydberg series converging to each ionization threshold $E_{N}=-2 / N^{2}$. This is indeed the case and the $N$ different series can be identified also in the exact helium quantum spectrum, see figure 34.6. The degeneracies between the different $N$ Rydberg series corresponding to the same principal quantum number $N$, are removed by the electron-electron interaction. We thus already have a rather good idea of the coarse structure of the spectrum.

In the next step, we may even speculate which parts of the $L=0$ spectrum can be reproduced by the semiclassical quantization of collinear helium. In the collinear helium, both classical electrons move back and forth along a common axis through the nucleus, so each has zero angular momentum. We therefore expect that collinear helium describes the Rydberg series with $l=l_{1}=l_{2}=0$. These series are the energetically lowest states for fixed ( $N, n$ ), corresponding to the Rydberg series on the outermost left side of the spectrum in figure 34.6. We will see in the next
section that this is indeed the case and that the collinear model holds down to the $N=1$ bound state series, including even the ground state of helium! We will also find a semiclassical quantum number corresponding to the angular momentum $l$ and show that the collinear model describes states for moderate angular momentum $l$ as long as $l \ll N$. .

### 34.5.2 Semiclassical spectral determinant for collinear helium

Nothing but lassitude can stop us now from calculating our first semiclassical eigenvalues. The only thing left to do is to set up the spectral determinant in terms of the periodic orbits of collinear helium and to write out the first few terms of its cycle expansion with the help of the binary symbolic dynamics. The semiclassical spectral determinant (30.12) has been written as product over all cycles of the classical systems. The energy dependence in collinear helium enters the classical dynamics only through simple scaling transformations described in sect. 7.3 .1 which makes it possible to write the semiclassical spectral determinant in the form

$$
\begin{equation*}
\operatorname{det}(\hat{H}-E)_{s c}=\exp \left(-\sum_{p} \sum_{r=1}^{\infty} \frac{1}{r} \frac{e^{i r\left(s S_{p}-m_{p} \frac{\pi}{2}\right)}}{\left(-\operatorname{det}\left(1-\mathbf{J}_{p \perp}^{r}\right)\right)^{1 / 2}\left|\operatorname{det}\left(1-\mathbf{J}_{p \|}^{r}\right)\right|^{1 / 2}}\right) \tag{34.10}
\end{equation*}
$$

with the energy dependence absorbed into the variable

$$
s=\frac{e^{2}}{\hbar} \sqrt{\frac{m_{e}}{-E}},
$$

obtained by using the scaling relation (34.2) for the action. As explained in sect. 34.3, the fact that the [ $4 \times 4$ ] Jacobian matrix decouples into two [ $2 \times 2$ ] submatrices corresponding to the dynamics in the collinear space and perpendicular to it makes it possible to write the denominator in terms of a product of two determinants. Stable and unstable degrees of freedom enter the trace formula in different ways, reflected by the absence of the modulus sign and the minus sign in front of $\operatorname{det}\left(1-\mathbf{J}_{\perp}\right)$. The topological index $m_{p}$ corresponds to the unstable dynamics in the collinear plane. Note that the factor $e^{i \pi N(E)}$ present in (30.12) is absent in (34.10). Collinear helium is an open system, that is, the eigenenergies are resonances corresponding to the complex zeros of the semiclassical spectral determinant and the mean energy staircase $\bar{N}(E)$ not defined. In order to obtain a spectral determinant as an infinite product of the form (30.18) we may proceed as in (15.9) by expanding the determinants in (34.10) in terms of the eigenvalues of the corresponding Jacobian matrices. The matrix representing displacements perpendicular to the collinear space has eigenvalues of the form $\exp ( \pm 2 \pi i \sigma)$, reflecting stable linearized dynamics. $\sigma$ is the full winding number along the orbit in the stable degree of freedom, multiplicative under multiple repetitions of this orbit.The eigenvalues corresponding to the unstable
dynamics along the collinear axis are paired as $\{\Lambda, 1 / \Lambda\}$ with $|\Lambda|>1$ and real. As in (15.9) and (30.18) we may thus write

$$
\begin{align*}
& {\left[-\operatorname{det}\left(1-\mathbf{J}_{\perp}^{r}\right)\left|\operatorname{det}\left(1-\mathbf{J}_{\|}^{r}\right)\right|\right]^{-1 / 2}}  \tag{34.11}\\
& =\left[-\left(1-\Lambda^{r}\right)\left(1-\Lambda^{-r}\right) \mid\left(1-e^{2 \pi i r \sigma}\right)\left(1-e^{-2 \pi i r \sigma}\right)\right]^{-1 / 2} \\
& =\sum_{k, \ell=0}^{\infty} \frac{1}{\left|\Lambda^{r}\right|^{1 / 2} \Lambda^{r k}} e^{-i r(\ell+1 / 2) \sigma} .
\end{align*}
$$

The $\pm$ sign corresponds to the hyperbolic/inverse hyperbolic periodic orbits with positive/negative eigenvalues $\Lambda$. Using the relation (34.12) we see that the sum over $r$ in (34.10) is the expansion of the logarithm, so the semiclassical spectral determinant can be rewritten as a product over dynamical zeta functions, as in (15.9):

$$
\begin{equation*}
\operatorname{det}(\hat{H}-E)_{s c}=\prod_{k=0}^{\infty} \prod_{m=0}^{\infty} \zeta_{k, m}^{-1}=\prod_{k=0}^{\infty} \prod_{m=0}^{\infty} \prod_{p}\left(1-t_{p}^{(k, m)}\right) \tag{34.12}
\end{equation*}
$$

where the cycle weights are given by

$$
\begin{equation*}
t_{p}^{(k, m)}=\frac{1}{|\Lambda|^{1 / 2} \Lambda^{k}} e^{i\left(s S_{p}-m_{p} \frac{\pi}{2}-4 \pi(\ell+1 / 2) \sigma_{p}\right)} \tag{34.13}
\end{equation*}
$$

and $m_{p}$ is the topological index for the motion in the collinear plane which equals twice the topological length of the cycle. The two independent directions perpendicular to the collinear axis lead to a twofold degeneracy in this degree of freedom which accounts for an additional factor 2 in front of the winding number $\sigma$. The values for the actions, winding numbers and stability indices of the shortest cycles in collinear helium are listed in table 34.3.

The integer indices $\ell$ and $k$ play very different roles in the semiclassical spectral determinant (34.12). A linearized approximation of the flow along a cycle corresponds to a harmonic approximation of the potential in the vicinity of the trajectory. Stable motion corresponds to a harmonic oscillator potential, unstable motion to an inverted harmonic oscillator. The index $\ell$ which contributes as a phase to the cycle weights in the dynamical zeta functions can therefore be interpreted as a harmonic oscillator quantum number; it corresponds to vibrational modes in the $\Theta$ coordinate and can in our simplified picture developed in sect. 34.5 .1 be related to the quantum number $l=l_{1}=l_{2}$ representing the single particle angular momenta. Every distinct $\ell$ value corresponds to a full spectrum which we obtain from the zeros of the semiclassical spectral determinant $1 / \zeta_{\ell}$ keeping $\ell$ fixed. The harmonic oscillator approximation will eventually break down with increasing off-line excitations and thus increasing $\ell$. The index $k$ corresponds to 'excitations' along the unstable direction and can be identified with local resonances of the inverted harmonic oscillator centered
on the given orbit. The cycle contributions $t_{p}^{(k, m)}$ decrease exponentially with increasing $k$. Higher $k$ terms in an expansion of the determinant give corrections which become important only for large negative imaginary $s$ values. As we are interested only in the leading zeros of (34.12), that is, the zeros closest to the real energy axis, it is sufficient to take only the $k=0$ terms into account.

Next, let us have a look at the discrete symmetries discussed in sect. 34.2. Collinear helium has a $C_{2}$ symmetry as it is invariant under reflection across the $r_{1}=r_{2}$ line corresponding to the electron-electron exchange symmetry. As explained in sects. 22.1.2 and 22.5, we may use this symmetry to factorize the semiclassical spectral determinant. The spectrum corresponding to the states symmetric or antisymmetric with respect to reflection can be obtained by writing the dynamical zeta functions in the symmetry factorized form

$$
\begin{equation*}
1 / \zeta^{(\ell)}=\prod_{a}\left(1-t_{a}\right)^{2} \prod_{\tilde{s}}\left(1-t_{\tilde{s}}^{2}\right) . \tag{34.14}
\end{equation*}
$$

Here, the first product is taken over all asymmetric prime cycles, that is, cycles that are not self-dual under the $C_{2}$ symmetry. Such cycles come in pairs, as two equivalent orbits are mapped into each other by the symmetry transformation. The second product runs over all self-dual cycles; these orbits cross the axis $r_{1}=r_{2}$ twice at a right angle. The self-dual cycles close in the fundamental domain $r_{1} \leq r_{2}$ already at half the period compared to the orbit in the full domain, and the cycle weights $t_{\tilde{s}}$ in (34.14) are the weights of fundamental domain cycles. The $C_{2}$ symmetry now leads to the factorization of (34.14) $1 / \zeta=\zeta_{+}^{-1} \zeta_{-}^{-1}$, with

$$
\begin{align*}
1 / \zeta_{+}^{(\ell)} & =\prod_{a}\left(1-t_{a}\right) \prod_{\tilde{s}}\left(1-t_{\tilde{s}}\right), \\
1 / \zeta_{-}^{(\ell)} & =\prod_{a}\left(1-t_{a}\right) \prod_{\tilde{s}}\left(1+t_{\tilde{s}}\right), \tag{34.15}
\end{align*}
$$

setting $k=0$ in what follows. The symmetric subspace resonances are given by the zeros of $1 / \zeta_{+}^{(\ell)}$, antisymmetric resonances by the zeros of $1 / \zeta_{-}^{(\ell)}$, with the two dynamical zeta functions defined as products over orbits in the fundamental domain. The symmetry properties of an orbit can be read off directly from its symbol sequence, as explained in sect. 34.2. An orbit with an odd number of 1's in the itinerary is self-dual under the $C_{2}$ symmetry and enters the spectral determinant in (34.15) with a negative or a positive sign, depending on the symmetry subspace under consideration.

### 34.5.3 Cycle expansion results

So far we have established a factorized form of the semiclassical spectral determinant and have thereby picked up two good quantum numbers; the
quantum number $m$ has been identified with an excitation of the bending vibrations, the exchange symmetry quantum number $\pm 1$ corresponds to states being symmetric or antisymmetric with respect to the electronelectron exchange. We may now start writing down the binary cycle expansion (18.5) and determine the zeros of spectral determinant. There is, however, still another problem: there is no cycle 0 in the collinear helium. The symbol sequence $\overline{0}$ corresponds to the limit of an outer electron fixed with zero kinetic energy at $r_{1}=\infty$, the inner electron bouncing back and forth into the singularity at the origin. This introduces intermittency in our system, a problem discussed in chapter 21 . We note that the behavior of cycles going far out in the channel $r_{1}$ or $r_{2} \rightarrow \infty$ is very different from those staying in the near core region. A cycle expansion using the binary alphabet reproduces states where both electrons are localized in the near core regions: these are the lowest states in each Rydberg series. The states converging to the various ionization thresholds $E_{N}=-2 / N^{2}$ correspond to eigenfunctions where the wave function of the outer electron is stretched far out into the ionization channel $r_{1}, r_{2} \rightarrow \infty$. To include those states, we have to deal with the dynamics in the limit of large $r_{1}, r_{2}$. This turns out to be equivalent to switching to a symbolic dynamics with an infinite alphabet. With this observation in mind, we may write the cycle expansion (....) for a binary alphabet without the $\overline{0}$ cycle as

$$
\begin{align*}
1 / \zeta^{\ell}(s)=1 & -t_{1}^{(\ell)}-t_{01}^{(\ell)}-\left[t_{001}^{(\ell)}+t_{011}^{(\ell)}-t_{01}^{(\ell)} t_{1}^{(\ell)}\right] \\
& -\left[t_{0001}^{(\ell)}+t_{0011}^{(\ell)}-t_{001}^{(\ell)} t_{1}^{(\ell)}+t_{0111}^{(\ell)}-t_{011}^{(\ell)} t_{1}^{(\ell)}\right]-\ldots \tag{34.16}
\end{align*}
$$

The weights $t_{p}^{(\ell)}$ are given in (34.12), with contributions of orbits and composite orbits of the same total symbol length collected within square brackets. The cycle expansion depends only on the classical actions, stability indices and winding numbers, given for orbits up to length 6 in table 34.3. To get reacquainted with the cycle expansion formula (34.16), consider a truncation of the series after the first term

$$
1 / \zeta^{(\ell)}(s) \approx 1-t_{1} .
$$

The quantization condition $1 / \zeta^{(\ell)}(s)=0$ leads to

$$
\begin{equation*}
E_{m, N}=-\frac{\left(S_{1} / 2 \pi\right)^{2}}{\left[m+\frac{1}{2}+2\left(N+\frac{1}{2}\right) \sigma_{1}\right]^{2}}, \quad m, N=0,1,2, \ldots, \tag{34.17}
\end{equation*}
$$

with $S_{1} / 2 \pi=1.8290$ for the action and $\sigma_{1}=0.5393$ for the winding number, see table 34.3, the 1 cycle in the fundamental domain. This cycle can be described as the asymmetric stretch orbit, see figure 34.5. The additional quantum number $N$ in (34.17) corresponds to the principal quantum number defined in sect. 34.5.1. The states described by the quantization

| $N$ | $n$ | $j=1$ | $j=4$ | $j=8$ | $j=12$ | $j=16$ | $-E_{\mathrm{qm}}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 1 | 3.0970 | 2.9692 | 2.9001 | 2.9390 | 2.9248 | 2.9037 |
| 2 | 2 | 0.8044 | 0.7714 | 0.7744 | 0.7730 | 0.7727 | 0.7779 |
| 2 | 3 | - | 0.5698 | 0.5906 | 0.5916 | 0.5902 | 0.5899 |
| 2 | 4 | - | - | - | 0.5383 | 0.5429 | 0.5449 |
| 3 | 3 | 0.3622 | 0.3472 | 0.3543 | 0.3535 | 0.3503 | 0.3535 |
| 3 | 4 | - | - | 0.2812 | 0.2808 | 0.2808 | 0.2811 |
| 3 | 5 | - | - | 0.2550 | 0.2561 | 0.2559 | 0.2560 |
| 3 | 6 | - | - | - | 0.2416 | 0.2433 | 0.2438 |
| 4 | 4 | 0.2050 | 0.1962 | 0.1980 | 0.2004 | 0.2012 | 0.2010 |
| 4 | 5 | - | 0.1655 | 0.1650 | 0.1654 | 0.1657 | 0.1657 |
| 4 | 6 | - | - | 0.1508 | 0.1505 | 0.1507 | 0.1508 |
| 4 | 7 | - | - | 0.1413 | 0.1426 | 0.1426 | 0.1426 |

Table 34.2: Collinear helium, real part of the symmetric subspace resonances obtained by a cycle expansion (34.16) up to cycle length $j$. The exact quantum energies [34.3] are in the last column. The states are labeled by their principal quantum numbers. A dash as an entry indicates a missing zero at that level of approximation.
condition (34.17) are those centered closest to the nucleus and correspond therefore to the lowest states in each Rydberg series (for a fixed $m$ and $N$ values), in figure 34.6. The simple formula (34.17) gives already a rather good estimate for the ground state of helium! Results obtained from (34.17) are tabulated in table 34.2, see the 3rd column under $j=1$ and the comparison with the full quantum calculations.

In order to obtain higher excited quantum states, we need to include more orbits in the cycle expansion (34.16), covering more of the phase space dynamics further away from the center. Taking longer and longer cycles into account, we indeed reveal more and more states in each $N$-series for fixed $m$. This is illustrated by the data listed in table 34.2 for symmetric
page 601 states obtained from truncations of the cycle expansion of $1 / \zeta_{+}$.

Results of the same quality are obtained for antisymmetric states by calculating the zeros of $1 / \zeta_{-}^{(\ell)}$. Repeating the calculation with $\ell=1$ or higher in (34.15) reveals states in the Rydberg series which are to the right of the energetically lowest series in figure 34.6.

## Commentary

Remark 34.1 Sources. The full 3-dimensional Hamiltonian after elimination of the center of mass coordinates, and an account of the finite nucleus mass effects is given in ref. [34.2]. The general two-body collision regularizing KustaanheimoStiefel transformation [34.5], a generalization of Levi-Civita's [34.13] Pauli matrix two-body collision regularization for motion in a plane, is due to Kustaanheimo [34.12] who realized that the correct higher-dimensional generalization of
the "square root removal" trick (7.14), by introducing a vector $Q$ with property $r=|Q|^{2}$, is the same as Dirac's trick of getting linear equation for spin $1 / 2$ fermions by means of spinors. Vector spaces equipped with a product and a known satisfy $|Q \cdot Q|=|Q|^{2}$ define normed algebras. They appear in various physical applications - as quaternions, octonions, spinors. The technique was originally developed in celestial mechanics [34.6] to obtain numerically stable solutions for planetary motions. The basic idea was in place as early as 1931, when H. Hopf [34.14] used a KS transformation in order to illustrate a Hopf's invariant. The KS transformation for the collinear helium was introduced in ref. [34.2].

Remark 34.2 Complete binary symbolic dynamics. No stable periodic orbit and no exception to the binary symbolic dynamics of the collinear helium cycles have been found in numerical investigations. A proof that all cycles are unstable, that they are uniquely labeled by the binary symbolic dynamcis, and that this dynamics is complete is, however, still missing. The conjectured Markov partition of the phase space is given by the triple collision manifold, that is, by those trajectories which start in or end at the singular point $r_{1}=r_{2}=0$. See also ref. [34.2].

Remark 34.3 Spin and particle exchange symmetry. In our presentation of collinear helium we have completely ignored all dynamical effects due to the spin of the particles involved, such as the electronic spin-orbit coupling. Electrons are fermions and that determines the symmetry properties of the quantum states. The total wave function, including the spin degrees of freedom, must be antisymmetric under the electron-electron exchange transformation. That means that a quantum state symmetric in the position variables must have an antisymmetric spin wave function, that is, the spins are antiparallel and the total spin is zero (singletstate). Antisymmetric states have symmetric spin wave function with total spin 1 (tripletstates). The threefold degeneracy of spin 1 states is lifted by the spin-orbit coupling.

Remark 34.4 Helium quantum numbers. The classification of the helium states in terms of single electron quantum numbers, sketched in sect. 34.5.1, prevailed until the 1960 's; a growing discrepancy between experimental results and theoretical predictions made it necessary to refine this picture. In particular, the different Rydberg series sharing a given $N$-quantum number correspond, roughly speaking, to a quantization of the inter electronic angle $\Theta$, see figure 34.1, and can not be described in terms of single electron quantum numbers $l_{1}, l_{2}$. The fact that something is slightly wrong with the single electron picture laid out in sect. 34.5.1 is highlighted when considering the collinear configuration where both electrons are on the same side of the nucleus. As both electrons again have angular momentum equal to zero, the corresponding quantum states should also belong to single electron quantum numbers $\left(l_{1}, l_{2}\right)=(0,0)$. However, the single electron picture breaks down completely in the limit $\Theta=0$ where electron-electron interaction becomes the dominant effect. The quantum states corresponding to this classical configuration are distinctively different from those obtained from the collinear dynamics with electrons on different sides of the nucleus. The Rydberg series related to the classical $\Theta=0$ dynamics are on the outermost rigth side in each $N$ subspectrum in figure 34.6 , and contain the energetically highest states for given $N, n$ quantum
numbers, see also remark 34.5. A detailed account of the historical development as well as a modern interpretation of the spectrum can be found in ref. [34.1].

Remark 34.5 Beyond the unstable collinear helium subspace. The semiclassical quantization of the chaotic collinear helium subspace is discussed in refs. [34.7, 34.8, 34.9]. Classical and semiclassical considerations beyond what has been discussed in sect. 34.5 follow several other directions, all outside the main of this book.

A classical study of the dynamics of collinear helium where both electrons are on the same side of the nucleus reveals that this configuration is fully stable both in the collinear plane and perpendicular to it. The corresponding quantum states can be obtained with the help of an approximate EBK-quantization which reveals helium resonances with extremely long lifetimes (quasi - bound states in the continuum). These states form the energetically highest Rydberg series for a given principal quantum number $N$, see figure 34.6. Details can be found in refs. [34.10, 34.11].

In order to obtain the Rydberg series structure of the spectrum, that is, the succession of states converging to various ionization thresholds, we need to take into account the dynamics of orbits which make large excursions along the $r_{1}$ or $r_{2}$ axis. In the chaotic collinear subspace these orbits are characterized by symbol sequences of form $\left(a 0^{n}\right)$ where $a$ stands for an arbitrary binary symbol sequence and $0^{n}$ is a succession of $n 0$ 's in a row. A summation of the form $\sum_{n=0}^{\infty} t_{a 0^{n}}$, where $t_{p}$ are the cycle weights in (34.12), and cycle expansion of indeed yield all Rydberg states up the various ionization thresholds, see ref. [34.4]. For a comprehensive overview on spectra of two-electron atoms and semiclassical treatments ref. [34.1].

## Résumé

We have covered a lot of ground starting with considerations of the classical properties of a three-body Coulomb problem, and ending with the semiclassical helium spectrum. We saw that the three-body problem restricted to the dynamics on a collinear appears to be fully chaotic; this implies that traditional semiclassical methods such as WKBquantization will not work and that we needed the full periodic orbit theory to obtain leads to the semiclassical spectrum of helium. As a piece of unexpected luck the symbolic dynamics is simple, and the semiclassical quantization of the collinear dynamics yields an important part of the helium spectrum, including the ground state, to a reasonable accuracy. A sceptic might say: "Why bother with all the semiclassical considerations? A straightforward numerical quantum calculation achieves the same goal with better precision." While this is true, the semiclassical analysis offers new insights into the structure of the spectrum. We discovered that the dynamics perpendicular to the collinear plane was stable, giving rise to an additional (approximate) quantum number $\ell$. We thus understood the origin of the different Rydberg series depicted in figure 34.6, a fact which is not at all obvious from a numerical solution of the quantum problem.

Having traversed the long road from the classical game of pinball all the
way to a credible helium spectrum computation, we could declare victory and fold down this enterprise. Nevertheless, there is still much to think about - what about such quintessentially quantum effects as diffraction, tunnelling, ...? As we shall now see, the periodic orbit theory has still much of interest to offer.

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## Exercises

Exercise 34.1 Kustaanheimo-Stiefel transformation. Check the KustaanheimoStiefel regularization for collinear helium; derive the Hamiltonian (7.17) and the collinear helium equations of motion (7.18).

Exercise 34.2 Helium in the plane. Starting with the helium Hamiltonian in the infinite nucleus mass approximation $m_{h e}=\infty$, and angular momentum $L=$ 0 , show that the three body problem can be written in terms of three independent coordinates only, the electron-nucleus distances $r_{1}$ and $r_{2}$ and the inter-electron angle $\Theta$, see figure 7.1.

Exercise 34.3 Helium trajectories. Do some trial integrations of the collinear helium equations of motion (7.18). Due to the energy conservation, only three of the phase space coordinates $\left(Q_{1}, Q_{2}, P_{1}, P_{2}\right)$ are independent. Alternatively, you can integrate in 4 dimensions and use the energy conservation as a check on the quality of your integrator.

The dynamics can be visualized as a motion in the original configuration space $\left(r_{1}, r_{2}\right), r_{i} \geq 0$ quadrant, or, better still, by an appropriately chosen 2-d Poincaré section, exercise 34.4. Most trajectories will run away, do not be surprised - the classical collinear helium is unbound. Try to guess approximately the shortest cycle of figure 34.4.

Exercise 34.4 A Poincaré section for collinear Helium. Construct a Poincaré section of figure 34.3b that reduces the helium flow to a map. Try to delineate regions which correspond to finite symbol sequences, that is initial conditions that follow the same topological itinerary in the figure 34.3a space for a finite number of bounces. Such rough partition can be used to initiate 2dimensional Newton-Raphson method searches for helium cycles, exercise 34.5.

Exercise 34.5 Collinear helium cycles. The motion in the ( $r_{1}, r_{2}$ ) plane is topologically similar to the pinball motion in a 3-disk system, except that the motion is in the Coulomb potential.

Just as in the 3-disk system the dynamics is simplified if viewed in the fundamental domain, in this case the region between $r_{1}$ axis and the $r_{1}=r_{2}$ diagonal. Modify your integration routine so the trajectory bounces off the diagonal as off a mirror. Miraculously, the symbolic dynamics for the survivors again turns out to be binary, with 0 symbol signifying a bounce off the $r_{1}$ axis, and 1 symbol for a bounce off the diagonal. Just as in the 3-disk game of pinball, we thus know what cycles need to be computed for the cycle expansion (34.16).

Guess some short cycles by requiring that topologically they correspond to sequences of bounces either returning to the same $r_{i}$ axis or reflecting off the diagonal. Now either Use special symmetries of orbits such as self-retracing to find all orbits up to length 5 by a 1-dimensional Newton search.

Exercise 34.6 Collinear helium cycle stabilities. Compute the eigenvalues for the cycles you found in exercise 34.5, as described in sect. 34.3. You may either integrate the reduced $2 \times 2$ matrix using equations (34.6) together with the generating function 1 given in local coordinates by (34.7) or integrate the full $4 \times 4$ Jacobian matrix, see sect. 20.1. Integration in 4 dimensions should give eigenvalues of the form $\left(1,1, \Lambda_{p}, 1 / \Lambda_{p}\right)$; The unit eigenvalues are due to the usual periodic orbit invariances; displacements along the orbit as well as perpendicular to the energy manifold are conserved; the latter one provides a check of the accuracy of your computation. Compare with table 34.3; you should get the actions and Lyapunov exponents right, but topological indices and stability angles we take on faith.

Exercise 34.7 Helium eigenenergies. Compute the lowest eigenenergies of singlet and triplet states of helium by substituting cycle data into the cycle expansion (34.16) for the symmetric and antisymmetric zeta functions (34.15). Probably the quickest way is to plot the magnitude of the zeta function as function of real energy and look for the minima. As the eigenenergies in general have a small imaginary part, a contour plot such as figure 18.1, can yield informed guesses. Better way would be to find the zeros by Newton method, sect. 18.1.4. How close are you to the cycle expansion and quantum results listed in table 34.2? You can find more quantum data in ref. [34.3].

## Chapter 35

## Diffraction distraction

(N. Whelan)

Diffraction effects characteristic to scattering off wedges are incorporated into the periodic orbit theory.

### 35.1 Quantum eavesdropping

As noted in chapter 34, the classical mechanics of the helium atom is undefined at the instant of a triple collision. This is a common phenomenon there is often some singularity or discontinuity in the classical mechanics of physical systems. This discontinuity can even be helpful in classifying the dynamics. The points in phase space which have a past or future at the discontinuity form manifolds which divide the phase space and provide the symbolic dynamics. The general rule is that quantum mechanics smoothes over these discontinuities in a process we interpret as diffraction. We solve the local diffraction problem quantum mechanically and then incorporate this into our global solution. By doing so, we reconfirm the central leitmotif of this treatise: think locally - act globally.

While being a well-motivated physical example, the helium atom is somewhat involved. In fact, so involved that we do not have a clue how to do it. In its place we illustrate the concept of diffractive effects with a pinball game. There are various classes of discontinuities which a billiard can have. There may be a grazing condition such that some trajectories hit a smooth surface while others are unaffected - this leads to the creeping described in chapter 32. There may be a vertex such that trajectories to one side bounce differently from those to the other side. There may be a point scatterer or a magnetic flux line such that we do not know how to continue classical mechanics through the discontinuities. In what follows, we specialize the discussion to the second example - that of vertices or wedges. To further simplify the discussion, we consider the special case of a half line which can be thought of as a wedge of angle zero.

Figure 35.1: Scattering of a plane wave off a half line.


We start by solving the problem of the scattering of a plane wave off a half line (see figure 35.1). This is the local problem whose solution we will use to construct a global solution of more complicated geometries. We define the vertex to be the origin and launch a plane wave at it from an angle $\alpha$. What is the total field? This is a problem solved by Sommerfeld in 1896 and our discussion closely follows his.

The total field consists of three parts - the incident field, the reflected field and the diffractive field. Ignoring the third of these for the moment, we see that the space is divided into three regions. In region I there is both an incident and a reflected wave. In region II there is only an incident field. In region III there is nothing so we call this the shadowed region. However, because of diffraction the field does enter this region. This accounts for why you can overhear a conversation if you are on the opposite side of a thick wall but with a door a few meters away. Traditionally such effects have been ignored in semiclassical calculations because they are relatively weak. However, they can be significant.

To solve this problem Sommerfeld worked by analogy with the full line case, so let us briefly consider that much simpler problem. There we know that the problem can be solved by images. An incident wave of amplitude $A$ is of the form

$$
\begin{equation*}
v(r, \psi)=A e^{-i k r \cos \psi} \tag{35.1}
\end{equation*}
$$

where $\psi=\phi-\alpha$ and $\phi$ is the angular coordinate. The total field is then given by the method of images as

$$
\begin{equation*}
v_{\mathrm{tot}}=v(r, \phi-\alpha)-v(r, \phi+\alpha), \tag{35.2}
\end{equation*}
$$

where the negative sign ensures that the boundary condition of zero field on the line is satisfied.

Sommerfeld then argued that $v(r, \psi)$ can also be given a complex integral representation

$$
\begin{equation*}
v(r, \psi)=A \int_{C} d \beta f(\beta, \psi) e^{-i k r \cos \beta} . \tag{35.3}
\end{equation*}
$$



Figure 35.2: The contour in the complex $\beta$ plane. The pole is at $\beta=-\psi$ (marked by $\times$ in the figure) and the integrand approaches zero in the shaded regions as the magnitude of the imaginary part of $\beta$ approaches infinity.

This is certainly correct if the function $f(\beta, \psi)$ has a pole of residue $1 / 2 \pi i$ at $\beta=-\psi$ and if the contour $C$ encloses that pole. One choice is

$$
\begin{equation*}
f(\beta, \psi)=\frac{1}{2 \pi} \frac{e^{i \beta}}{e^{i \beta}-e^{-i \psi}} . \tag{35.4}
\end{equation*}
$$

(We choose the pole to be at $\beta=-\psi$ rather than $\beta=\psi$ for reasons discussed later.) One valid choice for the contour is shown in figure 35.2. This encloses the pole and vanishes as $|\operatorname{Im} \beta| \rightarrow \infty$ (as denoted by the shading). The sections $D_{1}$ and $D_{2}$ are congruent because they are displaced by $2 \pi$. However, they are traversed in an opposite sense and cancel, so our contour consists of just the sections $C_{1}$ and $C_{2}$. The motivation for expressing the solution in this complicated manner should become clear soon.

What have we done? We extended the space under consideration by a factor of two and then constructed a solution by assuming that there is also a source in the unphysical space. We superimpose the solutions from the two sources and at the end only consider the solution in the physical space to be meaningful. Furthermore, we expressed the solution as a contour integral which reflects the $2 \pi$ periodicity of the problem. The half line scattering problem follows by analogy.

Whereas for the full line the field is periodic in $2 \pi$, for the half line it is periodic in $4 \pi$. This can be seen by the fact that the field can be expanded in a series of the form $\{\sin (\phi / 2), \sin (\phi), \sin (3 \phi / 2), \cdots\}$. As above, we extend the space by thinking of it as two sheeted. The physical sheet is as shown in figure 35.1 and the unphysical sheet is congruent to it. The sheets are glued together along the half line so that a curve in the physical space which intersects the half line is continued in the unphysical space and vice-versa.

The boundary conditions are that the total field is zero on both faces of the half line (which are physically distinct boundary conditions) and that as $r \rightarrow \infty$ the field is composed solely of plane waves and outgoing circular waves of the form $g(\phi) \exp (i k r) / \sqrt{k r}$. This last condition is a result of Huygens' principle.

We assume that the complete solution is also given by the method of images as

$$
\begin{equation*}
v_{\mathrm{tot}}=u(r, \phi-\alpha)-u(r, \phi+\alpha) \tag{35.5}
\end{equation*}
$$

where $u(r, \psi)$ is a $4 \pi$-periodic function to be determined. The second term is interpreted as an incident field from the unphysical space and the negative sign guarantees that the solution vanishes on both faces of the half line. Sommerfeld then made the ansatz that $u$ is as given in equation (35.3) with the same contour $C_{1}+C_{2}$ but with the $4 \pi$ periodicity accounted for by replacing equation (35.4) with

$$
\begin{equation*}
f(\beta, \psi)=\frac{1}{4 \pi} \frac{e^{i \beta / 2}}{e^{i \beta / 2}-e^{-i \psi / 2}} \tag{35.6}
\end{equation*}
$$

(We divide by $4 \pi$ rather than $2 \pi$ so that the residue is properly normalized.) The integral (35.3) can be thought of as a linear superposition of an infinity of plane waves each of which satisfies the Helmholtz equation $\left(\nabla^{2}+k^{2}\right) v=$ 0 , and so their combination also satisfies the Helmholtz equation. We will see that the diffracted field is an outgoing circular wave; this being a result of choosing the pole at $\beta=-\psi$ rather than $\beta=\psi$ in equation (35.4). Therefore, this ansatz is a solution of the equation and satisfies all boundary conditions and therefore constitutes a valid solution. By uniqueness this is the only solution.

In order to further understand this solution, it is useful to massage the contour. Depending on $\phi$ there may or may not be a pole between $\beta=-\pi$ and $\beta=\pi$. In region I , both functions $u(r, \phi \pm \alpha)$ have poles which correspond to the incident and reflected waves. In region II, only $u(r, \phi-\alpha)$ has a pole corresponding to the incident wave. In region III there are no poles because of the shadow. Once we have accounted for the geometrical waves (that is, the poles), we extract the diffracted waves by saddle point analysis at $\beta= \pm \pi$. We do this by deforming the contours $C$ so that they go through the saddles as shown in figure 35.2.

Contour $C_{1}$ becomes $E_{2}+F$ while contour $C_{2}$ becomes $E_{1}-F$ where the minus sign indicates that it is traversed in a negative sense. As a result, $F$ has no net contribution and the contour consists of just $E_{1}$ and $E_{2}$.

As a result of these machinations, the curves $E$ are simply the curves $D$ of figure 35.2 but with a reversed sense. Since the integrand is no longer $2 \pi$ periodic, the contributions from these curves no longer cancel. We evaluate


Figure 35.3: The contour used to evaluate the diffractive field after the contribution of possible poles has been explicitly evaluated. The curve $F$ is traversed twice in opposite directions and has no net contribution.
both stationary phase integrals to obtain

$$
\begin{equation*}
u(r, \psi) \approx-A \frac{e^{i \pi / 4}}{\sqrt{8 \pi}} \sec (\psi / 2) \frac{e^{i k r}}{\sqrt{k r}} \tag{35.7}
\end{equation*}
$$

so that the total diffracted field is

$$
\begin{equation*}
v_{\mathrm{diff}}=-A \frac{e^{i \pi / 4}}{\sqrt{8 \pi}}\left(\sec \left(\frac{\phi-\alpha}{2}\right)-\sec \left(\frac{\phi+\alpha}{2}\right)\right) \frac{e^{i k r}}{\sqrt{k r}} \tag{35.8}
\end{equation*}
$$

Note that this expression breaks down when $\phi \pm \alpha=\pi$. These angles correspond to the borders among the three regions of figure 35.1 and must be handled more carefully - we can not do a stationary phase integral in the vicinity of a pole. However, the integral representation (35.3) and (35.6) is uniformly valid.

We now turn to the simple task of translating this result into the language of semiclassical Green's functions. Instead of an incident plane wave, we assume a source at point $x^{\prime}$ and then compute the resulting field at the receiver position $x$. If $x$ is in region I , there is both a direct term, and a reflected term, if $x$ is in region II there is only a direct term and if $x$ is in region III there is neither. In any event these contributions to the semiclassical Green's function are known since the free space Green's function between two points $x_{2}$ and $x_{1}$ is

$$
\begin{equation*}
G_{\mathrm{f}}\left(x_{2}, x_{1}, k\right)=-\frac{i}{4} H_{0}^{(+)}(k d) \approx-\frac{1}{\sqrt{8 \pi k d}} \exp \{i(k d+\pi / 4)\}, \tag{35.9}
\end{equation*}
$$

where $d$ is the distance between the points. For a reflection, we need to multiply by -1 and the distance is the length of the path via the reflection
point. Most interesting for us, there is also a diffractive contribution to the Green's function. In equation (35.8), we recognize that the coefficient $A$ is simply the intensity at the origin if there were no scatterer. This is therefore replaced by the Green's function to go from the source to the vertex which we label $x_{V}$. Furthermore, we recognize that $\exp (i k r) / \sqrt{k r}$ is, within a proportionality constant, the semiclassical Green's function to go from the vertex to the receiver.

Collecting these facts, we say

$$
\begin{equation*}
G_{\mathrm{diff}}\left(x, x^{\prime}, k\right)=G_{\mathrm{f}}\left(x, x_{V}, k\right) d\left(\theta, \theta^{\prime}\right) G_{\mathrm{f}}\left(x_{V}, x^{\prime}, k\right), \tag{35.10}
\end{equation*}
$$

where, by comparison with equations (35.8) and (35.9), we have

$$
\begin{equation*}
d\left(\theta, \theta^{\prime}\right)=\sec \left(\frac{\theta-\theta^{\prime}}{2}\right)-\sec \left(\frac{\theta+\theta^{\prime}}{2}\right) . \tag{35.11}
\end{equation*}
$$

Here $\theta^{\prime}$ is the angle to the source as measured from the vertex and $\theta$ is the angle to the receiver. They were denoted as $\alpha$ and $\phi$ previously. Note that there is a symmetry between the source and receiver as we expect for a time-reversal invariant process. Also the diffraction coefficient $d$ does not depend on which face of the half line we use to measure the angles. As we will see, a very important property of $G_{\text {diff }}$ is that it is a simple multiplicative combination of other semiclassical Green's functions.

We now recover our classical perspective by realizing that we can still think of classical trajectories. In calculating the quantum Green's function, we sum over the contributions of various paths. These include the classical trajectories which connect the points and also paths which connect the points via the vertex. These have different weights as given by equations (35.9) and (35.10) but the concept of summing over classical paths is preserved.

For completeness, we remark that there is an exact integral representation for the Green's function in the presence of a wedge of arbitrary opening angle [35.15]. It can be written as

$$
\begin{equation*}
G\left(x, x^{\prime}, k\right)=g\left(r, r^{\prime}, k, \theta^{\prime}-\theta\right)-g\left(r, r^{\prime}, k, \theta^{\prime}+\theta\right) \tag{35.12}
\end{equation*}
$$

where $(r, \theta)$ and $\left(r^{\prime}, \theta^{\prime}\right)$ are the polar coordinates of the points $x$ and $x^{\prime}$ as measured from the vertex and the angles are measured from either face of the wedge. The function $g$ is given by

$$
\begin{equation*}
g\left(r, r^{\prime}, k, \psi\right)=\frac{i}{8 \pi \nu} \int_{C_{1}+C_{2}} d \beta \frac{H_{0}^{+}\left(k \sqrt{r^{2}+r^{\prime 2}-2 r r^{\prime} \cos \beta}\right)}{1-\exp \left(i \frac{\beta+\psi}{\nu}\right)} \tag{35.13}
\end{equation*}
$$

where $\nu=\gamma / \pi$ and $\gamma$ is the opening angle of the wedge. (ie $\gamma=2 \pi$ in the case of the half plane). The contour $C_{1}+C_{2}$ is the same as shown in figure 35.2.

Figure 35.4: The billiard considered here. The dynamics consists of free motion followed by specular reflections off the faces. The top vertex induces diffraction while the bottom one is a right angle and induces two specular geometric reflections.


The poles of this integral give contributions which can be identified with the geometric paths connecting $x$ and $x^{\prime}$. The saddle points at $\beta= \pm \pi$ give contributions which can be identified with the diffractive path connecting $x$ and $x^{\prime}$. The saddle point analysis allows us to identify the diffraction constant as

$$
\begin{equation*}
d\left(\theta, \theta^{\prime}\right)=-\frac{4 \sin \frac{\pi}{\nu}}{\nu} \frac{\sin \frac{\theta}{\nu} \sin \frac{\theta^{\prime}}{\nu}}{\left(\cos \frac{\pi}{\nu}-\cos \frac{\theta+\theta^{\prime}}{\nu}\right)\left(\cos \frac{\pi}{\nu}-\cos \frac{\theta-\theta^{\prime}}{\nu}\right)}, \tag{35.14}
\end{equation*}
$$

which reduces to (35.11) when $\nu=2$. Note that the diffraction coefficient vanishes identically if $\nu=1 / n$ where $n$ is any integer. This corresponds to wedge angles of $\gamma=\pi / n$ (eg. $\mathrm{n}=1$ corresponds to a full line and $\mathrm{n}=2$ corresponds to a right angle). This demonstration is limited by the fact that it came from a leading order asymptotic expansion but the result is quite general. For such wedge angles, we can use the method of images (we will require $2 n-1$ images in addition to the actual source point) to obtain the Green's function and there is no diffractive contribution to any order. Classically this corresponds to the fact that for such angles, there is no discontinuity in the dynamics. Trajectories going into the vertex can be continued out of them unambiguously. This meshes with the discussion in the introduction where we argued that diffractive effects are intimately linked with classical discontinuities.

The integral representation is also useful because it allows us to consider geometries such that the angles are near the optical boundaries or the wedge angle is close to $\pi / n$. For these geometries the saddle point analysis leading to (35.14) is invalid due to the existence of a nearby pole. In that event, we require a more sophisticated asymptotic analysis of the full integral representation.

### 35.2 An application

Although we introduced diffraction as a correction to the purely classical effects; it is instructive to consider a system which can be quantized solely in terms of periodic diffractive orbits. Consider the geometry shown in figure 35.4 The classical mechanics consists of free motion followed by specular reflections off faces. The upper vertex is a source of diffraction while the

Figure 35.5: The dashed line shows a simple periodic diffractive orbit $\gamma$. Between the vertex $V$ and a point $P$ close to the orbit there are two geometric legs labeled $\pm$. The origin of the coordinate system is chosen to be at $R$.

lower one is a right angle and induces no diffraction. This is an open system, there are no bound states - only scattering resonances. However, we can still test the effectiveness of the theory in predicting them. Formally, scattering resonances are the poles of the scattering $S$ matrix and by an identity of Balian and Bloch are also poles of the quantum Green's function. We demonstrate this fact in chapter 32 for 2-dimensional scatterers. The poles have complex wavenumber $k$, as for the 3 -disk problem.

Let us first consider how diffractive orbits arise in evaluating the trace of $G$ which we call $g(k)$. Specifying the trace means that we must consider all paths which close on themselves in the configuration space while stationary phase arguments for large wavenumber $k$ extract those which are periodic just as for classical trajectories. In general, $g(k)$ is given by the sum over all diffractive and geometric orbits. The contribution of the simple diffractive orbit labeled $\gamma$ shown in figure 35.5 to $g(k)$ is determined as follows.

We consider a point $P$ just a little off the path and determine the semiclassical Green's function to return to $P$ via the vertex using (35.9) and (35.10). To leading order in $y$ the lengths of the two geometric paths connecting $P$ and $V$ are $d_{ \pm}=(L \pm x)+y^{2} /(L \pm x)^{2} / 2$ so that the phase factor $i k\left(d_{+}+d_{-}\right)$equals $2 i k L+i k y^{2} /\left(L^{2}-x^{2}\right)$. The trace integral involves integrating over all points $P$ and is

$$
\begin{equation*}
g_{\gamma}(k) \approx-2 d_{\gamma} \frac{e^{i(2 k L+\pi / 2)}}{8 \pi k} \int_{0}^{L} \frac{d x}{\sqrt{L^{2}-x^{2}}} \int_{-\infty}^{\infty} d y e^{\left(i k y^{2} \frac{L}{L^{2}-x^{2}}\right)} . \tag{35.15}
\end{equation*}
$$

We introduced an overall negative sign to account for the reflection at the hard wall and multiplied by 2 to account for the two traversal senses, $V R P V$ and $V P R V$. In the spirit of stationary phase integrals, we have neglected the $y$ dependence everywhere except in the exponential. The diffraction constant $d_{\gamma}$ is the one corresponding to the diffractive periodic orbit. To evaluate the $y$ integral, we use the identity

$$
\begin{equation*}
\int_{-\infty}^{\infty} d \xi e^{i a \xi^{2}}=e^{i \pi / 4} \sqrt{\frac{\pi}{a}}, \tag{35.16}
\end{equation*}
$$

and thus obtain a factor which precisely cancels the $x$ dependence in the $x$ integral. This leads to the rather simple result

$$
\begin{equation*}
g_{\gamma} \approx-\frac{i l_{\gamma}}{2 k}\left\{\frac{d_{\gamma}}{\sqrt{8 \pi k l_{\gamma}}}\right\} e^{i\left(k l_{\gamma}+\pi / 4\right)} \tag{35.17}
\end{equation*}
$$

where $l_{\gamma}=2 L$ is the length of the periodic diffractive orbit. A more sophisticated analysis of the trace integral has been done [35.6] using the integral representation (35.13). It is valid in the vicinity of an optical boundary and also for wedges with opening angles close to $\pi / n$.

Consider a periodic diffractive orbit with $n_{\gamma}$ reflections off straight hard walls and $\mu_{\gamma}$ diffractions each with a diffraction constant $d_{\gamma, j}$. The total length of the orbit $L_{\gamma}=\sum l_{\gamma, j}$ is the sum of the various diffractive legs and $l_{\gamma}$ is the length of the corresponding prime orbit. For such an orbit, (35.17) generalizes to

$$
\begin{equation*}
g_{\gamma}(k)=-\frac{i l_{\gamma}}{2 k}\left\{\prod_{j=1}^{\mu_{\gamma}} \frac{d_{\gamma, j}}{\sqrt{8 \pi k l_{\gamma, j}}}\right\} \exp \left\{i\left(k L_{\gamma}+n_{\gamma} \pi-3 \mu_{\gamma} \pi / 4\right)\right\} . \tag{35.18}
\end{equation*}
$$

Each diffraction introduces a factor of $1 / \sqrt{k}$ and multi-diffractive orbits are thereby suppressed.

If the orbit $\gamma$ is prime then $L_{\gamma}=l_{\gamma}$. If $\gamma$ is the $r^{\prime}$ th repeat of a prime orbit $\beta$ we have $L_{\gamma}=r l_{\beta}, n_{\gamma}=r p_{\beta}$ and $\mu_{\gamma}=r \sigma_{\beta}$, where $l_{\beta}, p_{\beta}$ and $\sigma_{\beta}$ all refer to the prime orbit. We can then write

$$
\begin{equation*}
g_{\gamma}=g_{\beta, r}=-\frac{i l_{\beta}}{2 k} t_{\beta}^{r} \tag{35.19}
\end{equation*}
$$

where

$$
\begin{equation*}
t_{\beta}=\left\{\prod_{j=1}^{\sigma_{\beta}} \frac{d_{\beta, j}}{\sqrt{8 \pi k l_{\beta, j}}}\right\} \exp \left\{i\left(k l_{\beta}+p_{\beta} \pi-3 \sigma_{\beta} \pi / 4\right)\right\} \tag{35.20}
\end{equation*}
$$

It then makes sense to organize the sum over diffractive orbits as a sum over the prime diffractive orbits and a sum over the repetitions

$$
\begin{equation*}
g_{\mathrm{diff}}(k)=\sum_{\beta} \sum_{r=1}^{\infty} g_{\beta, r}=-\frac{i}{2 k} \sum_{\beta} l_{\beta} \frac{t_{\beta}}{1-t_{\beta}} . \tag{35.21}
\end{equation*}
$$

We cast this as a logarithmic derivative (15.7) by noting that $\frac{d t_{\beta}}{d k}=$ $i l_{\beta} t_{\beta}-\sigma_{\beta} t_{\beta} / 2 k$ and recognizing that the first term dominates in the semiclassical limit. It follows that

$$
\begin{equation*}
g_{\mathrm{diff}}(k) \approx \frac{1}{2 k} \frac{d}{d k}\left\{\ln \prod_{\beta}\left(1-t_{\beta}\right)\right\} . \tag{35.22}
\end{equation*}
$$

In the case that there are only diffractive periodic orbits - as in the geometry of figure 35.4 - the poles of $g(k)$ are the zeros of a dynamical zeta function

$$
\begin{equation*}
1 / \zeta(k)=\prod_{\beta}\left(1-t_{\beta}\right) . \tag{35.23}
\end{equation*}
$$

Figure 35.6: The two-node Markov graph with all the diffractive processes connecting the nodes.


For geometric orbits, this function would be evaluated with a cycle expansion as discussed in chapter 18. However, here we can use the multiplicative nature of the weights $t_{\beta}$ to find a closed form representation of the function using a Markov graph, as in sect. 11.5.1. This multiplicative property of the weights follows from the fact that the diffractive Green's function (35.10) is multiplicative in segment semiclassical Green's functions, unlike the geometric case.

There is a reflection symmetry in the problem which means that all resonances can be classified as even or odd. Because of this, the dynamical zeta function factorizes as $1 / \zeta=1 / \zeta_{+} \zeta_{-}$(as explained in sects. 22.5 and 22.1.2) and we determine $1 / \zeta_{+}$and $1 / \zeta_{-}$separately using the ideas of symmetry decomposition of chapter 22 .

In the Markov graph shown in figure 35.6, we enumerate all processes. We start by identifying the fundamental domain as just the right half of figure 35.4. There are two nodes which we call $A$ and $B$. To get to another node from $B$, we can diffract (always via the vertex) in one of three directions. We can diffract back to $B$ which we denote as process 1 . We can diffract to $B^{\prime}$ s image point $B^{\prime}$ and then follow this by a reflection. This process we denote as $\overline{2}$ where the bar indicates that it involves a reflection. Thirdly, we can diffract to node $A$. Starting at $A$ we can also diffract to a node in three ways. We can diffract to $B$ which we denote as 4 . We can diffract to $B^{\prime}$ followed by a reflection which we denote as $\overline{4}$. Finally, we can diffract back to $A$ which we denote as process 5 . Each of these processes has its own weight which we can determine from the earlier discussion. First though, we construct the dynamical zeta functions.

The dynamical zeta functions are determined by enumerating all closed loops which do not intersect themselves in figure 35.6. We do it first for $1 / \zeta_{+}$because that is simpler. In that case, the processes with bars are treated on an equal footing as the others. Appealing back to sect. 22.5 we find

$$
\begin{align*}
1 / \zeta_{+} & =1-t_{1}-t_{\overline{2}}-t_{5}-t_{3} t_{4}-t_{3} t_{\overline{4}}+t_{5} t_{1}+t_{5} t_{\overline{2}}, \\
& =1-\left(t_{1}+t_{\overline{2}}+t_{5}\right)-2 t_{3} t_{4}+t_{5}\left(t_{1}+t_{\overline{2}}\right) \tag{35.24}
\end{align*}
$$

where we have used the fact that $t_{4}=t_{\overline{4}}$ by symmetry. The last term has a positive sign because it involves the product of shorter closed loops. To calculate $1 / \zeta_{-}$, we note that the processes with bars have a relative negative sign due to the group theoretic weight. Furthermore, process 5 is a boundary orbit (see sect. 22.3.1) and only affects the even resonances the terms involving $t_{5}$ are absent from $1 / \zeta_{-}$. The result is

$$
1 / \zeta_{-}=1-t_{1}+t_{\overline{2}}-t_{3} t_{4}+t_{3} t_{\overline{4}}
$$

$$
\begin{equation*}
=1-\left(t_{1}-t_{\overline{2}}\right) \tag{35.25}
\end{equation*}
$$

Note that these expressions have a finite number of terms and are not in the form of a curvature expansion, as for the 3-disk problem.

It now just remains to fix the weights. We use equation (35.20) but note that each weight involves just one diffraction constant. It is then convenient to define the quantities

$$
\begin{equation*}
u_{A}^{2}=\frac{\exp \{i(2 k L+2 \pi)\}}{\sqrt{16 \pi k L}} \quad u_{B}^{2}=\frac{\exp \{i(2 k H+\pi)\}}{\sqrt{16 \pi k H}} \tag{35.26}
\end{equation*}
$$

The lengths $L$ and $H=L / \sqrt{2}$ are defined in figure 35.4; we set $L=1$ throughout. Bouncing inside the right angle at $A$ corresponds to two specular reflections so that $p=2$. We therefore explicitly include the factor $\exp (i 2 \pi)$ in (35.26) although it is trivially equal to one. Similarly, there is one specular reflection at point $B$ giving $p=1$ and therefore a factor of $\exp (i \pi)$. We have defined $u_{A}$ and $u_{B}$ because, together with some diffraction constants, they can be used to construct all of the weights. Altogether we define four diffraction coefficients: $d_{A B}$ is the constant corresponding to diffracting from $B$ to $A$ and is found from (35.11) with $\theta^{\prime}=3 \pi / 4$ and $\theta=\pi$ and equals $2 \sec (\pi / 8) \approx 2.165$. With analogous notation, we have $d_{A A}$ and $d_{B B}=d_{B^{\prime} B}$ which equal 2 and $1+\sqrt{2}$ respectively. $d_{i j}=d_{j i}$ due to the Green's function symmetry between source and receiver referred to earlier. Finally, there is the diffractive phase factor $s=\exp (-i 3 \pi / 4)$ each time there is a diffraction. The weights are then as follows:

$$
\begin{align*}
t_{1}=s d_{B B} u_{B}^{2} \quad t_{\overline{2}} & =s d_{B^{\prime} B} u_{B}^{2} \quad t_{3}=t_{4}=t_{\overline{4}}=s d_{A B} u_{A} u_{B} \\
t_{5} & =s d_{A A} u_{A}^{2} . \tag{35.27}
\end{align*}
$$

Each weight involves two $u$ 's and one $d$. The $u$ 's represent the contribution to the weight from the paths connecting the nodes to the vertex and the $d$ gives the diffraction constant connecting the two paths.

The equality of $d_{B B}$ and $d_{B^{\prime} B}$ implies that $t_{1}=t_{\overline{2}}$. From (35.25) this means that there are no odd resonances because 1 can never equal 0 . For the even resonances equation (35.24) is an implicit equation for $k$ which has zeros shown in figure 35.7.

For comparison we also show the result from an exact quantum calculation. The agreement is very good right down to the ground state as is so often the case with semiclassical calculations. In addition we can use our dynamical zeta function to find arbitrarily high resonances and the results actually improve in that limit. In the same limit, the exact numerical solution becomes more difficult to find so the dynamical zeta function approximation is particularly useful in that case.

In general a system will consist of both geometric and diffractive orbits. In that case, the full dynamical zeta function is the product of the

Figure 35.7: The even resonances of the wedge scatterer of figure 35.4 plotted in the complex $k$-plane, with $L=1$. The exact resonances are represented as circles and their semiclassical approximations as crosses.

geometric zeta function and the diffractive one. The diffractive weights are typically smaller by order $O(1 / \sqrt{k})$ but for small $k$ they can be numerically competitive so that there is a significant diffractive effect on the low-lying spectrum. It might be expected that higher in the spectrum, the effect of diffraction is weaker due to the decreasing weights. However, it should be pointed out that an analysis of the situation for creeping diffraction [35.7] concluded that the diffraction is actually more important higher in the spectrum due to the fact that an ever greater fraction of the orbits need to be corrected for diffractive effects. The equivalent analysis has not been done for edge diffraction but a similar conclusion can probably be expected.

To conclude this chapter, we return to the opening paragraph and discuss the possibility of doing such an analysis for helium. The important point which allowed us to successfully analyze the geometry of figure 35.4 is that when a trajectory is near the vertex, we can extract its diffraction constant without reference to the other facets of the problem. We say, therefore, that this is a "local" analysis for the purposes of which we have "turned off" the other aspects of the problem, namely sides $A B$ and $A B^{\prime}$. By analogy, for helium, we would look for some simpler description of the problem which applies near the three body collision. However, there is nothing to "turn off". The local problem is just as difficult as the global one since they are precisely the same problem, just related by scaling. Therefore, it is not at all clear that such an analysis is possible for helium.

## Commentary

Remark 35.1 Classical discontinuities. Various classes of discontinuities for billiard and potential problems discussed in the literature:

- a grazing condition such that some trajectories hit a smooth surface while others are unaffected, refs. [35.1, 35.2, 35.3, 35.7]
- a vertex such that trajectories to one side bounce differently from those to the other side, refs. [35.2, 35.4, 35.5, 35.8, 35.9].
- a point scatterer [35.10, 35.11] or a magnetic flux line [35.12, 35.13] such that we do not know how to continue classical mechanics through the discontinuities.

Remark 35.2 Geometrical theory of diffraction. In the above discussion we borrowed heavily from the ideas of Keller who was interested in extending the geometrical ray picture of optics to cases where there is a discontinuity. He maintained
that we could hang onto that ray-tracing picture by allowing rays to strike the vertex and then leave at any angle with amplitude (35.8). Both he and Sommerfeld were thinking of optics and not quantum mechanics and they did not phrase the results in terms of semiclassical Green's functions but the essential idea is the same.

Remark 35.3 Generalizations Consider the effect of replacing our half line by a wedge of angle $\gamma_{1}$ and the right angle by an arbitrary angle $\gamma_{2}$. If $\gamma_{2}>\gamma_{1}$ and $\gamma_{2} \geq \pi / 2$ this is an open problem whose solution is given by equations (35.24) and (35.25) (there will then be odd resonances) but with modified weights reflecting the changed geometry [35.8]. (For $\gamma_{2}<\pi / 2$, more diffractive periodic orbits appear and the dynamical zeta functions are more complicated but can be calculated with the same machinery.) When $\gamma_{2}=\gamma_{1}$, the problem in fact has bound states [35.21, 35.22]. This last case has been of interest in studying electron transport in mesoscopic devices and in microwave waveguides. However we can not use our formalism as it stands because the diffractive periodic orbits for this geometry lie right on the border between illuminated and shadowed regions so that equation (35.7) is invalid. Even the more uniform derivation of [35.6] fails for that particular geometry, the problem being that the diffractive orbit actually lives on the edge of a family of geometric orbits and this makes the analysis still more difficult.

Remark 35.4 Diffractive Green's functions. The result (35.17) is proportional to the length of the orbit times the semiclassical Green's function (35.9) to go from the vertex back to itself along the classical path. The multi-diffractive formula (35.18) is proportional to the total length of the orbit times the product of the semiclassical Green's functions to go from one vertex to the next along classical paths. This result generalizes to any system - either a pinball or a potential which contains point singularities such that we can define a diffraction constant as above. The contribution to the trace of the semiclassical Green's function coming from a diffractive orbit which hits the singularities is proportional to the total length (or period) of the orbit times the product of semiclassical Green's functions in going from one singularity to the next. This result first appeared in reference [35.2] and a derivation can be found in reference [35.9]. A similar structure also exists for creeping [35.2].

Remark 35.5 Diffractive orbits for hydrogenic atoms. An analysis in terms of diffractive orbits has been made in a different atomic physics system, the response of hydrogenic atoms to strong magnetic fields [35.23]. In these systems, a single electron is highly excited and takes long traversals far from the nucleus. Upon returning to a hydrogen nucleus, it is re-ejected with the reversed momentum as discussed in chapter 34. However, if the atom is not hydrogen but sodium or some other atom with one valence electron, the returning electron feels the charge distribution of the core electrons and not just the charge of the nucleus. This so-called quantum defect induces scattering in addition to the classical re-ejection present in the hydrogen atom. (In this case the local analysis consists of neglecting the magnetic field when the trajectory is near the nucleus.) This is formally similar to the vertex which causes both specular reflection and diffraction. There is then additional structure in the Fourier transform of the quantum spectrum correspond-
ing to the induced diffractive orbits, and this has been observed experimentally [35.24].

## Résumé

In this chapter we have discovered new types of periodic orbits contributing to the semiclassical traces and determinants. Unlike the periodic orbits we had seen so far, these are not true classical orbits. They are generated by singularities of the scattering potential. In these singular points the classical dynamics has no unique definition, and the classical orbits hitting the singularities can be continued in many different directions. While the classical mechanics does not know which way to go, quantum mechanics solves the dilemma by allowing us to continue in all possible directions. The likelihoods of different paths are given by the quantum mechanical weights called diffraction constants. The total contribution to a trace from such orbit is given by the product of transmission amplitudes between singularities and diffraction constants of singularities. The weights of diffractive periodic orbits are at least of order $1 / \sqrt{k}$ weaker than the weights associated with classically realizable orbits, and their contribution at large energies is therefore negligible. Nevertheless, they can strongly influence the low lying resonances or energy levels. In some systems, such as the $N$ disk scattering the diffraction effects do not only perturb semiclassical resonances, but can also create new low energy resonances. Therefore it is always important to include the contributions of diffractive periodic orbits when semiclassical methods are applied at low energies.

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## Exercises

(N. Whelan)

Exercise 35.1 Stationary phase integral. Evaluate the two stationary phase integrals corresponding to contours $E_{1}$ and $E_{2}$ of figure 35.3 and thereby verify (35.7).

Exercise 35.2 Scattering from a small disk Imagine that instead of a wedge, we have a disk whose radius $a$ is much smaller than the typical wavelengths we are considering. In that limit, solve the quantum scattering problem - find the scattered wave which result from an incident plane wave. You can do this by the method of partial waves - the analogous three dimensional problem is discussed in most quantum textbooks. You should find that only the $m=0$ partial wave contributes for small $a$. Following the discussion above, show that the diffraction constant is

$$
\begin{equation*}
d=\frac{2 \pi}{\log \left(\frac{2}{k a}\right)-\gamma_{e}+i \frac{\pi}{2}} \tag{35.28}
\end{equation*}
$$

where $\gamma_{e}=0.577 \cdots$ is Euler's constant. Note that in this limit d depends weakly on $k$ but not on the scattering angle.

Exercise 35.3 Several diffractive legs. Derive equation (35.18). The calculation involves considering slight variations of the diffractive orbit as in the simple case discussed above. Here it is more complicated because there are more diffractive arcs - however you should convince yourself that a slight variation of the diffractive orbit only affects one leg at a time.

Exercise 35.4 Unsymmetrized dynamical zeta function. Assume you know nothing about symmetry decomposition. Construct the three node Markov diagram for figure 35.1 by considering $A, B$ and $B^{\prime}$ to be physically distinct. Write down the corresponding dynamical zeta function and check explicitly that for $B=B^{\prime}$ it factorizes into the product of the the even and odd dynamical zeta functions. Why is there no term $t_{\overline{2}}$ in the full dynamical zeta function?

Exercise 35.5 Three point scatterers. Consider the limiting case of the three disk game of pinball of figure 1.1 where the disks are very much smaller than their spacing $R$. Use the results of exercise 35.2 to construct the desymmetrized dynamical zeta functions, as in sect. 22.6. You should find $1 / \zeta_{A_{1}}=1-2 t$ where $t=d e^{i(k R-3 \pi / 4)} / \sqrt{8 \pi k R}$. Compare this formula with that from chapter 11. By assuming that the real part of $k$ is much greater than the imaginary part show that the positions of the resonances are $k_{n} R=\alpha_{n}-i \beta_{n}$ where $\alpha_{n}=2 \pi n+3 \pi / 4$, $\beta_{n}=\log \left(\sqrt{2 \pi \alpha_{n}} / d\right)$ and $n$ is a non-negative integer. (See also reference [35.11].)

## Epilogue

> Nowadays, whatever the truth of the matter may be (and we will probably never know), the simplest solution is no longer emotionally satisfying. Everything we know about the world militates against it. The concepts of indeterminacy and chaos have filtered down to us from the higher sciences to confirm our nagging suspicions.
> L. Sante, "Review of 'American Tabloid' by James Ellroy", New York Review of Books (May 11, 1995)

A motion on a strange attractor can be approximated by shadowing long orbits by sequences of nearby shorter periodic orbits. This notion has here been made precise by approximating orbits by prime cycles, and evaluating associated curvatures. A curvature measures the deviation of a long cycle from its approximation by shorter cycles; the smoothness of the dynamical system implies exponential fall-off for (almost) all curvatures. We propose that the theoretical and experimental non-wandering sets be expressed in terms of the symbol sequences of short cycles (a topological characterization of the spatial layout of the non-wandering set) and their eigenvalues (metric structure)

## Cycles as the skeleton of chaos

We wind down this all-too-long treatise by asking: why cycle?
We tend to think of a dynamical system as a smooth system whose evolution can be followed by integrating a set of differential equations. Traditionally one used integrable motions as zeroth-order approximations to physical systems, and accounted for weak nonlinearities perturbatively. However, when the evolution is actually followed through to asymptotic times, one discovers that the strongly nonlinear systems show an amazingly rich structure which is not at all apparent in their formulation in terms of differential equations. In particular, the periodic orbits are important because they form the skeleton onto which all trajectories trapped for long times cling. This was already appreciated century ago by H. Poincaré, who, describing in Les méthodes nouvelles de la méchanique céleste his discovery of homoclinic tangles, mused that "the complexity of this figure will be striking, and I shall not even try to draw it". Today such drawings
are cheap and plentiful; but Poincaré went a step further and, noting that hidden in this apparent chaos is a rigid skeleton, a tree of cycles (periodic orbits) of increasing lengths and self-similar structure, suggested that the cycles should be the key to chaotic dynamics.

The zeroth-order approximations to harshly chaotic dynamics are very different from those for the nearly integrable systems: a good starting approximation here is the stretching and kneading of a baker's map, rather than the winding of a harmonic oscillator.

For low dimensional deterministic dynamical systems description in terms of cycles has many virtues:

1. cycle symbol sequences are topological invariants: they give the spatial layout of a non-wandering set
2. cycle eigenvalues are metric invariants: they give the scale of each piece of a non-wandering set
3. cycles are dense on the asymptotic non-wandering set
4. cycles are ordered hierarchically: short cycles give good approximations to a non-wandering set, longer cycles only refinements. Errors due to neglecting long cycles can be bounded, and typically fall off exponentially or super-exponentially with the cutoff cycle length
5. cycles are structurally robust: for smooth flows eigenvalues of short cycles vary slowly with smooth parameter changes
6. asymptotic averages (such as correlations, escape rates, quantum mechanical eigenstates and other "thermodynamic" averages) can be efficiently computed from short cycles by means of cycle expansions

Points 1, 2: That the cycle topology and eigenvalues are invariant properties of dynamical systems follows from elementary considerations. If the same dynamics is given by a map $f$ in one set of coordinates, and a map $g$ in the next, then $f$ and $g$ (or any other good representation) are related by a reparametrization and a coordinate transformation $f=h^{-1} \circ g \circ h$. As both $f$ and $g$ are arbitrary representations of the dynamical system, the explicit form of the conjugacy $h$ is of no interest, only the properties invariant under any transformation $h$ are of general import. The most obvious invariant properties are topological; a fixed point must be a fixed point in any representation, a trajectory which exactly returns to the initial point (a cycle) must do so in any representation. Furthermore, a good representation should not mutilate the data; $h$ must be a smooth transformation which maps nearby cycle points of $f$ into nearby cycle points of $g$. This smoothness guarantees that the cycles are not only topological invariants, but that their linearized neighborhoods are also metrically invariant. In particular, the cycle eigenvalues (eigenvalues of the Jacobian matrixs $d f^{n}(x) / d x$ of periodic orbits $\left.f^{n}(x)=x\right)$ are invariant.

Point 5: An important virtue of cycles is their structural robustness. Many quantities customarily associated with dynamical systems depend on the notion of "structural stability", that is, robustness of non-wandering set to small parameter variations.

Still, the sufficiently short unstable cycles are structurally robust in the sense that they are only slightly distorted by such parameter changes, and averages computed using them as a skeleton are insensitive to small deformations of the non-wandering set. In contrast, lack of structural stability wreaks havoc with long time averages such as Lyapunov exponents, for which there is no guarantee that they converge to the correct asymptotic value in any finite time numerical computation.

The main recent theoretical advance is point 4: we now know how to control the errors due to neglecting longer cycles. As we seen above, even though the number of invariants is infinite (unlike, for example, the number of Casimir invariants for a compact Lie group) the dynamics can be well approximated to any finite accuracy by a small finite set of invariants. The origin of this convergence is geometrical, as we shall see in appendix J.1.2, and for smooth flows the convergence of cycle expansions can even be superexponential.

The cycle expansions such as (18.5) outperform the pedestrian methods such as extrapolations from the finite cover sums (19.2) for a number of reasons. The cycle expansion is a better averaging procedure than the naive box counting algorithms because the strange attractor is here pieced together in a topologically invariant way from neighborhoods ("space average") rather than explored by a long ergodic trajectory ("time average"). The cycle expansion is co-ordinate and reparametrization invariant - a finite $n$th level sum (19.2) is not. Cycles are of finite period but infinite duration, so the cycle eigenvalues are already evaluated in the $n \rightarrow \infty$ limit, but for the sum (19.2) the limit has to be estimated by numerical extrapolations. And, crucially, the higher terms in the cycle expansion (18.5) are deviations of longer prime cycles from their approximations by shorter cycles. Such combinations vanish exactly in piecewise linear approximations and fall off exponentially for smooth dynamical flows.

In the above we have reviewed the general properties of the cycle expansions; those have been applied to a series of examples of low-dimensional chaos: 1-d strange attractors, the period-doubling repeller, the Hénon-type maps and the mode locking intervals for circle maps. The cycle expansions have also been applied to the irrational windings set of critical circle maps, to the Hamiltonian period-doubling repeller, to a Hamiltonian three-disk game of pinball, to the three-disk quantum scattering resonances and to the extraction of correlation exponents, Feasibility of analysis of experimental non-wandering set in terms of cycles is discussed in ref. [18.1].

## Homework assignment

> "Lo! thy dread empire Chaos is restor'd, Light dies before thy uncreating word; Thy hand, great Anarch, lets the curtain fall, And universal darkness buries all."
> Alexander Pope, The Dunciad

We conclude cautiously with a homework assignment posed May 22, 1990 (the original due date was May 22, 2000, but alas...):

1. Topology Develop optimal sequences ("continued fraction approximants") of finite subshift approximations to generic dynamical systems. Apply to (a) the Hénon map, (b) the Lorenz flow and (c) the Hamiltonian standard map.
2. Non-hyperbolicity Incorporate power-law (marginal stability orbits, "intermittency") corrections into cycle expansions. Apply to long-time tails in the Hamiltonian diffusion problem.
3. Phenomenology Carry through a convincing analysis of a genuine experimentally extracted data set in terms of periodic orbits.
4. Invariants Prove that the scaling functions, or the cycles, or the spectrum of a transfer operator are the maximal set of invariants of an (physically interesting) dynamically generated non-wandering set.
5. Field theory Develop a periodic orbit theory of systems with many unstable degrees of freedom. Apply to (a) coupled lattices, (b) cellular automata, (c) neural networks.
6. Tunneling Add complex time orbits to quantum mechanical cycle expansions (WKB theory for chaotic systems).
7. Unitarity Evaluate corrections to the Gutzwiller semiclassical periodic orbit sums. (a) Show that the zeros (energy eigenvalues) of the appropriate Selberg products are real. (b) Find physically realistic systems for which the "semiclassical" periodic orbit expansions yield the exact quantization.
8. Atomic spectra Compute the helium spectrum from periodic orbit expansions (already accomplished by Wintgen and Tanner!).
9. Symmetries Include fermions, gauge fields into the periodic orbit theory.
10. Quantum field theory Develop quantum theory of systems with infinitely many classically unstable degrees of freedom. Apply to (a) quark confinement (b) early universe (c) the brain.

## Conclusion

Good-bye. I am leaving because I am bored.
George Saunders' dying words

Nadie puede escribir un libro. Para Que un libro sea verdaderamente, Se requieren la aurora y el poniente Siglos, armas y el mar que une y separa.
Jorge Luis Borges El Hacedor, Ariosto y los arabes
The buttler did it.

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## Chaos: Classical and Quantum

## Part III: Material available on ChaosBook.org

## Appendix A

## A brief history of chaos

## Laws of attribution

1. Arnol'd's Law: everything that is discovered is named after someone else (including Arnol'd's law)
2. Berry's Law: sometimes, the sequence of antecedents seems endless. So, nothing is discovered for the first time.
3. Whiteheads's Law: Everything of importance has been said before by someone who did not discover it.
M.V. Berry

## A. 1 Chaos is born

(R. Mainieri)

Classical mechanics has not stood still since Newton. The formalism that we use today was developed by Euler and Lagrange. By the end of the 1800's the three problems that would lead to the notion of chaotic dynamics were already known: the three-body problem, the ergodic hypothesis, and nonlinear oscillators.

## A.1.1 Three-body problem

Trying to predict the motion of the Moon has preoccupied astronomers since antiquity. Accurate understanding of its motion was important for determining the longitude of ships while traversing open seas. Kepler's Rudolphine tables had been a great improvement over previous tables, and Kepler was justly proud of his achievements. He wrote in the introduction to the announcement of Kepler's third law, Harmonice Mundi (Linz, 1619)
in a style that would not fly with the contemporary Physical Review Letters editors:


#### Abstract

What I prophesied two-and-twenty years ago, as soon as I discovered the five solids among the heavenly orbits - what I firmly believed long before I had seen Ptolemy's Harmonics - what I had promised my friends in the title of this book, which I named before I was sure of my discovery - what sixteen years ago, I urged as the thing to be sought - that for which I joined Tycho Brahé, for which I settled in Prague, for which I have devoted the best part of my life to astronomical contemplations, at length I have brought to light, and recognized its truth beyond my most sanguine expectations. It is not eighteen months since I got the first glimpse of light, three months since the dawn, very few days since the unveiled sun, most admirable to gaze upon, burst upon me. Nothing holds me; I will indulge my sacred fury; I will triumph over mankind by the honest confession that I have stolen the golden vases of the Egyptians to build up a tabernacle for my God far away from the confines of Egypt. If you forgive me, I rejoice; if you are angry, I can bear it; the die is cast, the book is written, to be read either now or in posterity, I care not which; it may well wait a century for a reader, as God has waited six thousand years for an observer.


Bernoulli used Newton's work on mechanics to derive the elliptic orbits of Kepler and set an example of how equations of motion could be solved by integrating. But the motion of the Moon is not well approximated by an ellipse with the Earth at a focus; at least the effects of the Sun have to be taken into account if one wants to reproduce the data the classical Greeks already possessed. To do that one has to consider the motion of three bodies: the Moon, the Earth, and the Sun. When the planets are replaced by point particles of arbitrary masses, the problem to be solved is known as the three-body problem. The three-body problem was also a model to another concern in astronomy. In the Newtonian model of the solar system it is possible for one of the planets to go from an elliptic orbit around the Sun to an orbit that escaped its domain or that plunged right into it. Knowing if any of the planets would do so became the problem of the stability of the solar system. A planet would not meet this terrible end if solar system consisted of two celestial bodies, but whether such fate could befall in the three-body case remained unclear.

After many failed attempts to solve the three-body problem, natural philosophers started to suspect that it was impossible to integrate. The usual technique for integrating problems was to find the conserved quantities, quantities that do not change with time and allow one to relate the momenta and positions different times. The first sign on the impossibility of integrating the three-body problem came from a result of Burns that showed that there were no conserved quantities that were polynomial in the momenta and positions. Burns' result did not preclude the possibility of more complicated conserved quantities. This problem was settled by Poincaré and Sundman in two very different ways.

In an attempt to promote the journal Acta Mathematica, Mittag-Leffler got the permission of the King Oscar II of Sweden and Norway to establish
a mathematical competition. Several questions were posed (although the king would have preferred only one), and the prize of 2500 kroner would go to the best submission. One of the questions was formulated by Weierstrass:

Given a system of arbitrary mass points that attract each other according to Newton's laws, under the assumption that no two points ever collide, try to find a representation of the coordinates of each point as a series in a variable that is some known function of time and for all of whose values the series converges uniformly.

This problem, whose solution would considerably extend our understanding of the solar system, ...

Poincaré's submission won the prize. He showed that conserved quantities that were analytic in the momenta and positions could not exist. To show that he introduced methods that were very geometrical in spirit: the importance of phase flow, the role of periodic orbits and their cross sections, the homoclinic points.

The interesting thing about Poincaré's work was that it did not solve the problem posed. He did not find a function that would give the coordinates as a function of time for all times. He did not show that it was impossible either, but rather that it could not be done with the Bernoulli technique of finding a conserved quantity and trying to integrate. Integration would seem unlikely from Poincaré's prize-winning memoir, but it was accomplished by the Finnish-born Swedish mathematician Sundman. Sundman showed that to integrate the three-body problem one had to confront the two-body collisions. He did that by making them go away through a trick known as regularization of the collision manifold. The trick is not to expand the coordinates as a function of time $t$, but rather as a function of $\sqrt[3]{t}$. To solve the problem for all times he used a conformal map into a strip. This allowed Sundman to obtain a series expansion for the coordinates valid for all times, solving the problem that was proposed by Weirstrass in the King Oscar II's competition.

The Sundman's series are not used today to compute the trajectories of any three-body system. That is more simply accomplished by numerical methods or through series that, although divergent, produce better numerical results. The conformal map and the collision regularization mean that the series are effectively in the variable $1-e^{-\sqrt[3]{t}}$. Quite rapidly this gets exponentially close to one, the radius of convergence of the series. Many terms, more terms than any one has ever wanted to compute, are needed to achieve numerical convergence. Though Sundman's work deserves better credit than it gets, it did not live up to Weirstrass's expectations, and the series solution did not "considerably extend our understanding of the solar system." The work that followed from Poincaré did.

## A.1.2 Ergodic hypothesis

The second problem that played a key role in development of chaotic dynamics was the ergodic hypothesis of Boltzmann. Maxwell and Boltzmann
had combined the mechanics of Newton with notions of probability in order to create statistical mechanics, deriving thermodynamics from the equations of mechanics. To evaluate the heat capacity of even a simple system, Boltzmann had to make a great simplifying assumption of ergodicity: that the dynamical system would visit every part of the phase space allowed by conservations law equally often. This hypothesis was extended to other averages used in statistical mechanics and was called the ergodic hypothesis. It was reformulated by Poincaré to say that a trajectory comes as close as desired to any phase space point.

Proving the ergodic hypothesis turned out to be very difficult. By the end of our own century it has only been shown true for a few systems and wrong for quite a few others. Early on, as a mathematical necessity, the proof of the hypothesis was broken down into two parts. First one would show that the mechanical system was ergodic (it would go near any point) and then one would show that it would go near each point equally often and regularly so that the computed averages made mathematical sense. Koopman took the first step in proving the ergodic hypothesis when he noticed that it was possible to reformulate it using the recently developed methods of Hilbert spaces. This was an important step that showed that it was possible to take a finite-dimensional nonlinear problem and reformulate it as a infinite-dimensional linear problem. This does not make the problem easier, but it does allow one to use a different set of mathematical tools on the problem. Shortly after Koopman started lecturing on his method, von Neumann proved a version of the ergodic hypothesis, giving it the status of a theorem. He proved that if the mechanical system was ergodic, then the computed averages would make sense. Soon afterwards Birkhoff published a much stronger version of the theorem.

## A.1.3 Nonlinear oscillators

The third problem that was very influential in the development of the theory of chaotic dynamical systems was the work on the nonlinear oscillators. The problem is to construct mechanical models that would aid our understanding of physical systems. Lord Rayleigh came to the problem through his interest in understanding how musical instruments generate sound. In the first approximation one can construct a model of a musical instrument as a linear oscillator. But real instruments do not produce a simple tone forever as the linear oscillator does, so Lord Rayleigh modified this simple model by adding friction and more realistic models for the spring. By a clever use of negative friction he created two basic models for the musical instruments. These models have more than a pure tone and decay with time when not stroked. In his book The Theory of Sound Lord Rayleigh introduced a series of methods that would prove quite general, such as the notion of a limit cycle, a periodic motion a system goes to regardless of the initial conditions.

## A. 2 Chaos grows up

(R. Mainieri)

The theorems of von Neumann and Birkhoff on the ergodic hypothesis were published in 1912 and 1913. This line of enquiry developed in two directions. One direction took an abstract approach and considered dynamical systems as transformations of measurable spaces into themselves. Could we classify these transformations in a meaningful way? This lead Kolmogorov to the introduction of the concept of entropy for dynamical systems. With entropy as a dynamical invariant it became possible to classify a set of abstract dynamical systems known as the Bernoulli systems. The other line that developed from the ergodic hypothesis was in trying to find mechanical systems that are ergodic. An ergodic system could not have stable orbits, as these would break ergodicity. So in 1898 Hadamard published a paper with a playful title of "... billiards ...," where he showed that the motion of balls on surfaces of constant negative curvature is everywhere unstable. This dynamical system was to prove very useful and it was taken up by Birkhoff. Morse in 1923 showed that it was possible to enumerate the orbits of a ball on a surface of constant negative curvature. He did this by introducing a symbolic code to each orbit and showed that the number of possible codes grew exponentially with the length of the code. With contributions by Artin, Hedlund, and Hopf it was eventually proven that the motion of a ball on a surface of constant negative curvature was ergodic. The importance of this result escaped most physicists, one exception being Krylov, who understood that a physical billiard was a dynamical system on a surface of negative curvature, but with the curvature concentrated along the lines of collision. Sinai, who was the first to show that a physical billiard can be ergodic, knew Krylov's work well.

The work of Lord Rayleigh also received vigorous development. It prompted many experiments and some theoretical development by van der Pol, Duffing, and Hayashi. They found other systems in which the nonlinear oscillator played a role and classified the possible motions of these systems. This concreteness of experiments, and the possibility of analysis was too much of temptation for Mary Lucy Cartwright and J.E. Littlewood, who set out to prove that many of the structures conjectured by the experimentalists and theoretical physicists did indeed follow from the equations of motion. Birkhoff had found a "remarkable curve" in a two dimensional map; it appeared to be non-differentiable and it would be nice to see if a smooth flow could generate such a curve. The work of Cartwright and Littlewood lead to the work of Levinson, which in turn provided the basis for the horseshoe construction of S. Smale.

In Russia, Lyapunov paralleled the methods of Poincaré and initiated the strong Russian dynamical systems school. Andronov carried on with the study of nonlinear oscillators and in 1937 introduced together with Pontryagin the notion of coarse systems. They were formalizing the understanding garnered from the study of nonlinear oscillators, the understanding that many of the details on how these oscillators work do not affect the overall
picture of the phase space: there will still be limit cycles if one changes the dissipation or spring force function by a little bit. And changing the system a little bit has the great advantage of eliminating exceptional cases in the mathematical analysis. Coarse systems were the concept that caught Smale's attention and enticed him to study dynamical systems.

## A. 3 Chaos with us

(R. Mainieri)

In the fall of 1961 Steven Smale was invited to Kiev where he met Arnol'd, Anosov, Sinai, and Novikov. He lectured there, and spent a lot of time with Anosov. He suggested a series of conjectures, most of which Anosov proved within a year. It was Anosov who showed that there are dynamical systems for which all points (as opposed to a non-wandering set) admit the hyperbolic structure, and it was in honor of this result that Smale named these systems Axiom-A. In Kiev Smale found a receptive audience that had been thinking about these problems. Smale's result catalyzed their thoughts and initiated a chain of developments that persisted into the 1970's.

Smale collected his results and their development in the 1967 review article on dynamical systems, entitled "Differentiable dynamical systems". There are many great ideas in this paper: the global foliation of invariant sets of the map into disjoint stable and unstable parts; the existence of a horseshoe and enumeration and ordering of all its orbits; the use of zeta functions to study dynamical systems. The emphasis of the paper is on the global properties of the dynamical system, on how to understand the topology of the orbits. Smale's account takes you from a local differential equation (in the form of vector fields) to the global topological description in terms of horseshoes.

The path traversed from ergodicity to entropy is a little more confusing. The general character of entropy was understood by Weiner, who seemed to have spoken to Shannon. In 1948 Shannon published his results on information theory, where he discusses the entropy of the shift transformation. Kolmogorov went far beyond and suggested a definition of the metric entropy of an area preserving transformation in order to classify Bernoulli shifts. The suggestion was taken by his student Sinai and the results published in 1959. In 1960 Rohlin connected these results to measure-theoretical notions of entropy. The next step was published in 1965 by Adler and Palis, and also Adler, Konheim, McAndrew; these papers showed that one could define the notion of topological entropy and use it as an invariant to classify continuous maps. In 1967 Anosov and Sinai applied the notion of entropy to the study of dynamical systems. It was in the context of studying the entropy associated to a dynamical system that Sinai introduced Markov partitions in 1968.

Markov partitions allow one to relate dynamical systems and statisti-
cal mechanics; this has been a very fruitful relationship. It adds measure notions to the topological framework laid down in Smale's paper. Markov partitions divide the phase space of the dynamical system into nice little boxes that map into each other. Each box is labeled by a code and the dynamics on the phase space maps the codes around, inducing a symbolic dynamics. From the number of boxes needed to cover all the space, Sinai was able to define the notion of entropy of a dynamical system. In 1970 Bowen came up independently with the same ideas, although there was presumably some flow of information back and forth before these papers got published. Bowen also introduced the important concept of shadowing of chaotic orbits. We do not know whether at this point the relations with statistical mechanics were clear to every one. They became explicit in the work of Ruelle. Ruelle understood that the topology of the orbits could be specified by a symbolic code, and that one could associate an "energy" to each orbit. The energies could be formally combined in a "partition function" to generate the invariant measure of the system.

After Smale, Sinai, Bowen, and Ruelle had laid the foundations of the statistical mechanics approach to chaotic systems, research turned to studying particular cases. The simplest case to consider is one-dimensional maps. The topology of the orbits for parabola-like maps was worked out in 1973 by Metropolis, Stein, and Stein. The more general one-dimensional case was worked out in 1976 by Milnor and Thurston in a widely circulated preprint, whose extended version eventually got published in 1988.

A lecture of Smale and the results of Metropolis, Stein, and Stein inspired Feigenbaum to study simple maps. This lead him to the discovery of the universality in quadratic maps and the application of ideas from field-theory to dynamical systems. Feigenbaum's work was the culmination in the study of one-dimensional systems; a complete analysis of a nontrivial transition to chaos. Feigenbaum introduced many new ideas into the field: the use of the renormalization group which lead him to introduce functional equations in the study of dynamical systems, the scaling function which completed the link between dynamical systems and statistical mechanics, and the use of presentation functions as the dynamics of scaling functions.

The work in more than one dimension progressed very slowly and is still far from completed. The first result in trying to understand the topology of the orbits in two dimensions (the equivalent of Metropolis, Stein, and Stein, or Milnor and Thurston's work) was obtained by Thurston. Around 1975 Thurston was giving lectures "On the geometry and dynamics of diffeomorphisms of surfaces". Thurston's techniques exposed in that lecture have not been applied in physics, but much of the classification that Thurston developed can be obtained from the notion of a "pruning front" developed independently by Cvitanović.

Once one develops an understanding for the topology of the orbits of a dynamical system, one needs to be able to compute its properties. Ruelle had already generalized the zeta function introduced by Artin and Mazur so that it could be used to compute the average value of observables. The
difficulty with Ruelle's zeta function is that it does not converge very well. Starting out from Smale's observation that a chaotic dynamical system is dense with a set of periodic orbits, Cvitanović used these orbits as a skeleton on which to evaluate the averages of observables, and organized such calculations in terms of rapidly converging cycle expansions. This convergence is attained by using the shorter orbits used as a basis for shadowing the longer orbits.

This account is far from complete, but we hope that it will help get a sense of perspective on the field. It is not a fad and it will not die anytime soon.

Remark A. 1 Notion of global foliations. For each paper cited in dynamical systems literature, there are many results that went into its development. As an example, take the notion of global foliations that we attribute to Smale. As far as we can trace the idea, it goes back to René Thom; local foliations were already used by Hadamard. Smale attended a seminar of Thom in 1958 or 1959. In that seminar Thom was explaining his notion of transversality. One of Thom's disciples introduced Smale to Brazilian mathematician Peixoto. Peixoto (who had learned the results of the Andronov-Pontryagin school from Lefschetz) was the closest Smale had ever come until then to the Andronov-Pontryagin school. It was from Peixoto that Smale learned about structural stability, a notion that got him enthusiastic about dynamical systems, as it blended well with his topological background. It was from discussions with Peixoto that Smale got the problems in dynamical systems that lead him to his 1960 paper on Morse inequalities. The next year Smale published his result on the hyperbolic structure of the nonwandering set. Smale was not the first to consider a hyperbolic point, Poincaré had already done that; but Smale was the first to introduce a global hyperbolic structure. By 1960 Smale was already lecturing on the horseshoe as a structurally stable dynamical system with an infinity of periodic points and promoting his global viewpoint.

> (R. Mainieri)

Remark A. 2 Levels of ergodicity. In the mid 1970's A. Katok and Ya.B. Pesin tried to use geometry to establish positive Lyapunov exponents. A. Katok and J.M. Strelcyn carried out the program and developed a theory of general dynamical systems with singularities. They studied uniformly hyperbolic systems (as strong as Anosov's), but with sets of singularities. Under iterations a dense set of points hits the singularities. Even more important are the points that never hit the singularity set. In order to establish some control over how they approach the set, one looks at trajectories that apporach the set by some given $\epsilon^{n}$, or faster.

Ya.G. Sinai, L. Bunimovich and Chernov studied the geometry of billiards in a very detailed way. A. Katok and Ya.B. Pesin's idea was much more robust. Look at the discontinuity set (geometry of it matters not at all), take an $\epsilon$ neighborhood around it. Given that the Lebesgue measure is $\epsilon^{\alpha}$ and the stability grows not faster than (distance) ${ }^{n}$, A. Katok and J.-M. Strelcyn prove that the Lyapunov exponent is non-zero.

In mid 1980's Ya.B. Pesin studied the dissipative case. Now the problem has no invariant Lebesgue measure. Assuming uniform hyperbolicity, with singularities, and tying together Lebesgue measure and discontinutities, and given that the stability grows not faster than (distance) ${ }^{n}$, Ya.B. Pesin proved that the Lyapunov
exponent is non-zero, and that SRB measure exists. He also proved that the Lorenz, Lozi and Byelikh attractors satisfy these conditions.

In the the systems were uniformly hyperbolic, all trouble was in differentials. For the Hénon attractor, already the differentials are nonhyperbolic. The points do not separate uniformly, but the analogue of the singularity set can be obtained by excizing the regions that do not separate. Hence there are 3 levels of ergodic systems:

1. Anosov flow
2. Anosov flow + singularity set

- the Hamiltonian systems: general case A. Katok and J.-M. Strelcyn, billiards Ya.G. Sinai and L. Bunimovich.
- the dissipative case: Ya.B. Pesin

3. Hénon

- The first proof was given by M. Benedicks and L. Carleson [12.33].
- A more readable proof is given in M. Benedicks and L.-S. Young [3.10]
(based on Ya.B. Pesin's comments)


## A.3.1 Periodic orbit theory

The history of the periodic orbit theory is rich and curious, and the recent advances are to equal degree inspired by a century of separate development of three disparate subjects; 1. classical chaotic dynamics, initiated by Poincaré and put on its modern footing by Smale, Ruelle, and many others; 2. quantum theory initiated by Bohr, with the modern "chaotic" formulation by Gutzwiller; and 3. analytic number theory initiated by Riemann and formulated as a spectral problem by Selberg. Following totally different lines of reasoning and driven by very different motivations, the three separate roads all arrive at formally nearly identical trace formulas, zeta functions and spectral determinants.

That these topics should be related is far from obvious. Connection between dynamics and number theory arises from Selberg's observation that description of geodesic motion and wave mechanics on spaces of constant negative curvature is essentially a number-theoretic problem. A posteriori, one can say that zeta functions arise in both classical and quantum mechanics because in both the dynamical evolution can be described by the action of linear evolution (or transfer) operators on infinite-dimensional vector spaces. The spectra of these operators are given by the zeros of appropriate determinants. One way to evaluate determinants is to expand them in terms of traces, $\log$ det $=\operatorname{tr} \log$, and in this way the spectrum of an evolution operator becames related to its traces, that is, periodic orbits. A perhaps deeper way of restating this is to observe that the trace formulas perform the same service in all of the above problems; they relate the spectrum of lengths (local dynamics) to the spectrum of eigenvalues (global averages), and for nonlinear geometries they play a role analogous to that the Fourier transform plays for the circle.

## A. 4 Death of the Old Quantum Theory

In 1913 Otto Stern and Max Theodor Felix von Laue went up for a walk up the Uetliberg. On the top they sat down and talked about physics. In particular they talked about the new atom model of Bohr. There and then they made the "Uetli Schwur": If that crazy model of Bohr turned out to be right, then they would leave physics. It did and they didn't.
A. Pais, Inward Bound: of Matter and Forces in the Physical World

In an afternoon of May 1991 Dieter Wintgen is sitting in his office at the Niels Bohr Institute beaming with the unparalleled glee of a boy who has just committed a major mischief. The starting words of the manuscript he has just penned are

The failure of the Copenhagen School to obtain a reasonable ...

34 years old at the time, Dieter was a scruffy kind of guy, always in sandals and holed out jeans, a left winger and a mountain climber, working around the clock with his students Gregor and Klaus to complete the work that Bohr himself would have loved to see done back in 1916: a "planetary" calculation of the helium spectrum.

Never mind that the "Copenhagen School" refers not to the old quantum theory, but to something else. The old quantum theory was no theory at all; it was a set of rules bringing some order to a set of phenomena which defied logic of classical theory. The electrons were supposed to describe planetary orbits around the nucleus; their wave aspects were yet to be discovered. The foundations seemed obscure, but Bohr's answer for the once-ionized helium to hydrogen ratio was correct to five significant figures and hard to ignore. The old quantum theory marched on, until by 1924 it reached an impasse: the helium spectrum and the Zeeman effect were its death knell.

Since the late 1890's it had been known that the helium spectrum consists of the orthohelium and parahelium lines. In 1915 Bohr suggested that the two kinds of helium lines might be associated with two distinct shapes of orbits (a suggestion that turned out to be wrong). In 1916 he got Kramers to work on the problem, and wrote to Rutherford: "I have used all my spare time in the last months to make a serious attempt to solve the problem of ordinary helium spectrum ...I think really that at last I have a clue to the problem." To other colleagues he wrote that "the theory was worked out in the fall of 1916 " and of having obtained a "partial agreement with the measurements." Nevertheless, the Bohr-Sommerfeld theory, while by and large successful for hydrogen, was a disaster for neutral helium. Heroic efforts of the young generation, including Kramers and Heisenberg, were of no avail.

For a while Heisenberg thought that he had the ionization potential for helium, which he had obtained by a simple perturbative scheme. He wrote enthusiastic letters to Sommerfeld and was drawn into a collaboration with Max Born to compute the spectrum of helium using Born's systematic perturbative scheme. In first approximation, they reproduced the earlier calculations. The next level of corrections turned out to be larger than the computed effect. The concluding paragraph of Max Born's classic "Vorlesungen über Atommechanik" from 1925 sums it up in a somber tone:
(...) the systematic application of the principles of the quantum theory (...) gives results in agreement with experiment only in those cases where the motion of a single electron is considered; it fails even in the treatment of the motion of the two electrons in the helium atom.

This is not surprising, for the principles used are not really consistent. (...) A complete systematic transformation of the classical mechanics into a discontinuous mechanics is the goal towards which the quantum theory strives.

That year Heisenberg suffered a bout of hay fever, and the old quantum theory was dead. In 1926 he gave the first quantitative explanation of the helium spectrum. He used wave mechanics, electron spin and the Pauli exclusion principle, none of which belonged to the old quantum theory, and planetary orbits of electrons were cast away for nearly half a century.

Why did Pauli and Heisenberg fail with the helium atom? It was not the fault of the old quantum mechanics, but rather it reflected their lack of understanding of the subtleties of classical mechanics. Today we know what they missed in 1913-24: the role of conjugate points (topological indices) along classical trajectories was not accounted for, and they had no idea of the importance of periodic orbits in nonintegrable systems.

Since then the calculation for helium using the methods of the old quantum mechanics has been fixed. Leopold and Percival added the topological indices in 1980, and in 1991 Wintgen and collaborators orbits. Dieter had good reasons to gloat; while the rest of us were preparing to sharpen our pencils and supercomputers in order to approach the dreaded 3-body problem, they just went ahead and did it. What it took - and much else - is described in this book. One is also free to ponder what quantum theory would look like today if all this was worked out in 1917.

Remark A. 3 Sources. This tale, aside from a few personal recollections, is in large part lifted from Abraham Pais' accounts of the demise of the old quantum theory [A.7, A.8], as well as Jammer's account [A.3]. The helium spectrum is taken up in chapter 34. In August 1994 Dieter Wintgen died in a climbing accident in the Swiss Alps.

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## Appendix B

## Infinite-dimensional flows



Flows described by partial differential equations (PDEs) are considered infinite dimensional because if one writes them down as a set of ordinary differential equations (ODEs) then one needs an infinity of the ordinary kind to represent the dynamics of one equation of the partial kind. Even though the phase space is infinite-dimensional, for many systems of physical interest the global attractor is finite-dimensional. We illustrate how this works for dissipative systems with a concrete example, the Kuramoto-Sivashinsky system.

## B.0.1 Partial differential equations

First, a few words about partial differential equations in general. Many of the partial differential equations of mathematical physics can be written in the quasi-linear form

$$
\begin{equation*}
\partial_{t} u=A u+N(u), \tag{B.1}
\end{equation*}
$$

where $u$ is a function (possibly a vector function) of the coordinate $x$ and time $t, A$ is a linear operator, usually containing the Laplacian and a few other derivatives of $u$, and $N(u)$ is the nonlinear part of the equation (terms like $u \partial_{x} u$ in (2.18)).

Not all equations are stated in the form (B.1), but they can easily be so transformed, just as the ordinary differential equations can be rewritten as first-order systems. We will illustrate the method with a variant of the D'Alambert's wave equation describing a plucked string:

$$
\begin{equation*}
\partial_{t t} y=\left(c+\frac{1}{2}\left(\partial_{x} y\right)^{2}\right) \partial_{x x} y . \tag{B.2}
\end{equation*}
$$

Were the term $\partial_{x} y$ small, this equation would be just the ordinary wave equation. To rewrite the equation in the first order form (B.1), we need a
field $u=(y, w)$ that is two-dimensional,

$$
\partial_{t}\left[\begin{array}{c}
y  \tag{B.3}\\
w
\end{array}\right]=\left[\begin{array}{cc}
0 & 1 \\
c \partial_{x x} & 0
\end{array}\right]\left[\begin{array}{c}
y \\
w
\end{array}\right]+\left[\begin{array}{c}
0 \\
\partial_{x x} y\left(\partial_{x} y\right)^{2} / 2
\end{array}\right]
$$

The $[2 \times 2]$ matrix is the linear operator $A$ and the vector on the far right is the nonlinear function $N(u)$. Unlike ordinary functions, differentiations are part of the function. The nonlinear part can also be expressed as a function on the infinite set of numbers that represent the field, as exemplified by the Kuramoto-Sivashinsky system (2.18).

The usual technique for solving the linear part is to use Fourier methods. Just as in the ordinary differential equation case, one can integrate the linear part of

$$
\begin{equation*}
\partial_{t} u=A u \tag{B.4}
\end{equation*}
$$

to obtain

$$
\begin{equation*}
u(x, t)=e^{t A} u(x, 0) \tag{B.5}
\end{equation*}
$$

If $u$ is expressed as Fourier series $\sum_{k} a_{k} \exp (i k x)$, as we will do for the Kuramoto-Shivashinsky system, then we can determine the action of $e^{t A}$ on $u(x, 0)$. This can be done because differentiations in $A$ act rather simply on the exponentials. For example,

$$
\begin{equation*}
e^{t \partial_{x}} u(x, 0)=e^{t \partial_{x}} \sum_{k} a_{k} e^{i k x}=\sum_{k} a_{k} \frac{(i t)^{k}}{k!} e^{i k x} \tag{B.6}
\end{equation*}
$$

Depending on the behavior of the linear part, one distinguishes three classes of partial differential equations: diffusion, wave, and potential. The classification relies on the solution by a Fourier series, as in (B.5). In mathematical literature these equations are also called parabolic, hyperbolic and elliptic. If the nonlinear part $N(u)$ is as big as the linear part, the classification is not a good indication of behavior, and one can encounter features of one class of equations while studying the others.

In diffusion-type equations the modes of high frequency tend to become smooth, and all initial conditions tend to an attractor, called the inertial manifold. The Kuramoto-Sivashinsky system studied below is of this type. The solution being attracted to the inertial manifold does not mean that the amplitudes of all but a finite number of modes go to zero (alas were we so lucky), but that there is a finite set of modes that could be used to describe any solution of the inertial manifold. The only catch is that there is no simple way to discover what these inertial manifold modes might be.

In wave-like equations the high frequency modes do not die out and the solutions tend to be distributions. The equations can be solved by
variations on the WKB idea: the wave-like equations can be approximated by the trajectories of the wave fronts.

Elliptic equations have no time dependence and do not represent dynamical systems.

## Appendix C

## Stability of Hamiltonian flows

## C. 1 Symplectic invariance

(M.J. Feigenbaum and P. Cvitanović)

The symplectic structure of Hamilton's equations buys us much more than the incompressibility, or the phase space volume conservation alluded to in sect. 5.1. We assume you are at home with Hamiltonian formalism. If you would like to see the Hamilton's equations derived, Hamilton's original line of reasoning is retraced in sect. 28.1.1. The evolution equations for any $p, q$ dependent quantity $Q=Q(q, p)$ are given by (9.30).

In terms of the Poisson brackets, the time evolution equation for $Q=$ $Q(q, p)$ is given by (9.32). We now recast the symplectic condition (5.13) in a form convenient for using the symplectic constraints on J. Writing $x(t)=x^{\prime}=\left[p^{\prime}, q^{\prime}\right]$ and the Jacobian matrix and its inverse

$$
\mathbf{J}=\left(\begin{array}{cc}
\frac{\partial q^{\prime}}{\partial q} & \frac{\partial q^{\prime}}{\partial p}  \tag{C.1}\\
\frac{\partial p^{\prime}}{\partial q} & \frac{\partial p^{\prime}}{\partial p}
\end{array}\right), \quad \mathbf{J}^{-1}=\left(\begin{array}{cc}
\frac{\partial q}{\partial q^{\prime}} & \frac{\partial q}{\partial p^{\prime}} \\
\frac{\partial p}{\partial q^{\prime}} & \frac{\partial p}{\partial p^{\prime}}
\end{array}\right),
$$

we can spell out the symplectic invariance condition (5.13):

$$
\begin{align*}
& \frac{\partial q_{k}^{\prime}}{\partial q_{i}} \frac{\partial p_{k}^{\prime}}{\partial q_{j}}-\frac{\partial p_{k}^{\prime}}{\partial q_{i}} \frac{\partial q_{k}^{\prime}}{\partial q_{j}}=0 \\
& \frac{\partial q_{k}^{\prime}}{\partial p_{i}} \frac{\partial p_{k}^{\prime}}{\partial p_{j}}-\frac{\partial p_{k}^{\prime}}{\partial p_{i}} \frac{\partial q_{k}^{\prime}}{\partial p_{j}}=0 \\
& \frac{\partial q_{k}^{\prime}}{\partial q_{i}} \frac{\partial p_{k}^{\prime}}{\partial p_{j}}-\frac{\partial p_{k}^{\prime}}{\partial q_{i}} \frac{\partial q_{k}^{\prime}}{\partial p_{j}}=\delta_{i j} . \tag{C.2}
\end{align*}
$$

From (5.15) we obtain

$$
\begin{equation*}
\frac{\partial q_{i}}{\partial q_{j}^{\prime}}=\frac{\partial p_{j}^{\prime}}{\partial p_{i}}, \quad \frac{\partial p_{i}}{\partial p_{j}^{\prime}}=\frac{\partial q_{j}^{\prime}}{\partial q_{i}}, \quad \frac{\partial q_{i}}{\partial p_{j}^{\prime}}=-\frac{\partial q_{j}^{\prime}}{\partial p_{i}}, \quad \frac{\partial p_{i}}{\partial q_{j}^{\prime}}=-\frac{\partial p_{j}^{\prime}}{\partial q_{i}} . \tag{C.3}
\end{equation*}
$$

Taken together, (C.3) and (C.2) imply that the flow conserves the $\{p, q\}$ Poisson brackets

$$
\begin{align*}
& \left\{q_{i}, q_{j}\right\}=\frac{\partial q_{i}}{\partial p_{k}^{\prime}} \frac{\partial q_{j}}{\partial q_{k}^{\prime}}-\frac{\partial q_{j}}{\partial p_{k}^{\prime}} \frac{\partial q_{i}}{\partial q_{k}^{\prime}}=0 \\
& \left\{p_{i}, p_{j}\right\}=0, \quad\left\{p_{i}, q_{j}\right\}=\delta_{i j}, \tag{C.4}
\end{align*}
$$

that is, the transformations induced by a Hamiltonian flow are canonical, preserving the form of the equations of motion. The first two relations are symmetric under $i, j$ interchange and yield $D(D-1) / 2$ constraints each; the last relation yields $D^{2}$ constraints. Hence only $(2 D)^{2}-2 D(D-1) / 2-$ $D^{2}=2 D^{2}+D$ elements of $\mathbf{J}$ are linearly independent, as it behooves group elements of the symplectic group $S p(2 D)$.

We have now succeeded in making the full set of constraints explicit as we shall see in appendix D , this will enable us to implement dynamics in such a way that the symplectic invariance will be automatically preserved.

## C. 2 Monodromy matrix for Hamiltonian flows

It is not the Jacobian matrix of the flow, but the monodromy matrix, which enters the trace formula. This matrix gives the time dependence of a displacement perpendicular to the flow on the energy manifold. Indeed, we discover some trivial parts in the Jacobian matrix J. An initial displacement in the direction of the flow $x=\omega \nabla H(x)$ transfers according to $\delta x(t)=x_{t}(t) \delta t$ with $\delta t$ time independent. The projection of any displacement on $\delta x$ on $\nabla H(x)$ is constant, that is, $\nabla H(x(t)) \delta x(t)=\delta E$. We get the equations of motion for the monodromy matrix directly choosing a suitable local coordinate system on the orbit $x(t)$ in form of the (non singular) transformation $\mathbf{U}(x(t))$ :

$$
\begin{equation*}
\tilde{\mathbf{J}}(x(t))=\mathbf{U}^{-1}(x(t)) \mathbf{J}(x(t)) \mathbf{U}(x(0)) \tag{C.5}
\end{equation*}
$$

These lead to

$$
\left.\begin{array}{rl} 
& \dot{\mathbf{J}}
\end{array}=\tilde{\mathbf{L}} \tilde{\mathbf{J}}, ~=\dot{\mathbf{U}}\right)
$$

Note that the properties a) - c) are only fulfilled for $\tilde{\mathbf{J}}$ and $\tilde{\mathbf{L}}$, if $\mathbf{U}$ itself is symplectic.
Choosing $x_{E}=\nabla H(t) /|\nabla H(t)|^{2}$ and $x_{t}$ as local coordinates uncovers the
two trivial eigenvalues 1 of the transformed matrix in (C.5) at any time $t$. Setting $\mathbf{U}=\left(x_{t}^{T}, x_{E}^{T}, x_{1}^{T}, \ldots, x_{2 d-2}^{T}\right)$ gives

$$
\tilde{\mathbf{J}}=\left(\begin{array}{ccccc}
1 & * & * & \ldots & *  \tag{C.7}\\
0 & 1 & 0 & \ldots & 0 \\
0 & * & & & \\
\vdots & \vdots & & \mathbf{m} & \\
0 & * & & &
\end{array}\right) ; \quad \tilde{\mathbf{L}}=\left(\begin{array}{ccccc}
0 & * & * & \ldots & * \\
0 & 0 & 0 & \ldots & 0 \\
0 & * & & & \\
\vdots & \vdots & & 1 & \\
0 & * & & &
\end{array}\right)
$$

The matrix $\mathbf{m}$ is now the monodromy matrix and the equation of motion are given by

$$
\begin{equation*}
\dot{\mathrm{m}}=\mathrm{lm} . \tag{C.8}
\end{equation*}
$$

The vectors $x_{1}, \ldots, x_{2 d-2}$ must span the space perpendicular to the flow on the energy manifold.

For a system with two degrees of freedom, the matrix $\mathbf{U}(\mathbf{t})$ can be written down explicitly, that is,

$$
\mathbf{U}(t)=\left(x_{t}, x_{1}, x_{E}, x_{2}\right)=\left(\begin{array}{cccc}
\dot{x} & -\dot{y} & -\dot{u} / q^{2} & -\dot{v} / q^{2}  \tag{C.9}\\
\dot{y} & \dot{x} & -\dot{v} / q^{2} & \dot{u} / q^{2} \\
\dot{u} & \dot{v} & \dot{x} / q^{2} & -\dot{y} / q^{2} \\
\dot{v} & -\dot{u} & \dot{y} / q^{2} & \dot{x} / q^{2}
\end{array}\right)
$$

with $x^{T}=(x, y ; u, v)$ and $q=|\nabla H|=|\dot{x}|$. The matrix $\mathbf{U}$ is non singular and symplectic at every phase space point $x$ (except the equilibrium points $\dot{x}=0$ ). The matrix elements for $\mathbf{l}$ are given (C.11). One distinguishes 4 classes of eigenvalues of $\mathbf{m}$.

- stable or elliptic, if $\Lambda=e^{ \pm i \pi \nu}$ and $\left.\nu \in\right] 0,1[$.
- marginal, if $\Lambda= \pm 1$.
- hyperbolic, inverse hyperbolic, if $\Lambda=e^{ \pm \lambda}, \Lambda=-e^{ \pm \lambda} ; \lambda>0$ is called the Lyapunov exponent of the periodic orbit.
- loxodromic, if $\Lambda=e^{ \pm u \pm i \Psi}$ with $u$ and $\Psi$ real. This is the most general case possible only in systems with 3 or more degree of freedoms.

For 2 degrees of freedom, that is, $\mathbf{m}$ is a $(2 \times 2)$ matrix, the eigenvalues are determined by

$$
\begin{equation*}
\lambda=\frac{\operatorname{Tr}(\mathbf{m}) \pm \sqrt{\operatorname{Tr}(\mathbf{m})^{2}-4}}{2} \tag{C.10}
\end{equation*}
$$

that is, $\operatorname{Tr}(\mathbf{m})=2$ separates stable and unstable behavior.

The $\mathbf{l}$ matrix elements for the local transformation (C.9) are

$$
\begin{align*}
\tilde{\mathbf{l}}_{11}= & \frac{1}{q}\left[\left(h_{x}^{2}-h_{y}^{2}-h_{u}^{2}+h_{v}^{2}\right)\left(h_{x u}-h_{y v}\right)+2\left(h_{x} h_{y}-h_{u} h_{v}\right)\left(h_{x v}+h_{y u}\right)\right. \\
& \left.-\left(h_{x} h_{u}+h_{y} h_{v}\right)\left(h_{x x}+h_{y y}-h_{u u}-h_{v v}\right)\right] \\
\tilde{\mathbf{l}}_{12}= & \frac{1}{q^{2}}\left[\left(h_{x}^{2}+h_{v}^{2}\right)\left(h_{y y}+h_{u u}\right)+\left(h_{y}^{2}+h_{u}^{2}\right)\left(h_{x x}+h_{v v}\right)\right. \\
& \left.-2\left(h_{x} h_{u}+h_{y} h_{v}\right)\left(h_{x u}+h_{y v}\right)-2\left(h_{x} h_{y}-h_{u} h_{v}\right)\left(h_{x y}-h_{u v}\right)\right] \\
\tilde{\mathbf{l}}_{21}= & -\left(h_{x}^{2}+h_{y}^{2}\right)\left(h_{u u}+h_{v v}\right)-\left(h_{u}^{2}+h_{v}^{2}\right)\left(h_{x x}+h_{y y}\right) \\
& +2\left(h_{x} h_{u}-h_{y} h_{v}\right)\left(h_{x u}-h_{y v}\right)+2\left(h_{x} h_{v}+h_{y} h_{u}\right)\left(h_{x v}+h_{y u}\right) \\
\tilde{\mathbf{l}}_{22}= & -\tilde{\mathbf{l}}_{11}, \tag{C.11}
\end{align*}
$$

with $h_{i}, h_{i j}$ is the derivative of the Hamiltonian $H$ with respect to the phase space coordinates and $q=|\nabla H|^{2}$.

## Appendix D

## Implementing evolution

## D. 1 Koopmania

The way in which time evolution acts on densities may be rephrased in the language of functional analysis, by introducing the Koopman operator, whose action on a phase space function $a(x)$ is to replace it by its downstream value time $t$ later, $a(x) \rightarrow a(x(t))$ evaluated at the trajectory point $x(t)$ :

$$
\begin{equation*}
\mathcal{K}^{t} a(x)=a\left(f^{t}(x)\right) . \tag{D.1}
\end{equation*}
$$

Observable $a(x)$ has no explicit time dependence; all time dependence is carried in its evaluation at $x(t)$ rather than at $x=x(0)$.

Suppose we are starting with an initial density of representative points $\rho(x)$ : then the average value of $a(x)$ evolves as

$$
\langle a\rangle(t)=\frac{1}{\left|\rho_{\mathcal{M}}\right|} \int_{\mathcal{M}} d x a\left(f^{t}(x)\right) \rho(x)=\frac{1}{\left|\rho_{\mathcal{M}}\right|} \int_{\mathcal{M}} d x\left[\mathcal{K}^{t} a(x)\right] \rho(x) .
$$

An alternative point of view (analogous to the shift from the Heisenberg to the Schrödinger picture in quantum mechanics) is to push dynamical effects into the density. In contrast to the Koopman operator which advances the trajectory by time $t$, the Perron-Frobenius operator (9.10) depends on the trajectory point time $t$ in the past, so the Perron-Frobenius operator is the adjoint of the Koopman operator

$$
\begin{equation*}
\int_{\mathcal{M}} d x\left[\mathcal{K}^{t} a(x)\right] \rho(x)=\int_{\mathcal{M}} d x a(x)\left[\mathcal{L}^{t} \rho(x)\right] . \tag{D.2}
\end{equation*}
$$

Checking this is an easy change of variables exercise. For finite dimensional deterministic invertible flows the Koopman operator (D.1) is simply the
inverse of the Perron-Frobenius operator (9.6), so in what follows we shall not distinguish the two. However, for infinite dimensional flows contracting forward in time and for stochastic flows such inverses do not exist, and there you need to be more careful.

The family of Koopman's operators $\left\{\mathcal{K}^{t}\right\}_{t \in \mathbb{R}_{+}}$forms a semigroup parametrized by time
(a) $\mathcal{K}^{0}=I$
(b) $\mathcal{K}^{t} \mathcal{K}^{t^{\prime}}=\mathcal{K}^{t+t^{\prime}} \quad t, t^{\prime} \geq 0 \quad$ (semigroup property),
with the generator of the semigroup, the generator of infinitesimal time translations defined by

$$
\mathcal{A}=\lim _{t \rightarrow 0^{+}} \frac{1}{t}\left(\mathcal{K}^{t}-I\right) .
$$

(If the flow is finite-dimensional and invertible, $\mathcal{A}$ is a generator of a group). The explicit form of $\mathcal{A}$ follows from expanding dynamical evolution up to first order, as in (2.4):

$$
\begin{equation*}
\mathcal{A} a(x)=\lim _{t \rightarrow 0^{+}} \frac{1}{t}\left(a\left(f^{t}(x)\right)-a(x)\right)=v_{i}(x) \partial_{i} a(x) . \tag{D.3}
\end{equation*}
$$

Of course, that is nothing but the definition of the time derivative, so the equation of motion for $a(x)$ is

$$
\begin{equation*}
\left(\frac{d}{d t}-\mathcal{A}\right) a(x)=0 . \tag{D.4}
\end{equation*}
$$

The finite time Koopman operator (D.1) can be formally expressed by exponentiating the time evolution generator $\mathcal{A}$ as

$$
\begin{equation*}
\mathcal{K}^{t}=e^{t \mathcal{A}} . \tag{D.5}
\end{equation*}
$$

The generator $\mathcal{A}$ looks very much like the generator of translations. Indeed, for a constant velocity field dynamical evolution is nothing but a translation by time $\times$ velocity:

$$
\begin{equation*}
e^{t v \frac{\partial}{\partial x}} a(x)=a(x+t v) . \tag{D.6}
\end{equation*}
$$

As we will not need to implement a computational formula for general $e^{t \mathcal{A}}$ in what follows, we relegate making sense of such operators to appendix D.2. Here we limit ourselves to a brief remark about the notion of "spectrum" of a linear operator.

The Koopman operator $\mathcal{K}$ acts multiplicatively in time, so it is reasonable to suppose that there exist constants $M>0, \beta \geq 0$ such that $\left\|\mathcal{K}^{t}\right\| \leq M e^{t \beta}$ for all $t \geq 0$. What does that mean? The operator norm is define in the same spirit in which we defined the matrix norms in sect. K.2: We are assuming that no value of $\mathcal{K}^{t} \rho(x)$ grows faster than exponentially for any choice of function $\rho(x)$, so that the fastest possible growth can be bounded by $e^{t \beta}$, a reasonable expectation in the light of the simplest example studied so far, the exact escape rate (10.20). If that is so, multiplying $\mathcal{K}^{t}$ by $e^{-t \beta}$ we construct a new operator $e^{-t \beta} \mathcal{K}^{t}=e^{t(\mathcal{A}-\beta)}$ which decays exponentially for large $t,\left\|e^{t(\mathcal{A}-\beta)}\right\| \leq M$. We say that $e^{-t \beta} \mathcal{K}^{t}$ is an element of a bounded semigroup with generator $\mathcal{A}-\beta I$. Given this bound, it follows by the Laplace transform

$$
\begin{equation*}
\int_{0}^{\infty} d t e^{-s t} \mathcal{K}^{t}=\frac{1}{s-\mathcal{A}}, \quad \operatorname{Re} s>\beta \tag{D.7}
\end{equation*}
$$

that the resolvent operator $(s-\mathcal{A})^{-1}$ is bounded ("resolvent" $=$ able to cause separation into constituents)

$$
\left\|\frac{1}{s-\mathcal{A}}\right\| \leq \int_{0}^{\infty} d t e^{-s t} M e^{t \beta}=\frac{M}{s-\beta}
$$

If one is interested in the spectrum of $\mathcal{K}$, as we will be, the resolvent operator is a natural object to study. The main lesson of this brief aside is that for the continuous time flows the Laplace transform is the tool that brings down the generator in (9.27) into the resolvent form (9.29) and enables us to study its spectrum.

## D. 2 Implementing evolution

(R. Artuso and P. Cvitanović)

We now come back to the semigroup of operators $\mathcal{K}^{t}$. We have introduced the generator of the semigroup (9.25) as

$$
\mathcal{A}=\left.\frac{d}{d t} \mathcal{K}^{t}\right|_{t=0}
$$

If we now take the derivative at arbitrary times we get

$$
\begin{aligned}
\left(\frac{d}{d t} \mathcal{K}^{t} \psi\right)(x) & =\lim _{\eta \rightarrow 0} \frac{\psi\left(f^{t+\eta}(x)\right)-\psi\left(f^{t}(x)\right)}{\eta} \\
& =\left.v_{i}\left(f^{t}(x)\right) \frac{\partial}{\partial \tilde{x}_{i}} \psi(\tilde{x})\right|_{\tilde{x}=f^{t}(x)} \\
& =\left(\mathcal{K}^{t} \mathcal{A} \psi\right)(x)
\end{aligned}
$$

which can be formally integrated like an ordinary differential equation yield-
D. 1
page 665 ing

$$
\begin{equation*}
\mathcal{K}^{t}=e^{t \mathcal{A}} \tag{D.8}
\end{equation*}
$$

This guarantees that the Laplace transform manipulations in sect. 9.4 are correct. Though the formal expression of the semigroup (D.8) is quite simple one has to take care in implementing its action. If we express the exponential through the power series

$$
\begin{equation*}
\mathcal{K}^{t}=\sum_{k=0}^{\infty} \frac{t^{k}}{k!} \mathcal{A}^{k} \tag{D.9}
\end{equation*}
$$

we encounter the problem that the infinitesimal generator (9.25) contains non-commuting pieces, that is, there are $i, j$ combinations for which the commutator does not satisfy

$$
\left[\frac{\partial}{\partial x_{i}}, v_{j}(x)\right]=0
$$

To derive a more useful representation, we follow the strategy used for finitedimensional matrix operators in sects. 4.2 and 4.3 and use the semigroup property to write

$$
\mathcal{K}^{t}=\prod_{m=1}^{t / \delta \tau} \mathcal{K}^{\delta \tau}
$$

as the starting point for a discretized approximation to the continuous time dynamics, with time step $\delta \tau$. Omitting terms from the second order onwards in the expansion of $\mathcal{K}^{\delta \tau}$ yields an error of order $O\left(\delta \tau^{2}\right)$. This might be acceptable if the time step $\delta \tau$ is sufficiently small. In practice we write the Euler product

$$
\begin{equation*}
\mathcal{K}^{t}=\prod_{m=1}^{t / \delta \tau}\left(1+\delta \tau \mathcal{A}_{(m)}\right)+O\left(\delta \tau^{2}\right) \tag{D.10}
\end{equation*}
$$

where

$$
\left(\mathcal{A}_{(m)} \psi\right)(x)=\left.v_{i}\left(f^{m \delta \tau}(x)\right) \frac{\partial \psi}{\partial \tilde{x}_{i}}\right|_{\tilde{x}=f^{m \delta \tau}(x)}
$$

As far as the $x$ dependence is concerned, $e^{\delta \tau \mathcal{A}_{i}}$ acts as

$$
e^{\delta \tau \mathcal{A}_{i}}\left\{\begin{array}{c}
x_{1}  \tag{D.11}\\
\cdot \\
x_{i} \\
x_{d}
\end{array}\right\} \rightarrow\left\{\begin{array}{c}
x_{1} \\
\cdot \\
x_{i}+\delta \tau v_{i}(x) \\
x_{d}
\end{array}\right\}
$$

We see that the product form (D.10) of the operator is nothing else but a prescription for finite time step integration of the equations of motion - in this case the simplest Euler type integrator which advances the trajectory by $\delta \tau \times$ velocity at each time step.

## D.2.1 A symplectic integrator

- ${ }^{3}$
The procedure we described above is only a starting point for more sophisticated approximations. As an example on how to get a sharper bound on the error term consider the Hamiltonian flow $\mathcal{A}=\mathcal{B}+\mathcal{C}, \mathcal{B}=$ $p_{i} \frac{\partial}{\partial q_{i}}, \mathcal{C}=-\partial_{i} V(q) \frac{\partial}{\partial p_{i}}$. Clearly the potential and the kinetic parts do not commute. We make sense of the formal solution (D.10) by spliting it into infinitesimal steps and keeping terms up to $\delta \tau^{2}$ in


$$
\begin{equation*}
\mathcal{K}^{\delta \tau}=\hat{\mathcal{K}}^{\delta \tau}+\frac{1}{24}(\delta \tau)^{3}[\mathcal{B}+2 \mathcal{C},[\mathcal{B}, \mathcal{C}]]+\cdots, \tag{D.12}
\end{equation*}
$$

where

$$
\begin{equation*}
\hat{\mathcal{K}}^{\delta \tau}=e^{\frac{1}{2} \delta \tau \mathcal{B}} e^{\delta \tau \mathcal{C}} e^{\frac{1}{2} \delta \tau \mathcal{B}} \tag{D.13}
\end{equation*}
$$

The approximate infinitesimal Liouville operator $\hat{\mathcal{K}}^{\delta \tau}$ is of the form that now generates evolution as a sequence of mappings induced by (9.28), a free flight by $\frac{1}{2} \delta \tau \mathcal{B}$, scattering by $\delta \tau \partial V\left(q^{\prime}\right)$, followed again by $\frac{1}{2} \delta \tau \mathcal{B}$ free flight:

$$
\begin{align*}
e^{\frac{1}{2} \delta \tau \mathcal{B}}\left\{\begin{array}{l}
q \\
p
\end{array}\right\} & \rightarrow\left\{\begin{array}{l}
q^{\prime} \\
p^{\prime}
\end{array}\right\}=\left\{\begin{array}{c}
q-\frac{\delta \tau}{2} p \\
p
\end{array}\right\} \\
e^{\delta \tau \mathcal{C}}\left\{\begin{array}{l}
q^{\prime} \\
p^{\prime}
\end{array}\right\} & \rightarrow\left\{\begin{array}{l}
q^{\prime \prime} \\
p^{\prime \prime}
\end{array}\right\}=\left\{\begin{array}{c}
q^{\prime} \\
p^{\prime}+\delta \tau \partial V\left(q^{\prime}\right)
\end{array}\right\} \\
e^{\frac{1}{2} \delta \tau \mathcal{B}}\left\{\begin{array}{l}
q^{\prime \prime} \\
p^{\prime \prime}
\end{array}\right\} & \rightarrow\left\{\begin{array}{l}
q^{\prime \prime \prime} \\
p^{\prime \prime \prime}
\end{array}\right\}=\left\{\begin{array}{c}
q^{\prime}-\frac{\delta \tau}{2} p^{\prime \prime} \\
p^{\prime \prime}
\end{array}\right\} \tag{D.14}
\end{align*}
$$

Collecting the terms we obtain an integration rule for this type of symplectic flow which is better than the straight Euler integration (D.11) as it is accurate up to order $\delta \tau^{2}$ :

$$
\begin{align*}
q_{n+1} & =q_{n}-\delta \tau p_{n}-\frac{(\delta \tau)^{2}}{2} \partial V\left(q_{n}-\delta \tau p_{n} / 2\right) \\
p_{n+1} & =p_{n}+\delta \tau \partial V\left(q_{n}-\delta \tau p_{n} / 2\right) \tag{D.15}
\end{align*}
$$

The Jacobian matrix of one integration step is given by

$$
\mathbf{J}=\left(\begin{array}{cc}
1 & -\delta \tau / 2  \tag{D.16}\\
0 & 1
\end{array}\right)\left(\begin{array}{cc}
1 & 0 \\
\delta \tau \partial V\left(q^{\prime}\right) & 1
\end{array}\right)\left(\begin{array}{cc}
1 & -\delta \tau / 2 \\
0 & 1
\end{array}\right)
$$

Note that the billiard flow (6.11) is an example of such symplectic integrator. In that case the free flight is interupted by instantaneous wall reflections, and can be integrated out.

## Commentary

Remark D. 1 Koopman operators. The "Heisenberg picture" in dynamical system theory has been introduced by Koopman refs. [D.1, D.2], see also ref. [9.6]. Inspired by the contemporary advances in quantum mechanics, Koopman [D.1] observed in 1931 that $\mathcal{K}^{t}$ is unitary on $L^{2}(\mu)$ Hilbert spaces. The Liouville/Koopman operator is the classical analogue of the quantum evolution operator - the kernel of $\mathcal{L}^{t}(y, x)$ introduced in (9.15) (see also sect. 10.2) is the analogue of the Green's function. The relation between the spectrum of the Koopman operator and classical ergodicity was formalized by von Neumann [D.2]. We shall not use Hilbert spaces here and the operators that we shall study will not be unitary. For a discussion of the relation between the Perron-Frobenius operators and the Koopman operators for finite dimensional deterministic invertible flows, infinite dimensional contracting flows, and stochastic flows, see Lasota-Mackey [9.6] and Gaspard [1.4].

Remark D. 2 Symplectic integration. The reviews [D.5] and [D.6] offer a good starting point for exploring the symplectic integrators literature. For a higher order integrators of type (D.13), check ref. [D.7].

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## Exercises

Exercise D. 1 Exponential form of semigroup elements. Check that the Koopman operator and the evolution generator commute, $\mathcal{K}^{t} \mathcal{A}=\mathcal{A K}^{t}$, by considering the action of both operators on an arbitrary phase space function $a(x)$.

Exercise D. 2 Symplectic volume preservation. Check that the sequence of mappings (D.14) is volume preserving, $\operatorname{det} \hat{U}=1$.

Exercise D. 3 Noncommutativity. Check that the commutators in (D.12) are not vanishing by showing that

$$
[\mathcal{B}, \mathcal{C}]=-p\left(V^{\prime \prime} \frac{\partial}{\partial p}-V^{\prime} \frac{\partial}{\partial q}\right)
$$

Exercise D. 4 Symplectic leapfrog integrator. Implement (D.15) for 2dimensional Hamiltonian flows; compare it with Runge-Kutta integrator by integrating trajectories in some (chaotic) Hamiltonian flow.

## Appendix E

## Symbolic dynamics techniques

The kneading theory for unimodal mappings is developed in sect. E.1. The prime factorization for dynamical itineraries of sect. E. 2 illustrates the sense in which prime cycles are "prime" - the product structure of zeta functions is a consequence of the unique factorization property of symbol sequences.

## E. 1 Topological zeta functions for infinite subshifts

(P. Dahlqvist)


The Markov graph methods outlined in chapter 11 are well suited for symbolic dynamics of finite subshift type. A sequence of well defined rules leads to the answer, the topological zeta function, which turns out to be a polynomial. For infinite subshifts one would have to go through an infinite sequence of graph constructions and it is of course very difficult to make any asymptotic statements about the outcome. Luckily, for some simple systems the goal can be reached by much simpler means. This is the case for unimodal maps.

We will restrict our attention to the topological zeta function for unimodal maps with one external parameter $f_{\Lambda}(x)=\Lambda g(x)$. As usual, symbolic dynamics is introduced by mapping a time series $\ldots x_{i-1} x_{i} x_{i+1} \ldots$ onto a sequence of symbols $\ldots s_{i-1} s_{i} s_{i+1} \ldots$ where

$$
\begin{align*}
s_{i}=0 & x_{i}<x_{c} \\
s_{i}=C & x_{i}=x_{c} \\
s_{i}=1 & x_{i}>x_{c} \tag{E.1}
\end{align*}
$$

and $x_{c}$ is the critical point of the map (that is, maximum of $g$ ). In addition to the usual binary alphabet we have added a symbol $C$ for the critical point. The kneading sequence $K_{\Lambda}$ is the itinerary of the critical point. The crucial observation is that no periodic orbit can have a topological

| $I(C)$ | $\zeta_{\text {top }}^{-1}(z) /(1-z)$ | $I(C)$ | $\zeta_{\text {top }}^{-1}(z) /(1-z)$ |
| :--- | :--- | :--- | :--- |
| $1 C$ |  | $1001 C$ |  |
| $101 C$ |  | $100111 C$ |  |
| $1011101 C$ |  | $10011 C$ |  |
| $H^{\infty}(1)$ | $\prod_{n=0}^{\infty}\left(1-z^{2^{n}}\right)$ | $100110 C$ |  |
| $1011 C$ |  | $100 C$ |  |
| $101111 C$ |  | $100010 C$ |  |
| $101 \infty$ | $\left(1-2 z^{2}\right) /(1+z)$ | $10001 C$ |  |
| $1011111 C$ |  | $100011 C$ |  |
| $101111 C$ |  | $1000 C$ |  |
| $1011 C$ |  | $100001 C$ |  |
| $101101 C$ |  | $10000 C$ |  |
| $10 C$ | $\left(1-z-z^{2}\right)$ | $100000 C$ |  |
| $10010 C$ |  | $10^{\infty}$ | $(1-2 z) /(1-z)$ |
| $100101 C$ |  |  |  |

Table E.1: All ordered kneading sequences up to length seven, as well as some longer kneading sequences. Harmonic extension $H^{\infty}(1)$ is defined below.
coordinate (see sect. E.1.1) beyond that of the kneading sequence. The kneading sequence thus inserts a border in the list of periodic orbits (ordered according to maximal topological coordinate), cycles up to this limit are allowed, all beyond are pruned. All unimodal maps (obeying some further constraints) with the same kneading sequence thus have the same set of periodic orbitsand the same topological zeta function. The topological coordinate of the kneading sequence increases with increasing $\Lambda$.

The kneading sequence can be of one of three types

1. It maps to the critical point again, after $n$ iterations. If so, we adopt the convention to terminate the kneading sequence with a $C$, and refer to the kneading sequence as finite.
2. Preperiodic, that is, it is infinite but with a periodic tail.
3. Aperiodic.

As an archetype unimodal map we will choose the tent map

$$
x \mapsto f(x)=\left\{\begin{array}{ll}
\Lambda x & x \in[0,1 / 2]  \tag{E.2}\\
\Lambda(1-x) & x \in(1 / 2,1]
\end{array},\right.
$$

where the parameter $\Lambda \in(1,2]$. The topological entropy is $h=\log \Lambda$. This follows from the fact any trajectory of the map is bounded, the escape rate is strictly zero, and so the dynamical zeta function

$$
1 / \zeta(z)=\prod_{p}\left(1-\frac{z^{n_{p}}}{\left|\Lambda_{p}\right|}\right)=\prod_{p}\left(1-\left(\frac{z}{\Lambda}\right)^{n_{p}}\right)=1 / \zeta_{\mathrm{top}}(z / \Lambda)
$$

has its leading zero at $z=1$.

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The set of periodic points of the tent map is countable. A consequence of this fact is that the set of parameter values for which the kneading sequence is periodic or preperiodic are countable and thus of measure zero and consequently the kneading sequence is aperiodic for almost all $\Lambda$. For general unimodal maps the corresponding statement is that the kneading sequence is aperiodic for almost all topological entropies.

For a given periodic kneading sequence of period $n, \underline{K}_{\Lambda}=P C=$ $s_{1} s_{2} \ldots s_{n-1} C$ there is a simple expansion for the topological zeta function. Then the expanded zeta function is a polynomial of degree $n$

$$
\begin{equation*}
1 / \zeta_{\mathrm{top}}(z)=\prod_{p}\left(1-z_{p}^{n}\right)=(1-z) \sum_{i=0}^{n-1} a_{i} z^{i}, \quad a_{i}=\prod_{j=1}^{i}(-1)^{s_{j}} \tag{E.3}
\end{equation*}
$$

and $a_{0}=1$.
Aperiodic and preperiodic kneading sequences are accounted for by simply replacing $n$ by $\infty$.

Example. Consider as an example the kneading sequence $K_{\Lambda}=10 C$. From (E.3) we get the topological zeta function $1 / \zeta_{\text {top }}(z)=(1-z)(1-z-$ $z^{2}$ ), see table E.1. This can also be realized by redefining the alphabet. The only forbidden subsequence is 100 . All allowed periodic orbits, except $\overline{0}$, can can be built from a alphabet with letters $\underline{10}$ and $\underline{1}$. We write this alphabet as $\{\underline{10}, \underline{1} ; \overline{0}\}$, yielding the topological zeta function $1 / \zeta_{\text {top }}(z)=(1-z)(1-$ $\left.z-z^{2}\right)$. The leading zero is the inverse golden mean $z_{0}=(\sqrt{5}-1) / 2$.

Example. As another example we consider the preperiodic kneading sequence $K_{\Lambda}=101^{\infty}$. From (E.3) we get the topological zeta function $1 / \zeta_{\text {top }}(z)=(1-z)\left(1-2 z^{2}\right) /(1+z)$, see table E.1. This can again be realized by redefining the alphabet. There are now an infinite number of forbidden subsequences, namely $101^{2 n} 0$ where $n \geq 0$. These pruning rules are respected by the alphabet $\left\{\underline{01^{2 n+1}} ; \overline{1}, \overline{0}\right\}$, yielding the topological zeta function above. The pole in the zeta function $\zeta_{\text {top }}^{-1}(z)$ is a consequence of the infinite alphabet.

An important consequence of (E.3) is that the sequence $\left\{a_{i}\right\}$ has a periodic tail if and only if the kneading sequence has one (however, their period may differ by a factor of two). We know already that the kneading sequence is aperiodic for almost all $\Lambda$.

The analytic structure of the function represented by the infinite series $\sum a_{i} z_{i}$ with unity as radius of convergence, depends on whether the tail of $\left\{a_{i}\right\}$ is periodic or not. If the period of the tail is $N$ we can write

$$
1 / \zeta_{\text {top }}(z)=p(z)+q(z)\left(1+z^{N}+z^{2 N} \ldots\right)=p(z)+\frac{q(z)}{1-z^{N}}
$$

for some polynomials $p(z)$ and $q(z)$. The result is a set of poles spread out along the unit circle. This applies to the preperiodic case. An aperiodic sequence of coefficients would formally correspond to infinite $N$ and it is natural to assume that the singularities will fill the unit circle. There is indeed a theorem ensuring that this is the case [12.70], provided the $a_{i}$ 's can only take on a finite number of values. The unit circle becomes a natural boundary, already apparent in a finite polynomial approximations to the topological zeta function, as in figure 13.4. A function with a natural boundary lacks an analytic continuation outside it.

To conclude: The topological zeta function $1 / \zeta_{\text {top }}$ for unimodal maps has the unit circle as a natural boundary for almost all topological entropies and for the tent map (E.2), for almost all $\Lambda$.

Let us now focus on the relation between the analytic structure of the topological zeta function and the number of periodic orbits, or rather (13.6), the number $N_{n}$ of fixed points of $f^{n}(x)$. The trace formula is (see sect. 13.4)

$$
N_{n}=\operatorname{tr} T^{n}=\frac{1}{2 \pi i} \oint_{\gamma_{r}} d z z^{-n} \frac{d}{d z} \log \zeta_{\text {top }}^{-1}
$$

where $\gamma_{r}$ is a (circular) contour encircling the origin $z=0$ in clockwise direction. Residue calculus turns this into a sum over zeros $z_{0}$ and poles $z_{p}$ of $\zeta_{\text {top }}^{-1}$

$$
N_{n}=\sum_{z_{0}: r<\left|z_{0}\right|<R} z_{0}^{-n}-\sum_{z_{p}: r<\left|z_{p}\right|<R} z_{0}^{-n}+\frac{1}{2 \pi i} \oint_{\gamma_{R}} d z z^{-n} \frac{d}{d z} \log \zeta_{\text {top }}^{-1}
$$

and a contribution from a large circle $\gamma_{R}$. For meromorphic topological zeta functions one may let $R \rightarrow \infty$ with vanishing contribution from $\gamma_{R}$, and $N_{n}$ will be a sum of exponentials.

The leading zero is associated with the topological entropy, as discussed in chapter 13.

We have also seen that for preperiodic kneading there will be poles on the unit circle.

To appreciate the role of natural boundaries we will consider a (very) special example. Cascades of period doublings is a central concept for the description of unimodal maps. This motivates a close study of the function

$$
\begin{equation*}
\Xi(z)=\prod_{n=0}^{\infty}\left(1-z^{2^{n}}\right) \tag{E.4}
\end{equation*}
$$

This function will appear again when we derive (E.3).
The expansion of $\Xi(z)$ begins as $\Xi(z)=1-z-z^{2}+z^{3}-z^{4}+z^{5} \ldots$. The radius of convergence is obviously unity. The simple rule governing the

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expansion will effectively prohibit any periodicity among the coefficients making the unit circle a natural boundary.

It is easy to see that $\Xi(z)=0$ if $z=\exp \left(2 \pi m / 2^{n}\right)$ for any integer $m$ and $n$. (Strictly speaking we mean that $\Xi(z) \rightarrow 0$ when $z \rightarrow \exp \left(2 \pi m / 2^{n}\right)$ from inside). Consequently, zeros are dense on the unit circle. One can also show that singular points are dense on the unit circle, for instance $|\Xi(z)| \rightarrow \infty$ when $z \rightarrow \exp \left(2 \pi m / 3^{n}\right)$ for any integer $m$ and $n$.

As an example, the topological zeta function at the accumulation point of the first Feigenbaum cascade is $\zeta_{\text {top }}^{-1}(z)=(1-z) \Xi(z)$. Then $N_{n}=$ $2^{l+1}$ if $n=2^{l}$, otherwise $N_{n}=0$. The growth rate in the number of cycles is anything but exponential. It is clear that $N_{n}$ cannot be a sum of exponentials, the contour $\gamma_{R}$ cannot be pushed away to infinity, $R$ is restricted to $R \leq 1$ and $N_{n}$ is entirely determined by $\int_{\gamma_{R}}$ which picks up its contribution from the natural boundary.

We have so far studied the analytic structure for some special cases and we know that the unit circle is a natural boundary for almost all $\Lambda$. But how does it look out there in the complex plane for some typical parameter values? To explore that we will imagine a journey from the origin $z=$ 0 out towards the unit circle. While traveling we let the parameter $\Lambda$ change slowly. The trip will have a distinct science fiction flavor. The first zero we encounter is the one connected to the topological entropy. Obviously it moves smoothly and slowly. When we move outward to the unit circle we encounter zeros in increasing densities. The closer to the unit circle they are, the wilder and stranger they move. They move from and back to the horizon, where they are created and destroyed through bizarre bifurcations. For some special values of the parameter the unit circle suddenly gets transparent and and we get (infinitely) short glimpses of another world beyond the horizon.

We end this section by deriving eqs (E.5) and (E.6). The impenetrable prose is hopefully explained by the accompanying tables.

We know one thing from chapter 11, namely for that finite kneading sequence of length $n$ the topological polynomial is of degree $n$. The graph contains a node which is connected to itself only via the symbol 0 . This implies that a factor $(1-z)$ may be factored out and $\zeta_{\text {top }}(z)=(1-z) \sum_{i=0}^{n-1} a_{i} z^{i}$. The problem is to find the coefficients $a_{i}$.

The ordered list of (finite) kneading sequences table E. 1 and the ordered list of periodic orbits (on maximal form) are intimately related. In table E. 2 we indicate how they are nested during a period doubling cascade. Every finite kneading sequence $P C$ is bracketed by two periodic orbits, $\overline{P 1}$ and $\overline{P 0}$. We have $\overline{P 1}<P C<\overline{P 0}$ if $P$ contains an odd number of 1's, and $\overline{P 0}<P C<\overline{P 1}$ otherwise. From now on we will assume that $P$ contains an odd number of 1's. The other case can be worked out in complete analogy. The first and second harmonic of $P C$ are displayed in table E.2. The periodic orbit $\overline{P 1}$ (and the corresponding infinite kneading sequence) is sometimes referred to as the antiharmonic extension of $P C$ (denoted

| periodic orbits | finite kneading sequences |
| :--- | :--- |
| $\overline{P 1}=A^{\infty}(P)$ | $P C$ |
| $\overline{P 0}$ | $P 0 P C$ |
| $\overline{P 0 P 1}$ | $P 0 P 1 P 0 P C$ |
| $\downarrow$ | $\downarrow$ |
| $H^{\infty}(P)$ | $H^{\infty}(P)$ |

Table E.2: Relation between periodic orbits and finite kneading sequences in a harmonic cascade. The string $P$ is assumed to contain an odd number of 1's.
$\left.A^{\infty}(P)\right)$ and the accumulation point of the cascade is called the harmonic extension of $P C$ [11.8] (denoted $\left.H^{\infty}(P)\right)$.

A central result is the fact that a period doubling cascade of $P C$ is not interfered by any other sequence. Another way to express this is that a kneading sequence $P C$ and its harmonic are adjacent in the list of kneading sequences to any order.

| $I(C)$ |  | $\zeta_{\text {top }}^{-1}(z) /(1-z)$ |
| ---: | :--- | :--- |
| $P_{1}$ | $=100 C$ | $1-z-z^{2}-z^{3}$ |
| $H^{\infty}\left(P_{1}\right)$ | $=10001001100 \ldots$ | $1-z-z^{2}-z^{3}-z^{4}+z^{5}+z^{6}+z^{7}-z^{8} \ldots$ |
| $P^{\prime}$ | $=10001 C$ | $1-z-z^{2}-z^{3}-z^{4}+z^{5}$ |
| $A^{\infty}\left(P_{2}\right)$ | $=1000110001 \ldots$ | $1-z-z^{2}-z^{3}-z^{4}+z^{5}-z^{6}-z^{7}-z^{8} \ldots$ |
| $P_{2}$ | $=1000 C$ | $1-z-z^{2}-z^{3}-z^{4}$ |

Table E.3: Example of a step in the iterative construction of the list of kneading sequences $P C$.

Table E. 3 illustrates another central result in the combinatorics of kneading sequences. We suppose that $P_{1} C$ and $P_{2} C$ are neighbors in the list of order 5 (meaning that the shortest finite kneading sequence $P^{\prime} C$ between $P_{1} C$ and $P_{2} C$ is longer than 5.) The important result is that $P^{\prime}$ (of length $\left.n^{\prime}=6\right)$ has to coincide with the first $n^{\prime}-1$ letters of both $H^{\infty}\left(P_{1}\right)$ and $A^{\infty}\left(P_{2}\right)$. This is exemplified in the left column of table E.3. This fact makes it possible to generate the list of kneading sequences in an iterative way.

The zeta function at the accumulation point $H^{\infty}\left(P_{1}\right)$ is

$$
\begin{equation*}
\zeta_{P_{1}}^{-1}(z) \Xi\left(z^{n_{1}}\right) \tag{E.5}
\end{equation*}
$$

and just before $A^{\infty}\left(P_{2}\right)$

$$
\begin{equation*}
\zeta_{P_{2}}^{-1}(z) /\left(1-z^{n_{2}}\right) \tag{E.6}
\end{equation*}
$$

A short calculation shows that this is exactly what one would obtain by applying (E.3) to the antiharmonic and harmonic extensions directly, provided that it applies to $\zeta_{P_{1}}^{-1}(z)$ and $\zeta_{P_{2}}^{-1}(z)$. This is the key observation.

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Recall now the product representation of the zeta function $\zeta^{-1}=\prod_{p}(1-$ $\left.z^{n_{p}}\right)$. We will now make use of the fact that the zeta function associated with $P^{\prime} C$ is a polynomial of order $n^{\prime}$. There is no periodic orbit of length shorter than $n^{\prime}+1$ between $H^{\infty}\left(P_{1}\right)$ and $A^{\infty}\left(P_{2}\right)$. It thus follows that the coefficients of this polynomial coincides with those of (E.5) and (E.6), see Table E.3. We can thus conclude that our rule can be applied directly to $P^{\prime} C$.

This can be used as an induction step in proving that the rule can be applied to every finite and infinite kneading sequences.

Remark E. 1 How to prove things. The explicit relation between the kneading sequence and the coefficients of the topological zeta function is not commonly seen in the literature. The result can proven by combining some theorems of Milnor and Thurston [12.15]. That approach is hardly instructive in the present context. Our derivation was inspired by Metropolis, Stein and Stein classical paper [11.8]. For further detail, consult [12.69].

## E.1.1 Periodic orbits of unimodal maps

A periodic point (or a cycle point) $x_{i}$ belonging to a cycle of period $n$ is a real solution of

$$
\begin{equation*}
f^{n}\left(x_{i}\right)=f\left(f\left(\ldots f\left(x_{i}\right) \ldots\right)\right)=x_{i}, \quad i=0,1,2, \ldots, n-1 \tag{E.7}
\end{equation*}
$$

The $n$th iterate of a unimodal map crosses the diagonal at most $2^{n}$ times. Similarly, the backward and the forward Smale horseshoes intersect at most $2^{n}$ times, and therefore there will be $2^{n}$ or fewer periodic points of length $n$. A cycle of length $n$ corresponds to an infinite repetition of a length $n$ symbol string, customarily indicated by a line over the string:

$$
S=\left(s_{1} s_{2} s_{3} \ldots s_{n}\right)^{\infty}=\overline{s_{1} s_{2} s_{3} \ldots s_{n}} .
$$

If $\overline{s_{1} s_{2} \ldots s_{n}}$ is the symbol string associated with $x_{0}$, its cyclic permutation $\overline{s_{k} s_{k+1} \ldots s_{n} s_{1} \ldots s_{k-1}}$ corresponds to the point $x_{k-1}$ in the same cycle. A cycle $p$ is called prime if its itinerary $S$ cannot be written as a repetition of a shorter block $S^{\prime}$.

Each cycle yields $n$ rational values of $\gamma$. The repeating string $s_{1}, s_{2}, \ldots s_{n}$ contains an odd number " 1 "s, the string of well ordered symbols $w_{1} w_{2} \ldots w_{n}$ has to be of the double length before it repeats itself. The value $\gamma$ is a geometrical sum which we can write as the finite sum

$$
\gamma\left(\overline{s_{1} s_{2} \ldots s_{n}}\right)=\frac{2^{2 n}}{2^{2 n}-1} \sum_{t=1}^{2 n} w_{t} / 2^{t}
$$

Using this we can calculate the $\hat{\gamma}(S)$ for all short cycles.
Here we give explicit formulas for the topological coordinate of a periodic point, given its itinerary. For the purpose of what follows it is convenient to compactify the itineraries by replacing the binary alphabet $s_{i}=\{0,1\}$ by the infinite alphabet

$$
\begin{equation*}
\left\{a_{1}, a_{2}, a_{3}, a_{4}, \cdots ; \overline{0}\right\}=\{1,10,100,1000, \ldots ; \overline{0}\} \tag{E.8}
\end{equation*}
$$

In this notation the itinerary $S=a_{i} a_{j} a_{k} a_{l} \cdots$ and the corresponding topological coordinate (??) are related by $\gamma(S)=.1^{i} 0^{j} 1^{k} 0^{l} \cdots$. For example:

$$
\begin{aligned}
& S=111011101001000 \ldots=a_{1} a_{1} a_{2} a_{1} a_{1} a_{2} a_{3} a_{4} \ldots \\
& \gamma(S)=.101101001110000 \ldots=.1^{1} 0^{1} 1^{2} 0^{1} 1^{1} 0^{2} 1^{3} 0^{4} \ldots
\end{aligned}
$$

Cycle points whose itineraries start with $w_{1}=w_{2}=\ldots=w_{i}=0, w_{i+1}=1$ remain on the left branch of the tent map for $i$ iterations, and satisfy $\gamma(0 \ldots 0 S)=\gamma(S) / 2^{i}$.

A periodic point (or a cycle point) $x_{i}$ belonging to a cycle of period $n$ is a real solution of

$$
\begin{equation*}
f^{n}\left(x_{i}\right)=f\left(f\left(\ldots f\left(x_{i}\right) \ldots\right)\right)=x_{i}, \quad i=0,1,2, \ldots, n-1 . \tag{E.9}
\end{equation*}
$$

The $n$th iterate of a unimodal map has at most $2^{n}$ monotone segments, and therefore there will be $2^{n}$ or fewer periodic points of length $n$. A periodic orbit of length $n$ corresponds to an infinite repetition of a length $n$ symbol string, customarily indicated by a line over the string:

$$
S=\left(s_{1} s_{2} s_{3} \ldots s_{n}\right)^{\infty}=\overline{s_{1} s_{2} s_{3} \ldots s_{n}} .
$$

As all itineraries are infinite, we shall adopt convention that a finite string itinerary $S=s_{1} s_{2} s_{3} \ldots s_{n}$ stands for infinite repetition of a finite block, and routinely omit the overline. If $\overline{s_{1} s_{2} \ldots s_{n}}$ is the symbol string associated with $x_{0}$, its cyclic permutation $\overline{s_{k} s_{k+1} \ldots s_{n} s_{1} \ldots s_{k-1}}$ corresponds to the point $x_{k-1}$ in the same cycle. A periodic orbit $p$ is called prime if its itinerary $S$ cannot be written as a repetition of a shorter block $S^{\prime}$.

Periodic points correspond to rational values of $\gamma$, but we have to distinguish even and odd cycles. The even (odd) cycles contain even (odd) number of $a_{i}$ in the repeating block, with periodic points given by

$$
\gamma\left(a_{i} a_{j} \cdots a_{k} a_{\ell}\right)=\left\{\begin{array}{lc}
\frac{2^{n}}{2^{n}-1} \cdot 1^{i} 0^{j} \cdots 1^{k} & \text { even }  \tag{E.10}\\
\frac{1}{2^{n}+1}\left(1+2^{n} \times \cdot 1^{i} 0^{j} \cdots 1^{\ell}\right) & \text { odd }
\end{array},\right.
$$

where $n=i+j+\cdots+k+\ell$ is the cycle period. The maximal value cycle point is given by the cyclic permutation of $S$ with the largest $a_{i}$ as the first
symbol, followed by the smallest available $a_{j}$ as the next symbol, and so on. For example:

$$
\begin{aligned}
& \hat{\gamma}(1)=\gamma\left(a_{1}\right)=.10101 \ldots=\overline{10}=2 / 3 \\
& \hat{\gamma}(10)=\gamma\left(a_{2}\right)=.1^{2} 0^{2} \ldots=\overline{1100}=4 / 5 \\
& \hat{\gamma}(100)=\gamma\left(a_{3}\right)=.1^{3} 0^{3} \ldots=\overline{. \overline{111000}}=8 / 9 \\
& \hat{\gamma}(101)=\gamma\left(a_{2} a_{1}\right)=.1^{2} 0^{1} \ldots=\overline{110}=6 / 7
\end{aligned}
$$

An example of a cycle where only the third symbol determines the maximal value cycle point is

$$
\hat{\gamma}(1101110)=\gamma\left(a_{2} a_{1} a_{2} a_{1} a_{1}\right)=\overline{11011010010010}=100 / 129 .
$$

Maximal values of all cycles up to length 5 are given in table!?

## E. 2 Prime factorization for dynamical itineraries

The Möbius function is not only a number-theoretic function, but can be used to manipulate ordered sets of noncommuting objects such as symbol strings. Let $\mathcal{P}=\left\{p_{1}, p_{2}, p_{3}, \cdots\right\}$ be an ordered set of prime strings, and

$$
\mathcal{N}=\{n\}=\left\{p_{1}^{k_{1}} p_{2}^{k_{2}} p_{3}^{k_{3}} \cdots p_{j}^{k_{j}}\right\}
$$

$j \in \mathbb{N}, k_{i} \in \mathbb{Z}_{+}$, be the set of all strings $n$ obtained by the ordered concatenation of the "primes" $p_{i}$. By construction, every string $n$ has a unique prime factorization. We say that a string has a divisor $d$ if it contains $d$ as a substring, and define the string division $n / d$ as $n$ with the substring $d$ deleted. Now we can do things like this: defining $t_{n}:=t_{p_{1}}^{k_{1}} t_{p_{2}}^{k_{2}} \cdots t_{p_{j}}^{k_{j}}$ we can write the inverse dynamical zeta function (18.2) as

$$
\prod_{p}\left(1-t_{p}\right)=\sum_{n} \mu(n) t_{n}
$$

and, if we care (we do in the case of the Riemann zeta function), the dynamical zeta function as .

$$
\begin{equation*}
\prod_{p} \frac{1}{1-t_{p}}=\sum_{n} t_{n} \tag{E.11}
\end{equation*}
$$

A striking aspect of this formula is its resemblance to the factorization of natural numbers into primes: the relation of the cycle expansion (E.11)

| factors | string | factors | string | factors | string | factors | string |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $p_{1}$ | 0 | $p_{1}^{4}$ | 0000 | $p_{1}^{5}$ | 00000 | $p_{1}^{2} p_{5}$ | 00101 |
| $p_{2}$ | 1 | $p_{1}^{3} p_{2}$ | 0001 | $p_{1}^{4} p_{2}$ | 00001 | $p_{1} p_{2} p_{5}$ | 01101 |
|  |  | $p_{1}^{2} p_{2}^{2}$ | 0011 | $p_{1}^{3} p_{2}^{2}$ | 00011 | $p_{2}^{2} p_{5}$ | 11101 |
| $p_{1}^{2}$ | 00 | $p_{1} p_{2}^{3}$ | 0111 | $p_{1}^{2} p_{2}^{3}$ | 00111 | $p_{3} p_{5}$ | 10101 |
| $p_{1} p_{2}$ | 01 | $p_{2}^{4}$ | 1111 | $p_{1} p_{2}^{4}$ | 01111 | $p_{1} p_{6}$ | 01000 |
| $p_{2}^{2}$ | 11 | $p_{1}^{2} p_{3}$ | 0010 | $p_{2}^{5}$ | 11111 | $p_{2} p_{6}$ | 11000 |
| $p_{3}$ | 10 | $p_{1} p_{2} p_{3}$ | 0110 | $p_{1}^{3} p_{3}$ | 00010 | $p_{1} p_{7}$ | 01001 |
|  |  | $p_{2}^{2} p_{3}$ | 1110 | $p_{1}^{2} p_{2} p_{3}$ | 00110 | $p_{2} p_{7}$ | 11001 |
| $p_{1}^{3}$ | 000 | $p_{3}^{2}$ | 1010 | $p_{1} p_{2}^{2} p_{3}$ | 01110 | $p_{1} p_{8}$ | 01011 |
| $p_{1}^{2} p_{2}$ | 001 | $p_{1} p_{4}$ | 0100 | $p_{2}^{3} p_{3}$ | 11110 | $p_{2} p_{8}$ | 11011 |
| $p_{1} p_{2}^{2}$ | 011 | $p_{2} p_{4}$ | 1100 | $p_{1} p_{3}^{2}$ | 01010 | $p_{9}$ | 10000 |
| $p_{2}^{3}$ | 111 | $p_{1} p_{5}$ | 0101 | $p_{2} p_{3}^{2}$ | 11010 | $p_{10}$ | 10001 |
| $p_{1} p_{3}$ | 010 | $p_{2} p_{5}$ | 1101 | $p_{1}^{2} p_{4}$ | 00100 | $p_{11}$ | 10010 |
| $p_{2} p_{3}$ | 110 | $p_{6}$ | 1000 | $p_{1} p_{2} p_{4}$ | 01100 | $p_{12}$ | 10011 |
| $p_{4}$ | 100 | $p_{7}$ | 1001 | $p_{2}^{2} p_{4}$ | 11100 | $p_{13}$ | 10110 |
| $p_{5}$ | 101 | $p_{8}$ | 1011 | $p_{3} p_{4}$ | 10100 | $p_{14}$ | 10111 |

Table E.4: Factorization of all periodic points strings up to length 5 into ordered concatenations $p_{1}^{k_{1}} p_{2}^{k_{2}} \cdots p_{n}^{k_{n}}$ of prime strings $p_{1}=0, p_{2}=1, p_{3}=10, p_{4}=100, \ldots$ , $p_{14}=10111$.
to the product over prime cycles is analogous to the Riemann zeta (exercise 15.9) represented as a sum over natural numbers vs. its Euler product representation.

We now implement this factorization explicitly by decomposing recursively binary strings into ordered concatenations of prime strings. There are 2 strings of length 1 , both prime: $p_{1}=0, p_{2}=1$. There are 4 strings of length 2: $00,01,11,10$. The first three are ordered concatenations of primes: $00=p_{1}^{2}, 01=p_{1} p_{2}, 11=p_{2}^{2}$; by ordered concatenations we mean that $p_{1} p_{2}$ is legal, but $p_{2} p_{1}$ is not. The remaining string is the only prime of length $2, p_{3}=10$. Proceeding by discarding the strings which are concatenations of shorter primes $p_{1}^{k_{1}} p_{2}^{k_{2}} \cdots p_{j}^{k_{j}}$, with primes lexically ordered, we generate the standard list of primes, in agreement with table 11.1: 0,1 , 10, 101, 100, 1000, 1001, 1011, 10000, 10001, 10010, 10011, 10110, 10111, 100000, 100001, 100010, 100011, 100110, 100111, 101100, 101110, 101111, .... This factorization is illustrated in table E.4.

## E.2.1 Prime factorization for spectral determinants

Following sect. E.2, the spectral determinant cycle expansions is obtained by expanding $F$ as a multinomial in prime cycle weights $t_{p}$

$$
\begin{equation*}
F=\prod_{p} \sum_{k=0}^{\infty} C_{p^{k}} t_{p}^{k}=\sum_{k_{1} k_{2} k_{3} \cdots=0}^{\infty} \tau_{p_{1}^{k_{1}} p_{2}^{k_{2}} p_{3}^{k_{3}} \ldots} \tag{E.12}
\end{equation*}
$$

where the sum goes over all pseudocycles. In the above we have defined

$$
\begin{equation*}
\tau_{p_{1}^{k_{1}} p_{2}^{k_{2}} p_{3}^{k_{3} \ldots}}=\prod_{i=1}^{\infty} C_{p_{i} k_{i}} k_{p_{i}}^{k_{i}} . \tag{E.13}
\end{equation*}
$$

A striking aspect of the spectral determinant cycle expansion is its re- page 264 semblance to the factorization of natural numbers into primes: as we already noted in sect. E.2, the relation of the cycle expansion (E.12) to the product formula (15.9) is analogous to the Riemann zeta represented as a sum over natural numbers vs. its Euler product representation.

This is somewhat unexpected, as the cycle weights factorize exactly with respect to $r$ repetitions of a prime cycle, $t_{p p \ldots p}=t_{p}^{r}$, but only approximately (shadowing) with respect to subdividing a string into prime substrings, $t_{p_{1} p_{2}} \approx t_{p_{1}} t_{p_{2}}$.

The coefficients $C_{p^{k}}$ have a simple form only in 1-d, given by the Euler formula (16.32). In higher dimensions $C_{p^{k}}$ can be evaluated by expanding (15.9), $F(z)=\prod_{p} F_{p}$, where

$$
F_{p}=1-\left(\sum_{r=1}^{\infty} \frac{t_{p}^{r}}{r d_{p, r}}\right)+\frac{1}{2}\left(\sum_{r=1}^{\infty} \frac{t_{p}^{r}}{r d_{p, r}}\right)^{2}-\ldots .
$$

Expanding and recollecting terms, and suppressing the $p$ cycle label for the moment, we obtain

$$
\begin{align*}
F_{p} & =\sum_{r=1}^{\infty} C_{k} t^{k}, \quad C_{k}=(-)^{k} c_{k} / D_{k}, \\
D_{k} & =\prod_{r=1}^{k} d_{r}=\prod_{a=1}^{d} \prod_{r=1}^{k}\left(1-u_{a}^{r}\right) \tag{E.14}
\end{align*}
$$

where evaluation of $c_{k}$ requires a certain amount of not too luminous algebra:

$$
\begin{aligned}
c_{0}= & 1 \\
c_{1}= & 1 \\
c_{2}= & \frac{1}{2}\left(\frac{d_{2}}{d_{1}}-d_{1}\right)=\frac{1}{2}\left(\prod_{a=1}^{d}\left(1+u_{a}\right)-\prod_{a=1}^{d}\left(1-u_{a}\right)\right) \\
c_{3}= & \frac{1}{3!}\left(\frac{d_{2} d_{3}}{d_{1}^{2}}+2 d_{1} d_{2}-3 d_{3}\right) \\
= & \frac{1}{6}\left(\prod_{a=1}^{d}\left(1+2 u_{a}+2 u_{a}^{2}+u_{a}^{3}\right)\right. \\
& \left.+2 \prod_{a=1}^{d}\left(1-u_{a}-u_{a}^{2}+u_{a}^{3}\right)-3 \prod_{a=1}^{d}\left(1-u_{a}^{3}\right)\right)
\end{aligned}
$$

etc.. For example, for a general 2-dimensional map we have

$$
F_{p}=1-\frac{1}{D_{1}} t+\frac{u_{1}+u_{2}}{D_{2}} t^{2}-\frac{u_{1} u_{2}\left(1+u_{1}\right)\left(1+u_{2}\right)+u_{1}^{3}+u_{2}^{3}}{D_{3}} t^{3}+\ldots \text { (1) }
$$

We discuss the convergence of such cycle expansions in sect. J.4.
With $\tau_{p_{1}^{k_{1}} p_{2}^{k_{2} \ldots p_{n}^{k_{n}}}}$ defined as above, the prime factorization of symbol strings is unique in the sense that each symbol string can be written as a unique concatenation of prime strings, up to a convention on ordering of primes. This factorization is a nontrivial example of the utility of generalized Möbius inversion, sect. E.2.

How is the factorization of sect. E. 2 used in practice? Suppose we have computed (or perhaps even measured in an experiment) all prime cycles up to length $n$, that is, we have a list of $t_{p}$ 's and the corresponding Jacobian matrix eigenvalues $\Lambda_{p, 1}, \Lambda_{p, 2}, \ldots \Lambda_{p, d}$. A cycle expansion of the Selberg product is obtained by generating all strings in order of increasing length $j$ allowed by the symbolic dynamics and constructing the multinomial

$$
\begin{equation*}
F=\sum_{n} \tau_{n} \tag{E.16}
\end{equation*}
$$

where $n=s_{1} s_{2} \cdots s_{j}$, $s_{i}$ range over the alphabet, in the present case $\{0,1\}$. Factorizing every string $n=s_{1} s_{2} \cdots s_{j}=p_{1}^{k_{1}} p_{2}^{k_{2}} \cdots p_{j}^{k_{j}}$ as in table E.4, and substituting $\tau_{p_{1}^{k_{1}} p_{2}^{k_{2}} \ldots}$ we obtain a multinomial approximation to $F$. For example, $\tau_{001001010101}=\tau_{\underline{001} \underline{001} \underline{\underline{0101} \underline{01}}=}=\tau_{001^{2}} \tau_{01^{3}}$, and $\tau_{01^{3}}, \tau_{001^{2}}$ are known functions of the corresponding cycle eigenvalues. The zeros of $F$ can now be easily determined by standard numerical methods. The fact that as far as the symbolic dynamics is concerned, the cycle expansion of a Selberg product is simply an average over all symbolic strings makes Selberg products rather pretty.

To be more explicit, we illustrate the above by expressing binary strings as concatenations of prime factors. We start by computing $N_{n}$, the number of terms in the expansion (E.12) of the total cycle length $n$. Setting $C_{p^{k}} t_{p}^{k}=$ $z^{n_{p} k}$ in (E.12), we obtain

$$
\sum_{n=0}^{\infty} N_{n} z^{n}=\prod_{p} \sum_{k=0}^{\infty} z^{n_{p} k}=\frac{1}{\prod_{p}\left(1-z^{n_{p}}\right)} .
$$

So the generating function for the number of terms in the Selberg product is the topological zeta function. For the complete binary dynamics we have $N_{n}=2^{n}$ contributing terms of length $n$ :

$$
\zeta_{\text {top }}=\frac{1}{\prod_{p}\left(1-z^{n_{p}}\right)}=\frac{1}{1-2 z}=\sum_{n=0}^{\infty} 2^{n} z^{n}
$$

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Hence the number of distinct terms in the expansion (E.12) is the same as the number of binary strings, and conversely, the set of binary strings of length $n$ suffices to label all terms of the total cycle length $n$ in the expansion (E.12).

## Appendix F

## Counting itineraries

## F. 1 Counting curvatures

One consequence of the finitness of topological polynomials is that the contributions to curvatures at every order are even in number, half with positive and half with negative sign. For instance, for complete binary labeling (18.5),

$$
\begin{align*}
c_{4}= & -t_{0001}-t_{0011}-t_{0111}-t_{0} t_{01} t_{1} \\
& +t_{0} t_{001}+t_{0} t_{011}+t_{001} t_{1}+t_{011} t_{1} \tag{F.1}
\end{align*}
$$

We see that $2^{3}$ terms contribute to $c_{4}$, and exactly half of them appear with a negative sign - hence if all binary strings are admissible, this term vanishes in the counting expression.

Such counting rules arise from the identity

$$
\begin{equation*}
\prod_{p}\left(1+t_{p}\right)=\prod_{p} \frac{1-t_{p}^{2}}{1-t_{p}} . \tag{F.2}
\end{equation*}
$$

Substituting $t_{p}=z^{n_{p}}$ and using (13.15) we obtain for unrestricted symbol dynamics with $N$ letters

$$
\prod_{p}^{\infty}\left(1+z^{n_{p}}\right)=\frac{1-N z^{2}}{1-N z}=1+N z+\sum_{k=2}^{\infty} z^{k}\left(N^{k}-N^{k-1}\right)
$$

The $z^{n}$ coefficient in the above expansion is the number of terms contributing to $c_{n}$ curvature, so we find that for a complete symbolic dynamics of $N$ symbols and $n>1$, the number of terms contributing to $c_{n}$ is $(N-1) N^{k-1}$ (of which half carry a minus sign).

We find that for complete symbolic dynamics of $N$ symbols and $n>1$, the number of terms contributing to $c_{n}$ is $(N-1) N^{n-1}$. So, superficially, not much is gained by going from periodic orbits trace sums which get $N^{n}$ contributions of $n$ to the curvature expansions with $N^{n}(1-1 / N)$. However, the point is not the number of the terms, but the cancellations between them.

## Exercises

Exercise F. 1 Lefschetz zeta function. Elucidate the relation betveen the topological zeta function and the Lefschetz zeta function.

Exercise F. 2 Counting the 3-disk pinball counterterms. Verify that the number of terms in the 3-disk pinball curvature expansion (18.31) is given by

$$
\begin{aligned}
\prod_{p}\left(1+t_{p}\right) & =\frac{1-3 z^{4}-2 z^{6}}{1-3 z^{2}-2 z^{3}}=1+3 z^{2}+2 z^{3}+\frac{z^{4}\left(6+12 z+2 z^{2}\right)}{1-3 z^{2}-2 z^{3}} \\
& =1+3 z^{2}+2 z^{3}+6 z^{4}+12 z^{5}+20 z^{6}+48 z^{7}+84 z^{8}+184 z^{9}+(\mathrm{F} .3)
\end{aligned}
$$

This means that, for example, $c_{6}$ has a total of 20 terms, in agreement with the explicit 3-disk cycle expansion (18.32).

Exercise F. 3 Cycle expansion denominators**. Prove that the denominator of $c_{k}$ is indeed $D_{k}$, as asserted (E.14).

Exercise F. 4 Counting subsets of cycles. The techniques developed above can be generalized to counting subsets of cycles. Consider the simplest example of a dynamical system with a complete binary tree, a repeller map (11.8) with two straight branches, which we label 0 and 1 . Every cycle weight for such map factorizes, with a factor $t_{0}$ for each 0 , and factor $t_{1}$ for each 1 in its symbol string. The transition matrix traces (13.5) collapse to $\operatorname{tr}\left(T^{k}\right)=\left(t_{0}+t_{1}\right)^{k}$, and $1 / \zeta$ is simply

$$
\begin{equation*}
\prod_{p}\left(1-t_{p}\right)=1-t_{0}-t_{1} \tag{F.4}
\end{equation*}
$$

Substituting into the identity

$$
\prod_{p}\left(1+t_{p}\right)=\prod_{p} \frac{1-t_{p}^{2}}{1-t_{p}}
$$

we obtain

$$
\begin{align*}
\prod_{p}\left(1+t_{p}\right) & =\frac{1-t_{0}^{2}-t_{1}^{2}}{1-t_{0}-t_{1}}=1+t_{0}+t_{1}+\frac{2 t_{0} t_{1}}{1-t_{0}-t_{1}} \\
& =1+t_{0}+t_{1}+\sum_{n=2}^{\infty} \sum_{k=1}^{n-1} 2\binom{n-2}{k-1} t_{0}^{k} t_{1}^{n-k} \tag{F.5}
\end{align*}
$$

Hence for $n \geq 2$ the number of terms in the expansion ?! with $k 0$ 's and $n-k$ 1 's in their symbol sequences is $2\binom{n-2}{k-1}$. This is the degeneracy of distinct cycle eigenvalues in fig.?!; for systems with non-uniform hyperbolicity this degeneracy is lifted (see fig. ?!).

In order to count the number of prime cycles in each such subset we denote with $M_{n, k} \quad(n=1,2, \ldots ; k=\{0,1\}$ for $n=1 ; k=1, \ldots, n-1$ for $n \geq 2)$ the
number of prime $n$-cycles whose labels contain $k$ zeros, use binomial string counting and Möbius inversion and obtain

$$
\begin{aligned}
M_{1,0} & =M_{1,1}=1 \\
n M_{n, k} & =\sum_{\left.m\right|_{\frac{n}{k}}} \mu(m)\binom{n / m}{k / m}, \quad n \geq 2, k=1, \ldots, n-1
\end{aligned}
$$

where the sum is over all $m$ which divide both $n$ and $k$.

## Appendix G

## Finding cycles

(C. Chandre)

## G. 1 Newton-Raphson method

## G.1.1 Contraction rate

Consider a $d$-dimensional map $x^{\prime}=f(x)$ with an unstable fixed point $x_{*}$. The Newton-Raphson algorithm is obtained by iterating the following map

$$
x^{\prime}=g(x)=x-(J(x)-1)^{-1}(f(x)-x) .
$$

The linearization of $g$ near $x_{*}$ leads to

$$
x_{*}+\epsilon^{\prime}=x_{*}+\epsilon-\left(J\left(x_{*}\right)-\mathbf{1}\right)^{-1}\left(f\left(x_{*}\right)+J\left(x_{*}\right) \epsilon-x_{*}-\epsilon\right)+O\left(\|\epsilon\|^{2}\right),
$$

where $\epsilon=x-x_{*}$. Therefore,

$$
x^{\prime}-x_{*}=O\left(\left(x-x_{*}\right)^{2}\right) .
$$

After $n$ steps and if the initial guess $x_{0}$ is close to $x_{*}$, the error decreases super-exponentially

$$
g^{n}\left(x_{0}\right)-x_{*}=O\left(\left(x_{0}-x_{*}\right)^{2^{n}}\right) .
$$

## G.1.2 Computation of the inverse

The Newton-Raphson method for finding $n$-cycles of $d$-dimensional mappings using the multi-shooting method reduces to the following equation

$$
\left(\begin{array}{cccc}
\mathbf{1} & & & -D f\left(x_{n}\right)  \tag{G.1}\\
-D f\left(x_{1}\right) & \mathbf{1} & & \\
& \cdots & \mathbf{1} & \\
& & -D f\left(x_{n-1}\right) & \mathbf{1}
\end{array}\right)\left(\begin{array}{c}
\delta_{1} \\
\delta_{2} \\
\cdots \\
\delta_{n}
\end{array}\right)=-\left(\begin{array}{c}
F_{1} \\
F_{2} \\
\cdots \\
F_{n}
\end{array}\right),(
$$

where $D f(x)$ is the $[d \times d]$ Jacobian matrix of the map evaluated at the point $x$, and $\delta_{m}=x_{m}^{\prime}-x_{m}$ and $F_{m}=x_{m}-f\left(x_{m-1}\right)$ are $d$-dimensional vectors. By some starightforward algebra, the vectors $\delta_{m}$ are expressed as functions of the vectors $F_{m}$ :

$$
\begin{equation*}
\delta_{m}=-\sum_{k=1}^{m} \beta_{k, m-1} F_{k}-\beta_{1, m-1}\left(\mathbf{1}-\beta_{1, n}\right)^{-1}\left(\sum_{k=1}^{n} \beta_{k, n} F_{k}\right), \tag{G.2}
\end{equation*}
$$

for $m=1, \ldots, n$, where $\beta_{k, m}=D f\left(x_{m}\right) D f\left(x_{m-1}\right) \cdots D f\left(x_{k}\right)$ for $k<m$ and $\beta_{k, m}=\mathbf{1}$ for $k \geq m$. Therefore, finding $n$-cycles by a Newton-Raphson method with multiple shooting requires the inversing of a $[d \times d$ ] matrix 1 - $D f\left(x_{n}\right) D f\left(x_{n-1}\right) \cdots D f\left(x_{1}\right)$.

## G. 2 Hybrid Newton-Raphson / relaxation method



Consider a $d$-dimensional map $x^{\prime}=f(x)$ with an unstable fixed point $x_{*}$. The transformed map is the following one:

$$
x^{\prime}=g(x)=x+\gamma C(f(x)-x),
$$

where $\gamma>0$ and $C$ is a $d \times d$ invertible constant matrix. We notice that $x_{*}$ is also a fixed point of $g$. Consider the stability matrix at the fixed point $x_{*}$

$$
A_{g}=\left.\frac{d g}{d x}\right|_{x=x_{*}}=1+\gamma C\left(A_{f}-1\right) .
$$

The matrix $C$ is constructed such that the eigenvalues of $A_{g}$ are of modulus less than one. Assume that $A_{f}$ is diagonalizable: In the basis of diagonalization, the matrix writes:

$$
\tilde{A}_{g}=1+\gamma \tilde{C}\left(\tilde{A}_{f}-1\right),
$$

where $\tilde{A}_{f}$ is diagonal with elements $\mu_{i}$. We restrict the set of matrices $\tilde{C}$ to diagonal matrices with $\tilde{C}_{i i}=\epsilon_{i}$ where $\epsilon_{i}= \pm 1$. Thus $\tilde{A}_{g}$ is diagonal with eigenvalues $\gamma_{i}=1+\gamma \epsilon_{i}\left(\mu_{i}-1\right)$. The choice of $\gamma$ and $\epsilon_{i}$ is such that $\left|\gamma_{i}\right|<1$. It is easy to see that if $\operatorname{Re}\left(\mu_{i}\right)<1$ one has to choose $\epsilon_{i}=1$, and if $\operatorname{Re}\left(\mu_{i}\right)>1, \epsilon_{i}=-1$. If $\lambda$ is chosen such that

$$
0<\gamma<\min _{i=1, \ldots, d} \frac{2\left|\operatorname{Re}\left(\mu_{i}\right)-1\right|}{\left|\mu_{i}-1\right|^{2}}
$$

all the eigenvalues of $A_{g}$ have modulus less that one. The contraction rate at the fixed point for the map $g$ is then $\max _{i}\left|1+\gamma \epsilon_{i}\left(\mu_{i}-1\right)\right|$. We notice that if $\operatorname{Re}\left(\mu_{i}\right)=1$, it is not possible to stabilize $x_{*}$ by the set of matrices $\gamma C$.
From the construction of $C$, we see that $2^{d}$ choices of matrices are possible. For example, for two-dimensional systems, these matrices are

$$
C \in\left\{\left(\begin{array}{ll}
1 & 0 \\
0 & 1
\end{array}\right),\left(\begin{array}{cc}
-1 & 0 \\
0 & 1
\end{array}\right),\left(\begin{array}{cc}
1 & 0 \\
0 & -1
\end{array}\right),\left(\begin{array}{cc}
-1 & 0 \\
0 & -1
\end{array}\right)\right\} .
$$

## G.2. HYBRID NEWTON-RAPHSON / RELAXATION METHOD

For 2-dimensional dissipative maps, the eigenvalues satisfy $\operatorname{Re}\left(\mu_{1}\right) \operatorname{Re}\left(\mu_{2}\right) \leq$ det $D f<1$. The case $\left(\operatorname{Re}\left(\mu_{1}\right)>1, \operatorname{Re}\left(\mu_{2}\right)>1\right)$ which is stabilized by $\left(\begin{array}{cc}-1 & 0 \\ 0 & -1\end{array}\right)$ has to be discarded. The minimal set is reduced to three matrices.

## Appendix H

## Applications

Man who says it cannot be done should not interrupt man doing it.
Sayings of Vattay Gábor

In this appendix we show that the multidimensional Lyapunov exponents and relaxation exponents (dynamo rates) of vector fields can be expressed in terms of leading eigenvalues of appropriate evolution operators.

## H. 1 Evolution operator for Lyapunov exponents



Lyapunov exponents were introduced and computed for 1-d maps in sect. 10.3.2. For higher-dimensional flows only the Jacobian matrices are multiplicative, not individual eigenvalues, and the construction of the evolution operator for evaluation of the Lyapunov spectra requires the extension of evolution equations to the flow in the tangent space. We now develop the requisite theory.

Here we construct a multiplicative evolution operator (H.4) whose spectral determinant (H.8) yields the leading Lyapunov exponent of a $d$-dimensional flow (and is entire for Axiom A flows).

The key idea is to extending the dynamical system by the tangent space of the flow, suggested by the standard numerical methods for evaluation of Lyapunov exponents: start at $x_{0}$ with an initial infinitesimal tangent space vector $\eta(0) \in \mathbf{T} \mathcal{M}_{x}$, and let the flow transport it along the trajectory $x(t)=f^{t}\left(x_{0}\right)$.

The dynamics in the $(x, \eta) \in U \times T U_{x}$ space is governed by the system of equations of variations [7.1]:

$$
\dot{x}=\mathbf{v}(x), \quad \dot{\eta}=\mathbf{D v}(x) \eta
$$

Here $\mathbf{D v}(x)$ is the derivative matrix of the flow. We write the solution as

$$
\begin{equation*}
x(t)=f^{t}\left(x_{0}\right), \quad \eta(t)=\mathbf{J}^{t}\left(x_{0}\right) \cdot \eta_{0}, \tag{H.1}
\end{equation*}
$$

with the tangent space vector $\eta$ transported by the stability matrix $\mathbf{J}^{t}\left(x_{0}\right)=$ $\partial x(t) / \partial x_{0}$.

As explained in sect. 4.1, the growth rate of this vector is multiplicative along the trajectory and can be represented as $\eta(t)=|\eta(t)| /|\eta(0)| \mathbf{u}(t)$ where $\mathbf{u}(t)$ is a "unit" vector in some norm $\|$.$\| . For asymptotic times$ and for almost every initial $\left(x_{0}, \eta(0)\right)$, this factor converges to the leading eigenvalue of the linearized stability matrix of the flow.

We implement this multiplicative evaluation of stability eigenvalues by adjoining the $d$-dimensional transverse tangent space $\eta \in \mathbf{T} \mathcal{M}_{x} ; \eta(x) \mathbf{v}(x)=$ 0 to the $(d+1)$-dimensional dynamical evolution space $x \in \mathcal{M} \subset \mathbb{R}^{d+1}$. In order to determine the length of the vector $\eta$ we introduce a homogeneous differentiable scalar function $g(\eta)=\|\eta\|$. It has the property $g(\Lambda \eta)=|\Lambda| g(\eta)$ for any $\Lambda$. An example is the projection of a vector to its $d$ th component

$$
g\left(\begin{array}{c}
\eta_{1} \\
\eta_{2} \\
\cdots \\
\eta_{d}
\end{array}\right)=\left|\eta_{d}\right| .
$$

Any vector $\eta \in T U_{x}$ can now be represented by the product $\eta=\Lambda \mathbf{u}$, where $\mathbf{u}$ is a "unit" vector in the sense that its norm is $\|\mathbf{u}\|=1$, and the factor

$$
\begin{equation*}
\Lambda^{t}\left(x_{0}, \mathbf{u}_{0}\right)=g(\eta(t))=g\left(\mathbf{J}^{t}\left(x_{0}\right) \cdot \mathbf{u}_{0}\right) \tag{H.2}
\end{equation*}
$$

is the multiplicative "stretching" factor.
Unlike the leading eigenvalue of the Jacobian the stretching factor is multiplicative along the trajectory:

$$
\Lambda^{t^{\prime}+t}\left(x_{0}, \mathbf{u}_{0}\right)=\Lambda^{t^{\prime}}(x(t), \mathbf{u}(t)) \Lambda^{t}\left(x_{0}, \mathbf{u}_{0}\right)
$$

The $\mathbf{u}$ evolution constrained to $E \mathbf{T}_{g, x}$, the space of unit transverse tangent vectors, is given by rescaling of (H.1):

$$
\begin{equation*}
\mathbf{u}^{\prime}=R^{t}(x, \mathbf{u})=\frac{1}{\Lambda^{t}(x, \mathbf{u})} \mathbf{J}^{t}(x) \cdot \mathbf{u} . \tag{H.3}
\end{equation*}
$$

Eqs. (H.1), (H.2) and (H.3) enable us to define a multiplicative evolution operator on the extended space $U \times E \mathbf{T}_{g, x}$

$$
\begin{equation*}
\mathcal{L}^{t}\left(x^{\prime}, \mathbf{u}^{\prime} ; x, \mathbf{u}\right)=\delta\left(x^{\prime}-f^{t}(x)\right) \frac{\delta\left(\mathbf{u}^{\prime}-R^{t}(x, \mathbf{u})\right)}{\left|\Lambda^{t}(x, \mathbf{u})\right|^{\beta-1}} \tag{H.4}
\end{equation*}
$$

where $\beta$ is a variable.
To evaluate the expectation value of $\log \left|\Lambda^{t}(x, \mathbf{u})\right|$ which is the Lyapunov exponent we again have to take the proper derivative of the leading eigenvalue of (H.4). In order to derive the trace formula for the operator (H.4) we need to evaluate $\operatorname{Tr} \mathcal{L}^{t}=\int d x d \mathbf{u} \mathcal{L}^{t}(\mathbf{u}, x ; \mathbf{u}, x)$. The $\int d x$ integral yields a weighted sum over prime periodic orbits $p$ and their repetitions $r$ :

$$
\begin{align*}
\operatorname{Tr} \mathcal{L}^{t} & =\sum_{p} T_{p} \sum_{r=1}^{\infty} \frac{\delta\left(t-r T_{p}\right)}{\left|\operatorname{det}\left(1-\mathbf{J}_{p}^{r}\right)\right|} \Delta_{p, r}, \\
\Delta_{p, r} & =\int_{g} d \mathbf{u} \frac{\delta\left(\mathbf{u}-R^{T_{p} r}\left(x_{p}, \mathbf{u}\right)\right)}{\left|\Lambda^{T_{p} r}\left(x_{p}, \mathbf{u}\right)\right|^{\beta-1}}, \tag{H.5}
\end{align*}
$$

where $\mathbf{J}_{p}$ is the prime cycle $p$ transverse stability matrix. As we shall see below, $\Delta_{p, r}$ is intrinsic to cycle $p$, and independent of any particular cycle point $x_{p}$.

We note next that if the trajectory $f^{t}(x)$ is periodic with period $T$, the tangent space contains $d$ periodic solutions

$$
\mathbf{e}_{i}(x(T+t))=\mathbf{e}_{i}(x(t)), \quad i=1, \ldots, d,
$$

corresponding to the $d$ unit eigenvectors $\left\{\mathbf{e}_{1}, \mathbf{e}_{2}, \cdots, \mathbf{e}_{d}\right\}$ of the transverse stability matrix, with "stretching" factors (H.2) given by its eigenvalues

$$
\mathbf{J}_{p}(x) \cdot \mathbf{e}_{i}(x)=\Lambda_{p, i} \mathbf{e}_{i}(x), \quad i=1, \ldots, d . \quad \text { (no summation on } i \text { ) }
$$

The $\int d \mathbf{u}$ integral in (H.5) picks up contributions from these periodic solutions. In order to compute the stability of the $i$ th eigendirection solution, it is convenient to expand the variation around the eigenvector $\mathbf{e}_{i}$ in the stability matrix eigenbasis $\delta \mathbf{u}=\sum \delta u_{\ell} \mathbf{e}_{\ell}$. The variation of the map (H.3) at a complete period $t=T$ is then given by

$$
\begin{align*}
\delta R^{T}\left(\mathbf{e}_{i}\right) & =\frac{\mathbf{J} \cdot \delta \mathbf{u}}{g\left(\mathbf{J} \cdot \mathbf{e}_{i}\right)}-\frac{\mathbf{J} \cdot \mathbf{e}_{i}}{g\left(\mathbf{J} \cdot \mathbf{e}_{i}\right)^{2}}\left(\frac{\partial g\left(\mathbf{e}_{i}\right)}{\partial \mathbf{u}} \cdot \mathbf{J} \cdot \delta \mathbf{u}\right) \\
& =\sum_{k \neq i} \frac{\Lambda_{p, k}}{\Lambda_{p, i}}\left(\mathbf{e}_{k}-\mathbf{e}_{i} \frac{\partial g\left(\mathbf{e}_{i}\right)}{\partial u_{k}}\right) \delta u_{k} . \tag{H.6}
\end{align*}
$$

The $\delta u_{i}$ component does not contribute to this sum since $g\left(\mathbf{e}_{i}+d u_{i} \mathbf{e}_{i}\right)=$ $1+d u_{i}$ implies $\partial g\left(\mathbf{e}_{i}\right) / \partial u_{i}=1$. Indeed, infinitesimal variations $\delta \mathbf{u}$ must satisfy

$$
g(\mathbf{u}+\delta \mathbf{u})=g(\mathbf{u})=1 \quad \Longrightarrow \quad \sum_{\ell=1}^{d} \delta u_{\ell} \frac{\partial g(\mathbf{u})}{\partial u_{\ell}}=0
$$

so the allowed variations are of form

$$
\delta \mathbf{u}=\sum_{k \neq i}\left(\mathbf{e}_{k}-\mathbf{e}_{i} \frac{\partial g\left(\mathbf{e}_{i}\right)}{\partial u_{k}}\right) c_{k}, \quad\left|c_{k}\right| \ll 1,
$$

and in the neighborhood of the $\mathbf{e}_{i}$ eigenvector the $\int d \mathbf{u}$ integral can be expressed as

$$
\int_{g} d \mathbf{u}=\int \prod_{k \neq i} d c_{k}
$$

Inserting these variations into the $\int d \mathbf{u}$ integral we obtain

$$
\begin{aligned}
\int_{g} d \mathbf{u} & \delta\left(\mathbf{e}_{i}+\delta \mathbf{u}-R^{T}\left(\mathbf{e}_{i}\right)-\delta R^{T}\left(\mathbf{e}_{i}\right)+\ldots\right) \\
= & \int \prod_{k \neq i} d c_{k} \delta\left(\left(1-\Lambda_{k} / \Lambda_{i}\right) c_{k}+\ldots\right) \\
& =\prod_{k \neq i} \frac{1}{\left|1-\Lambda_{k} / \Lambda_{i}\right|},
\end{aligned}
$$

and the $\int d \mathbf{u}$ trace (H.5) becomes

$$
\begin{equation*}
\Delta_{p, r}=\sum_{i=1}^{d} \frac{1}{\left|\Lambda_{p, i}^{r}\right|^{\beta-1}} \prod_{k \neq i} \frac{1}{\left|1-\Lambda_{p, k}^{r} / \Lambda_{p, i}^{r}\right|} . \tag{H.7}
\end{equation*}
$$

The corresponding spectral determinant is obtained by observing that the Laplace transform of the trace (14.19) is a logarithmic derivative $\operatorname{Tr} \mathcal{L}(s)=$ $-\frac{d}{d s} \log F(s)$ of the spectral determinant:

$$
\begin{equation*}
F(\beta, s)=\exp \left(-\sum_{p, r} \frac{e^{s T_{p} r}}{r\left|\operatorname{det}\left(1-\mathbf{J}_{p}^{r}\right)\right|} \Delta_{p, r}(\beta)\right) \tag{H.8}
\end{equation*}
$$

This determinant is the central result of this section. Its zeros correspond to the eigenvalues of the evolution operator (H.4), and can be evaluated by the cycle expansion methods.

The leading zero of (H.8) is called "pressure" (or free energy)

$$
\begin{equation*}
P(\beta)=s_{0}(\beta) \tag{H.9}
\end{equation*}
$$

The average Lyapunov exponent is then given by the first derivative of the pressure at $\beta=1$ :

$$
\begin{equation*}
\bar{\lambda}=P^{\prime}(1) . \tag{H.10}
\end{equation*}
$$

The simplest application of (H.8) is to 2-dimensional hyperbolic Hamiltonian maps. The stability eigenvalues are related by $\Lambda_{1}=1 / \Lambda_{2}=\Lambda$, and the spectral determinant is given by

$$
\begin{align*}
F(\beta, z) & =\exp \left(-\sum_{p, r} \frac{z^{r n_{p}}}{r\left|\Lambda_{p}^{r}\right|} \frac{1}{\left(1-1 / \Lambda_{p}^{r}\right)^{2}} \Delta_{p, r}(\beta)\right) \\
\Delta_{p, r}(\beta) & =\frac{\left|\Lambda_{p}^{r}\right|^{1-\beta}}{1-1 / \Lambda_{p}^{2 r}}+\frac{\left|\Lambda_{p}^{r}\right|^{\beta-3}}{1-1 / \Lambda_{p}^{2 r}} . \tag{H.11}
\end{align*}
$$

The dynamics (H.3) can be restricted to a $u$ unit eigenvector neighborhood corresponding to the largest eigenvalue of the Jacobi matrix. On this neighborhood the largest eigenvalue of the Jacobi matrix is the only fixed point, and the spectral determinant obtained by keeping only the largest term the $\Delta_{p, r}$ sum in (H.7) is also entire.

In case of maps it is practical to introduce the logarithm of the leading zero and to call it "pressure"

$$
\begin{equation*}
P(\beta)=\log z_{0}(\beta) . \tag{H.12}
\end{equation*}
$$

The average of the Lyapunov exponent of the map is then given by the first derivative of the pressure at $\beta=1$ :

$$
\begin{equation*}
\bar{\lambda}=P^{\prime}(1) . \tag{H.13}
\end{equation*}
$$

By factorizing the determinant (H.11) into products of zeta functions we can conclude that the leading zero of the (H.4) can also be recovered from the leading zeta function

$$
\begin{equation*}
1 / \zeta_{0}(\beta, z)=\exp \left(-\sum_{p, r} \frac{z^{r n_{p}}}{r\left|\Lambda_{p}^{r}\right|^{\beta}}\right) . \tag{H.14}
\end{equation*}
$$

This zeta function plays a key role in thermodynamic applications as we will will see in Chapter 20.

## H. 2 Advection of vector fields by chaotic flows

Fluid motions can move embedded vector fields around. An example is the magnetic field of the Sun which is "frozen" in the fluid motion. A passively evolving vector field $\mathbf{V}$ is governed by an equation of the form

$$
\begin{equation*}
\partial_{t} \mathbf{V}+\mathbf{u} \cdot \nabla \mathbf{V}-\mathbf{V} \cdot \nabla \mathbf{u}=0 \tag{H.15}
\end{equation*}
$$

where $\mathbf{u}(x, t)$ represents the velocity field of the fluid. The strength of the vector field can grow or decay during its time evolution. The amplification of the vector field in such a process is called the "dynamo effect". In a strongly chaotic fluid motion we can characterize the asymptotic behavior of the field with an exponent

$$
\begin{equation*}
\mathbf{V}(x, t) \sim \mathbf{V}(x) e^{\nu t} \tag{H.16}
\end{equation*}
$$

where $\nu$ is called the fast dynamo rate. The goal of this section is to show that periodic orbit theory can be developed for such a highly non-trivial system as well.

We can write the solution of (H.15) formally, as shown by Cauchy. Let $\mathbf{x}(t, \mathbf{a})$ be the position of the fluid particle that was at the point a at $t=0$. Then the field evolves according to

$$
\begin{equation*}
\mathbf{V}(\mathbf{x}, t)=\mathbf{J}(\mathbf{a}, t) \mathbf{V}(\mathbf{a}, 0) \tag{H.17}
\end{equation*}
$$

where $\mathbf{J}(\mathbf{a}, t)=\partial(\mathbf{x}) / \partial(\mathbf{a})$ is the Jacobian matrix of the transformation that moves the fluid into itself $\mathbf{x}=\mathbf{x}(\mathbf{a}, t)$.

We write $\mathbf{x}=f^{t}(\mathbf{a})$, where $f^{t}$ is the flow that maps the initial positions of the fluid particles into their positions at time $t$. Its inverse, $\mathbf{a}=f^{-t}(\mathbf{x})$, maps particles at time $t$ and position $\mathbf{x}$ back to their initial positions. Then we can write (H.17)

$$
\begin{equation*}
V_{i}(\mathbf{x}, t)=\sum_{j} \int d^{3} \mathbf{a} \mathcal{L}_{i j}^{t}(\mathbf{x}, \mathbf{a}) V_{j}(\mathbf{a}, 0) \tag{H.18}
\end{equation*}
$$

with

$$
\begin{equation*}
\mathcal{L}_{i j}^{t}(\mathbf{x}, \mathbf{a})=\delta\left(\mathbf{a}-f^{-t}(\mathbf{x})\right) \frac{\partial x_{i}}{\partial a_{j}} \tag{H.19}
\end{equation*}
$$

For large times, the effect of $\mathcal{L}^{t}$ is dominated by its leading eigenvalue, $e^{\nu_{0} t}$ with $\operatorname{Re}\left(\nu_{0}\right)>\operatorname{Re}\left(\nu_{i}\right), i=1,2,3, \ldots$. In this way the transfer operator furnishes the fast dynamo rate, $\nu:=\nu_{0}$.

The trace of the transfer operator is the sum over all periodic orbit contributions, with each cycle weighted by its intrinsic stability

$$
\begin{equation*}
\operatorname{Tr} \mathcal{L}^{t}=\sum_{p} T_{p} \sum_{r=1}^{\infty} \frac{\operatorname{tr} \mathbf{J}_{p}^{r}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{-r}\right)\right|} \delta\left(t-r T_{p}\right) . \tag{H.20}
\end{equation*}
$$

We can construct the corresponding spectral determinant as usual

$$
\begin{equation*}
F(s)=\exp \left[-\sum_{p} \sum_{r=1}^{\infty} \frac{1}{r} \frac{\operatorname{tr} \mathbf{J}_{p}^{r}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{-r}\right)\right|} e^{s r T_{p}}\right] \tag{H.21}
\end{equation*}
$$

Note that in this formuli we have omitted a term arising from the Jacobian transformation along the orbit which would give 1 $+\operatorname{tr} \mathbf{J}_{p}^{r}$ in the numerator rather than just the trace of $\mathbf{J}_{p}^{r}$. Since the extra term corresponds to advection along the orbit, and this does not evolve the magnetic field, we have chosen to ignore it. It is also interesting to note that the negative powers of the Jacobian occur in the denominator, since we have $f^{-t}$ in (H.19).

In order to simplify $F(s)$, we factor the denominator cycle stability determinants into products of expanding and contracting eigenvalues. For a 3 -dimensional fluid flow with cycles possessing one expanding eigenvalue $\Lambda_{p}$ (with $\left|\Lambda_{p}\right|>1$ ), and one contracting eigenvalue $\lambda_{p}$ (with $\left|\lambda_{p}\right|<1$ ) the determinant may be expanded as follows:

$$
\begin{equation*}
\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{-r}\right)\right|^{-1}=\left|\left(1-\Lambda_{p}^{-r}\right)\left(1-\lambda_{p}^{-r}\right)\right|^{-1}=\left|\lambda_{p}\right|^{r} \sum_{j=0}^{\infty} \sum_{k=0}^{\infty} \Lambda_{p}^{-j r} \lambda_{p}^{k r} \tag{H.22}
\end{equation*}
$$

With this decomposition we can rewrite the exponent in (H.21) as

$$
\begin{equation*}
\sum_{p} \sum_{r=1}^{\infty} \frac{1}{r} \frac{\left(\lambda_{p}^{r}+\Lambda_{p}^{r}\right) e^{s r T_{p}}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{-r}\right)\right|}=\sum_{p} \sum_{j, k=0}^{\infty} \sum_{r=1}^{\infty} \frac{1}{r}\left(\left|\lambda_{p}\right| \Lambda_{p}^{-j} \lambda_{p}^{k} e^{s T_{p}}\right)^{r}\left(\lambda_{p}^{r}+\Lambda_{p}^{r}\right), \tag{H.23}
\end{equation*}
$$

which has the form of the expansion of a logarithm:

$$
\sum_{p} \sum_{j, k}\left[\log \left(1-e^{s T_{p}}\left|\lambda_{p}\right| \Lambda_{p}^{1-j} \lambda_{p}^{k}\right)+\log \left(1-e^{s T_{p}}\left|\lambda_{p}\right| \Lambda_{p}^{-j} \lambda_{p}^{1+k}\right)\right]
$$

The spectral determinant is therefore of the form,

$$
\begin{equation*}
F(s)=F_{e}(s) F_{c}(s) \tag{H.25}
\end{equation*}
$$

where

$$
\begin{align*}
& F_{e}(s)=\prod_{p} \prod_{j, k=0}^{\infty}\left(1-t_{p}^{(j k)} \Lambda_{p}\right),  \tag{H.26}\\
& F_{c}(s)=\prod_{p} \prod_{j, k=0}^{\infty}\left(1-t_{p}^{(j k)} \lambda_{p}\right), \tag{H.27}
\end{align*}
$$

with

$$
\begin{equation*}
t_{p}^{(j k)}=e^{s T_{p}}\left|\lambda_{p}\right| \frac{\lambda_{p}^{k}}{\Lambda_{p}^{j}} \tag{H.28}
\end{equation*}
$$

The two factors present in $F(s)$ correspond to the expanding and contracting exponents. (Had we not neglected a term in (H.21), there would be a third factor corresponding to the translation.)

For 2- $d$ Hamiltonian volume preserving systems, $\lambda=1 / \Lambda$ and (H.26) reduces to

$$
\begin{equation*}
F_{e}(s)=\prod_{p} \prod_{k=0}^{\infty}\left(1-\frac{t_{p}}{\Lambda_{p}^{k-1}}\right)^{k+1}, \quad t_{p}=\frac{e^{s T_{p}}}{\left|\Lambda_{p}\right|} . \tag{H.29}
\end{equation*}
$$

With $\sigma_{p}=\Lambda_{p} /\left|\Lambda_{p}\right|$, the Hamiltonian zeta function (the $j=k=0$ part of the product (H.27)) is given by

$$
\begin{equation*}
1 / \zeta_{d y n}(s)=\prod_{p}\left(1-\sigma_{p} e^{s T_{p}}\right) . \tag{H.30}
\end{equation*}
$$

This is a curious formula - the zeta function depends only on the return times, not on the eigenvalues of the cycles. Furthermore, the identity,

$$
\frac{\Lambda+1 / \Lambda}{|(1-\Lambda)(1-1 / \Lambda)|}=\sigma+\frac{2}{|(1-\Lambda)(1-1 / \Lambda)|}
$$

when substituted into (H.25), leads to a relation between the vector and scalar advection spectral determinants:

$$
\begin{equation*}
F_{d y n}(s)=F_{0}^{2}(s) / \zeta_{d y n}(s) . \tag{H.31}
\end{equation*}
$$

The spectral determinants in this equation are entire for hyperbolic (axiom A) systems, since both of them correspond to multiplicative operators.

In the case of a flow governed by a map, we can adapt the formulas (H.29) and (H.30) for the dynamo determinants by simply making the substitution

$$
\begin{equation*}
z^{n_{p}}=e^{s T_{p}} \tag{H.32}
\end{equation*}
$$

where $n_{p}$ is the integer order of the cycle. Then we find the spectral determinant $F_{e}(z)$ given by equation (H.29) but with

$$
\begin{equation*}
t_{p}=\frac{z^{n_{p}}}{\left|\Lambda_{p}\right|} \tag{H.33}
\end{equation*}
$$

for the weights, and

$$
\begin{equation*}
1 / \zeta_{\text {dyn }}(z)=\Pi_{p}\left(1-\sigma_{p} z^{n_{p}}\right) \tag{H.34}
\end{equation*}
$$

for the zeta-function

For maps with finite Markov partition the inverse zeta function (H.34) reduces to a polynomial for $z$ since curvature terms in the cycle expansion vanish. For example, for maps with complete binary partition, and with the fixed point stabilities of opposite signs, the cycle expansion reduces to

$$
\begin{equation*}
1 / \zeta_{d y n}(s)=1 \tag{H.35}
\end{equation*}
$$

For such maps the dynamo spectral determinant is simply the square of the scalar advection spectral determinant, and therefore all its zeros are double. In other words, for flows governed by such discrete maps, the fast dynamo rate equals the scalar advection rate.

In contrast, for three-dimensional flows, the dynamo effect is distinct from the scalar advection. For example, for flows with finite symbolic dynamical grammars, (H.31) implies that the dynamo zeta function is a ratio of two entire determinants:

$$
\begin{equation*}
1 / \zeta_{d y n}(s)=F_{d y n}(s) / F_{0}^{2}(s) \tag{H.36}
\end{equation*}
$$

This relation implies that for flows the zeta function has double poles at the zeros of the scalar advection spectral determinant, with zeros of the dynamo spectral determinant no longer coinciding with the zeros of the scalar advection spectral determinant; Usually the leading zero of the dynamo spectral determinant is larger than the scalar advection rate, and the rate of decay of the magnetic field is no longer governed by the scalar advection.

## Commentary

Remark H. 1 Dynamo zeta. The dynamo zeta (H.34) has been introduced by Aurell and Gilbert [H.3] and reviewed in ref. [H.4]. Our exposition follows ref. [17.10].

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## Exercises

Exercise H. 1 Stretching factor. Prove the multiplicative property of the stretching factor (H.2). Why should we extend the phase space with the tangent space?

Exercise H. 2 Dynamo rate. Suppose that the fluid dynamics is highly dissipative and can be well approximated by the piecewise linear map

$$
f(x)= \begin{cases}1+a x & \text { if } \quad x<0  \tag{H.37}\\ 1-b x & \text { if } \quad x>0\end{cases}
$$

on an appropriate surface of section $(a, b>2)$. Suppose also that the return time is constant $T_{a}$ for $x<0$ and $T_{b}$ for $x>0$. Show that the dynamo zeta is

$$
\begin{equation*}
1 / \zeta_{d y n}(s)=1-e^{s T_{a}}+e^{s T_{b}} \tag{H.38}
\end{equation*}
$$

Show also that the escape rate is the leading zero of

$$
\begin{equation*}
1 / \zeta_{0}(s)=1-e^{s T_{a}} / a-e^{s T_{b}} / b \tag{H.39}
\end{equation*}
$$

Calculate the dynamo and the escape rates analytically if $b=a^{2}$ and $T_{b}=2 T_{a}$. Do the calculation for the case when you reverse the signs of the slopes of the map. What is the difference?

## Appendix I

## Discrete symmetries

## I. 1 Preliminaries and definitions

(A. Wirzba and P. Cvitanović)

In the following we will define what we mean by the concepts group, representation, symmetry of a dynamical system, and invariance.

Group axioms. First, we define a group in abstract terms: A group $G$ is a set of elements $g_{1}, g_{2}, g_{3}, \ldots$ for which a law of composition or group multiplication is given such that the product $g_{2} \circ g_{1}$ (which we will also just abbreviate as $g_{2} g_{1}$ ) of any two elements satisfies the following conditions:

1. If $g_{1}, g_{2} \in G$, then $g_{2} \circ g_{1} \in G$.
2. The group multiplication is associative: $g_{3} \circ\left(g_{2} \circ g_{1}\right)=\left(g_{3} \circ g_{2}\right) \circ g_{1}$.
3. The group $G$ contains an element $e$ called identity such that $g \circ e=$ $e \circ g=g$ for every element $g \in G$.
4. For every element $g \in G$, there exists an unique element $h \in G$ such that $h \circ g=g \circ h=e$. The element $h$ is called inverse of $g$, and is denoted by $h=g^{-1}$.

A finite group is a group with a finite number of elements

$$
G=\left\{e, g_{2}, \ldots, g_{|G|}\right\}
$$

where $|G|$, the number of elements, will be referred to as order of the group.

Matrix group on vector space. We will now apply these abstract group definitions to the set of $[d \times d]$-dimensional non-singular matrices $\mathbf{A}, \mathbf{B}, \mathbf{C}, \ldots$ acting in a $d$-dimensional vector space $V \in \mathbb{R}^{d}$, that is, the product of matrices $\mathbf{A}$ and $\mathbf{B}$ gives the single matrix $\mathbf{C}$, such that

$$
\begin{equation*}
\mathbf{C} v=\mathbf{B}(\mathbf{A} v) \in V, \quad \forall v \in V . \tag{I.1}
\end{equation*}
$$

The identity of the group is the unit matrix $\mathbf{1}$ which leaves all vectors in $V$ unchanged. Every matrix in the group has a unique inverse.

Linear representation of a group. Let us now map the abstract group $G$ homeomorphically on a group of matrices $\mathrm{D}(G)$ in the vector space $V$, that is, in such a way that the group properties, especially the group multiplication, are preserved:

1. Any $g \in G$ is mapped to a matrix $\mathbf{D}(g) \in \mathrm{D}(G)$.
2. The group product $g_{2} \circ g_{1} \in G$ is mapped onto the matrix product $\mathbf{D}\left(g_{2} \circ g_{1}\right)=\mathbf{D}\left(g_{2}\right) \mathbf{D}\left(g_{1}\right)$.
3. The associativity is preserved: $\mathbf{D}\left(g_{3} \circ\left(g_{2} \circ g_{1}\right)\right)=\mathbf{D}\left(g_{3}\right)\left(\mathbf{D}\left(g_{2}\right) \mathbf{D}\left(g_{1}\right)\right)=$ $\left(\mathbf{D}\left(g_{3}\right)\left(\mathbf{D}\left(g_{2}\right)\right) \mathbf{D}\left(g_{1}\right)\right.$.
4. The identity element $e \in G$ is mapped onto the unit matrix $\mathrm{D}(e)=\mathbf{1}$ and the inverse element $g^{-1} \in G$ is mapped onto the inverse matrix $\mathbf{D}\left(g^{-1}\right)=[\mathbf{D}(g)]^{-1} \equiv \mathbf{D}^{-\mathbf{1}}(g)$.

We call the so defined matrix group $\mathrm{D}(G)$ a linear or matrix representation of the group $G$ in the representation space $V$. Note that the matrix operation on a vector is by definition linear. We use the specification linear in order to discriminate the matrix representations from other operator representations that do not have to be linear, in general. Throughout this appendix we only consider linear representations.

If the dimensionality of $V$ is $d$, we say the representation is an $d$ dimensional representation or has the degree $d$. The matrices $\mathbf{D}(g) \in \mathrm{D}(G)$ are non-singular $[d \times d]$ matrices, which we will also just abbreviate as $\mathbf{g}$, that is, $x^{\prime}=\mathbf{g} x$ corresponds to the normal matrix operation $x_{i}^{\prime}=\sum_{j=1}^{d}(\mathbf{g})_{i j} x_{j}=$ $\sum_{j=1}^{d} g_{i j} x_{j}$.

Character of a representation. The character of $\chi_{\alpha}(g)$ of an $d$-dimensional representation $\mathbf{D}(g)$ of the group element $g$ of a discrete group $G$ is defined as trace

$$
\chi_{\alpha}(g)=\sum_{i=1}^{n} \mathbf{D}_{i i}(g) \equiv \operatorname{tr} \mathbf{D}(g) .
$$

Note especially that $\chi(e)=n$, since $\mathbf{D}_{i j}(e)=\delta_{i j}$ for $1 \leq i, j \leq n$.

Faithful representations. If the homomorphismus mapping $G$ on $\mathrm{D}(G)$ becomes an isomorphism, the representation is said to be faithful. In this case the order of the group of matrices $D(G)$ is equal to the order $|G|$ of the group. In general, however, there will be several elements of $G$ that will be mapped on the unit matrix $\mathbf{D}(e)=\mathbf{1}$. This property can be used to define a subgroup $H \subset G$ of the group $G$ consisting of all elements $h \in G$ that are mapped to the unit matrix of a given representation. Then the cosidered representation is a faithful representation of the factor group $G / H$.

Equivalent representations. From this remarks it should be clear that the representation of a group is by no means unique. If the basis in the $d$-dimensional vector space $V$ is changed, the matrices $\mathbf{D}(g)$ have to be replaced by their transformations $\mathbf{D}^{\prime}(g)$. In this case, however, the new matrices $\mathbf{D}^{\prime}(g)$ and the old matrices $\mathbf{D}(g)$ are related by an equivalence transformation through a non-singular matrix $\mathbf{C}$

$$
\mathbf{D}^{\prime}(g)=\mathbf{C} \mathbf{D}(g) \mathbf{C}^{-1} .
$$

Thus, the group of matrices $\mathbf{D}^{\prime}(g)$ form an equivalent representation $\mathrm{D}^{\prime}(G)$ to the representation $\mathrm{D}(G)$ of the group $G$. The equivalent representations have the same structure, although the matrices look different. Because of the cylic nature of the trace and because equivalent representations have the same dimension, the character of equivalent representations is the same

$$
\chi(g)=\sum_{i=1}^{n} \mathbf{D}_{i i}^{\prime}(g)=\operatorname{tr}\left(\mathbf{D}^{\prime}(g)\right)=\operatorname{tr}\left(\mathbf{C D}(g) \mathbf{C}^{-1}\right)
$$

Regular representation of a group. The regular representation of a group is a special representation that is defined as follows: If we define the elements of a finite group as $g_{1}, g_{2}, \ldots, g_{|G|}$, the multiplying from the left by any element $g_{\nu}$ permutes the $g_{1}, g_{2}, \ldots, g_{|G|}$ among themselves. We can represent the element $g_{\nu}$ by the permutations of the $|G|$ "coordinates" $g_{1}, g_{2}, \ldots, g_{|G|}$. Thus for $i, j=1, \ldots,|G|$, we define the regular representation

$$
\mathbf{D}_{i j}\left(g_{\nu}\right)= \begin{cases}\delta_{j l_{i}} & \text { if } g_{\nu} g_{i}=g_{l_{i}} \text { with } l_{i}=1, \ldots,|G|, \\ 0 & \text { otherwise }\end{cases}
$$

In this regular representation the diagonal elements of all matrices are zero except for the element $g_{\nu_{0}}$ with $g_{\nu_{0}} g_{i}=g_{i}$, that is, for the identity element $e$. So in the regular representation the character is given by

$$
\chi(g)=\left\{\begin{array}{l}
1 \text { for } g=e, \\
0 \text { for } g \neq e .
\end{array}\right.
$$

Passive and active coordinate transformations. We have to discriminate between active and passive coordinate transformations. An active (coordinate) transformation corresponds to an non-singular $d \times d$ matrix that actively shifts/changes the vector $x \in \mathcal{M}$

$$
x \rightarrow \mathbf{T} x .
$$

The corresponding passive coordinate transformation changes the coordinate system with respect to which the vector $f(x) \in \mathcal{M}$ is measured. Thus it is given by

$$
f(x) \rightarrow \mathbf{T}^{-1} f(x)=f\left(\mathbf{T}^{-1} x\right) .
$$

Note that the combination of an passive and active coordinate transformation results to the identity

$$
f(x)=\mathbf{T}^{-1} f(\mathbf{T} x) .
$$

On the other hand, the evolution operator $\mathcal{L}(x, y)$ satisfies the following identity

$$
\mathcal{L}(x, y)=\left|\operatorname{det}\left(\frac{\partial \mathbf{T} x}{\partial x}\right)\right| \mathcal{L}(\mathbf{T} x, \mathbf{T} y)=|\operatorname{det} \mathbf{T}| \mathcal{L}(\mathbf{T} x, \mathbf{T} y) .
$$

Note the appearance of $\operatorname{det} \mathbf{T}$ instead of $\operatorname{det} \mathbf{T}^{-1}$ and therefore the contravariant transformation property of $\mathcal{L}(x, y)$ in correspondence to maps $f(x)$. If the coordinate transformation $\mathbf{T}$ belongs to the linear non-singular representation of a discrete (that is, finite) symmetry group $G$, then $|\operatorname{det} \mathbf{T}|=$ 1 , since for any element $g$ of a finite group $G$, where exists a number $m$ such that

$$
g^{n} \equiv \underbrace{g \circ g \circ \ldots \circ g}_{m \text { times }}=e .
$$

Thus $\mathbf{T}$ corresponds to the $m$ th root of $\mathbf{1}$ and the modulus of its determinant is unity.

Symmetry of dynamical system. A dynamical system $(\mathcal{M}, f)$ is invariant under a discrete symmetry group $G=\left\{e, g_{2}, \ldots, g_{|G|}\right\}$, if the map $f: \mathcal{M} \rightarrow \mathcal{M}$ (or the continous flow $f^{t}$ ) from the $d$-dimensional manifold $\mathcal{M}$ into itself (with $d$ finite) is invariant

$$
f(\mathbf{g} x)=\mathbf{g} f(x)
$$

for any coordinate $x \in \mathcal{M}$ and any finite non-singular linear representation (that is, a non-singular $d \times d$ matrix) $\mathbf{g}$ of any element $g \in G$. So a symmetry for a dynamical system $(\mathcal{M}, f)$ has to satisfy the two conditions

1) $\mathrm{g} x \in \mathcal{M} \quad \forall x \in \mathcal{M}$ and $\forall g \in G$,
2) $[\mathbf{D}(g), f]=0 \quad \forall f: \mathcal{M} \rightarrow \mathcal{M}$ and $\forall g \in G$.

Group integration. Note the following laws

$$
\frac{1}{|G|} \sum_{g \in G}=1
$$

and therefore

$$
\frac{1}{|G|} \sum_{g \in G} \mathbf{D}\left(g_{i}\right)=\mathbf{D}\left(g_{i_{0}}\right), \quad i_{0} \text { fixed }
$$

However,

$$
\frac{1}{|G|} \sum_{g \in G} \mathbf{D}(g)=\mathbf{0},
$$

where $\mathbf{0}$ is the zero matrix of same dimension as the representations $\mathbf{D}(g) \in$ $D(G)$. In particular,

$$
\frac{1}{|G|} \sum_{g \in G} \chi_{\alpha}(g)=\frac{1}{|G|} \sum_{g \in G} \sum_{i=1}^{d_{\alpha}} D(g)_{i i}=0 .
$$

Furthermore, if we consider all non-equilavent irreducible representations of a group $G$, then the quantities $D_{i j}^{(\alpha)}(g)$ for fixed $\alpha, i$ and $j$

Orthonormalitity of characters. But what can we say about

$$
\frac{1}{|G|} \sum_{g \in G} \chi_{\alpha}(h g) \chi_{\alpha}\left(g^{-1} k^{-1}\right) \quad \text { with } h, k \in G \text { fixed ? }
$$

Note the following relation

$$
\delta_{a b} \delta_{c d}=\frac{1}{n} \delta_{a d} \delta_{c b}+\left(\delta_{a b} \delta_{c d}-\frac{1}{n} \delta_{a d} \delta_{c b}\right) .
$$

Projection operators. The projection operator onto the $\alpha$ irreducible subspace of dimension $d_{\alpha}$ is given by

$$
\mathbf{P}_{\alpha}=\frac{d_{\alpha}}{|G|} \sum_{g \in G} \chi_{\alpha}(g) \mathbf{g}^{-1}
$$

Note that $\mathbf{P}_{\alpha}$ is a $[d \times d]$-dimensional matrix as the representation $\mathbf{g}$.

Irreducible subspaces of the evolution operator.

$$
\mathcal{L}=\sum_{\alpha} \operatorname{tr} \mathcal{L}_{\alpha}
$$

with

$$
\mathcal{L}_{\alpha}(y, x)=\frac{d_{\alpha}}{|G|} \sum_{g \in G} \chi_{\alpha}(g) \mathcal{L}\left(\mathrm{g}^{-1} y, x\right),
$$

where the prefactor $d_{\alpha}$ reflects the fact that a $d_{\alpha}$-dimensional representation occurs $d_{\alpha}$ times.

## I. $2 C_{4 v}$ factorization

If an $N$-disk arrangement has $C_{N}$ symmetry, and the disk visitation sequence is given by disk labels $\left\{\epsilon_{1} \epsilon_{2} \epsilon_{3} \ldots\right\}$, only the relative increments $\rho_{i}=\epsilon_{i+1}-\epsilon_{i} \bmod N$ matter. Symmetries under reflections across axes increase the group to $C_{N v}$ and add relations between symbols: $\left\{\epsilon_{i}\right\}$ and $\left\{N-\epsilon_{i}\right\}$ differ only by a reflection. As a consequence of this reflection increments become decrements until the next reflection and vice versa. Consider four equal disks placed on the vertices of a square (figure I.1a). The symmetry group consists of the identity $\mathbf{e}$, the two reflections $\sigma_{x}, \sigma_{y}$ across $x$, $y$ axes, the two diagonal reflections $\sigma_{13}, \sigma_{24}$, and the three rotations $C_{4}, C_{2}$ and $C_{4}^{3}$ by angles $\pi / 2, \pi$ and $3 \pi / 2$. We start by exploiting the $C_{4}$ subgroup symmetry in order to replace the absolute labels $\epsilon_{i} \in\{1,2,3,4\}$ by relative increments $\rho_{i} \in\{1,2,3\}$. By reflection across diagonals, an increment by 3 is equivalent to an increment by 1 and a reflection; this new symbol will be called $\underline{1}$. Our convention will be to first perform the increment and then to change the orientation due to the reflection. As an example, consider the fundamental domain cycle 112. Taking the disk $1 \rightarrow$ disk 2 segment as the starting segment, this symbol string is mapped into the disk visitation sequence $1_{+1} 2_{+1} 3_{+2} 1 \ldots=\overline{123}$, where the subscript indicates the increments (or decrements) between neighboring symbols; the period of the cycle $\overline{112}$ is thus 3 in both the fundamental domain and the full space. Similarly, the cycle $\overline{112}$ will be mapped into $1_{+1} 2_{-1} 1_{-2} 3_{-1} 2_{+1} 3_{+2} 1=\overline{121323}$ (note that


Figure I.1: (a) The symmetries of four disks on a square. (b) The symmetries of four disks on a rectangle. The fundamental domains are indicated by the shaded wedges.
the fundamental domain symbol $\underline{1}$ corresponds to a flip in orientation after the second and fifth symbols); this time the period in the full space is twice that of the fundamental domain. In particular, the fundamental domain fixed points correspond to the following 4-disk cycles:

| 4-disk |  | reduced |
| :--- | :--- | ---: |
| 12 | $\leftrightarrow$ | $\frac{1}{1}$ |
| 1234 | $\leftrightarrow$ |  |
| 13 | $\leftrightarrow$ | 2 |

Conversions for all periodic orbits of reduced symbol period less than 5 are listed in table I.1.

This symbolic dynamics is closely related to the group-theoretic structure of the dynamics: the global 4-disk trajectory can be generated by mapping the fundamental domain trajectories onto the full 4 -disk space by the accumulated product of the $C_{4 v}$ group elements $g_{1}=C, g_{2}=C^{2}$, $g_{\underline{1}}=\sigma_{\text {diag }} C=\sigma_{\text {axis }}$, where $C$ is a rotation by $\pi / 2$. In the $\overline{112}$ example worked out above, this yields $g_{\underline{1} 12}=g_{2} g_{1} g_{\underline{1}}=C^{2} C \sigma_{\text {axis }}=\sigma_{\text {diag }}$, listed in the last column of table I.1. Our convention is to multiply group elements in the reverse order with respect to the symbol sequence. We need these group elements for our next step, the dynamical zeta function factorizations.

The $C_{4 v}$ group has four one-dimensional representations, either symmetric $\left(A_{1}\right)$ or antisymmetric ( $A_{2}$ ) under both types of reflections, or symmetric under one and antisymmetric under the other $\left(B_{1}, B_{2}\right)$, and a degenerate pair of two-dimensional representations $E$. Substituting the $C_{4 v}$ characters

| $C_{4 v}$ | $A_{1}$ | $A_{2}$ | $B_{1}$ | $B_{2}$ | $E$ |
| :---: | ---: | ---: | ---: | ---: | ---: |
| $e$ | 1 | 1 | 1 | 1 | 2 |
| $C_{2}$ | 1 | 1 | 1 | 1 | -2 |
| $C_{4}, C_{4}^{3}$ | 1 | 1 | -1 | -1 | 0 |
| $\sigma_{\text {axes }}$ | 1 | -1 | 1 | -1 | 0 |
| $\sigma_{\text {diag }}$ | 1 | -1 | -1 | 1 | 0 |


|  |  |  | $\tilde{p}$ | $p$ | $\mathbf{h}_{\tilde{p}}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | $\begin{array}{\|l\|} \hline 0001 \\ 0002 \end{array}$ | $\begin{aligned} & 12121414 \\ & 12124343 \end{aligned}$ | $\sigma_{24}$$\sigma_{y}$ |
| $\tilde{p}$ | $p$ | $\mathbf{h}_{\tilde{p}}$ |  |  |  |
| 0 | 12 | $\sigma_{x}$ | 0011 | 12123434 | $C_{2}$ |
| 1 | 1234 | $C_{4}$ | 0012 | 1212414134342323 | $C_{4}^{3}$ |
| 2 | 13 | $C_{2}, \sigma_{13}$ | 0021 (a) | 1213414234312324 | $C_{4}^{3}$ |
| 01 | 1214 | $\sigma_{24}$ | 0022 | 1213 | $e$ |
| 02 | 1243 | $\sigma_{y}$ | 0102 (a) | 1214232134324143 | $C_{4}$ |
| 12 | 12413423 | $C_{4}^{3}$ | 0111 | 12143234 | $\sigma_{13}$ |
| 001 | 121232343414 | $\mathrm{C}_{4}$ | 0112 (b) | 12142123 | $\sigma_{x}$ |
| 002 | 121343 | $C_{2}$ | 0121 (b) | 12132124 | $\sigma_{x}$ |
| 011 | 121434 | $\sigma_{y}$ | 0122 | 12131413 | $\sigma_{24}$ |
| 012 | 121323 | $\sigma_{13}$ | 0211 | 12432134 | $\sigma_{x}$ |
| 021 | 124324 | $\sigma_{13}$ | 0212 | 12431423 | $\sigma_{24}$ |
| 022 | 124213 | $\sigma_{x}$ | 0221 | 12421424 | $\sigma_{24}$ |
| 112 | 123 | $e$ | 0222 | 12424313 | $\sigma_{y}$ |
| 122 | 124231342413 | $C_{4}$ | 1112 | 1234234134124123 | $C_{4}$ |
| 1122 12313413 <br> 1222 1242413134242313 |  |  |  |  | $C_{2}$ |
|  |  |  |  |  | $C_{4}^{3}$ |

Table I.1: $C_{4 v}$ correspondence between the ternary fundamental domain prime cycles $\tilde{p}$ and the full 4-disk $\{1,2,3,4\}$ labeled cycles $p$, together with the $C_{4 v}$ transformation that maps the end point of the $\tilde{p}$ cycle into an irreducible segment of the $p$ cycle. For typographical convenience, the symbol 1 of sect. I. 2 has been replaced by 0 , so that the ternary alphabet is $\{0,1,2\}$. The degeneracy of the $p$ cycle is $m_{p}=8 n_{\tilde{p}} / n_{p}$. Orbit $\overline{2}$ is the sole boundary orbit, invariant both under a rotation by $\pi$ and a reflection across a diagonal. The two pairs of cycles marked by $(a)$ and $(b)$ are related by time reversal, but cannot be mapped into each other by $C_{4 v}$ transformations.
into (22.15) we obtain:

$$
\begin{aligned}
& \begin{array}{llllll}
h_{\tilde{p}} & A_{1} & A_{2} & B_{1} & B_{2} & E
\end{array} \\
& e:\left(1-t_{\tilde{p}}\right)^{8}=\left(1-t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)^{4} \\
& C_{2}:\left(1-t_{\tilde{p}}^{2}\right)^{4}=\left(1-t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1+t_{\tilde{p}}\right)^{4} \\
& C_{4}, C_{4}^{3}:\left(1-t_{\tilde{p}}^{4}\right)^{2}=\left(1-t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1+t_{\tilde{p}}\right)\left(1+t_{\tilde{p}}\right)\left(1+t_{\tilde{p}}^{2}\right)^{2} \\
& \sigma_{\text {axes }}:\left(1-t_{\tilde{p}}^{2}\right)^{4}=\left(1-t_{\tilde{p}}\right)\left(1+t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1+t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}^{2}\right)^{2} \\
& \sigma_{\text {diag }}:\left(1-t_{\tilde{p}}^{2}\right)^{4}=\left(1-t_{\tilde{p}}\right)\left(1+t_{\tilde{p}}\right)\left(1+t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}^{2}\right)^{2}
\end{aligned}
$$

The possible irreducible segment group elements $\mathbf{h}_{\tilde{p}}$ are listed in the first column; $\sigma_{\text {axes }}$ denotes a reflection across either the x-axis or the y-axis, and $\sigma_{\text {diag }}$ denotes a reflection across a diagonal (see figure I.1a). In addition, degenerate pairs of boundary orbits can run along the symmetry lines in the full space, with the fundamental domain group theory weights $\mathbf{h}_{p}=$ $\left(C_{2}+\sigma_{x}\right) / 2$ (axes) and $\mathbf{h}_{p}=\left(C_{2}+\sigma_{13}\right) / 2$ (diagonals) respectively:

$$
\begin{array}{rccc} 
& A_{1} \quad A_{2} \quad B_{1} \quad B_{2} \quad E \\
\text { axes: } & \left(1-t_{\tilde{\tilde{p}}} \tilde{p}^{2}=\right. & \left(1-t_{\tilde{p}}\right)\left(1-0 t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1-0 t_{\tilde{p}}\right)\left(1+t_{\tilde{p}}\right)^{2} \\
\text { diagonals: }\left(1-t_{\tilde{p}}^{2}\right)^{2}= & \left.\left.\left(1-t_{\tilde{p}}\right)\left(1-0 t_{\tilde{p}}\right)\left(1-0 t_{\tilde{p}}\right)\left(1-t_{\tilde{p}}\right)\left(1+t_{(1)}\right)\right)^{2}\right)
\end{array}
$$

(we have assumed that $t_{\tilde{p}}$ does not change sign under reflections across symmetry axes). For the 4 -disk arrangement considered here only the diagonal
orbits $\overline{13}, \overline{24}$ occur; they correspond to the $\overline{2}$ fixed point in the fundamental domain.

The $A_{1}$ subspace in $C_{4 v}$ cycle expansion is given by

$$
\begin{align*}
1 / \zeta_{A_{1}}= & \left(1-t_{0}\right)\left(1-t_{1}\right)\left(1-t_{2}\right)\left(1-t_{01}\right)\left(1-t_{02}\right)\left(1-t_{12}\right) \\
& \left(1-t_{001}\right)\left(1-t_{002}\right)\left(1-t_{011}\right)\left(1-t_{012}\right)\left(1-t_{021}\right)\left(1-t_{022}\right)\left(1-t_{112}\right) \\
& \left(1-t_{122}\right)\left(1-t_{0001}\right)\left(1-t_{0002}\right)\left(1-t_{0011}\right)\left(1-t_{0012}\right)\left(1-t_{0021}\right) \ldots \\
= & 1-t_{0}-t_{1}-t_{2}-\left(t_{01}-t_{0} t_{1}\right)-\left(t_{02}-t_{0} t_{2}\right)-\left(t_{12}-t_{1} t_{2}\right) \\
& -\left(t_{001}-t_{0} t_{01}\right)-\left(t_{002}-t_{0} t_{02}\right)-\left(t_{011}-t_{1} t_{01}\right) \\
& -\left(t_{022}-t_{2} t_{02}\right)-\left(t_{112}-t_{1} t_{12}\right)-\left(t_{122}-t_{2} t_{12}\right) \\
& -\left(t_{012}+t_{021}+t_{0} t_{1} t_{2}-t_{0} t_{12}-t_{1} t_{02}-t_{2} t_{01}\right) \ldots \tag{I.3}
\end{align*}
$$

(for typographical convenience, $\underline{1}$ is replaced by 0 in the remainder of this section). For one-dimensional representations, the characters can be read off the symbol strings: $\chi_{A_{2}}\left(\mathbf{h}_{\tilde{\mathbf{p}}}\right)=(-1)^{n_{0}}, \chi_{B_{1}}\left(\mathbf{h}_{\tilde{\mathbf{p}}}\right)=(-1)^{n_{1}}, \chi_{B_{2}}\left(\mathbf{h}_{\tilde{\mathbf{p}}}\right)=$ $(-1)^{n_{0}+n_{1}}$, where $n_{0}$ and $n_{1}$ are the number of times symbols 0,1 appear in the $\tilde{p}$ symbol string. For $B_{2}$ all $t_{p}$ with an odd total number of 0 's and 1's change sign:

$$
\begin{align*}
1 / \zeta_{B_{2}}= & \left(1+t_{0}\right)\left(1+t_{1}\right)\left(1-t_{2}\right)\left(1-t_{01}\right)\left(1+t_{02}\right)\left(1+t_{12}\right) \\
& \left(1+t_{001}\right)\left(1-t_{002}\right)\left(1+t_{011}\right)\left(1-t_{012}\right)\left(1-t_{021}\right)\left(1+t_{022}\right)\left(1-t_{112}\right) \\
& \left(1+t_{122}\right)\left(1-t_{0001}\right)\left(1+t_{0002}\right)\left(1-t_{0011}\right)\left(1+t_{0012}\right)\left(1+t_{0021}\right) \ldots \\
= & 1+t_{0}+t_{1}-t_{2}-\left(t_{01}-t_{0} t_{1}\right)+\left(t_{02}-t_{0} t_{2}\right)+\left(t_{12}-t_{1} t_{2}\right) \\
& +\left(t_{001}-t_{0} t_{01}\right)-\left(t_{002}-t_{0} t_{02}\right)+\left(t_{011}-t_{1} t_{01}\right) \\
& +\left(t_{022}-t_{2} t_{02}\right)-\left(t_{112}-t_{1} t_{12}\right)+\left(t_{122}-t_{2} t_{12}\right) \\
& -\left(t_{012}+t_{021}+t_{0} t_{1} t_{2}-t_{0} t_{12}-t_{1} t_{02}-t_{2} t_{01}\right) \ldots \tag{I.4}
\end{align*}
$$

The form of the remaining cycle expansions depends crucially on the special role played by the boundary orbits: by (I.2) the orbit $t_{2}$ does not contribute to $A_{2}$ and $B_{1}$,

$$
\begin{align*}
1 / \zeta_{A_{2}}= & \left(1+t_{0}\right)\left(1-t_{1}\right)\left(1+t_{01}\right)\left(1+t_{02}\right)\left(1-t_{12}\right) \\
& \left(1-t_{001}\right)\left(1-t_{002}\right)\left(1+t_{011}\right)\left(1+t_{012}\right)\left(1+t_{021}\right)\left(1+t_{022}\right)\left(1-t_{112}\right) \\
& \left(1-t_{122}\right)\left(1+t_{0001}\right)\left(1+t_{0002}\right)\left(1-t_{0011}\right)\left(1-t_{0012}\right)\left(1-t_{0021}\right) \ldots \\
= & 1+t_{0}-t_{1}+\left(t_{01}-t_{0} t_{1}\right)+t_{02}-t_{12} \\
& -\left(t_{001}-t_{0} t_{01}\right)-\left(t_{002}-t_{0} t_{02}\right)+\left(t_{011}-t_{1} t_{01}\right) \\
& +t_{022}-t_{122}-\left(t_{112}-t_{1} t_{12}\right)+\left(t_{012}+t_{021}-t_{0} t_{12}-t_{1} t_{02}\right) \ldots \text { (I.5) } \tag{I.5}
\end{align*}
$$

and

$$
\begin{aligned}
1 / \zeta_{B_{1}}= & \left(1-t_{0}\right)\left(1+t_{1}\right)\left(1+t_{01}\right)\left(1-t_{02}\right)\left(1+t_{12}\right) \\
& \left(1+t_{001}\right)\left(1-t_{002}\right)\left(1-t_{011}\right)\left(1+t_{012}\right)\left(1+t_{021}\right)\left(1-t_{022}\right)\left(1-t_{112}\right)
\end{aligned}
$$

$$
\begin{align*}
& \left(1+t_{122}\right)\left(1+t_{0001}\right)\left(1-t_{0002}\right)\left(1-t_{0011}\right)\left(1+t_{0012}\right)\left(1+t_{0021}\right) \ldots \\
= & 1-t_{0}+t_{1}+\left(t_{01}-t_{0} t_{1}\right)-t_{02}+t_{12} \\
& +\left(t_{001}-t_{0} t_{01}\right)-\left(t_{002}-t_{0} t_{02}\right)-\left(t_{011}-t_{1} t_{01}\right) \\
& -t_{022}+t_{122}-\left(t_{112}-t_{1} t_{12}\right)+\left(t_{012}+t_{021}-t_{0} t_{12}-t_{1} t_{02}\right) \ldots(\text { I. } 6 \tag{I.6}
\end{align*}
$$

In the above we have assumed that $t_{2}$ does not change sign under $C_{4 v}$ reflections. For the mixed-symmetry subspace $E$ the curvature expansion is given by

$$
\begin{align*}
1 / \zeta_{E}= & 1+t_{2}+\left(-t_{0}{ }^{2}+t_{1}^{2}\right)+\left(2 t_{002}-t_{2} t_{0}{ }^{2}-2 t_{112}+t_{2} t_{1}{ }^{2}\right) \\
& +\left(2 t_{0011}-2 t_{0022}+2 t_{2} t_{002}-t_{01}{ }^{2}-t_{02}{ }^{2}+2 t_{1122}-2 t_{2} t_{112}\right. \\
& \left.+t_{12}{ }^{2}-t_{0}{ }^{2} t_{1}^{2}\right)+\left(2 t_{00002}-2 t_{00112}+2 t_{2} t_{0011}-2 t_{00121}-2 t_{00211}\right. \\
& +2 t_{00222}-2 t_{2} t_{0022}+2 t_{01012}+2 t_{01021}-2 t_{01102}-t_{2} t_{01}{ }^{2}+2 t_{002022} \\
& -t_{2} t_{02}^{2}+2 t_{11112}-2 t_{11222}+2 t_{2} t_{1122}-2 t_{12122}+t_{2} t_{12}{ }^{2}-t_{2} t_{0}{ }^{2} t_{1}{ }^{2} \\
& \left.+2 t_{002}\left(-t_{0}{ }^{2}+t_{1}{ }^{2}\right)-2 t_{112}\left(-t_{0}{ }^{2}+t_{1}{ }^{2}\right)\right) \tag{I.7}
\end{align*}
$$

A quick test of the $\zeta=\zeta_{A_{1}} \zeta_{A_{2}} \zeta_{B_{1}} \zeta_{B_{2}} \zeta_{E}^{2}$ factorization is afforded by the topological polynomial; substituting $t_{p}=z^{n_{p}}$ into the expansion yields

$$
1 / \zeta_{A_{1}}=1-3 z, \quad 1 / \zeta_{A_{2}}=1 / \zeta_{B_{1}}=1, \quad 1 / \zeta_{B_{2}}=1 / \zeta_{E}=1+z
$$

in agreement with (13.42).

Remark I. 1 Labelling conventions While there is a variety of labelling conventions [23.15, 22.13] for the reduced $C_{4 v}$ dynamics, we prefer the one introduced here because of its close relation to the group-theoretic structure of the dynamics: the global 4 -disk trajectory can be generated by mapping the fundamental domain trajectories onto the full 4 -disk space by the accumulated product of the $C_{4 v}$ group elements.

## I. $3 C_{2 v}$ factorization

An arrangement of four identical disks on the vertices of a rectangle has $C_{2 v}$ symmetry (figure I.1b). $C_{2 v}$ consists of $\left\{e, \sigma_{x}, \sigma_{y}, C_{2}\right\}$, that is, the reflections across the symmetry axes and a rotation by $\pi$.

This system affords a rather easy visualization of the conversion of a 4 -disk dynamics into a fundamental domain symbolic dynamics. An orbit leaving the fundamental domain through one of the axis may be folded back by a reflection on that axis; with these symmetry operations $g_{0}=$ $\sigma_{x}$ and $g_{1}=\sigma_{y}$ we associate labels 1 and 0 , respectively. Orbits going to the diagonally opposed disk cross the boundaries of the fundamental domain twice; the product of these two reflections is just $C_{2}=\sigma_{x} \sigma_{y}$, to

|  |  |  | $\tilde{p}$ | $p$ | g |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | 0001 | $\begin{aligned} & 14143232 \\ & 14142323 \end{aligned}$ | $C_{2}$ |
| $\tilde{p}$ | $p$ | g |  |  |  |
| 0 | 14 | $\sigma_{y}$ | 0011 | 1412 | $e$ |
| 1 | 12 | $\sigma_{x}$ | 0012 | 14124143 | $\sigma_{y}$ |
| 2 | 13 | $C_{2}$ | 0021 | 14134142 | $\sigma_{y}$ |
| 01 | 1432 | $\mathrm{C}_{2}$ | 0022 | 1413 | $e^{\text {a }}$ |
| 02 | 1423 | $\sigma_{x}$ | 0102 | 14324123 | $\sigma_{y}$ |
| 12 | 1243 | $\sigma_{y}$ | 0111 | 14343212 | $C_{2}$ |
| 001 | 141232 | $\sigma_{x}$ | 0112 | 14342343 | $\sigma_{x}$ |
| 002 | 141323 | $C_{2}$ | 0121 | 14312342 | $\sigma_{x}$ |
| 011 | 143412 | $\sigma_{y}$ | 0122 | 14313213 | $C_{2}$ |
| 012 | 143 | $e$ | 0211 | 14212312 | $\sigma_{x}$ |
| 021 | 142 | $e$ | 0212 | 14213243 | $C_{2}$ |
| 022 | 142413 | $\sigma_{y}$ | 0221 | 14243242 | $C_{2}$ |
| 112 | 121343 | $C_{2}$ | 0222 | 14242313 | $\sigma_{x}$ |
| 122 | 124213 | $\sigma_{x}$ | 1112 | 12124343 | $\sigma_{y}$ |
|  |  |  | 1122 | 1213 | $e^{\text {a }}$ |
|  |  |  | 1222 | 12424313 | $\sigma_{y}$ |

Table I.2: $C_{2 v}$ correspondence between the ternary $\{0,1,2\}$ fundamental domain prime cycles $\tilde{p}$ and the full 4-disk $\{1,2,3,4\}$ cycles $p$, together with the $C_{2 v}$ transformation that maps the end point of the $\tilde{p}$ cycle into an irreducible segment of the $p$ cycle. The degeneracy of the $p$ cycle is $m_{p}=4 n_{\tilde{p}} / n_{p}$. Note that the 012 and 021 cycles are related by time reversal, but cannot be mapped into each other by $C_{2 v}$ transformations. The full space orbit listed here is generated from the symmetry reduced code by the rules given in sect. I.3, starting from disk 1.
which we assign the label 2. For example, a ternary string $0010201 \ldots$ is converted into $12143123 \ldots$, and the associated group-theory weight is given by $\ldots g_{1} g_{0} g_{2} g_{0} g_{1} g_{0} g_{0}$.

Short ternary cycles and the corresponding 4-disk cycles are listed in table I.2. Note that already at length three there is a pair of cycles $(012=143$ and $021=142$ ) related by time reversal, but not by any $C_{2 v}$ symmetries.

The above is the complete description of the symbolic dynamics for 4 sufficiently separated equal disks placed at corners of a rectangle. However, if the fundamental domain requires further partitioning, the ternary description is insufficient. For example, in the stadium billiard fundamental domain one has to distinguish between bounces off the straight and the curved sections of the billiard wall; in that case five symbols suffice for constructing the covering symbolic dynamics.

The group $C_{2 v}$ has four one-dimensional representations, distinguished by their behavior under axis reflections. The $A_{1}$ representation is symmetric with respect to both reflections; the $A_{2}$ representation is antisymmetric with respect to both. The $B_{1}$ and $B_{2}$ representations are symmetric under one and antisymmetric under the other reflection. The character table is

| $C_{2 v}$ | $A_{1}$ | $A_{2}$ | $B_{1}$ | $B_{2}$ |
| :---: | :---: | ---: | ---: | ---: |
| $e$ | 1 | 1 | 1 | 1 |
| $C_{2}$ | 1 | 1 | -1 | -1 |
| $\sigma_{x}$ | 1 | -1 | 1 | -1 |
| $\sigma_{y}$ | 1 | -1 | -1 | 1 |

Substituted into the factorized determinant (22.14), the contributions of periodic orbits split as follows

$$
\begin{array}{rccccc}
g_{\tilde{p}} & & A_{1} & A_{2} & B_{1} & B_{2} \\
e: & \left(1-t_{\tilde{p}}\right)^{4} & =\left(1-t_{\tilde{p}}\right) & \left(1-t_{\tilde{p}}\right) & \left(1-t_{\tilde{p}}\right) & \left(1-t_{\tilde{p}}\right) \\
C_{2}: & \left(1-t_{\tilde{p}}^{2}\right)^{2} & =\left(1-t_{\tilde{p}}\right) & \left(1-t_{\tilde{p}}\right) & \left(1-t_{\tilde{p}}\right) & \left(1-t_{\tilde{p}}\right) \\
\sigma_{x}: & \left(1-t_{\tilde{\tilde{2}})^{2}}=\left(\begin{array}{ll}
\left(1-t_{\tilde{\tilde{p}}}\right) & \left(1+t_{\tilde{p}}\right) \\
\sigma_{y}: & \left(1-t_{\tilde{p}}^{2}\right)^{2}
\end{array}=\left(1-t_{\tilde{\tilde{p}}}\right)\right.\right. & \left(1+t_{\tilde{p}}\right) & \left(1-t_{\tilde{p}}\right) & \left(1+t_{\tilde{p}}\right) & \left(1+t_{\tilde{p}}\right) \\
\left(1-t_{\tilde{p}}\right)
\end{array}
$$

Cycle expansions follow by substituting cycles and their group theory factors from table I.2. For $A_{1}$ all characters are +1 , and the corresponding cycle expansion is given in (I.3). Similarly, the totally antisymmetric subspace factorization $A_{2}$ is given by (I.4), the $B_{2}$ factorization of $C_{4 v}$. For $B_{1}$ all $t_{p}$ with an odd total number of 0 's and 2 's change sign:

$$
\begin{align*}
1 / \zeta_{B_{1}}= & \left(1+t_{0}\right)\left(1-t_{1}\right)\left(1+t_{2}\right)\left(1+t_{01}\right)\left(1-t_{02}\right)\left(1+t_{12}\right) \\
& \left(1-t_{001}\right)\left(1+t_{002}\right)\left(1+t_{011}\right)\left(1-t_{012}\right)\left(1-t_{021}\right)\left(1+t_{022}\right)\left(1+t_{112}\right) \\
& \left(1-t_{122}\right)\left(1+t_{0001}\right)\left(1-t_{0002}\right)\left(1-t_{0011}\right)\left(1+t_{0012}\right)\left(1+t_{0021}\right) \ldots \\
= & 1+t_{0}-t_{1}+t_{2}+\left(t_{01}-t_{0} t_{1}\right)-\left(t_{02}-t_{0} t_{2}\right)+\left(t_{12}-t_{1} t_{2}\right) \\
& -\left(t_{001}-t_{0} t_{01}\right)+\left(t_{002}-t_{0} t_{02}\right)+\left(t_{011}-t_{1} t_{01}\right) \\
& +\left(t_{022}-t_{2} t_{02}\right)+\left(t_{112}-t_{1} t_{12}\right)-\left(t_{122}-t_{2} t_{12}\right) \\
& -\left(t_{012}+t_{021}+t_{0} t_{1} t_{2}-t_{0} t_{12}-t_{1} t_{02}-t_{2} t_{01}\right) \ldots \tag{I.8}
\end{align*}
$$

For $B_{2}$ all $t_{p}$ with an odd total number of 1 's and 2 's change sign:

$$
\begin{align*}
1 / \zeta_{B_{2}}= & \left(1-t_{0}\right)\left(1+t_{1}\right)\left(1+t_{2}\right)\left(1+t_{01}\right)\left(1+t_{02}\right)\left(1-t_{12}\right) \\
& \left(1+t_{001}\right)\left(1+t_{002}\right)\left(1-t_{011}\right)\left(1-t_{012}\right)\left(1-t_{021}\right)\left(1-t_{022}\right)\left(1+t_{112}\right) \\
& \left(1+t_{122}\right)\left(1+t_{0001}\right)\left(1+t_{0002}\right)\left(1-t_{0011}\right)\left(1-t_{0012}\right)\left(1-t_{0021}\right) \ldots \\
= & 1-t_{0}+t_{1}+t_{2}+\left(t_{01}-t_{0} t_{1}\right)+\left(t_{02}-t_{0} t_{2}\right)-\left(t_{12}-t_{1} t_{2}\right) \\
& +\left(t_{001}-t_{0} t_{01}\right)+\left(t_{002}-t_{0} t_{02}\right)-\left(t_{011}-t_{1} t_{01}\right) \\
& -\left(t_{022}-t_{2} t_{02}\right)+\left(t_{112}-t_{1} t_{12}\right)+\left(t_{122}-t_{2} t_{12}\right) \\
& -\left(t_{012}+t_{021}+t_{0} t_{1} t_{2}-t_{0} t_{12}-t_{1} t_{02}-t_{2} t_{01}\right) \ldots \tag{I.9}
\end{align*}
$$

Note that all of the above cycle expansions group long orbits together with their pseudoorbit shadows, so that the shadowing arguments for convergence still apply.

The topological polynomial factorizes as

$$
\frac{1}{\zeta_{A_{1}}}=1-3 z \quad, \quad \frac{1}{\zeta_{A_{2}}}=\frac{1}{\zeta_{B_{1}}}=\frac{1}{\zeta_{B_{2}}}=1+z
$$

consistent with the 4-disk factorization (13.42).

Remark I. $2 C_{2 v}$ symmetry $\quad C_{2 v}$ is the symmetry of several systems studied in the literature, such as the stadium billiard [6.9], and the 2-dimensional anisotropic Kepler potential [30.6].

## I. 4 Hénon map symmetries

We note here a few simple symmetries of the Hénon map (3.12). For $b \neq 0$ the Hénon map is reversible: the backward iteration of (3.13) is given by

$$
\begin{equation*}
x_{n-1}=-\frac{1}{b}\left(1-a x_{n}^{2}-x_{n+1}\right) . \tag{I.10}
\end{equation*}
$$

Hence the time reversal amounts to $b \rightarrow 1 / b, a \rightarrow a / b^{2}$ symmetry in the parameter plane, together with $x \rightarrow-x / b$ in the coordinate plane, and there is no need to explore the $(a, b)$ parameter plane outside the strip $b \in\{-1,1\}$. For $b=-1$ the map is orientation and area preserving (see (20.1) below),

$$
\begin{equation*}
x_{n-1}=1-a x_{n}^{2}-x_{n+1}, \tag{I.11}
\end{equation*}
$$

the backward and the forward iteration are the same, and the non-wandering set is symmetric across the $x_{n+1}=x_{n}$ diagonal. This is one of the simplest models of a Poincaré return map for a Hamiltonian flow. For the orientation reversing $b=1$ case we have

$$
\begin{equation*}
x_{n-1}=1-a x_{n}^{2}+x_{n+1}, \tag{I.12}
\end{equation*}
$$

and the non-wandering set is symmetric across the $x_{n+1}=-x_{n}$ diagonal.

## I. 5 Symmetries of the symbol square

## $\bullet$ advanced section •

Depending on the type of dynamical system, the symbol square might have a variety of symmetries. Under the time reversal

$$
\cdots s_{-2} s_{-1} s_{0} \cdot s_{1} s_{2} s_{3} \cdots \rightarrow \cdots s_{3} s_{2} s_{1} \cdot s_{0} s_{-1} s_{-2} \cdots
$$

the points in the symbol square for an orientation preserving map are symmetric across the diagonal $\gamma=\delta$, and for the orientation reversing case they are symmetric with respect to the $\gamma=1-\delta$ diagonal. Consequently the periodic orbits appear either in dual pairs $p=s_{1} s_{2} s_{3} \ldots s_{n}$, $\bar{p}=s_{n} s_{n-1} s_{n-2} \ldots s_{1}$, or are self-dual under time reversal, $S_{p}=S_{\bar{p}}$. For the orientation preserving case a self-dual cycle of odd period has at least one point on the symmetry diagonal. In particular, all fixed points lie on the symmetry diagonal. Determination of such symmetry lines can be of considerable practical utility, as it reduces some of the periodic orbit searches to 1-dimensional searches.

Remark I. 3 Symmetries of the symbol square. For a more detailed discussion of the symbolic dynamics symmetries, see refs. [3.5, 12.46].

## Appendix J

## Convergence of spectral determinants

## J. 1 Curvature expansions: geometric picture

If you has some experience with numerical estimates of fractal dimensions, you will note that the numerical convergence of cycle expansions for systems such as the 3 -disk game of pinball, table 18.2, is very impressive; only three input numbers (the two fixed points $\overline{0}, \overline{1}$ and the 2-cycle $\overline{10}$ ) already yield the escape rate to 4 significant digits! We have omitted an infinity of unstable cycles; so why does approximating the dynamics by a finite number of cycles work so well?

Looking at the cycle expansions simply as sums of unrelated contributions is not specially encouraging: the cycle expansion (18.2) is not absolutely convergent in the sense of Dirichlet series of sect. 18.5, so what one makes of it depends on the way the terms are arranged.

The simplest estimate of the error introduced by approximating smooth flow by periodic orbits is to think of the approximation as a tessalation of a smooth curve by piecewise linear tiles, figure 1.9.

## J.1.1 Tessalation of a smooth flow by cycles

One of the early high accuracy computations of $\pi$ was due to Euler. Euler computed the circumference of the circee of unit radius by inscribing into it a regular polygon with N sides; the error of such computation is proportional to $1-\cos (2 \pi / N) \propto N^{-2}$. In a periodic orbit tessalation of a smooth flow, we cover the phase space by $e^{h n}$ tiles at the $n$th level of resolution, where $h$ is the topological entropy, the growth rate of the number of tiles. Hence we expect the error in approximating a smooth flow by $e^{h n}$ linear segments to be exponentially small, of order $N^{-2} \propto e^{-2 h n}$.

## J.1.2 Shadowing and convergence of curvature expansions

We have shown in chapter 13 that if the symbolic dynamics is defined by a finite grammar, a finite number of cycles, let us say the first $k$ terms in the cycle expansion are necessary to correctly count the pieces of the Cantor set generated by the dynamical system.

They are composed of products of non-intersecting loops on the Markov graph, see (13.13). We refer to this set of non-intersecting loops as the fundamental cycles of the strange set. It is only after these terms have been included that the cycle expansion is expected to converge smoothly, that is, only for $n>k$ are the curvatures $c_{n}$ in (9.2??) a measure of the variation of the quality of a linearized covering of the dynamical Cantor set by the length $n$ cycles, and expected to fall off rapidly with $n$.

The rate of fall-off of the cycle expansion coefficients can be estimated by observing that for subshifts of finite type the contributions from longer orbits in curvature expansions such as (18.5) can always be grouped into shadowing combinations of pseudo-cycles. For example, a cycle with itinerary $\overline{a b}=s_{1} s_{2} \cdots s_{n}$ will appear in combination of form

$$
1 / \zeta=1-\cdots-\left(t_{a b}-t_{a} t_{b}\right)-\cdots,
$$

with $\overline{a b}$ shadowed by cycle $\bar{a}$ followed by cycle $\bar{b}$, where $a=s_{1} s_{2} \cdots s_{m}$, $b=s_{m+1} \cdots s_{n-1} s_{n}$, and $s_{k}$ labels the Markov partition $\mathcal{M}_{s_{k}}$ (11.4) that the trajectory traverses at the $k$ th return. If the two trajectories coincide in the first $m$ symbols, at the $m$ th return to a Poincaré section they can land anywhere in the phase space $\mathcal{M}$

$$
\left|f^{T_{a}}\left(x_{a}\right)-f^{T_{a \ldots}}\left(x_{a \ldots}\right)\right| \approx 1,
$$

where we have assumed that the $\mathcal{M}$ is compact, and that the maximal possible separation across $\mathcal{M}$ is $O(1)$. Here $x_{a}$ is a point on the $\bar{a}$ cycle of period $T_{a}$, and $x_{a \ldots . .}$ is a nearby point whose trajectory tracks the cycle $\bar{a}$ for the first $m$ Poincaré section returns completed at the time $T_{a \ldots \ldots}$. An estimate of the maximal separation of the initial points of the two neighboring trajectories is achieved by Taylor expanding around $x_{a \ldots}=x_{\bar{a}}+\delta x_{a \ldots}$

$$
f^{T_{a}}\left(x_{\bar{a}}\right)-f^{T_{a \ldots}}\left(x_{a \ldots}\right) \approx \frac{\partial f^{T_{a}}\left(x_{\bar{a}}\right)}{\partial x} \cdot \delta x_{a \ldots}=\mathbf{J}_{a} \cdot \delta x_{a \ldots},
$$

hence the hyperbolicity of the flow forces the initial points of neighboring trajectories that track each other for at least $m$ consecutive symbols to lie exponentially close

$$
\left|\delta x_{a \ldots}\right| \propto \frac{1}{\left|\Lambda_{a}\right|} .
$$

Similarly, for any observable (10.1) integrated along the two nearby trajectories

$$
A^{T_{a \ldots}}\left(x_{a \ldots}\right) \approx A^{T_{a}}\left(x_{\bar{a}}\right)+\left.\frac{\partial A^{T_{a}}}{\partial x}\right|_{x=x_{\bar{a}}} \cdot \delta x_{a \ldots}
$$

so

$$
\left|A^{T_{a \ldots}}\left(x_{a \ldots}\right)-A^{T_{a}}\left(x_{\bar{a}}\right)\right| \propto \frac{T_{a} \mathrm{Const}}{\left|\Lambda_{a}\right|}
$$

As the time of return is itself an integral along the trajectory, return times of nearby trajectories are exponentially close

$$
\left|T_{a \ldots}-T_{a}\right| \propto \frac{T_{a} \text { Const }}{\left|\Lambda_{a}\right|}
$$

and so are the trajectory stabilities

$$
\left|A^{T_{a \ldots}}\left(x_{a \ldots}\right)-A^{T_{a}}\left(x_{\bar{a}}\right)\right| \propto \frac{T_{a} \text { Const }}{\left|\Lambda_{a}\right|}
$$

Substituting $t_{a b}$ one finds

$$
\frac{t_{a b}-t_{a} t_{b}}{t_{a b}}=1-e^{-s\left(T_{a}+T_{b}-T_{a b}\right)}\left|\frac{\Lambda_{a} \Lambda_{b}}{\Lambda_{a b}}\right|
$$

Since with increasing $m$ segments of $\overline{a b}$ come closer to $\bar{a}$, the differences in action and the ratio of the eigenvalues converge exponentially with the eigenvalue of the orbit $\bar{a}$,

$$
T_{a}+T_{b}-T_{a b} \approx \mathrm{Const} \times \Lambda_{a}^{-j}, \quad\left|\Lambda_{a} \Lambda_{b} / \Lambda_{a b}\right| \approx \exp \left(-\mathrm{Const} / \Lambda_{a b}\right)
$$

Expanding the exponentials one thus finds that this term in the cycle expansion is of the order of

$$
\begin{equation*}
t_{a^{j} b}-t_{a} t_{a^{j-1} b} \approx \text { Const } \times t_{a^{j} b} \Lambda_{a}^{-j} \tag{J.1}
\end{equation*}
$$

Even though the number of terms in a cycle expansion grows exponentially, the shadowing cancellations improve the convergence by an exponential factor compared to trace formulas, and extend the radius of convergence of the periodic orbit sums. Table J. 1 shows some examples of such compensations between long cycles and their pseudo-cycle shadows.

It is crucial that the curvature expansion is grouped (and truncated) by topologically related cycles and pseudo-cycles; truncations that ignore

| $n$ | $t_{a b}-t_{a} t_{b}$ | $T_{a b}-\left(T_{a}+T_{b}\right)$ | $\log$ | $\frac{\Lambda_{a} \Lambda_{b}}{\Lambda_{a b}}$ | $a b-a \cdot b$ |
| :---: | :---: | :---: | ---: | :---: | :---: |
| 2 | $-5.23465150784 \times 10^{4}$ | $4.85802927371 \times 10^{2}$ | $-6.3 \times 10^{2}$ | $01-0 \cdot 1$ |  |
| 3 | $-7.96028600139 \times 10^{6}$ | $5.21713101432 \times 10^{3}$ | $-9.8 \times 10^{3}$ | $001-0 \cdot 01$ |  |
| 4 | $-1.03326529874 \times 10^{7}$ | $5.29858199419 \times 10^{4}$ | $-1.3 \times 10^{3}$ | $0001-0 \cdot 001$ |  |
| 5 | $-1.27481522016 \times 10^{9}$ | $5.35513574697 \times 10^{5}$ | $-1.6 \times 10^{4}$ | $00001-0 \cdot 0001$ |  |
| 6 | $-1.52544704823 \times 10^{11}$ | $5.40999882625 \times 10^{6}$ | $-1.8 \times 10^{5}$ | $000001-0 \cdot 00001$ |  |
| 2 | $-5.23465150784 \times 10^{4}$ | $4.85802927371 \times 10^{2}$ | $-6.3 \times 10^{2}$ | $01-0 \cdot 1$ |  |
| 3 | $5.30414752996 \times 10^{6}$ | $-3.67093656690 \times 10^{3}$ | $7.7 \times 10^{3}$ | $011-01 \cdot 1$ |  |
| 4 | $-5.40934261680 \times 10^{8}$ | $3.14925761316 \times 10^{4}$ | $-9.2 \times 10^{4}$ | $0111-011 \cdot 1$ |  |
| 5 | $4.99129508833 \times 10^{10}$ | $-2.67292822795 \times 10^{5}$ | $1.0 \times 10^{4}$ | $01111-0111 \cdot 1$ |  |
| 6 | $-4.39246000586 \times 10^{12}$ | $2.27087116266 \times 10^{6}$ | $-1.0 \times 10^{5}$ | $011111-01111 \cdot 1$ |  |

Table J.1: Demonstration of shadowing in curvature combinations of cycle weights of form $t_{a b}-t_{a} t_{b}$, the 3-disk fundamental domain cycles at $R: d=6$, table 31.3. The ratio $\Lambda_{a} \Lambda_{b} / \Lambda_{a b}$ is approaching unity exponentially fast.
topology, such as inclusion of all cycles with $T_{p}<T_{\max }$, will contain orbits unmatched by shadowed orbits, and exhibit a mediocre convergence compared with the curvature expansions.

Note that the existence of a pole at $z=1 / c$ implies that the cycle expansions have a finite radius of convergence, and that analytic continuations will be required for extraction of the non-leading zeros of $1 / \zeta$. Preferably, one should work with cycle expansions of Selberg products, as discussed in sect. 18.1.3.

## J.1.3 No shadowing, poorer convergence

Conversely, if the dynamics is not of a finite subshift type, there is no finite topological polynomial, there are no "curvature" corrections, and the convergence of the cycle expansions will be poor.

## J. 2 On importance of pruning

If the grammar is not finite and there is no finite topological polynomial, there will be no "curvature" expansions, and the convergence will be poor. That is the generic case, and one strategy for dealing with it is to find a good sequence of approximate but finite grammars; for each approximate grammar cycle expansions yield exponentially accurate eigenvalues, with successive approximate grammars converging toward the desired infinite grammar system.

When the dynamical system's symbolic dynamics does not have a finite grammar, and we are not able to arrange its cycle expansion into curvature combinations (18.5), the series is truncated as in sect. 18.4, by including all pseudo-cycles such that $\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right| \leq\left|\Lambda_{P}\right|$, where $P$ is the most unstable prime cycle included into truncation. The truncation error should then be of order $O\left(e^{h T_{P}} T_{P} /\left|\Lambda_{P}\right|\right)$, with $h$ the topological entropy, and $e^{h T_{P}}$ roughly the
number of pseudo-cycles of stability $\approx\left|\Lambda_{P}\right|$. In this case the cycle averaging formulas do not converge significantly better than the approximations such as the trace formula (19.18).

Numerical results (see for example the plots of the accuracy of the cycle expansion truncations for the Hénon map in ref. [18.3]) indicate that the truncation error of most averages tracks closely the fluctuations due to the irregular growth in the number of cycles. It is not known whether one can exploit the sum rules such as the mass flow conservation (19.11) to improve the accuracy of dynamical averaging.

## J. 3 Ma-the-matical caveats

"Lo duca e io per quel cammino ascoso intrammo a ritornar nel chiaro monde; e sanza cura aver d'alcun riposa
salimmo sù, el primo e io secondo, tanto ch'i' vidi de le cose belle che porta ' 1 ciel, per un perutgio tondo."
Dante


The periodic orbit theory is learned in stages. At first glance, it seems totally impenetrable. After basic exercises are gone through, it seems totally trivial; all that seems to be at stake are elementary manipulations with traces, determinants, derivatives. But if start thinking about you will get a more and more uncomfortable feeling that from the mathematical point of view, this is a perilous enterprise indeed. In chapter 16 we shall explain which parts of this enterprise are really solid; here you give a fortaste of what objections a mathematician might rise.

Birkhoff's 1931 ergodic theorem states that the time average (10.4) exists almost everywhere, and, if the flow is ergodic, it implies that $\langle a(x)\rangle=$ $\langle a\rangle$ is a constant for almost all $x$. The problem is that the above cycle averaging formulas implicitly rely on ergodic hypothesis: they are strictly correct only if the dynamical system is locally hyperbolic and globally mixing. If one takes a $\beta$ derivative of both sides

$$
\rho_{\beta}(y) e^{t s(\beta)}=\int_{\mathcal{M}} d x \delta\left(y-f^{t}(x)\right) e^{\beta \cdot A^{t}(x)} \rho_{\beta}(x),
$$

and integrates over $y$

$$
\begin{aligned}
\left.\int_{\mathcal{M}} d y \frac{\partial}{\partial \beta} \rho_{\beta}(y)\right|_{\beta=0}+ & \left.t \frac{\partial s}{\partial \beta}\right|_{\beta=0} \int_{\mathcal{M}} d y \rho_{0}(y)= \\
& \int_{\mathcal{M}} d x A^{t}(x) \rho_{0}(x)+\left.\int_{\mathcal{M}} d x \frac{\partial}{\partial \beta} \rho_{\beta}(x)\right|_{\beta=0}
\end{aligned}
$$

one obtains in the long time limit

$$
\begin{equation*}
\left.\frac{\partial s}{\partial \beta}\right|_{\beta=0}=\int_{\mathcal{M}} d y \rho_{0}(x)\langle a(x)\rangle . \tag{J.2}
\end{equation*}
$$

This is the expectation value (10.12) only if the time average (10.4) equals the space average (10.9), $\langle a(x)\rangle=\langle a\rangle$, for all $x$ except a subset $x \in \mathcal{M}$ of zero measure; if the phase space is foliated into non-communicating subspaces $\mathcal{M}=\mathcal{M}_{1}+\mathcal{M}_{2}$ of finite measure such that $f^{t}\left(\mathcal{M}_{1}\right) \cap \mathcal{M}_{2}=\emptyset$ for all $t$, this fails. In other words, we have tacitly assumed metric indecomposability or transitivity. We have also glossed over the nature of the "phase space" $\mathcal{M}$. For example, if the dynamical system is open, such as the 3 -disk game of pinball, $\mathcal{M}$ in the expectation value integral (10.22) is a Cantor set, the closure of the union of all periodic orbits. Alternatively, $\mathcal{M}$ can be considered continuous, but then the measure $\rho_{0}$ in (J.2) is highly singular. The beauty of the periodic orbit theory is that instead of using an arbitrary coordinatization of $\mathcal{M}$ it partitions the phase space by the intrinsic topology of the dynamical flow and builds the correct measure from cycle invariants, the stability eigenvalues of periodic orbits.

Were we to restrict the applications of the formalism only to systems which have been rigorously proven to be ergodic, we might as well fold up the shop right now. For example, even for something as simple as the Hénon mapping we do not know whether the asymptotic time attractor is strange or periodic. Physics applications require a more pragmatic attitude. In the cycle expansions approach we construct the invariant set of the given dynamical system as a closure of the union of periodic orbits, and investigate how robust are the averages computed on this set. This turns out to depend very much on the observable being averaged over; dynamical averages exhibit "phase transitions", and the above cycle averaging formulas apply in the "hyperbolic phase" where the average is dominated by exponentially many exponentially small contributions, but fail in a phase dominated by few marginally stable orbits. Here the noise - always present, no matter how weak - helps us by erasing an infinity of small traps that the deterministic dynamics might fall into.

Still, in spite of all the caveats, periodic orbit theory is a beautiful theory, and the cycle averaging formulas are the most elegant and powerful tool available today for evaluation of dynamical averages for low dimensional chaotic deterministic systems.

## J. 4 Estimate of the $n$th cumulant

An immediate consequence of the exponential spacing of the eigenvalues is that the convergence of the Selberg product expansion (E.12) as function of the topological cycle length, $F(z)=\sum_{n} C_{n} z^{n}$, is faster than exponential. Consider a $d$-dimensional map for which all Jacobian matrix eigenvalues are equal: $u_{p}=\Lambda_{p, 1}=\Lambda_{p, 2}=\cdots=\Lambda_{p, d}$. The stability eigenvalues are
generally not isotropic; however, to obtain qualitative bounds on the spectrum, we replace all stability eigenvalues with the least expanding one. In this case the $p$ cycle contribution to the product (15.9) reduces to

$$
\begin{align*}
F_{p}(z) & =\prod_{k_{1} \cdots k_{d}=0}^{\infty}\left(1-t_{p} u_{p}^{k_{1}+k_{2}+\cdots+k_{d}}\right) \\
& =\prod_{k=0}^{\infty}\left(1-t_{p} u_{p}^{k}\right)^{m_{k}} ; \quad m_{k}=\binom{d-1+k}{d-1}=\frac{(k+d-1)!}{k!(d-1)!} \\
& =\prod_{k=0}^{\infty} \sum_{\ell=0}^{m_{k}}\binom{m_{k}}{\ell}\left(-u_{p}^{k} t_{p}\right)^{\ell} \tag{J.3}
\end{align*}
$$

In one dimension the expansion can be given in closed form (16.32), and the coefficients $C_{k}$ in (E.12) are given by

$$
\begin{equation*}
\tau_{p^{k}}=(-1)^{k} \frac{u_{p}^{\frac{k(k-1)}{2}}}{\prod_{j=1}^{k}\left(1-u_{p}^{j}\right)} t_{p}^{k} . \tag{J.4}
\end{equation*}
$$

Hence the coefficients in the $F(z)=\sum_{n} C_{n} z^{n}$ expansion of the spectral determinant (18.8) fall off faster than exponentially, as $\left|C_{n}\right| \approx u^{n(n-1) / 2}$. In contrast, the cycle expansions of dynamical zeta functions fall of "only" exponentially; in numerical applications, the difference is dramatic.

In higher dimensions the expansions are not quite as compact. The leading power of $u$ and its coefficient are easily evaluated by use of binomial expansions (J.3) of the $\left(1+t u^{k}\right)^{m_{k}}$ factors. More precisely, the leading $u^{n}$ terms in $t^{k}$ coefficients are of form

$$
\begin{aligned}
\prod_{k=0}^{\infty}\left(1+t u^{k}\right)^{m_{k}} & =\ldots+u^{m_{1}+2 m_{2}+\ldots+j m_{j}} t^{1+m_{1}+m_{2}+\ldots+m_{j}}+\ldots \\
& =\ldots+\left(u^{\frac{m_{d}}{d+1}} t\right)^{\binom{d+m}{m}}+\ldots \approx \ldots+u^{\frac{d}{(d-1)!}} n^{\frac{d+1}{d}} t^{n}+\ldots
\end{aligned}
$$

Hence the coefficients in the $F(z)$ expansion fall off faster than exponentially, as $u^{n^{1+1 / d}}$. The Selberg products are entire functions in any dimension, provided that the symbolic dynamics is a finite subshift, and all cycle eigenvalues are sufficiently bounded away from 1 .

The case of particular interest in many applications are the 2-d Hamiltonian mappings; their symplectic structure implies that $u_{p}=\Lambda_{p, 1}=1 / \Lambda_{p, 2}$, and the Selberg product (15.13) In this case the expansion corresponding to (16.32) is given by (16.33) and the coefficients fall off asymptotically as $C_{n} \approx u^{n^{3 / 2}}$.

## Appendix K

## Infinite dimensional operators

(A. Wirzba)

This appendix taken from ref. [K.1] summarizes the definitions and properties for trace-class and Hilbert-Schmidt matrices, the determinants over infinite dimensional matrices and possible regularization schemes for matrices or operators which are not of trace-class.

## K. 1 Matrix-valued functions

(P. Cvitanović)

As a preliminary we summarize some of the properties of functions of finitedimensional matrices.

The derivative of a matrix is a matrix with elements

$$
\begin{equation*}
A^{\prime}(x)=\frac{d A(x)}{d x}, \quad A_{i j}^{\prime}(x)=\frac{d}{d x} A_{i j}(x) . \tag{K.1}
\end{equation*}
$$

Derivatives of products of matrices are evaluated by the chain rule

$$
\begin{equation*}
\frac{d}{d x}(A \mathbf{B})=\frac{d A}{d x} \mathbf{B}+A \frac{d \mathbf{B}}{d x} . \tag{K.2}
\end{equation*}
$$

A matrix and its derivative matrix in general do not commute

$$
\begin{equation*}
\frac{d}{d x} A^{2}=\frac{d A}{d x} A+A \frac{d A}{d x} \tag{K.3}
\end{equation*}
$$

The derivative of the inverse of a matrix, follows from $\frac{d}{d x}\left(A A^{-1}\right)=0$ :

$$
\begin{equation*}
\frac{d}{d x} A^{-1}=-\frac{1}{A} \frac{d A}{d x} \frac{1}{A} . \tag{K.4}
\end{equation*}
$$

A function of a single variable that can be expressed in terms of additions and multiplications generalizes to a matrix-valued function by replacing the variable by the matrix.

In particular, the exponential of a constant matrix can be defined either by its series expansion, or as a limit of an infinite product:

$$
\begin{align*}
e^{A} & =\sum_{k=0}^{\infty} \frac{1}{k!} A^{k}, \quad A^{0}=\mathbf{1}  \tag{K.5}\\
& =\lim _{N \rightarrow \infty}\left(\mathbf{1}+\frac{1}{N} A\right)^{N} \tag{K.6}
\end{align*}
$$

The first equation follows from the second one by the binomial theorem, so these indeed are equivalent definitions. That the terms of order $O\left(N^{-2}\right)$ or smaller do not matter follows from the bound

$$
\left(1+\frac{x-\epsilon}{N}\right)^{N}<\left(1+\frac{x+\delta x_{N}}{N}\right)^{N}<\left(1+\frac{x+\epsilon}{N}\right)^{N}
$$

where $\left|\delta x_{N}\right|<\epsilon$. If $\lim \delta x_{N} \rightarrow 0$ as $N \rightarrow \infty$, the extra terms do not contribute.

Consider now the determinant

$$
\operatorname{det}\left(e^{A}\right)=\lim _{N \rightarrow \infty}(\operatorname{det}(\mathbf{1}+A / N))^{N} .
$$

To the leading order in $1 / N$

$$
\operatorname{det}(\mathbf{1}+A / N)=1+\frac{1}{N} \operatorname{tr} A+O\left(N^{-2}\right) .
$$

hence

$$
\begin{equation*}
\operatorname{det} e^{A}=\lim _{N \rightarrow \infty}\left(1+\frac{1}{N} \operatorname{tr} A+O\left(N^{-2}\right)\right)^{N}=e^{\operatorname{tr} A} \tag{K.7}
\end{equation*}
$$

Due to non-commutativity of matrices, generalization of a function of several variables to a function is not as straightforward. Expression involving several matrices depend on their commutation relations. For example, the commutator expansion

$$
\begin{equation*}
e^{t \mathbf{A}} \mathbf{B} e^{-t \mathbf{A}}=\mathbf{B}+t[\mathbf{A}, \mathbf{B}]+\frac{t^{2}}{2}[\mathbf{A},[\mathbf{A}, \mathbf{B}]]+\frac{t^{3}}{3!}[\mathbf{A},[\mathbf{A},[\mathbf{A}, \mathbf{B}]]]+\cdots(\mathrm{K} \tag{K.8}
\end{equation*}
$$

sometimes used to establish the equivalence of the Heisenberg and Schrödinger pictures of quantum mechanics follows by recursive evaluation of $t$ derivaties

$$
\frac{d}{d t}\left(e^{t \mathbf{A}} \mathbf{B} e^{-t \mathbf{A}}\right)=e^{t \mathbf{A}}[\mathbf{A}, \mathbf{B}] e^{-t \mathbf{A}}
$$

Manipulations of such ilk yield

$$
e^{(\mathbf{A}+\mathbf{B}) / N}=e^{\mathbf{A} / N} e^{\mathbf{B} / N}-\frac{1}{2 N^{2}}[\mathbf{A}, \mathbf{B}]+O\left(N^{-3}\right),
$$

and the Trotter product formula: if $\mathbf{B}, \mathbf{C}$ and $\mathbf{A}=\mathbf{B}+\mathbf{C}$ are matrices, then

$$
\begin{equation*}
e^{\mathbf{A}}=\lim _{N \rightarrow \infty}\left(e^{\mathbf{B} / N} e^{\mathbf{C} / N}\right)^{N} \tag{K.9}
\end{equation*}
$$

## K. 2 Operator norms

> (R. Mainieri and P. Cvitanović)

- 

The limit used in the above definition involves matrices - operators in vector spaces - rather than numbers, and its convergence can be checked using tools familiar from calculus. We briefly review those tools here, as throughout the text we will have to consider many different operators and how they converge.

The $n \rightarrow \infty$ convergence of partial products

$$
\mathbf{E}_{n}=\prod_{0 \leq m<n}\left(\mathbf{1}+\frac{t}{m} A\right)
$$

can be verified using the Cauchy criterion, which states that the sequence $\left\{\mathbf{E}_{n}\right\}$ converges if the differences $\left\|\mathbf{E}_{k}-\mathbf{E}_{j}\right\| \rightarrow 0$ as $k, j \rightarrow \infty$. To make sense of this we need to define a sensible norm $\|\cdots\|$. Norm of a matrix is based on the Euclidean norm for a vector: the idea is to assign to a matrix $\mathbf{M}$ a norm that is the largest possible change it can cause to the length of a unit vector $\hat{n}$ :

$$
\begin{equation*}
\|\mathbf{M}\|=\sup _{\hat{n}}\|\mathbf{M} \hat{n}\|, \quad\|\hat{n}\|=1 \tag{K.10}
\end{equation*}
$$

We say that $\|\cdot\|$ is the operator norm induced by the vector norm $\|\cdot\|$. Constructing a norm for a finite-dimensional matrix is easy, but had $\mathbf{M}$ been an operator in an infinite-dimensional space, we would also have to
specify the space $\hat{n}$ belongs to. In the finite-dimensional case, the sum of the absolute values of the components of a vector is also a norm; the induced operator norm for a matrix $\mathbf{M}$ with components $M_{i j}$ in that case can be defined by

$$
\begin{equation*}
\|\mathbf{M}\|=\max _{i} \sum_{j}\left|M_{i j}\right| . \tag{K.11}
\end{equation*}
$$

The operator norm (K.11) and the vector norm (K.10) are only rarely distinguished by different notation, a bit of notational laziness that we shall uphold.

Now that we have learned how to make sense out of norms of operators, we can check that

$$
\begin{equation*}
\left\|e^{t A}\right\| \leq e^{t\|A\|} \tag{K.12}
\end{equation*}
$$

As $\|A\|$ is a number, the norm of $e^{t A}$ is finite and therefore well defined. In particular, the exponential of a matrix is well defined for all values of $t$, and the linear differential equation (4.8) has a solution for all times.

## K. 3 Trace class and Hilbert-Schmidt class

This section is mainly an extract from ref. [K.9]. Refs. [K.7, K.10, K.11, K.14] should be consulted for more details and proofs. The trace class and Hilbert-Schmidt property will be defined here for linear, in general nonhermitian operators $\mathbf{A} \in \mathcal{L}(\mathcal{H}): \mathcal{H} \rightarrow \mathcal{H}$ (where $\mathcal{H}$ is a separable Hilbert space). The transcription to matrix elements (used in the prior chapters) is simply $a_{i j}=\left\langle\phi_{i}, \mathbf{A} \phi_{j}\right\rangle$ where $\left\{\phi_{n}\right\}$ is an orthonormal basis of $\mathcal{H}$ and $\langle$,$\rangle is$ the inner product in $\mathcal{H}$ (see sect. K. 5 where the theory of von Koch matrices of ref. [K.12] is discussed). So, the trace is the generalization of the usual notion of the sum of the diagonal elements of a matrix; but because infinite sums are involved, not all operators will have a trace:

## Definition:

(a) An operator $\mathbf{A}$ is called trace class, $\mathbf{A} \in \mathcal{J}_{1}$, if and only if, for every orthonormal basis, $\left\{\phi_{n}\right\}$ :

$$
\begin{equation*}
\sum_{n}\left|\left\langle\phi_{n}, \mathbf{A} \phi_{n}\right\rangle\right|<\infty . \tag{K.13}
\end{equation*}
$$

The family of all trace class operators is denoted by $\mathcal{J}_{1}$.
(b) An operator $\mathbf{A}$ is called Hilbert-Schmidt, $\mathbf{A} \in \mathcal{J}_{2}$, if and only if, for every orthonormal basis, $\left\{\phi_{n}\right\}$ :

$$
\begin{equation*}
\sum_{n}\left\|\mathbf{A} \phi_{n}\right\|^{2}<\infty \tag{K.14}
\end{equation*}
$$

The family of all Hilbert-Schmidt operators is denoted by $\mathcal{J}_{2}$.

Bounded operators are dual to trace class operators. They satisfy the the following condition: $|\langle\psi, B \phi\rangle| \leq C\|\psi\|\| \| \phi \|$ with $C<\infty$ and $\psi, \phi \in \mathcal{H}$. If they have eigenvalues, these are bounded too. The family of bounded operators is denoted by $\mathcal{B}(\mathcal{H})$ with the norm $\|B\|=\sup _{\phi \neq 0} \frac{\|\mathbf{B} \phi\|}{\|\phi\|}$ for $\phi \in \mathcal{H}$. Examples for bounded operators are unitary operators and especially the unit matrix. In fact, every bounded operator can be written as linear combination of four unitary operators.

A bounded operator $\mathbf{C}$ is compact, if it is the norm limit of finite rank operators.

An operator $\mathbf{A}$ is called positive, $\mathbf{A} \geq 0$, if $\langle\mathbf{A} \phi, \phi\rangle \geq 0 \forall \phi \in \mathcal{H}$. Notice that $\mathbf{A}^{\dagger} \mathbf{A} \geq 0$. We define $|\mathbf{A}|=\sqrt{\mathbf{A}^{\dagger} \mathbf{A}}$.

The most important properties of the trace and Hilbert-Schmidt classes are summarized in (see refs. [K.7, K.9]):
(a) $\mathcal{J}_{1}$ and $\mathcal{J}_{2}$ are $*$ ideals., that is, they are vector spaces closed under scalar multiplication, sums, adjoints, and multiplication with bounded operators.
(b) $\mathbf{A} \in \mathcal{J}_{1}$ if and only if $\mathbf{A}=\mathbf{B C}$ with $\mathbf{B}, \mathbf{C} \in \mathcal{J}_{2}$.
(c) $\mathcal{J}_{1} \subset \mathcal{J}_{2} \subset$ Compact operators.
(d) For any operator $\mathbf{A}$, we have $\mathbf{A} \in \mathcal{J}_{2}$ if $\sum_{n}\left\|\mathbf{A} \phi_{n}\right\|^{2}<\infty$ for a single basis.
For any operator $\mathbf{A} \geq 0$ we have $\mathbf{A} \in \mathcal{J}_{1}$ if $\sum_{n}\left|\left\langle\phi_{n}, \mathbf{A} \phi_{n}\right\rangle\right|<\infty$ for a single basis.
(e) If $\mathbf{A} \in \mathcal{J}_{1}, \operatorname{Tr}(\mathbf{A})=\sum\left\langle\phi_{n}, \mathbf{A} \phi_{n}\right\rangle$ is independent of the basis used.
(f) $\operatorname{Tr}$ is linear and obeys $\operatorname{Tr}\left(\mathbf{A}^{\dagger}\right)=\overline{\operatorname{Tr}(\mathbf{A})} ; \operatorname{Tr}(\mathbf{A B})=\operatorname{Tr}(\mathbf{B A})$ if either $\mathbf{A} \in \mathcal{J}_{1}$ and $\mathbf{B}$ bounded, $\mathbf{A}$ bounded and $\mathbf{B} \in \mathcal{J}_{1}$ or both $\mathbf{A}, \mathbf{B} \in \mathcal{J}_{2}$.
(g) $\mathcal{J}_{2}$ endowed with the inner product $\langle\mathbf{A}, \mathbf{B}\rangle_{2}=\operatorname{Tr}\left(\mathbf{A}^{\dagger} \mathbf{B}\right)$ is a Hilbert space. If $\|\mathbf{A}\|_{2}=\left[\operatorname{Tr}\left(\mathbf{A}^{\dagger} \mathbf{A}\right)\right]^{\frac{1}{2}}$, then $\|\mathbf{A}\|_{2} \geq\|\mathbf{A}\|$ and $\mathcal{J}_{2}$ is the || $\|_{2}$-closure of the finite rank operators.
(h) $\mathcal{J}_{1}$ endowed with the norm $\|\mathbf{A}\|_{1}=\operatorname{Tr}\left(\sqrt{\mathbf{A}^{\dagger} \mathbf{A}}\right)$ is a Banach space. $\|\mathbf{A}\|_{1} \geq\|\mathbf{A}\|_{2} \geq\|\mathbf{A}\|$ and $\mathcal{J}_{1}$ is the $\left\|\|_{1}\right.$-norm closure of the finite rank operators. The dual space of $\mathcal{J}_{1}$ is $\mathcal{B}(\mathcal{H})$, the family of bounded operators with the duality $\langle\mathbf{B}, \mathbf{A}\rangle=\operatorname{Tr}(\mathbf{B A})$.
(i) If $\mathbf{A}, \mathbf{B} \in \mathcal{J}_{2}$, then $\|\mathbf{A B}\|_{1} \leq\|\mathbf{A}\|_{2}\|\mathbf{B}\|_{2}$. If $\mathbf{A} \in \mathcal{J}_{2}$ and $\mathbf{B} \in \mathcal{B}(\mathcal{H})$, then $\|\mathbf{A B}\|_{2} \leq\|\mathbf{A}\|_{2}\|\mathbf{B}\|$. If $\mathbf{A} \in \mathcal{J}_{1}$ and $\mathbf{B} \in \mathcal{B}(\mathcal{H})$, then $\|\mathbf{A B}\|_{1} \leq$ $\|\mathbf{A}\|_{1}\|\mathbf{B}\|$.

Note the most important property for proving that an operator is trace class is the decomposition (b) into two Hilbert-Schmidt ones, as the HilbertSchmidt property can easily be verified in one single orthonormal basis (see (d)). Property (e) ensures then that the trace is the same in any basis.

Properties (a) and (f) show that trace class operators behave in complete analogy to finite rank operators. The proof whether a matrix is trace-class (or Hilbert-Schmidt) or not simplifies enormously for diagonal matrices, as then the second part of property (d) is directly applicable: just the moduli of the eigenvalues (or - in case of Hilbert-Schmidt - the squares of the eigenvalues) have to be summed up in order to answer that question. A good strategy in checking the trace-class character of a general matrix $\mathbf{A}$ is therefore the decomposition of that matrix into two matrices $\mathbf{B}$ and $\mathbf{C}$ where one, say C, should be chosen to be diagonal and either just barely of Hilbert-Schmidt character leaving enough freedom for its partner $\mathbf{B}$ or of trace-class character such that one only has to show the boundedness for B.

## K. 4 Determinants of trace class operators

This section is mainly based on refs. [K.8, K.10] which should be consulted for more details and proofs. See also refs. [K.11, K.14].

Pre-definitions (Alternating algebra and Fock spaces):
Given a Hilbert space $\mathcal{H}, \otimes^{n} \mathcal{H}$ is defined as the vector space of multi-linear functionals on $\mathcal{H}$ with $\phi_{1} \otimes \cdots \otimes \phi_{n} \in \otimes^{n} \mathcal{H}$ in case $\phi_{1}, \ldots, \phi_{n} \in \mathcal{H} . \bigwedge^{n}(\mathcal{H})$ is defined as the subspace of $\otimes^{n} \mathcal{H}$ spanned by the wedge-product

$$
\begin{equation*}
\phi_{1} \wedge \cdots \wedge \phi_{n}=\frac{1}{\sqrt{n!}} \sum_{\pi \in \mathcal{P}_{n}} \epsilon(\pi)\left[\phi_{\pi(1)} \otimes \cdots \otimes \phi_{\pi(n)}\right] \tag{K.15}
\end{equation*}
$$

where $\mathcal{P}_{n}$ is the group of all permutations of $n$ letters and $\epsilon(\pi)= \pm 1$ depending on whether $\pi$ is an even or odd permutation, respectively. The inner product in $\bigwedge^{n}(\mathcal{H})$ is given by

$$
\begin{equation*}
\left(\phi_{1} \wedge \cdots \wedge \phi_{n}, \eta_{1} \wedge \cdots \wedge \eta_{n}\right)=\operatorname{det}\left\{\left(\phi_{i}, \eta_{j}\right)\right\} \tag{K.16}
\end{equation*}
$$

where $\operatorname{det}\left\{a_{i j}\right\}=\sum_{\pi \in \mathcal{P}_{n}} \epsilon(\pi) a_{1 \pi(1)} \cdots a_{n \pi(n)} \cdot \bigwedge^{n}(\mathbf{A})$ is defined as functor (a functor satisfies $\left.\bigwedge^{n}(\mathbf{A B})=\bigwedge^{n}(\mathbf{A}) \bigwedge^{n}(\mathbf{B})\right)$ on $\Lambda^{n}(\mathcal{H})$ with

$$
\begin{equation*}
\bigwedge^{n}(\mathbf{A})\left(\phi_{1} \wedge \cdots \wedge \phi_{n}\right)=\mathbf{A} \phi_{1} \wedge \cdots \wedge \mathbf{A} \phi_{n} \tag{K.17}
\end{equation*}
$$

When $n=0, \bigwedge^{n}(\mathcal{H})$ is defined to be $C$ and $\bigwedge^{n}(\mathbf{A})$ as 1: $C \rightarrow C$.
Properties: If $\mathbf{A}$ trace class, that is, $\mathbf{A} \in \mathcal{J}_{1}$, then for any $k, \bigwedge^{k}(\mathbf{A})$ is trace class, and for any orthonormal basis $\left\{\phi_{n}\right\}$ the cumulant

$$
\begin{equation*}
\operatorname{Tr}\left(\bigwedge^{k}(\mathbf{A})\right)=\sum_{i_{1}<\cdots<i_{k}}\left(\left(\phi_{i_{1}} \wedge \cdots \wedge \phi_{i_{k}}\right),\left(\mathbf{A} \phi_{i_{1}} \wedge \cdots \wedge \mathbf{A} \phi_{i_{k}}\right)\right)<\phi \tag{k.18}
\end{equation*}
$$

is independent of the basis (with the understanding that $\operatorname{Tr} \bigwedge^{0}(\mathbf{A}) \equiv 1$ ).

Definition: Let $\mathbf{A} \in \mathcal{J}_{1}$, then $\operatorname{det}(\mathbf{1}+\mathbf{A})$ is defined as

$$
\begin{equation*}
\operatorname{det}(\mathbf{1}+\mathbf{A})=\sum_{k=0}^{\infty} \operatorname{Tr}\left(\bigwedge^{k}(\mathbf{A})\right) \tag{K.19}
\end{equation*}
$$

## Properties:

Let $\mathbf{A}$ be a linear operator on a separable Hilbert space $\mathcal{H}$ and $\left\{\phi_{j}\right\}_{1}^{\infty}$ an orthonormal basis.
(a) $\sum_{k=0}^{\infty} \operatorname{Tr}\left(\bigwedge^{k}(\mathbf{A})\right)$ converges for each $\mathbf{A} \in \mathcal{J}_{1}$.
(b) $|\operatorname{det}(\mathbf{1}+\mathbf{A})| \leq \prod_{j=1}^{\infty}\left(1+\mu_{j}(\mathbf{A})\right)$ where $\mu_{j}(\mathbf{A})$ are the singular values of $\mathbf{A}$, that is, the eigenvalues of $|\mathbf{A}|=\sqrt{\mathbf{A}^{\dagger} \mathbf{A}}$.
(c) $|\operatorname{det}(\mathbf{1}+\mathbf{A})| \leq \exp \left(\|\mathbf{A}\|_{1}\right)$.
(d) For any $\mathbf{A}_{1}, \ldots, \mathbf{A}_{n} \in \mathcal{J}_{1},\left\langle z_{1}, \ldots, z_{n}\right\rangle \mapsto \operatorname{det}\left(\mathbf{1}+\sum_{i=1}^{n} z_{i} \mathbf{A}_{i}\right)$ is an entire analytic function.
(e) If $\mathbf{A}, \mathbf{B} \in \mathcal{J}_{1}$, then

$$
\begin{align*}
\operatorname{det}(\mathbf{1}+\mathbf{A}) \operatorname{det}(\mathbf{1}+\mathbf{B}) & =\operatorname{det}(\mathbf{1}+\mathbf{A}+\mathbf{B}+\mathbf{A B}) \\
& =\operatorname{det}((\mathbf{1}+\mathbf{A})(\mathbf{1}+\mathbf{B})) \\
& =\operatorname{det}((\mathbf{1}+\mathbf{B})(\mathbf{1}+\mathbf{A})) \tag{K.20}
\end{align*}
$$

If $\mathbf{A} \in \mathcal{J}_{1}$ and $\mathbf{U}$ unitary, then

$$
\begin{equation*}
\operatorname{det}\left(\mathbf{U}^{-1}(\mathbf{1}+\mathbf{A}) \mathbf{U}\right)=\operatorname{det}\left(\mathbf{1}+\mathbf{U}^{-1} \mathbf{A} \mathbf{U}\right)=\operatorname{det}(\mathbf{1}+\mathbf{A}) \tag{K.21}
\end{equation*}
$$

(f) If $\mathbf{A} \in \mathcal{J}_{1}$, then $(\mathbf{1}+\mathbf{A})$ is invertible if and only if $\operatorname{det}(\mathbf{1}+\mathbf{A}) \neq 0$.
(g) If $\lambda \neq 0$ is an $n$-times degenerate eigenvalue of $\mathbf{A} \in \mathcal{J}_{1}$, then $\operatorname{det}(\mathbf{1}+$ $z \mathbf{A}$ ) has a zero of order $n$ at $z=-1 / \lambda$.
(h) For any $\epsilon$, there is a $C_{\epsilon}(\mathbf{A})$, depending on $\mathbf{A} \in \mathcal{J}_{1}$, so that $\mid \operatorname{det}(\mathbf{1}+$ $z \mathbf{A}) \mid \leq C_{\epsilon}(\mathbf{A}) \exp (\epsilon|z|)$.
(i) For any $\mathbf{A} \in \mathcal{J}_{1}$,

$$
\begin{equation*}
\operatorname{det}(\mathbf{1}+\mathbf{A})=\prod_{j=1}^{N(\mathbf{A})}\left(1+\lambda_{j}(\mathbf{A})\right) \tag{K.22}
\end{equation*}
$$

where here and in the following $\left\{\lambda_{j}(\mathbf{A})\right\}_{j=1}^{N(\mathbf{A})}$ are the eigenvalues of A counted with algebraic multiplicity .
(j) Lidskii's theorem: For any $\mathbf{A} \in \mathcal{J}_{1}$,

$$
\begin{equation*}
\operatorname{Tr}(\mathbf{A})=\sum_{j=1}^{N(\mathbf{A})} \lambda_{j}(\mathbf{A})<\infty . \tag{K.23}
\end{equation*}
$$

(k) If $\mathbf{A} \in \mathcal{J}_{1}$, then

$$
\begin{aligned}
\operatorname{Tr}\left(\bigwedge^{k}(\mathbf{A})\right) & =\sum_{j=1}^{N\left(\wedge^{k}(\mathbf{A})\right)} \lambda_{j}\left(\bigwedge^{k}(\mathbf{A})\right) \\
& =\sum_{1 \leq j_{1}<\cdots<j_{k} \leq N(\mathbf{A})} \lambda_{j_{1}}(\mathbf{A}) \cdots \lambda_{j_{k}}(\mathbf{A})<\infty
\end{aligned}
$$

(1) If $\mathbf{A} \in \mathcal{J}_{1}$, then

$$
\operatorname{det}(1+z \mathbf{A})=\sum_{k=0}^{\infty} z^{k} \sum_{1 \leq j_{1}<\cdots<j_{k} \leq N(\mathbf{A})} \lambda_{j_{1}}(\mathbf{A}) \cdots \lambda_{j_{k}}(\mathbf{A})<\phi(\mathbb{K} .24)
$$

(m) If $\mathbf{A} \in \mathcal{J}_{1}$, then for $|z|$ small (that is, $|z| \max \left|\lambda_{j}(\mathbf{A})\right|<1$ ) the series $\sum_{k=1}^{\infty} z^{k} \operatorname{Tr}\left((-\mathbf{A})^{k}\right) / k$ converges and

$$
\begin{align*}
\operatorname{det}(1+z \mathbf{A}) & =\exp \left(-\sum_{k=1}^{\infty} \frac{z^{k}}{k} \operatorname{Tr}\left((-\mathbf{A})^{k}\right)\right) \\
& =\exp (\operatorname{Tr} \ln (\mathbf{1}+z \mathbf{A})) \tag{K.25}
\end{align*}
$$

(n) The Plemelj-Smithies formula: Define $\alpha_{m}(\mathbf{A})$ for $\mathbf{A} \in \mathcal{J}_{1}$ by

$$
\begin{equation*}
\operatorname{det}(\mathbf{1}+z \mathbf{A})=\sum_{m=0}^{\infty} z^{m} \frac{\alpha_{m}(\mathbf{A})}{m!} \tag{K.26}
\end{equation*}
$$

Then $\alpha_{m}(\mathbf{A})$ is given by the $m \times m$ determinant:

$$
\alpha_{m}(\mathbf{A})=\left|\begin{array}{ccccc}
\operatorname{Tr}(\mathbf{A}) & m-1 & 0 & \cdots & 0 \\
\operatorname{Tr}\left(\mathbf{A}^{2}\right) & \operatorname{Tr}(\mathbf{A}) & m-2 & \cdots & 0 \\
\operatorname{Tr}\left(\mathbf{A}^{3}\right) & \operatorname{Tr}\left(\mathbf{A}^{2}\right) & \operatorname{Tr}(\mathbf{A}) & \cdots & 0 \\
\vdots & \vdots & \vdots & \vdots & \vdots \\
& & & & 1 \\
\operatorname{Tr}\left(\mathbf{A}^{m}\right) & \operatorname{Tr}\left(\mathbf{A}^{(m-1)}\right) & \operatorname{Tr}\left(\mathbf{A}^{(m-2)}\right) & \cdots & \operatorname{Tr}(\mathbf{A})
\end{array}\right|
$$

with the understanding that $\alpha_{0}(\mathbf{A}) \equiv 1$ and $\alpha_{1}(\mathbf{A}) \equiv \operatorname{Tr}(\mathbf{A})$. Thus the cumulants $c_{m}(\mathbf{A}) \equiv \alpha_{m}(\mathbf{A}) / m$ ! satisfy the following recursion relation

$$
\begin{array}{rlr}
c_{m}(\mathbf{A}) & =\frac{1}{m} \sum_{k=1}^{m}(-1)^{k+1} c_{m-k}(\mathbf{A}) \operatorname{Tr}\left(\mathbf{A}^{k}\right) \quad \text { for } m \geq 1 \\
c_{0}(\mathbf{A}) & \equiv 1 . \tag{K.28}
\end{array}
$$

Note that in the context of quantum mechanics formula (K.26) is the quantum analog to the curvature expansion of the semiclassical zeta function with $\operatorname{Tr}\left(\mathbf{A}^{m}\right)$ corresponding to the sum of all periodic orbits (prime and also repeated ones) of total topological length $m$, that is, let $c_{m}$ (s.c.) denote the $m^{\text {th }}$ curvature term, then the curvature expansion of the semiclassical zeta function is given by the recursion relation

$$
\begin{array}{rlr}
c_{m}(\text { s.c. }) & =\frac{1}{m} \sum_{k=1}^{m}(-1)^{k+m+1} c_{m-k} \text { (s.c.) } \sum_{\substack{p ; r>0 \\
\text { with }[p] r=k}}[p] \frac{t_{p}(k)^{r}}{1-\left(\frac{1}{\Lambda_{p}}\right)^{r}} & \text { for } m \geq 1 \\
\left.c_{0} \text { (s.c. }\right) & \equiv 1 . & \tag{K.29}
\end{array}
$$

In fact, in the cumulant expansion (K.26) as well as in the curvature expansion there are large cancellations involved. Let us order - without lost of generality - the eigenvalues of the operator $\mathbf{A} \in \mathcal{J}_{1}$ as follows:

$$
\left|\lambda_{1}\right| \geq\left|\lambda_{2}\right| \geq \cdots \geq\left|\lambda_{i-1}\right| \geq\left|\lambda_{i}\right| \geq\left|\lambda_{i+1}\right| \geq \cdots
$$

(This is always possible because of $\sum_{i=1}^{N(\mathbf{A})}\left|\lambda_{i}\right|<\infty$.) Then, in the standard (Plemelj-Smithies) cumulant evaluation of the determinant, eq. (K.26), we have enormous cancellations of big numbers, e.g. at the $k^{\text {th }}$ cumulant order $(k>3)$, all the intrinsically large 'numbers' $\lambda_{1}^{k}, \lambda_{1}^{k-1} \lambda_{2}, \ldots, \lambda_{1}^{k-2} \lambda_{2} \lambda_{3}$, $\ldots$ and many more have to cancel out exactly until only $\sum_{1 \leq j_{1}<\cdots<j_{k} \leq N(\mathbf{A})} \lambda_{j_{1}} \cdots \lambda_{j_{k}}$ is finally left over. Algebraically, the fact that there are these large cancellations is of course of no importance. However, if the determinant is calculated numerically, the big cancellations might spoil the result or even the convergence. Now, the curvature expansion of the semiclassical zeta function, as it is known today, is the semiclassical approximation to the curvature expansion (unfortunately) in the Plemelj-Smithies form. As the exact quantum mechanical result is approximated semiclassically, the errors introduced in the approximation might lead to big effects as they are done with respect to large quantities which eventually cancel out and not - as it would be of course better - with respect to the small surviving cumulants. Thus it would be very desirable to have a semiclassical analog to the reduced cumulant expansion (K.24) or even to (K.22) directly. It might not be possible to find a direct semiclassical analog for the individual eigenvalues $\lambda_{j}$. Thus the direct construction of the semiclassical equivalent to (K.22) is rather unlikely. However, in order to have a semiclassical "cumulant" summation without large cancellations - see (K.24) - it would be just sufficient to find the semiclassical analog of each complete cumulant (K.24) and not of the single eigenvalues. Whether this will eventually be possible is still an open question.

## K. 5 Von Koch matrices

Implicitly, many of the above properties are based on the theory of von Koch matrices [K.11, K.12, K.13]: An infinite matrix $\mathbf{1}-\mathbf{A}=\left\|\delta_{j k}-a_{j k}\right\|_{1}^{\infty}$,
consisting of complex numbers, is called a matrix with an absolutely convergent determinant, if the series $\sum\left|a_{j_{1} k_{1}} a_{j_{2} k_{2}} \cdots a_{j_{n}, k_{n}}\right|$ converges, where the sum extends over all pairs of systems of indices $\left(j_{1}, j_{2}, \cdots, j_{n}\right)$ and $\left(k_{1}, k_{2}, \cdots, k_{n}\right)$ which differ from each other only by a permutation, and $j_{1}<j_{2}<\cdots j_{n}(n=1,2, \cdots)$. Then the limit

$$
\lim _{n \rightarrow \infty} \operatorname{det}\left\|\delta_{j k}-a_{j k}\right\|_{1}^{n}=\operatorname{det}(\mathbf{1}-\mathbf{A})
$$

exists and is called the determinant of the matrix $\mathbf{1 - A}$. It can be represented in the form

$$
\operatorname{det}(\mathbf{1}-\mathbf{A})=1-\sum_{j=1}^{\infty} a_{j j}+\frac{1}{2!} \sum_{j, k=1}^{\infty}\left|\begin{array}{ll}
a_{j j} & a_{j k} \\
a_{k j} & a_{k k}
\end{array}\right|-\frac{1}{3!} \sum_{j, k, m=1}^{\infty}\left|\begin{array}{ccc}
a_{j j} & a_{j k} & a_{j m} \\
a_{k j} & a_{k k} & a_{k m} \\
a_{m j} & a_{m k} & a_{m m}
\end{array}\right|+\cdots
$$

where the series on the r.h.s. will remain convergent even if the numbers $a_{j k}$ $(j, k=1,2, \cdots)$ are replaced by their moduli and if all the terms obtained by expanding the determinants are taken with the plus sign. The matrix $\mathbf{1} \mathbf{- \mathbf { A }}$ is called von Koch matrix, if both conditions

$$
\begin{align*}
\sum_{j=1}^{\infty}\left|a_{j j}\right| & <\infty  \tag{K.30}\\
\sum_{j, k=1}^{\infty}\left|a_{j k}\right|^{2} & <\infty \tag{K.31}
\end{align*}
$$

are fulfilled. Then the following holds (see ref. [K.11, K.13]): (1) Every von Koch matrix has an absolutely convergent determinant. If the elements of a von Koch matrix are functions of some parameter $\mu\left(a_{j k}=a_{j k}(\mu)\right.$, $j, k=1,2, \cdots)$ and both series in the defining condition converge uniformly in the domain of the parameter $\mu$, then as $n \rightarrow \infty$ the determinant $\operatorname{det} \| \delta_{j k}-$ $a_{j k}(\mu) \|_{1}^{n}$ tends to the determinant $\operatorname{det}(\mathbf{1}+\mathbf{A}(\mu))$ uniformly with respect to $\mu$, over the domain of $\mu$. (2) If the matrices $\mathbf{1}-\mathbf{A}$ and $\mathbf{1}-\mathbf{B}$ are von Koch matrices, then their product $\mathbf{1}-\mathbf{C}=(\mathbf{1}-\mathbf{A})(\mathbf{1}-\mathbf{B})$ is a von Koch matrix, and

$$
\begin{equation*}
\operatorname{det}(\mathbf{1}-\mathbf{C})=\operatorname{det}(\mathbf{1}-\mathbf{A}) \operatorname{det}(\mathbf{1}-\mathbf{B}) . \tag{K.32}
\end{equation*}
$$

Note that every trace-class matrix $\mathbf{A} \in \mathcal{J}_{1}$ is also a von Koch matrix (and that any matrix satisfying condition (K.31) is Hilbert-Schmidt and vice versa). The inverse implication, however, is not true: von Koch matrices are not automatically trace-class. The caveat is that the definition of von Koch matrices is basis-dependent, whereas the trace-class property is basis-independent. As the traces involve infinite sums, the basisindependence is not at all trivial. An example for an infinite matrix which
is von Koch, but not trace-class is the following:

$$
\mathbf{A}_{i j}=\left\{\begin{array}{lll}
2 / j & \text { for } & i-j=-1 \\
2 / i & \text { for } & i-j=+1 \\
0 & \text { else }, & \text { and } \\
i \text { even }, \\
0
\end{array}\right.
$$

that is,

$$
\mathbf{A}=\left(\begin{array}{ccccccc}
0 & 1 & 0 & 0 & 0 & 0 & \cdots  \tag{K.33}\\
1 & 0 & 0 & 0 & 0 & 0 & \cdots \\
0 & 0 & 0 & 1 / 2 & 0 & 0 & \cdots \\
0 & 0 & 1 / 2 & 0 & 0 & 0 & \cdots \\
0 & 0 & 0 & 0 & 0 & 1 / 3 & \ddots \\
0 & 0 & 0 & 0 & 1 / 3 & 0 & \ddots \\
\vdots & \vdots & \vdots & \vdots & \ddots & \ddots & \ddots
\end{array}\right)
$$

Obviously, condition (K.30) is fulfilled by definition. Secondly, condition (K.31) is satisfied as $\sum_{n=1}^{\infty} 2 / n^{2}<\infty$. However, the sum over the moduli of the eigenvalues is just twice the harmonic series $\sum_{n=1}^{\infty} 1 / n$ which does not converge. The matrix (K.33) violates the trace-class definition (K.13), as in its eigenbasis the sum over the moduli of its diagonal elements is infinite. Thus the absolute convergence is traded for a conditional convergence, since the sum over the eigenvalues themselves can be arranged to still be zero, if the eigenvalues with the same modulus are summed first. Absolute convergence is of course essential, if sums have to be rearranged or exchanged. Thus, the trace-class property is indispensable for any controlled unitary transformation of an infinite determinant, as then there will be necessarily a change of basis and in general also a re-ordering of the corresponding traces. Therefore the claim that a Hilbert-Schmidt operator with a vanishing trace is automatically trace-class is false. In general, such an operator has to be regularized in addition (see next chapter).

## K. 6 Regularization

Many interesting operators are not of trace class (although they might be in some $\mathcal{J}_{p}$ with $p>1$ - an operator $A$ is in $\mathcal{J}_{p}$ iff $\operatorname{Tr}|A|^{p}<\infty$ in any orthonormal basis). In order to compute determinants of such operators, an extension of the cumulant expansion is needed which in fact corresponds to a regularization procedure [K.8, K.10]:
E.g. let $\mathbf{A} \in \mathcal{J}_{p}$ with $p \leq n$. Define

$$
\begin{equation*}
R_{n}(z \mathbf{A})=(\mathbf{1}+z \mathbf{A}) \exp \left(\sum_{k=1}^{n-1} \frac{(-z)^{k}}{k} \mathbf{A}^{k}\right)-\mathbf{1} \tag{K.34}
\end{equation*}
$$

as the regulated version of the operator $z \mathbf{A}$. Then the regulated operator $R_{n}(z \mathbf{A})$ is trace class, that is, $R_{n}(z \mathbf{A}) \in \mathcal{J}_{1}$. Define now $\operatorname{det}_{n}(\mathbf{1}+z \mathbf{A})=$
$\operatorname{det}\left(\mathbf{1}+R_{n}(z \mathbf{A})\right)$. Then the regulated determinant

$$
\operatorname{det}_{n}(\mathbf{1}+z \mathbf{A})=\prod_{j=1}^{N(z \mathbf{A})}\left[\left(1+z \lambda_{j}(\mathbf{A})\right) \exp \left(\sum_{k=1}^{n-1} \frac{\left(-z \lambda_{j}(\mathbf{A})\right)^{k}}{k}\right)\right]<\alpha(\mathrm{K} .35)
$$

exists and is finite. The corresponding Plemelj-Smithies formula now reads [K.10]:

$$
\begin{equation*}
\operatorname{det}_{n}(\mathbf{1}+z \mathbf{A})=\sum_{m=0}^{\infty} z^{m} \frac{\alpha_{m}^{(n)}(\mathbf{A})}{m!} . \tag{K.36}
\end{equation*}
$$

with $\alpha_{m}^{(n)}(\mathbf{A})$ given by the $m \times m$ determinant:

$$
\alpha_{m}^{(n)}(\mathbf{A})=\left|\begin{array}{ccccc}
\sigma_{1}^{(n)} & m-1 & 0 & \cdots & 0  \tag{K.37}\\
\sigma_{2}^{(n)} & \sigma_{1}^{(n)} & m-2 & \cdots & 0 \\
\sigma_{3}^{(n)} & \sigma_{2}^{(n)} & \sigma_{1}^{(n)} & \cdots & 0 \\
\vdots & \vdots & \vdots & \vdots & \vdots \\
& & & & 1 \\
\sigma_{m}^{(n)} & \sigma_{m-1}^{(n)} & \sigma_{m-2}^{(n)} & \cdots & \sigma_{1}^{(n)}
\end{array}\right|
$$

where

$$
\sigma_{k}^{(n)}= \begin{cases}\operatorname{Tr}\left(\mathbf{A}^{k}\right) & k \geq n \\ 0 & k \leq n-1\end{cases}
$$

As Simon [K.10] says simply, the beauty of (K.37) is that we get $\operatorname{det}_{n}(\mathbf{1}+$ A) from the standard Plemelj-Smithies formula (K.26) by simply setting $\operatorname{Tr}(\mathbf{A}), \operatorname{Tr}\left(\mathbf{A}^{2}\right), \ldots, \operatorname{Tr}\left(\mathbf{A}^{n-1}\right)$ to zero.

See also ref. [K.15] where $\left\{\lambda_{j}\right\}$ are the eigenvalues of an elliptic (pseudo)differential operator $\mathbf{H}$ of order $m$ on a compact or bounded manifold of dimension $d, 0<\lambda_{0} \leq \lambda_{1} \leq \cdots$ and $\lambda_{k} \uparrow+\infty$. and the Fredholm determinant

$$
\begin{equation*}
\Delta(\lambda)=\prod_{k=0}^{\infty}\left(1-\frac{\lambda}{\lambda_{k}}\right) \tag{K.38}
\end{equation*}
$$

is regulated in the case $\mu \equiv d / m>1$ as Weierstrass product

$$
\begin{equation*}
\Delta(\lambda)=\prod_{k=0}^{\infty}\left[\left(1-\frac{\lambda}{\lambda_{k}}\right) \exp \left(\frac{\lambda}{\lambda_{k}}+\frac{\lambda^{2}}{2 \lambda_{k}^{2}}+\cdots+\frac{\lambda^{[\mu]}}{[\mu] \lambda_{k}^{[\mu]}}\right)\right] \tag{K.39}
\end{equation*}
$$

where [ $\mu$ ] denotes the integer part of $\mu$. This is, see ref. [K.15], the unique entire function of order $\mu$ having zeros at $\left\{\lambda_{k}\right\}$ and subject to the normalization conditions

$$
\begin{equation*}
\ln \Delta(0)=\frac{d}{d \lambda} \ln \Delta(0)=\cdots=\frac{d^{[\mu]}}{d \lambda^{[\mu]}} \ln \Delta(0)=0 . \tag{K.40}
\end{equation*}
$$

Clearly eq. (K.39) is the same as (K.35); one just has to identify $z=-\lambda$, $\mathbf{A}=1 / \mathbf{H}$ and $n-1=[\mu]$. An example is the regularization of the spectral determinant

$$
\begin{equation*}
\Delta(E)=\operatorname{det}[(E-\mathbf{H})] \tag{K.41}
\end{equation*}
$$

which - as it stands - would only make sense for a finite dimensional basis (or finite dimensional matrices). In ref. [K.16] the regulated spectral determinant for the example of the hyperbola billiard in two dimensions (thus $d=2, m=2$ and hence $\mu=1$ ) is given as

$$
\begin{equation*}
\Delta(E)=\operatorname{det}[(E-\mathbf{H}) \Omega(E, \mathbf{H})] \tag{K.42}
\end{equation*}
$$

where

$$
\begin{equation*}
\Omega(E, \mathbf{H})=-\mathbf{H}^{-1} e^{E \mathbf{H}^{-1}} \tag{K.43}
\end{equation*}
$$

such that the spectral determinant in the eigenbasis of $\mathbf{H}$ (with eigenvalues $E_{n} \neq 0$ ) reads

$$
\begin{equation*}
\Delta(E)=\prod_{n}\left(1-\frac{E}{E_{n}}\right) e^{E / E_{n}}<\infty . \tag{K.44}
\end{equation*}
$$

Note that $\mathbf{H}^{-1}$ is for this example of Hilbert-Schmidt character.

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## Appendix L

## Statistical mechanics recycled

A spin system with long-range interactions can be converted into a chaotic dynamical system that is differentiable and low-dimensional. The thermodynamic limit quantities of the spin system are then equivalent to long time averages of the dynamical system. In this way the spin system averages can be recast as the cycle expansions. If the resulting dynamical system is analytic, the convergence to the thermodynamic limit is faster than with the standard transfer matrix techniques.

## L. 1 The thermodynamic limit

There are two motivations to recycle statistical mechanics: one gets better control over the thermodynamic limit and one gets detailed information on how one is converging to it. From this information, most other quantities of physical interst can be computed.

In statistical mechanics one computes the averages of observables. These are functions that return a number for every state of the system; they are an abstraction of the process of measuring the pressure or temperature of a gas. The average of an observable is computed in the thermodynamic limit - the limit of system with an arbitrarily large number of particles. The thermodynamic limit is an essential step in the computation of averages, as it is only then that one observes the bulk properties of matter.

Without the thermodynamic limit many of the thermodynamic properties of matter could not be derived within the framework of statistical mechanics. There would be no extensive quantities, no equivalence of ensembles, and no phase transitions. From experiments it is known that certain quantities are extensive, that is, they are proportional to the size of the system. This is not true for an interacting set of particles. If two systems interacting via pairwise potentials are brought close together, work will be required to join them, and the final total energy will not be the sum of the
energies of each of the parts. To avoid the conflict between the experiments and the theory of Hamiltonian systems, one needs systems with an infinite number of particles. In the canonical ensemble the probability of a state is given by the Boltzman factor which does not impose the conservation of energy; in the microcanonical ensemble energy is conserved but the Boltzmann factor is no longer exact. The equality between the ensembles only appears in the limit of the number of particles going to infinity at constant density. The phase transitions are interpreted as points of non-analyticity of the free energy in the thermodynamic limit. For a finite system the partition function cannot have a zero as a function of the inverse temperature $\beta$, as it is a finite sum of positive terms.

The thermodynamic limit is also of central importance in the study of field theories. A field theory can be first defined on a lattice and then the lattice spacing is taken to zero as the correlation length is kept fixed. This continuum limit corresponds to the thermodynamic limit. In lattice spacing units the correlation length is going to infinity, and the interacting field theory can be thought of as a statistical mechanics model at a phase transition.

For general systems the convergence towards the thermodynamic limit is slow. If the thermodynamic limit exists for an interaction, the convergence of the free energy per unit volume $f$ is as an inverse power in the linear dimension of the system.

$$
\begin{equation*}
f(\beta) \rightarrow \frac{1}{n} \tag{L.1}
\end{equation*}
$$

where $n$ is proportional to $V^{1 / d}$, with $V$ the volume of the $d$-dimensional system. Much better results can be obtained if the system can be described by a transfer matrix. A transfer matrix is concocted so that the trace of its $n$th power is exactly the partition function of the system with one of the dimensions proportional to $n$. When the system is described by a transfer matrix then the convergence is exponential,

$$
\begin{equation*}
f(\beta) \rightarrow e^{-\alpha n} \tag{L.2}
\end{equation*}
$$

and may only be faster than that if all long-range correlations of the system are zero - that is, when there are no interactions. The coefficient $\alpha$ depends only on the inverse correlation length of the system.

One of the difficulties in using the transfer matrix techniques is that they seem at first limited to systems with finite range interactions. Phase transitions can happen only when the interaction is long range. One can try to approximate the long range interaction with a series of finite range interactions that have an ever increasing range. The problem with this approach is that in a formally defined transfer matrix, not all the eigenvalues of the matrix correspond to eigenvalues of the system (in the sense that the rate of decay of correlations is not the ratio of eigenvalues).

Knowledge of the correlations used in conjunction with finite size scaling to obtain accurate estimates of the parameters of systems with phase transitions. (Accurate critical exponents are obtained by series expansions or transfer matrices, and seldomly by renormalization group arguments or Monte Carlo.) In a phase transition the coefficient $\alpha$ of the exponential convergence goes to zero and the convergence to the thermodynamic limit is power-law.

The computation of the partition function is an example of a functional integral. For most interactions these integrals are ill-defined and require some form of normalization. In the spin models case the functional integral is very simple, as "space" has only two points and only "time" being infinite has to be dealt with. The same problem occurs in the computation of the trace of transfer matrices of systems with infinite range interactions. If one tries to compute the partition function $Z_{n}$

$$
Z_{n}=\operatorname{tr} T^{n}
$$

when $T$ is an infinite matrix, the result may be infinite for any $n$. This is not to say that $Z_{n}$ is infinite, but that the relation between the trace of an operator and the partition function breaks down. We could try regularizing the expression, but as we shall see below, that is not necessary, as there is a better physical solution to this problem.

What will described here solves both of these problems in a limited context: it regularizes the transfer operator in a physically meaningful way, and as a a consequence, it allows for the faster than exponential convergence to the thermodynamic limit and complete determination of the spectrum. The steps to achieve this are:

- Redefine the transfer operator so that there are no limits involved except for the thermodynamic limit.
- Note that the divergences of this operator come from the fact that it acts on a very large space. All that is needed is the smallest subspace containing the eigenvector corresponding to the largest eigenvalue (the Gibbs state).
- Rewrite all observables as depending on a local effective field. The eigenvector is like that, and the operator restricted to this space is trace-class.
- Compute the spectrum of the transfer operator and observe the magic.


## L. 2 Ising models

The Ising model is a simple model to study the cooperative effects of many small interacting magnetic dipoles. The dipoles are placed on a lattice and their interaction is greatly simplified. There can also be a field that includes the effects of an external magnetic field and the average effect of the dipoles
among themselves. We will define a general class of Ising models (also called spin systems) where the dipoles can be in one of many possible states and the interactions extend beyond the nearest neighboring sites of the lattice. But before we extend the Ising model, we will examine the simplest model in that class.

## L.2.1 Ising model

One of the simplest models in statistical mechanics is the Ising model. One imagines that one has a one-dimensional lattice with small magnets at each site that can point either up or down.


Each little magnet interacts only with its neighbors. If they both point in the same direction, then they contribute an energy $-J$ to the total energy of the system; and if they point in opposite directions, then they contribute $+J$. The signs are chsen so that they prefer to be aligned. Let us suppose that we have $n$ small magnets arranged in a line: A line is drawn between two sites to indicate that there is an interaction between the small magnets that are located on that site

(This figure can be thought of as a graph, with sites being vertices and interacting magnets indicated by edges.) To each of the sites we associate a variable, that we call a spin, that can be in either of two states: up ( $\uparrow$ ) or down $(\downarrow)$. This represents the two states of the small magnet on that site, and in general we will use the notation $\Sigma_{0}$ to represent the set of possible values of a spin at any site; all sites assume the same set of values. A configuration consists of assigning a value to the spin at each site; a typical configuration is


The set of all configurations for a lattice with $n$ sites is called $\Omega_{0}^{n}$ and is formed by the Cartesian product $\Omega_{0} \times \Omega_{0} \cdots \times \Omega_{0}$, the product repeated $n$ times. Each configuration $\sigma \in \Omega^{n}$ is a string of $n$ spins

$$
\begin{equation*}
\sigma=\left\{\sigma_{0}, \sigma_{1}, \ldots \sigma_{n}\right\} \tag{L.5}
\end{equation*}
$$

In the example configuration (L.4) there are two pairs of spins that have the same orientation and six that have the opposite orientation. Therefore
the total energy $H$ of the configuration is $J \times 6-J \times 2=4 J$. In general we can associate an energy $H$ to every configuration

$$
\begin{equation*}
H(\sigma)=\sum_{i} J \delta\left(\sigma_{i}, \sigma_{i+1}\right), \tag{L.6}
\end{equation*}
$$

where

$$
\delta\left(\sigma_{1}, \sigma_{2}\right)=\left\{\begin{array}{ll}
+1 & \text { if } \sigma_{1}=\sigma_{2}  \tag{L.7}\\
-1 & \text { if } \sigma_{1} \neq \sigma_{2}
\end{array} .\right.
$$

One of the problems that was avoided when computing the energy was what to do at the boundaries of the one-dimensional chain. Notice that as written, (L.6) requires the interaction of $\operatorname{spin} n$ with spin $n+1$. In the absence of phase transitions the boundaries do not matter much to the thermodynamic limit and we will connect the first site to the last, implementing periodic boundary conditions.

Thermodynamic quantities are computed from the partition function $Z^{(n)}$ as the size $n$ of the system becomes very large. For example, the free energy per site $f$ at inverse temperature $\beta$ is given by

$$
\begin{equation*}
-\beta f(\beta)=\lim _{n \rightarrow \infty} \frac{1}{n} \ln Z^{(n)} \tag{L.8}
\end{equation*}
$$

The partition function $Z^{(n)}$ is computed by a sum that runs over all the possible configurations on the one-dimensional chain. Each configuration contributes with its Gibbs factor $\exp (-\beta H(\sigma))$ and the partition function $Z^{(n)}$ is

$$
\begin{equation*}
Z^{(n)}(\beta)=\sum_{\sigma \in \Omega_{0}^{n}} e^{-\beta H(\sigma)} \tag{L.9}
\end{equation*}
$$

The partition function can be computed using transfer matrices. This is a method that generalizes to other models. At first, it is a little mysterious that matrices show up in the study of a sum. To see where they come from, we can try and build a configuration on the lattice site by site. The frst thing to do is to expand out the sum for the energy of the configuration

$$
\begin{equation*}
Z^{(n)}(\beta)=\sum_{\sigma \in \Omega^{n}} e^{\beta J \delta\left(\sigma_{1}, \sigma_{2}\right)} e^{\beta J \delta\left(\sigma_{2}, \sigma_{3}\right)} \cdots e^{\beta J \delta\left(\sigma_{n}, \sigma_{1}\right)} . \tag{L.10}
\end{equation*}
$$

Let us use the configuration in (L.4). The first site is $\sigma_{1}=\uparrow$. As the second site is $\uparrow$, we know that the first term in (L.10) is a term $e^{\beta J}$. The third spin is $\downarrow$, so the second term in (L.10) is $e^{-\beta J}$. If the third spin had been $\uparrow$, then the term would have been $e^{\beta J}$ but it would not depend on the value of the first spin $\sigma_{1}$. This means that the configuration can be built site by site
and that to compute the Gibbs factor for the configuration just requires knowing the last spin added. We can then think of the configuration as being a weighted random walk where each step of the walk contributes according to the last spin added. The random walk take place on the Markov graph


Choose one of the two sites as a starting point. Walk along any allowed edge making your choices randomly and keep track of the accumulated weight as you perform the $n$ steps. To implement the periodic boundary conditions make sure that you return to the starting node of the Markov graph. If the walk is carried out in all possible $2^{n}$ ways then the sum of all the weights is the partition function. To perform the sum we consider the matrix

$$
T(\beta)=\left[\begin{array}{cc}
e^{\beta J} & e^{-\beta J}  \tag{L.11}\\
e^{-\beta J} & e^{\beta J}
\end{array}\right] .
$$

As in chapter 11 the sum of all closed walks is given by the trace of powers of the matrix. These powers can easily be re-expressed in terms of the two eigenvalues $\lambda_{1}$ and $\lambda_{2}$ of the transfer matrix:

$$
\begin{equation*}
Z^{(n)}(\beta)=\operatorname{tr} T^{n}(\beta)=\lambda_{1}(\beta)^{n}+\lambda_{2}(\beta)^{n} \tag{L.12}
\end{equation*}
$$

## L.2.2 Averages of observables

Averages of observables can be re-expressed in terms of the eigenvectors of the transfer matrix. Alternatively, one can introduce a modified transfer matrix and compute the averages through derivatives. Sounds familiar?

## L.2.3 General spin models

The more general version of the Ising model - the spin models - will be defined on a regular lattice, $\mathbb{Z}^{D}$. At each lattice site there will be a spin variable that can assumes a finite number of states identified by the set $\Omega_{0}$.

The transfer operator $\mathcal{T}$ was introduced by Kramers and Wannier [L.12] to study the Ising model on a strip and concocted so that the trace of its $n$th power is the partition function $Z_{n}$ of system when one of its dimensions is $n$. The method can be generalized to deal with any finite-range interaction. If the range of the interaction is $L$, then $\mathcal{T}$ is a matrix of size $2^{L} \times 2^{L}$. The longer the range, the larger the matrix.

## L. 3 Fisher droplet model

In a series of articles [L.20], Fisher introduced the droplet model. It is a model for a system containing two phases: gas and liquid. At high temperatures, the typical state of the system consists of droplets of all sizes floating in the gas phase. As the temperature is lowered, the droplets coalesce, forming larger droplets, until at the transition temperature, all droplets form one large one. This is a first order phase transition.

Although Fisher formulated the model in three-dimensions, the analytic solution of the model shows that it is equivalent to a one-dimensional lattice gas model with long range interactions. Here we will show how the model can be solved for an arbitrary interaction, as the solution only depends on the asymptotic behavior of the interaction.

The interest of the model for the study of cycle expansions is its relation to intermittency. By having an interaction that behaves asymptotically as the scaling function for intermittency, one expects that the analytic structure (poles and cuts) will be same.

Fisher used the droplet model to study a first order phase transition [L.20]. Gallavotti [L.21] used it to show that the zeta functions cannot in general be extended to a meromorphic functions of the entire complex plane. The droplet model has also been used in dynamical systems to explain features of mode locking, see Artuso [L.22]. In computing the zeta function for the droplet model we will discover that at low temperatures the cycle expansion has a limited radius of convergence, but it is possible to factorize the expansion into the product of two functions, each of them with a better understood radius of convergence.

## L.3.1 Solution

The droplet model is a one-dimensional lattice gas where each site can have two states: empty or occupied. We will represent the empty state by 0 and the occupied state by 1 . The configurations of the model in this notation are then strings of zeros and ones. Each configuration can be viewed as groups of contiguous ones separated by one or more zeros. The contiguous ones represent the droplets in the model. The droplets do not interact with each other, but the individual particles within each droplet do.

To determine the thermodynamics of the system we must assign an energy to every configuration. At very high temperatures we would expect a gaseous phase where there are many small droplets, and as we decrease the temperature the droplets would be expected to coalesce into larger ones until at some point there is a phase transition and the configuration is dominated by one large drop. To construct a solvable model and yet one with a phase transition we need long range interaction among all the particles of a droplet. One choice is to assign a fixed energy $\theta_{n}$ for the interactions of the particles of a cluster of size $n$. In a given droplet one has to consider
all the possible clusters formed by contiguous particles. Consider for example the configuration 0111010. It has two droplets, one of size three and another of size one. The droplet of size one has only one cluster of size one and therefore contributes to the energy of the configuration with $\theta_{1}$. The cluster of size three has one cluster of size three, two clusters of size two, and three clusters of size one; each cluster contributing a $\theta_{n}$ term to the energy. The total energy of the configuration is then

$$
\begin{equation*}
H(0111010)=4 \theta_{1}+2 \theta_{2}+1 \theta_{3} . \tag{L.13}
\end{equation*}
$$

If there where more zeros around the droplets in the above configuration the energy would still be the same. The interaction of one site with the others is assumed to be finite, even in the ground state consisting of a single droplet, so there is a restriction on the sum of the cluster energies given by

$$
\begin{equation*}
a=\sum_{n>0} \theta_{n}<\infty . \tag{L.14}
\end{equation*}
$$

The configuration with all zeros does not contribute to the energy.
Once we specify the function $\theta_{n}$ we can computed the energy of any configuration, and from that determine the thermodynamics. Here we will evaluate the cycle expansion for the model by first computing the generating function

$$
\begin{equation*}
G(z, \beta)=\sum_{n>0} z^{n} \frac{Z_{n}(\beta)}{n} \tag{L.15}
\end{equation*}
$$

and then considering its exponential, the cycle expansion. Each partition function $Z_{n}$ must be evaluated with periodic boundary conditions. So if we were computing $Z_{3}$ we must consider all eight binary sequences of three bits, and when computing the energy of a configuration, say 011, we should determine the energy per three sites of the long chain

$$
\ldots 011011011011 \ldots
$$

In this case the energy would be $\theta_{2}+2 \theta_{1}$. If instead of 011 we had considered one of its rotated shifts, 110 or 101, the energy of the configuration would have been the same. To compute the partition function we only need to consider one of the configurations and multiply by the length of the configuration to obtain the contribution of all its rotated shifts. The factor $1 / n$ in the generating function cancels this multiplicative factor. This reduction will not hold if the configuration has a symmetry, as for example 0101 which has only two rotated shift configurations. To compensate this we replace the $1 / n$ factor by a symmetry factor $1 / s(b)$ for each configuration $b$. The evaluation of $G$ is now reduced to summing over all configurations that are
not rotated shift equivalent, and we call these the basic configurations and the set of all of them $B$. We now need to evaluate

$$
\begin{equation*}
G(z, \beta)=\sum_{b \in B} \frac{z^{|b|}}{s(b)} e^{-\beta H(b)} . \tag{L.16}
\end{equation*}
$$

The notation $|\cdot|$ represents the cardinality of the set.
Any basic configuration can be built by considering the set of droplets that form it. The smallest building block has size two, as we must also put a zero next to the one so that when two different blocks get put next to each other they do not coalesce. The first few building blocks are

```
size droplets
2 01
3 001 011
4 0001 0011 0111
```

Each droplet of size $n$ contributes with energy

$$
\begin{equation*}
W_{n}=\sum_{1 \leq k \leq n}(n-k+1) \theta_{k} . \tag{L.18}
\end{equation*}
$$

So if we consider the sum

$$
\begin{align*}
\sum_{n \geq 1} \frac{1}{n} & \left(z^{2} e^{-\beta H(01)}+z^{3}\left(e^{-\beta H(001)}+e^{-\beta H(011)}\right)+\right. \\
& \left.+z^{4}\left(e^{-\beta H(0001)}+e^{-\beta H(0011)}+e^{-\beta H(0111)}\right)+\cdots\right)^{n} \tag{L.19}
\end{align*}
$$

then the power in $n$ will generate all the configurations that are made from many droplets, while the $z$ will keep track of the size of the configuration. The factor $1 / n$ is there to avoid the over-counting, as we only want the basic configurations and not its rotated shifts. The $1 / n$ factor also gives the correct symmetry factor in the case the configuration has a symmetry. The sum can be simplified by noticing that it is a logarithmic series

$$
\begin{equation*}
-\ln \left(1-\left(z^{2} e^{-\beta W_{1}}+z^{3}\left(e^{-\beta W_{1}}+e^{-\beta W_{2}}\right)+\cdots\right),\right. \tag{L.20}
\end{equation*}
$$

where the $H(b)$ factors have been evaluated in terms of the droplet energies $W_{n}$. A proof of the equality of (L.19) and (L.20) can be given, but we there was not enough space on the margin to write it down. The series that is subtracted from one can be written as a product of two series and the logarithm written as

$$
\begin{equation*}
-\ln \left(1-\left(z^{1}+z^{2}+z^{3}+\cdots\right)\left(z e^{-\beta W_{1}}+z^{2} e^{-\beta W_{2}}+\cdots\right)\right) \tag{L.21}
\end{equation*}
$$

The product of the two series can be directly interpreted as the generating function for sequences of droplets. The first series adds one or more zeros to a configuration and the second series add a droplet.

There is a whole class of configurations that is not included in the above sum: the configurations formed from a single droplet and the vacuum configuration. The vacuum is the easiest, as it has zero energy it only contributes a $z$. The sum of all the null configurations of all sizes is

$$
\begin{equation*}
\sum_{n>0} \frac{z^{n}}{n} \tag{L.22}
\end{equation*}
$$

The factor $1 / n$ is here because the original $G$ had them and the null configurations have no rotated shifts. The single droplet configurations also do not have rotated shifts so their sum is

$$
\begin{equation*}
\sum_{n>0} \frac{z^{n} e^{-\beta H}(\overbrace{11 \ldots 11}^{n})}{n} . \tag{L.23}
\end{equation*}
$$

Because there are no zeros in the above configuration clusters of all size exist and the energy of the configuration is $n \sum \theta_{k}$ which we denote by $n a$.

From the three sums (L.21), (L.22), and (L.23) we can evaluate the generating function $G$ to be

$$
\begin{equation*}
G(z, \beta)=-\ln (1-z)-\ln \left(1-z e^{-\beta a}\right)-\ln \left(1-\frac{z}{1-z} \sum_{n \geq 1} z^{n} e^{-\beta W_{n}}\right) . \tag{L.24}
\end{equation*}
$$

The cycle expansion $\zeta^{-1}(z, \beta)$ is given by the exponential of the generating function $e^{-G}$ and we obtain

$$
\begin{equation*}
\zeta^{-1}(z, \beta)=\left(1-z e^{-\beta a}\right)\left(1-z\left(1+\sum_{n \geq 1} z^{n} e^{-\beta W_{n}}\right)\right) \tag{L.25}
\end{equation*}
$$

To pursue this model further we need to have some assumptions about the interaction strengths $\theta_{n}$. We will assume that the interaction strength decreases with the inverse square of the size of the cluster, that is, $\theta_{n}=$ $-1 / n^{2}$. With this we can estimate that the energy of a droplet of size $n$ is asymptotically

$$
\begin{equation*}
W_{n} \sim-n+\ln n+\mathcal{O}\left(\frac{1}{n}\right) . \tag{L.26}
\end{equation*}
$$

If the power chosen for the polynomially decaying interaction had been other than inverse square we would still have the droplet term proportional
to $n$, but there would be no logarithmic term, and the $\mathcal{O}$ term would be of a different power. The term proportional to $n$ survives even if the interactions falls off exponentially, and in this case the correction is exponentially small in the asymptotic formula. To simplify the calculations we are going to assume that the droplet energies are exactly

$$
\begin{equation*}
W_{n}=-n+\ln n \tag{L.27}
\end{equation*}
$$

in a system of units where the dimensional constants are one. To evaluate the cycle expansion (L.25) we need to evaluate the constant $a$, the sum of all the $\theta_{n}$. One can write a recursion for the $\theta_{n}$

$$
\begin{equation*}
\theta_{n}=W_{n}-\sum_{1 \leq k<n}(n-k+1) \theta_{k} \tag{L.28}
\end{equation*}
$$

and with an initial choice for $\theta_{1}$ evaluate all the others. It can be verified that independent of the choice of $\theta_{1}$ the constant $a$ is equal to the number that multiplies the $n$ term in (L.27). In the units used

$$
\begin{equation*}
a=-1 \tag{L.29}
\end{equation*}
$$

For the choice of droplet energy (L.27) the sum in the cycle expansion can be expressed in terms of a special function: the Lerch transcendental $\phi_{L}$. It is defined by

$$
\begin{equation*}
\phi_{L}(z, s, c)=\sum_{n \geq 0} \frac{z^{n}}{(n+c)^{s}} \tag{L.30}
\end{equation*}
$$

excluding from the sum any term that has a zero denominator. The Lerch function converges for $|z|<1$. The series can be analytically continued to the complex plane and it will have a branch point at $z=1$ with a cut chosen along the positive real axis. In terms of Lerch transcendental function we can write the cycle expansion (L.25) using (L.27) as

$$
\begin{equation*}
\zeta^{-1}(z, \beta)=\left(1-z e^{\beta}\right)\left(1-z\left(1+\phi_{L}\left(z e^{\beta}, \beta, 1\right)\right)\right) \tag{L.31}
\end{equation*}
$$

This serves as an example of a zeta function that cannot be extended to a meromorphic function of the complex plane as one could conjecture.

The thermodynamics for the droplet model comes from the smallest root of (L.31). The root can come from any of the two factors. For large value of $\beta$ (low temperatures) the smallest root is determined from the $\left(1-z e^{\beta}\right)$ factor, which gave the contribution of a single large drop. For small $\beta$ (large temperatures) the root is determined by the zero of the other factor, and it corresponds to the contribution from the gas phase of the droplet model.

The transition occurs when the smallest root of each of the factors become numerically equal. This determines the critical temperature $\beta_{c}$ through the equation

$$
\begin{equation*}
1-e^{-\beta_{c}}\left(1+\zeta_{R}\left(\beta_{c}\right)\right)=0 \tag{L.32}
\end{equation*}
$$

which can be solved numerically. One finds that $\beta_{c}=1.40495$. The phase transition occurs because the roots from two different factors get swapped in their roles as the smallest root. This in general leads to a first order phase transition. For large $\beta$ the Lerch transcendental is being evaluated at the branch point, and therefore the cycle expansion cannot be an analytic function at low temperatures. For large temperatures the smallest root is within the radius of convergence of the series for the Lerch transcendental, and the cycle expansion has a domain of analyticity containing the smallest root.

As we approach the phase transition point as a function of $\beta$ the smallest root and the branch point get closer together until at exactly the phase transition they collide. This is a sufficient condition for the existence of a first order phase transitions. In the literature of zeta functions [L.23] there have been speculations on how to characterize a phase transition within the formalism. The solution of the Fisher droplet model suggests that for first order phase transitions the factorized cycle expansion will have its smallest root within the radius of convergence of one of the series except at the phase transition when the root collides with a singularity. This does not seem to be the case for second order phase transitions.

The analyticity of the cycle expansion can be restored if we consider separate cycle expansions for each of the phases of the system. If we separate the two terms of $\zeta^{-1}$ in (L.31), each of them is an analytic function and contains the smallest root within the radius of convergence of the series for the relevant $\beta$ values.

## L. 4 Scaling functions

> "Clouds are not spheres, mountains are not cones, coastlines are not circles and bark is not smooth, nor does lightning travel in straight line."
> B.B. Mandelbrot

There is a relation between general spin models and dynamical system. If one thinks of the boxes of the Markov partition of a hyperbolic system as the states of a spin system, then computing averages in the dynamical system is carrying out a sum over all possible states. One can even construct the natural measure of the dynamical system from a translational invariant "interaction function" call the scaling function.

There are many routes that lead to an explanation of what a scaling function is and how to compute it. The shortest is by breaking away from

Figure L.1: Construction of the steps of the scaling function from a Cantor set. From one level to the next in the construction of the Cantor set the covers are shrunk, each parent segment into two children segments. The shrinkage of the last level of the construction is plotted and by removing the gaps one has an approximation to the scaling function of the Cantor set.

the historical development and considering first the presentation function of a fractal. The presentation function is a simple chaotic dynamical system (hyperbolic, unlike the circle map) that generates the fractal and is closely related to the definition of fractals of Hutchinson [L.24] and the iterated dynamical systems introduced by Barnsley and collaborators [H.19]. From the presentation function one can derive the scaling function, but we will not do it in the most elegant fashion, rather we will develop the formalism in a form that is directly applicable to the experimental data.

In the upper part of figure L. 1 we have the successive steps of the construction similar to the middle third Cantor set. The construction is done in levels, each level being formed by a collection of segments. From one level to the next, each "parent" segment produces smaller "children" segments by removing the middle section. As the construction proceeds, the segments better approximate the Cantor set. In the figure not all the segments are the same size, some are larger and some are smaller, as is the case with multifractals. In the middle third Cantor set, the ratio between a segment and the one it was generated from is exactly $1 / 3$, but in the case shown in the figure the ratios differ from $1 / 3$. If we went through the last level of the construction and made a plot of the segment number and its ratio to its parent segment we would have a scaling function, as indicated in the figure. A function giving the ratios in the construction of a fractal is the basic idea for a scaling function. Much of the formalism that we will introduce is to be able to give precise names to every segments and to arrange the "lineage" of segments so that the children segments have the correct parent. If we do not take these precautions, the scaling function would be a "wild function", varying rapidly and not approximated easily by simple functions.

To describe the formalism we will use a variation on the quadratic map that appears in the theory of period doubling. This is because the combinatorial manipulations are much simpler for this map than they are for the circle map. The scaling function will be described for a one dimensional map $F$ as shown in figure L.2. Drawn is the map

$$
\begin{equation*}
F(x)=5 x(1-x) \tag{L.33}
\end{equation*}
$$

restricted to the unit interval. We will see that this map is also a presentation function.

Figure L.2: A Cantor set presentation function. The Cantor set is the set of all points that under iteration do not leave the interval $[0,1]$. This set can be found by backwards iterating the gap between the two branches of the map. The dotted lines can be used to find these backward images. At each step of the construction one is left with a set of segments that form a cover of the Cantor set.


It has two branches separated by a gap: one over the left portion of the unit interval and one over the right. If we choose a point $x$ at random in the unit interval and iterate it under the action of the map $F$, (L.33), it will hop between the branches and eventually get mapped to minus infinity. An orbit point is guaranteed to go to minus infinity if it lands in the gap. The hopping of the point defines the orbit of the initial point $x: x \mapsto x_{1} \mapsto$ $x_{2} \mapsto \cdots$. For each orbit of the map $F$ we can associate a symbolic code. The code for this map is formed from 0s and 1 s and is found from the orbit by associating a 0 if $x_{t}<1 / 2$ and a 1 if $x_{t}>1 / 2$, with $t=0,1,2, \ldots$.

Most initial points will end up in the gap region between the two branches. We then say that the orbit point has escaped the unit interval. The points that do not escape form a Cantor set $\mathcal{C}$ (or Cantor dust) and remain trapped in the unit interval for all iterations. In the process of describing all the points that do not escape, the map $F$ can be used as a presentation of the Cantor set $\mathcal{C}$, and has been called a presentation function by Feigenbaum [24.13].

How does the map $F$ "present" the Cantor set? The presentation is done in steps. First we determine the points that do not escape the unit interval in one iteration of the map. These are the points that are not part of the gap. These points determine two segments, which are an approximation to the Cantor set. In the next step we determine the points that do not escape in two iterations. These are the points that get mapped into the gap in one iteration, as in the next iteration they will escape; these points form the two segments $\Delta_{0}^{(1)}$ and $\Delta_{1}^{(1)}$ at level 1 in figure L.2. The processes can be continued for any number of iterations. If we observe carefully what is being done, we discover that at each step the pre-images of the gap (backward iterates) are being removed from the unit interval. As the map has two branches, every point in the gap has two pre-images, and therefore the whole gap has two pre-images in the form of two smaller gaps. To generate all the gaps in the Cantor set one just has to iterate the gap backwards. Each iteration of the gap defines a set of segments, with the $n$th iterate defining the segments $\Delta_{k}^{(n)}$ at level $n$. For this map there will be $2^{n}$ segments at level $n$, with the first few drawn in figure L.2. As $n \rightarrow \infty$ the segments that remain for at least $n$ iterates converge to the Cantor set $\mathcal{C}$.

The segments at one level form a cover for the Cantor set and it is from a cover that all the invariant information about the set is extracted (the cover generated from the backward iterates of the gap form a Markov partition for the map as a dynamical system). The segments $\left\{\Delta_{k}^{(n)}\right\}$ at level $n$ are a refinement of the cover formed by segments at level $n-1$. From successive covers we can compute the trajectory scaling function, the spectrum of scalings $f(\alpha)$, and the generalized dimensions.

To define the scaling function we must give labels (names) to the segments. The labels are chosen so that the definition of the scaling function allows for simple approximations. As each segment is generated from an inverse image of the unit interval, we will consider the inverse of the presentation function $F$. Because $F$ does not have a unique inverse, we have to consider restrictions of $F$. Its restriction to the first half of the segment, from 0 to $1 / 2$, has a unique inverse, which we will call $F_{0}^{-1}$, and its restriction to the second half, from $1 / 2$ to 1 , also has a unique inverse, which we will call $F_{1}^{-1}$. For example, the segment labeled $\Delta^{(2)}(0,1)$ in figure L. 2 is formed from the inverse image of the unit interval by mapping $\Delta^{(0)}$, the unit interval, with $F_{1}^{-1}$ and then $F_{0}^{-1}$, so that the segment

$$
\begin{equation*}
\Delta^{(2)}(0,1)=F_{0}^{-1}\left(F_{1}^{-1}\left(\Delta^{(0)}\right)\right) . \tag{L.34}
\end{equation*}
$$

The mapping of the unit interval into a smaller interval is what determines its label. The sequence of the labels of the inverse maps is the label of the segment:

$$
\Delta^{(n)}\left(\epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n}\right)=F_{\epsilon_{1}}^{-1} \circ F_{\epsilon_{2}}^{-1} \circ \cdots F_{\epsilon_{n}}^{-1}\left(\Delta^{(0)}\right)
$$

The scaling function is formed from a set of ratios of segments length. We use $|\cdot|$ around a segment $\Delta^{(n)}(\epsilon)$ to denote its size (length), and define

$$
\sigma^{(n)}\left(\epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n}\right)=\frac{\left|\Delta^{(n)}\left(\epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n}\right)\right|}{\left|\Delta^{(n-1)}\left(\epsilon_{2}, \ldots, \epsilon_{n}\right)\right|} .
$$

We can then arrange the ratios $\sigma^{(n)}\left(\epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n}\right)$ next to each other as piecewise constant segments in increasing order of their binary label $\epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n}$ so that the collection of steps scan the unit interval. As $n \rightarrow \infty$ this collection of steps will converge to the scaling function.

## L. 5 Geometrization

The $\mathcal{L}$ operator is a generalization of the transfer matrix. It gets more by considering less of the matrix: instead of considering the whole matrix it is possible to consider just one of the rows of the matrix. The $\mathcal{L}$ operator also makes explicit the vector space in which it acts: that of the
observable functions. Observables are functions that to each configuration of the system associate a number: the energy, the average magnetization, the correlation between two sites. It is in the average of observables that one is interested in. Like the transfer matrix, the $\mathcal{L}$ operator considers only semi-infinite systems, that is, only the part of the interaction between spins to the right is taken into account. This may sound un-symmetric, but it is a simple way to count each interaction only once, even in cases where the interaction includes three or more spin couplings. To define the $\mathcal{L}$ operator one needs the interaction energy between one spin and all the rest to its right, which is given by the function $\phi$. The $\mathcal{L}$ operators defined as

$$
\mathcal{L} g(\sigma)=\sum_{\sigma_{0} \in \Omega_{0}} g\left(\sigma_{0} \sigma\right) e^{-\beta \phi\left(\sigma_{0} \sigma\right)} .
$$

To each possible value in $\Omega_{0}$ that the spin $\sigma_{0}$ can assume, an average of the observable $g$ is computed weighed by the Boltzmann factor $e^{-\beta \phi}$. The formal relations that stem from this definition are its relation to the free energy when applied to the observable $\iota$ that returns one for any configuration:

$$
-\beta f(\beta)=\lim _{n \rightarrow \infty} \frac{1}{n} \ln \left\|\mathcal{L}^{n} \iota\right\|
$$

and the thermodynamic average of an observable

$$
\langle g\rangle=\lim _{n \rightarrow \infty} \frac{\left\|\mathcal{L}^{n} g\right\|}{\left\|\mathcal{L}^{\mathcal{L}^{n}}\right\|}
$$

Both relations hold for almost all configurations. These relations are part of theorem of Ruelle that enlarges the domain of the Perron-Frobenius theorem and sharpens its results. The theorem shows that just as the transfer matrix, the largest eigenvalue of the $\mathcal{L}$ operator is related to the free-energy of the spin system. It also hows that there is a formula for the eigenvector related to the largest eigenvalue. This eigenvector $|\rho\rangle$ (or the corresponding one for the adjoint $\mathcal{L}^{*}$ of $\mathcal{L}$ ) is the Gibbs state of the system. From it all averages of interest in statistical mechanics can be computed from the formula

$$
\langle g\rangle=\langle\rho| g|\rho\rangle .
$$

The Gibbs state can be expressed in an explicit form in terms of the interactions, but it is of little computational value as it involves the Gibbs state for a related spin system. Even then it does have an enormous theoretical value. Later we will see how the formula can be used to manipulate the space of observables into a more convenient space.

The geometrization of a spin system converts the shift dynamics (necessary to define the Ruelle operator) into a smooth dynamics. This is equivalent to the mathematical problem in ergodic theory of finding a smooth embedding for a given Bernoulli map.

The basic idea for the dynamics is to establish the a set of maps $F_{\sigma_{k}}$ such that

$$
F_{\sigma_{k}}(0)=0
$$

and

$$
F_{\sigma_{1}} \circ F_{\sigma_{2}} \circ \cdots \circ F_{\sigma_{n}}(0)=\phi\left(+, \sigma_{1}, \sigma_{2}, \ldots, \sigma_{n},-,-, \ldots\right) .
$$

This is a formal relation that expresses how the interaction is to be converted into a dynamical systems. In most examples $F_{\sigma_{k}}$ is a collection of maps from a subset of $R^{D}$ to itself.

If the interaction is complicated, then the dimension of the set of maps may be infinite. If the resulting dynamical system is infinite have we gained anything from the transformation? The gain in this case is not in terms of added speed of convergence to the thermodynamic limit, but in the fact that the Ruelle operator is of trace-class and all eigenvalues are related to the spin system and not artifacts of the computation.

The construction of the higher dimensional system is done by borrowing the phase space reconstruction technique from dynamical systems. Phase space reconstruction can be done in several ways: by using delay coordinates, by using derivatives of the position, or by considering the value of several independent observables of the system. All these may be used in the construction of the equivalent dynamics. Just as in the study of dynamical systems, the exact method does not matter for the determination of the thermodynamics $(f(\alpha)$ spectra, generalized dimension), also in the construction of the equivalent dynamics the exact choice of observable does not matter.

We will only consider configurations for the half line. This is bescause for translational invariant interactions the thermodynamic limit on half line is the same as in the whole line. One can prove this by considering the difference in a thermodynamic average in the line and in the semiline and compare the two as the size of the system goes to infinity.

When the interactions are long range in principle one has to specify the boundary conditions to be able to compute the interaction energy of a configuration in a finite box. If there are no phase transitions for the interaction, then which boundary conditions are chosen is irrelevant in the thermodynamic limit. When computing quantities with the transfer matrix, the long rrange interaction is truncated at some finite range and the truncated interaction is then use to evaluate the transfer matrix. With the Ruelle operator the interaction is never truncated, and the boundary must be specified.

The interaction $\phi(\sigma)$ is any function that returns a number on a configuration. In general it is formed from pairwise spin interactions

$$
\phi(\sigma)=\sum_{n>0} \delta_{\sigma_{0}, \sigma_{n}} J(n)
$$

with different choices of $J(n)$ leading to differnt models. If $J(n)=1$ only if $n=1$ and ) otherwise, then one has the nearest neighbor Ising model. If $J(n)=n^{-2}$, then one has the inverse square model relevant in the study of the Kondo problem.

Let us say that each site of the lattice can assume two values,+- and the set of all possible configurations of the semiline is the set $\Omega$. Then an observable $g$ is a function from the set of configurations $\Omega$ to the reals. Each configuration is indexed by the integers from 0 up, and it is useful to think of the configuration as a string of spins. One can append a spin $\eta_{0}$ to its begining, $\eta \vee \sigma$, in which case $\eta$ is at site 0 , $\omega_{0}$ at site 1 , and so on.

The Ruelle operator $\mathcal{L}$ is defined as

$$
\mathcal{L} g(\eta)=\sum_{\omega_{0} \in \Omega_{0}} g\left(\omega_{0} \vee \eta\right) e^{-\beta \phi\left(\omega_{0} \vee \eta\right)}
$$

This is a positive and bounded operator over the space of bounded observables. There is a generalization of the Perron-Frobenius theorem by Ruelle that establishes that the largest eigenvalue of $\mathcal{L}$ is isolated from the rest of the spectrum and gives the thermodynamics of the spin system just as the largest eigenvalue of the transfer matrix does. Ruelle alos gave a formula for the eigenvector related to the largest eigenvalue.

The difficulty with it is that the relation between the partition function and the trace of its $n$th power, $\operatorname{tr} \mathcal{L}^{n}=Z_{n}$ no longer holds. The reason is that the trace of the Ruelle operator is ill-defined, it is infinite.

We now introduce a special set of observables $\left\{x_{1}(\sigma), \ldots, x_{1}(\sigma)\right\}$. The idea is to choose the observables in such a way that from their values on a particular configuration $\sigma$ the configuration can be reconstructed. We also introduce the interaction observables $h_{\sigma_{0}}$

To geometrize spin systems, the interactions are assumed to be translationally invariant. The spins $\sigma_{k}$ will only assume a finite number of values. For simplicity, we will take the interaction $\phi$ among the spins to depend only on pairwise interactions,

$$
\begin{equation*}
\phi(\sigma)=\phi\left(\sigma_{0}, \sigma_{1}, \sigma_{2}, \ldots\right)=J_{0} \sigma_{0}+\sum_{n>0} \delta_{\sigma_{0}, \sigma_{n}} J_{1}(n), \tag{L.35}
\end{equation*}
$$

and limit $\sigma_{k}$ to be in $\{+,-\}$. For the one-dimensional Ising model, $J_{0}$ is the external magnetic field and $J_{1}(n)=1$ if $n=1$ and 0 otherwise.

For an exponentially decaying interaction $J_{1}(n)=e^{-\alpha n}$. Two- and threedimensional models can be considered in this framework. For example, a strip of spins of $L \times \infty$ with helical boundary conditions is modeled by the potential $J_{1}(n)=\delta_{n, 1}+\delta_{n, L}$.

The transfer operator $\mathcal{T}$ was introduced by Kramers and Wannier [L.12] to study the Ising model on a strip and concocted so that the trace of its $n$th power is the partition function $Z_{n}$ of system when one of its dimensions is $n$. The method can be generalized to deal with any finite-range interaction. If the range of the interaction is $L$, then $\mathcal{T}$ is a matrix of size $2^{L} \times 2^{L}$. The longer the range, the larger the matrix. When the range of the interaction is infinite one has to define the $\mathcal{T}$ operator by its action on an observable $g$. Just as the observables in quantum mechanics, $g$ is a function that associates a number to every state (configuration of spins). The energy density and the average magnetization are examples of observables. From this equivalent definition one can recover the usual transfer matrix by making all quantities finite range. For a semi-infinite configuration $\sigma=\left\{\sigma_{0}, \sigma_{1}, \ldots\right\}$ :

$$
\begin{equation*}
\mathcal{T} g(\sigma)=g(+\vee \sigma) e^{-\beta \phi(+\vee \sigma)}+g(-\vee \sigma) e^{-\beta \phi(-\vee \sigma)} \tag{L.36}
\end{equation*}
$$

By $+\vee \sigma$ we mean the configuration obtained by prepending + to the beginning of $\sigma$ resulting in the configuration $\left\{+, \sigma_{0}, \sigma_{1}, \ldots\right\}$. When the range becomes infinite, $\operatorname{tr} \mathcal{T}^{n}$ is infinite and there is no longer a connection between the trace and the partition function for a system of size $n$ (this is a case where matrices give the wrong intuition). Ruelle [L.13] generalized the Perron-Frobenius theorem and showed that even in the case of infinite range interactions the largest eigenvalue of the $\mathcal{T}$ operator is related to the free-energy of the spin system and the corresponding eigenvector is related to the Gibbs state. By applying $\mathcal{T}$ to the constant observable $u$, which returns 1 for any configuration, the free energy per site $f$ is computed as

$$
\begin{equation*}
-\beta f(\beta)=\lim _{n \rightarrow \infty} \frac{1}{n} \ln \left\|\mathcal{T}^{n} u\right\| \tag{L.37}
\end{equation*}
$$

To construct a smooth dynamical system that reproduces the properties of $\mathcal{T}$, one uses the phase space reconstruction technique of Packard et al. [L.6] and Takens [L.7], and introduces a vector of state observables $x(\sigma)=\left\{x_{1}(\sigma), \ldots, x_{D}(\sigma)\right\}$. To avoid complicated notation we will limit the discussion to the example $x(\sigma)=\left\{x_{+}(\sigma), x_{-}(\sigma)\right\}$, with $x_{+}(\sigma)=\phi(+\vee \sigma)$ and $x_{-}(\sigma)=\phi(-\vee \sigma)$; the more general case is similar and used in a later example. The observables are restricted to those $g$ for which, for all configurations $\sigma$, there exist an analytic function $G$ such that $G\left(x_{1}(\sigma), \ldots, x_{D}(\sigma)\right)=$ $g(\sigma)$. This at first seems a severe restriction as it may exclude the eigenvector corresponding to the Gibbs state. It can be checked that this is not the case by using the formula given by Ruelle [L.14] for this eigenvector. A simple example where this formalism can be carried out is for the interaction $\phi(\sigma)$ with pairwise exponentially decaying potential $J_{1}(n)=a^{n}$ (with $|a|<1$ ). In this case $\phi(\sigma)=\sum_{n>0} \delta_{\sigma_{0}, \sigma_{n}} a^{n}$ and the state observables
are $x_{+}(\sigma)=\sum_{n>0} \delta_{+, \sigma_{n}} a^{n}$ and $x_{-}(\sigma)=\sum_{n>0} \delta_{-, \sigma_{n}} a^{n}$. In this case the observable $x_{+}$gives the energy of + spin at the origin, and $x_{-}$the energy of $\mathrm{a}-$ spin.

Using the observables $x_{+}$and $x_{-}$, the transfer operator can be reexpressed as

$$
\begin{equation*}
\mathcal{T} G(x(\sigma))=\sum_{\eta \in\{+,-\}} G\left(x_{+}(\eta \vee \sigma), x_{-}(\eta \vee \sigma)\right) e^{-\beta x_{\eta}(\sigma)} \tag{L.38}
\end{equation*}
$$

In this equation the only reference to the configuration $\sigma$ is when computing the new observable values $x_{+}(\eta \vee \sigma)$ and $x_{-}(\eta \vee \sigma)$. The iteration of the function that gives these values in terms of $x_{+}(\sigma)$ and $x_{-}(\sigma)$ is the dynamical system that will reproduce the properties of the spin system. For the simple exponentially decaying potential this is given by two maps, $F_{+}$and $F_{-}$. The map $F_{+}$takes $\left\{x_{+}(\sigma), x_{+}(\sigma)\right\}$ into $\left\{x_{+}(+\vee \sigma), x_{-}(+\vee \sigma)\right\}$ which is $\left\{a\left(1+x_{+}\right), a x_{-}\right\}$and the map $F_{-}$takes $\left\{x_{+}, x_{-}\right\}$into $\left\{a x_{+}, a\left(1+x_{-}\right)\right\}$. In a more general case we have maps $F_{\eta}$ that take $x(\sigma)$ to $x(\eta \vee \sigma)$.

We can now define a new operator $\mathcal{L}$

$$
\begin{equation*}
\mathcal{L} G(x) \stackrel{\text { def }}{=} \mathcal{T} G(x(\sigma))=\sum_{\eta \in\{+,-\}} G\left(F_{\eta}(x)\right) e^{-\beta x_{\eta}} \tag{L.39}
\end{equation*}
$$

where all dependencies on $\sigma$ have disappeared - if we know the value of the state observables $x$, the action of $\mathcal{L}$ on $G$ can be computed.

A dynamical system is formed out of the maps $F_{\eta}$. They are chosen so that one of the state variables is the interaction energy. One can consider the two maps $F_{+}$and $F_{-}$as the inverse branches of a hyperbolic map $f$, that is, $f^{-1}(x)=\left\{F_{+}(x), F_{-}(x)\right\}$. Studying the thermodynamics of the interaction $\phi$ is equivalent to studying the long term behavior of the orbits of the map $f$, achieving the transformation of the spin system into a dynamical system.

Unlike the original transfer operator, the $\mathcal{L}$ operator - acting in the space of observables that depend only on the state variables - is of traceclass (its trace is finite). The finite trace gives us a chance to relate the trace of $\mathcal{L}^{n}$ to the partition function of a system of size $n$. We can do better. As most properties of interest (thermodynamics, fall-off of correlations) are determined directly from its spectrum, we can study instead the zeros of the Fredholm determinant $\operatorname{det}(1-z \mathcal{L})$ by the technique of cycle expansions developed for dynamical systems [18.2]. A cycle expansion consists of finding a power series expansion for the determinant by writing $\operatorname{det}(1-z \mathcal{L})=\exp (\operatorname{tr} \ln (1-z \mathcal{L}))$. The logarithm is expanded into a power series and one is left with terms of the form $\operatorname{tr} \mathcal{L}^{n}$ to evaluate. For evaluating the trace, the $\mathcal{L}$ operator is equivalent to

$$
\begin{equation*}
\mathcal{L} G(x)=\int_{\mathbf{R}^{D}} d y \delta(y-f(x)) e^{-\beta y} G(y) \tag{L.40}
\end{equation*}
$$

from which the trace can be computed:

$$
\begin{equation*}
\operatorname{tr} \mathcal{L}^{n}=\sum_{x=f^{(\circ n)}(x)} \frac{e^{-\beta H(x)}}{\left|\operatorname{det}\left(1-\partial_{x} f^{(o n)}(x)\right)\right|} \tag{L.41}
\end{equation*}
$$

with the sum running over all the fixed points of $f^{(\circ n)}$ (all spin configurations of a given length). Here $f^{(o n)}$ is $f$ composed with itself $n$ times, and $H(x)$ is the energy of the configuration associated with the point $x$. In practice the map $f$ is never constructed and the energies are obtained directly from the spin configurations.

To compute the value of $\operatorname{tr} \mathcal{L}^{n}$ we must compute the value of $\partial_{x} f^{(\circ n)}$; this involves a functional derivative. To any degree of accuracy a number $x$ in the range of possible interaction energies can be represented by a finite string of spins $\epsilon$, such as $x=\phi\left(+, \epsilon_{0}, \epsilon_{1}, \ldots,-,-, \ldots\right)$. By choosing the sequence $\epsilon$ to have a large sequence of spins -, the number $x$ can be made as small as needed, so in particular we can represent a small variation by $\phi(\eta)$. As $x_{+}(\epsilon)=\phi(+\vee \epsilon)$, from the definition of a derivative we have:

$$
\begin{equation*}
\partial_{x} f(x)=\lim _{m \rightarrow \infty} \frac{\phi\left(\epsilon \vee \eta^{(m)}\right)-\phi(\epsilon)}{\phi\left(\eta^{(m)}\right)} \tag{L.42}
\end{equation*}
$$

where $\eta^{(m)}$ is a sequence of spin strings that make $\phi\left(\eta^{(m)}\right)$ smaller and smaller. By substituting the definition of $\phi$ in terms of its pairwise interaction $J(n)=n^{s} a^{n \gamma}$ and taking the limit for the sequences $\eta^{(m)}=$ $\left\{+,-,-\ldots, \eta_{m+1}, \eta_{m+2}, \ldots\right\}$ one computes that the limit is $a$ if $\gamma=1$, 1 if $\gamma<1$, and 0 if $\gamma>1$. It does not depend on the positive value of $s$. When $\gamma<1$ the resulting dynamical system is not hyperbolic and the construction for the operator $\mathcal{L}$ fails, so one cannot apply it to potentials such as $(1 / 2)^{\sqrt{n}}$. One may solve this problem by investigating the behavior of the formal dynamical system as $\gamma \rightarrow 0$.

The manipulations have up to now assumed that the map $f$ is smooth. If the dimension $D$ of the embedding space is too small, $f$ may not be smooth. Determining under which conditions the embedding is smooth is a complicated question [L.15]. But in the case of spin systems with pairwise interactions it is possible to give a simple rule. If the interaction is of the form

$$
\begin{equation*}
\phi(\sigma)=\sum_{n \geq 1} \delta_{\sigma_{0}, \sigma_{n}} \sum_{k} p_{k}(n) a_{k}^{n^{\gamma}} \tag{L.43}
\end{equation*}
$$

where $p_{k}$ are polynomials and $\left|a_{k}\right|<1$, then the state observables to use are $x_{s, k}(\sigma)=\sum \delta_{+, \sigma_{n}} n^{s} a_{k}^{n}$. For each $k$ one uses $x_{0, k}, x_{1, k}, \ldots$ up to the largest power in the polynomial $p_{k}$. An example is the interaction with $J_{1}(n)=n^{2}(3 / 10)^{n}$. It leads to a 3 -dimensional system with variables $x_{0,0}$, $x_{1,0}$, and $x_{2,0}$. The action of the map $F_{+}$for this interaction is illustrated

Figure L.3: The spin adding map $F_{+}$for the potential $J(n)=\sum n^{2} a^{\alpha n}$. The action of the map takes the value of the interaction energy between + and the semi-infinite configuration $\left\{\sigma_{1}, \sigma_{2}, \sigma_{3}, \ldots\right\}$ and returns the interaction energy between + and the configuration $\left\{+, \sigma_{1}, \sigma_{2}, \sigma_{3}, \ldots\right\}$.



Figure L.4: Number of digits for the Fredholm method (•) and the transfer function method $(\times)$. The size refers to the largest cycle considered in the Fredholm expansions, and the truncation length in the case of the transfer matrix.
figure L.3. Plotted are the pairs $\{\phi(+\vee \sigma), \phi(+\vee+\vee \sigma)\}$. This can be seen as the strange attractor of a chaotic system for which the variables $x_{0,0}$, $x_{1,0}$, and $x_{2,0}$ provide a good (analytic) embedding.

The added smoothness and trace-class of the $\mathcal{L}$ operator translates into faster convergence towards the thermodynamic limit. As the reconstructed dynamics is analytic, the convergence towards the thermodynamic limit is faster than exponential [2.18, L.16]. We will illustrate this with the polynomial-exponential interactions (L.43) with $\gamma=1$, as the convergence is certainly faster than exponential if $\gamma>1$, and the case of $a^{n}$ has been studied in terms of another Fredholm determinant by Gutzwiller [L.17]. The convergence is illustrated in figure L. 4 for the interaction $n^{2}(3 / 10)^{n}$. Plotted in the graph, to illustrate the transfer matrix convergence, are the number of decimal digits that remain unchanged as the range of the interaction is increased. Also in the graph are the number of decimal digits that remain unchanged as the largest power of $\operatorname{tr} \mathcal{L}^{n}$ considered. The plot is effectively a logarithmic plot and straight lines indicate exponentially fast convergence. The curvature indicates that the convergence is faster than exponential. By fitting, one can verify that the free energy is converging to its limiting value as $\exp \left(-n^{(4 / 3)}\right)$. Cvitanović [2.18] has estimated that the Fredholm determinant of a map on a $D$ dimensional space should converge as $\exp \left(-n^{(1+1 / D)}\right)$, which is confirmed by these numerical simulations.

## Commentary

Remark L. 1 Presentation functions. The best place to read about Feigenbaum's work is in his review article published in Los Alamos Science (reproduced in various
reprint collections and conference proceedings, such as ref. [18.5]). Feigenbaum's Journal of Statistical Physics article [24.13] is the easiest place to learn about presentation functions.

Remark L. 2 Interactions are smooth In most computational schemes for thermodynamic quantities the translation invariance and the smoothness of the basic interaction are never used. In Monte Carlo schemes, aside from the periodic boundary conditions, the interaction can be arbitrary. In principle for each configuration it could be possible to have a different energy. Schemes such as the Sweneson-Wang cluster flipping algorithm use the fact that interaction is local and are able to obtain dramatic speed-ups in the equilibration time for the dynamical Monte Carlo simulation. In the geometrization program for spin systems, the interactions are assumed translation invariant and smooth. The smoothness means that any interaction can be decomposed into a series of terms that depend only on the spin arrangement and the distance between spins:

$$
\phi\left(\sigma_{0}, \sigma_{1}, \sigma_{2}, \ldots\right)=J_{0} \sigma_{0}+\sum \delta\left(\sigma_{0}, \sigma_{n}\right) J_{1}(n)+\sum \delta\left(\sigma_{0}, \sigma_{n_{1}}, \sigma_{n_{2}}\right) J_{2}\left(n_{1}, n_{2}\right)+\cdots
$$

where the $J_{k}$ are symmetric functions of their arguments and the $\delta$ are arbitrary discrete functions. This includes external constant fields ( $J_{0}$ ), but it excludes site dependent fields such as a random external magnetic field.

## Résumé

The geometrization of spin systems strengthens the connection between statistical mechanics and dynamical systems. It also further establishes the value of the Fredholm determinant of the $\mathcal{L}$ operator as a practical computational tool with applications to chaotic dynamics, spin systems, and semiclassical mechanics. The example above emphasizes the high accuracy that can be obtained: by computing the shortest 14 periodic orbits of period 5 or less it is possible to obtain three digit accuracy for the free energy. For the same accuracy with a transfer matrix one has to consider a $256 \times 256$ matrix. This make the method of cycle expansions practical for analytic calculations.

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## Exercises

Exercise L. 1 Not all Banach spaces are also Hilbert If we are given a norm $\|\cdot\|$ of a Banach space $B$, it may be possible to find an inner product $\langle\cdot, \cdot\rangle($ so that $B$ is also a Hilbert space $H$ ) such that for all vectors $f \in B$, we have

$$
\|f\|=\langle f, f\rangle^{1 / 2} .
$$

This is the norm induced by the scalar product. If we cannot find the inner product how do we know that we just are not being clever enough? By checking the parallelogram law for the norm. A Banach space can be made into a Hilbert space if and only if the norm satisfies the parallelogram law. The parallelogram law says that for any two vectors $f$ and $g$ the equality

$$
\|f+g\|^{2}+\|f-g\|^{2}=2\|f\|^{2}+2\|g\|^{2},
$$

must hold.
Consider the space of bounded observables with the norm given by $\|a\|=$ $\sup _{\sigma \in \Omega^{\mathbb{N}}}|a(\sigma)|$. Show that ther is no scalar product that will induce this norm.

Exercise L. 2 Automaton for a droplet Find the Markov graph and the weights on the edges so that the energies of configurations for the dropolet model are correctly generated. For any string starting in zero and ending in zero your diagram should yield a configuration the weight $e^{H(\sigma)}$, with $H$ computed along the lines of (L.13) and (L.18).

Hint: the Markov graph is infinite.

Exercise L. 3 Spectral determinant for $a^{n}$ interactions Compute the spectral determinant for one-dimensional Ising model with the interaction

$$
\phi(\sigma)=\sum_{k>0} a^{k} \delta\left(\sigma_{0}, \sigma_{k}\right) .
$$

Take $a$ as a number smaller than $1 / 2$.
(a) What is the dynamical system this generates? That is, find $F_{+}$and $F_{-}$ as used in (L.39).
(b) Show that

$$
\frac{d}{d x} F_{\{+ \text {or }-\}}=\left[\begin{array}{ll}
a & 0 \\
0 & a
\end{array}\right]
$$

## Exercise L. 4 Ising model on a thin strip Compute the transfer matrix

 for the Ising model defined on the graph

Assume that whenever there is a bond connecting two sites, there is a contribution $J \delta\left(\sigma_{i}, \sigma_{j}\right)$ to the energy.

Exercise L. 5 Infinite symbolic dynamics Let $\sigma$ be a function that returns zeo or one for every infinite binary string: $\sigma:\{0,1\}^{\mathbb{N}} \rightarrow\{0,1\}$. Its value is represented by $\sigma\left(\epsilon_{1}, \epsilon_{2}, \ldots\right)$ where the $\epsilon_{i}$ are either 0 or 1 . We will now define an operator $\mathcal{T}$ that acts on observables on the space of binary strings. A function $a$ is an observable if it has bounded variation, that is, if

$$
\|a\|=\sup _{\left\{\epsilon_{i}\right\}}\left|a\left(\epsilon_{1}, \epsilon_{2}, \ldots\right)\right|<\infty .
$$

For these functions

$$
\mathcal{T} a\left(\epsilon_{1}, \epsilon_{2}, \ldots\right)=a\left(0, \epsilon_{1}, \epsilon_{2}, \ldots\right) \sigma\left(0, \epsilon_{1}, \epsilon_{2}, \ldots\right)+a\left(1, \epsilon_{1}, \epsilon_{2}, \ldots\right) \sigma\left(1, \epsilon_{1}, \epsilon_{2}, \ldots\right) .
$$

The function $\sigma$ is assumed such that any of $\mathcal{T}$ 's "matrix representations" in (a) have the Markov property (the matrix, if read as an adjacency graph, corresponds to a graph where one can go from any node to any other node).
(a) (easy) Consider a finite version $T_{n}$ of the operator $\mathcal{T}$ :

$$
\begin{aligned}
& T_{n} a\left(\epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n}\right)= \\
& \quad a\left(0, \epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n-1}\right) \sigma\left(0, \epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n-1}\right)+ \\
& a\left(1, \epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n-1}\right) \sigma\left(1, \epsilon_{1}, \epsilon_{2}, \ldots, \epsilon_{n-1}\right) .
\end{aligned}
$$

Show that $T_{n}$ is a $2^{n} \times 2^{n}$ matrix. Show that its trace is bounded by a number independent of $n$.
(b) (medium) With the operator norm induced by the function norm, show that $\mathcal{T}$ is a bounded operator.
(c) (hard) Show that $\mathcal{T}$ is not trace-class. (Hint: check if $\mathcal{T}$ is compact).

Classes of operators are nested; trace-class $\leq$ compact $\leq$ bounded.

## Appendix M

# Noise/quantum corrections 

(G. Vattay)


The Gutzwiller trace formula is only a good approximation to the quantum mechanics when $\hbar$ is small. Can we improve the trace formula by adding quantum corrections to the semiclassical terms? A similar question can be posed when the classical deterministic dynamics is disturbed by some way Gaussian white noise with strength $D$. The deterministic dynamics then can be considered as the weak noise limit $D \rightarrow 0$. The effect of the noise can be taken into account by adding noise corrections to the classical trace formula. A formal analogy exists between the noise and the quantum problem. This analogy allows us to treat the noise and quantum corrections together.

## M. 1 Periodic orbits as integrable systems

From now on, we use the language of quantum mechanics, since it is more convenient to visualize the results there. Where it is necessary we will discuss the difference between noise and quantum cases.

First we would like to introduce periodic orbits from an unusual point of view, which can convince you, that chaotic and integrable systems are in fact not as different from each other, than we might think. If we start orbits in the neighborhood of a periodic orbit and look at the picture on the Poincaré section we can see a regular picture. For stable periodic orbits the points form small ellipses around the center and for unstable orbits they form hyperbolas (See Fig. M.1).

The motion close to a periodic orbits is regular in both cases. This is due to the fact, that we can linearize the Hamiltonian close to an orbit, and

Figure M.1: Poincaré section close to a stable and an unstable periodic orbit
linear systems are always integrable. The linearized Hamilton's equations close to the periodic orbit $\left(q_{p}(t)+q, p_{p}(t)+p\right)$ look like

$$
\begin{align*}
\dot{q} & =+\partial_{p q}^{2} H\left(q_{p}(t), p_{p}(t)\right) q+\partial_{p p}^{2} H\left(q_{p}(t), p_{p}(t)\right) p,  \tag{M.1}\\
\dot{p} & =-\partial_{q q}^{2} H\left(q_{p}(t), p_{p}(t)\right) q-\partial_{q p}^{2} H\left(q_{p}(t), p_{p}(t)\right) p, \tag{M.2}
\end{align*}
$$

where the new coordinates $q$ and $p$ are relative to a periodic orbit. This linearized equation can be regarded as a $d$ dimensional oscillator with time periodic frequencies. These equations are representing the equation of motion in a redundant way since more than one combination of $q, p$ and $t$ determines the same point of the phase space. This can be cured by an extra restriction on the variables, a constraint the variables should fulfill. This constraint can be derived from the time independence or stacionarity of the full Hamiltonian

$$
\begin{equation*}
\partial_{t} H\left(q_{p}(t)+q, p_{p}(t)+p\right)=0 . \tag{M.3}
\end{equation*}
$$

Using the linearized form of this constraint we can eliminate one of the linearized equations. It is very useful, although technically difficult, to do one more transformation and to introduce a coordinate, which is parallel with the Hamiltonian flow $\left(x_{\|}\right)$and others which are orthogonal. In the orthogonal directions we again get linear equations. These equations with $x_{\|}$dependent rescaling can be transformed into normal coordinates, so that we get tiny oscillators in the new coordinates with constant frequencies. This result has first been derived by Poincaré for equilibrium points and later it was extended for periodic orbits by V.I. Arnol'd and co-workers. In the new coordinates, the Hamiltonian reads as

$$
\begin{equation*}
H_{0}\left(x_{\|}, p_{\|}, x_{n}, p_{n}\right)=\frac{1}{2} p_{\|}^{2}+U\left(x_{\|}\right)+\sum_{n=1}^{d-1} \frac{1}{2}\left(p_{n}^{2} \pm \omega_{n}^{2} x_{n}^{2}\right) \tag{M.4}
\end{equation*}
$$

which is the general form of the Hamiltonian in the neighborhood of a periodic orbit. The $\pm$ sign denotes, that for stable modes the oscillator potential is positive while for an unstable mode it is negative. For the unstable modes, $\omega$ is the Lyapunov exponent of the orbit

$$
\begin{equation*}
\omega_{n}=\ln \Lambda_{p, n} / T_{p} \tag{M.5}
\end{equation*}
$$

where $\Lambda_{p, n}$ is the expanding eigenvalue of the Jacobi matrix. For the stable directions the eigenvalues of the Jacobi matrix are connected with $\omega$ as

$$
\begin{equation*}
\Lambda_{p, n}=e^{-i \omega_{n} T_{p}} \tag{M.6}
\end{equation*}
$$

The Hamiltonian close to the periodic orbit is integrable and can be quantized by the Bohr-Sommerfeld rules. The result of the Bohr-Sommerfeld
quantization for the oscillators gives the energy spectra

$$
\begin{align*}
& E_{n}=\hbar \omega_{n}\left(j_{n}+\frac{1}{2}\right) \text { for stable modes }  \tag{M.7}\\
& E_{n}=-i \hbar \omega_{n}\left(j_{n}+\frac{1}{2}\right) \text { for unstable modes }
\end{align*}
$$

where $j_{n}=0,1, \ldots$. It is convenient to introduce the index $s_{n}=1$ for stable and $s_{n}=-i$ for unstable directions. The parallel mode can be quantized implicitly trough the classical action function of the mode:

$$
\begin{equation*}
\frac{1}{2 \pi} \oint p_{\|} d x_{\|}=\frac{1}{2 \pi} S_{\|}\left(E_{m}\right)=\hbar\left(m+\frac{m_{p} \pi}{2}\right) \tag{M.8}
\end{equation*}
$$

where $m_{p}$ is the topological index of the motion in the parallel direction. This latter condition can be rewritten by a very useful trick into the equivalent form

$$
\begin{equation*}
\left(1-e^{i S_{\|}\left(E_{m}\right) / \hbar-i m_{p} \pi / 2}\right)=0 \tag{M.9}
\end{equation*}
$$

The eigenenergies of a semiclassically quantized periodic orbit are all the possible energies

$$
\begin{equation*}
E=E_{m}+\sum_{n=1}^{d-1} E_{n} \tag{M.10}
\end{equation*}
$$

This relation allows us to change in (M.9) $E_{m}$ with the full energy minus the oscillator energies $E_{m}=E-\sum_{n} E_{n}$. All the possible eigenenergies of the periodic orbit then are the zeroes of the expression

$$
\begin{equation*}
\Delta_{p}(E)=\prod_{j_{1}, \ldots, j_{d-1}}\left(1-e^{i S_{\|}\left(E-\sum_{n} \hbar s_{n} \omega_{n}\left(j_{n}+1 / 2\right)\right) / \hbar-i m_{p} \pi / 2}\right) \tag{M.11}
\end{equation*}
$$

If we Taylor expand the action around $E$ to first order

$$
\begin{equation*}
S_{\|}(E+\epsilon) \approx S_{\|}(E)+T(E) \epsilon \tag{M.12}
\end{equation*}
$$

where $T(E)$ is the period of the orbit, and use the relations of $\omega$ and the eigenvalues of the Jacobi matrix, we get the expression of the Selberg product

$$
\begin{equation*}
\Delta_{p}(E)=\prod_{j_{1}, \ldots, j_{d-1}}\left(1-\frac{e^{i S_{p}(E) / \hbar-i m_{p} \pi / 2}}{\prod_{n} \Lambda_{p, n}^{\left(1 / 2+j_{n}\right)}}\right) . \tag{M.13}
\end{equation*}
$$

If we use the right convention for the square root we get exactly the $d$ dimensional expression of the Selberg product formula we derived from the Gutzwiller trace formula in ? . Just here we derived it in a different way! The function $\Delta_{p}(E)$ is the semiclassical zeta function for one prime orbit.

Now, if we have many prime orbits and we would like to construct a function which is zero, whenever the energy coincides with the BS quantized energy of one of the periodic orbits, we have to take the product of these determinants:

$$
\begin{equation*}
\Delta(E)=\prod_{p} \Delta_{p}(E) \tag{M.14}
\end{equation*}
$$

The miracle of the semiclassical zeta function is, that if we take infinitely many periodic orbits, the infinite product will have zeroes not at these energies, but close to the eigenenergies of the whole system !

So we learned, that both stable and unstable orbits are integrable systems and can be individually quantized semiclassically by the old BohrSommerfeld rules. So we almost completed the program of Sommerfeld to quantize general systems with the method of Bohr. Let us have a remark here. In addition to the Bohr-Sommerfeld rules, we used the unjustified approximation (M.12). Sommerfeld would never do this ! At that point we loose some important precision compared to the $B S$ rules and we get somewhat worse results than a semiclassical formula is able to do. We will come back to this point later when we discuss the quantum corrections. To complete the program of full scale Bohr-Sommerfeld quantization of chaotic systems we have to go beyond the linear approximation around the periodic orbit.

The Hamiltonian close to a periodic orbit in the parallel and normal coordinates can be written as the 'harmonic' plus 'anharmonic' perturbation

$$
\begin{equation*}
H\left(x_{\|}, p_{\|}, x_{n}, p_{n}\right)=H_{0}\left(x_{\|}, p_{\|}, x_{n}, p_{n}\right)+H_{A}\left(x_{\|}, x_{n}, p_{n}\right), \tag{M.15}
\end{equation*}
$$

where the anharmonic part can be written as a sum of homogeneous polynomials of $x_{n}$ and $p_{n}$ with $x_{\|}$dependent coefficients:

$$
\begin{align*}
H_{A}\left(x_{\|}, x_{n}, p_{n}\right) & =\sum_{k=3} H^{k}\left(x_{\|}, x_{n}, p_{n}\right)  \tag{M.16}\\
H^{k}\left(x_{\|}, x_{n}, p_{n}\right) & =\sum_{\sum l_{n}+m_{n}=k} H_{l_{n}, m_{n}}^{k}\left(x_{\|}\right) x_{n}^{l_{n}} p_{n}^{m_{n}} \tag{M.17}
\end{align*}
$$

This classical Hamiltonian is hopeless from Sommerfeld's point of view, since it is non integrable. However, Birkhoff in $1927^{3}$ introduced the concept of normal form, which helps us out from this problem by giving successive integrable approximation to a non-integrable problem. Let's learn a bit more about it!

[^4]
## M. 2 The Birkhoff normal form

Birkhoff studied the canonical perturbation theory close to an equilibrium point of a Hamiltonian. Equilibrium point is where the potential has a minimum $\nabla U=0$ and small perturbations lead to oscillatory motion. We can linearize the problem and by introducing normal coordinates $x_{n}$ and conjugate momentums $p_{n}$ the quadratic part of the Hamiltonian will be a set of oscillators

$$
\begin{equation*}
H_{0}\left(x_{n}, p_{n}\right)=\sum_{n=1}^{d} \frac{1}{2}\left(p_{n}^{2}+\omega_{n}^{2} x_{n}^{2}\right) \tag{M.18}
\end{equation*}
$$

The full Hamiltonian can be rewritten with the new coordinates

$$
\begin{equation*}
H\left(x_{n}, p_{n}\right)=H_{0}\left(x_{n}, p_{n}\right)+H_{A}\left(x_{n}, p_{n}\right), \tag{M.19}
\end{equation*}
$$

where $H_{A}$ is the anharmonic part of the potential in the new coordinates. The anharmonic part can be written as a series of homogeneous polynomials

$$
\begin{align*}
H_{A}\left(x_{n}, p_{n}\right) & =\sum_{j=3}^{\infty} H^{j}\left(x_{n}, p_{n}\right)  \tag{M.20}\\
H^{j}\left(x_{n}, p_{n}\right) & =\sum_{|||+|m|=j} h_{l m}^{j} x^{l} p^{m} \tag{M.21}
\end{align*}
$$

where $h_{l m}^{j}$ are real constants and we used the multi-indices $l:=\left(l_{1}, \ldots, l_{d}\right)$ with definitions

$$
|l|=\sum l_{n}, x^{l}:=x_{1}^{l_{1}} x_{2}^{l_{2}} \ldots x_{d}^{l_{d}} .
$$

Birkhoff showed, that that by successive canonical transformations one can introduce new momentums and coordinates such, that in the new coordinates the anharmonic part of the Hamiltonian up to any given $n$ polynomial will depend only on the variable combination

$$
\begin{equation*}
\tau_{n}=\frac{1}{2}\left(p_{n}^{2}+\omega_{n}^{2} x_{n}^{2}\right), \tag{M.22}
\end{equation*}
$$

where $x_{n}$ and $p_{n}$ are the new coordinates and momentums, but $\omega_{n}$ is the original frequency. This is called the Birkhoff normal form of degree $N$ :

$$
\begin{equation*}
H\left(x_{n}, p_{n}\right)=\sum_{j=2}^{N} H^{j}\left(\tau_{1}, \ldots, \tau_{d}\right) \tag{M.23}
\end{equation*}
$$

where $H^{j}$ are homogeneous degree $j$ polynomials of $\tau$-s. This is an integrable Hamiltonian, the non-integrability is pushed into the remainder, which consists of polynomials of degree higher than $N$. We run into trouble only when the oscillator frequencies are commensurate e.g. it is possible to find a set of integers $m_{n}$ such that the linear combination

$$
\sum_{n=1}^{d} \omega_{n} m_{n}
$$

vanishes. This extra problem has been solved by Gustavson in 1966 and we call the the object Birkhoff-Gustavson normal form. The procedure of the successive canonical transformations can be computerized and can be carried out up to high orders ( $\sim 20$ ).

Of course, we pay a price for forcing the system to be integrable up to degree $N$. For a non-integrable system the high order terms behave quete widely and the series is not convergent. Therefore we have to use this tool carefully. Now, we learned how to approximate a non-integrable system with a sequence of integrable systems and we can go back and carry out the BS quantization.

## M. 3 Bohr-Sommerfeld quantization of periodic orbits

There is some difference between equilibrium points and periodic orbits. The Hamiltonian (M.4) is not a sum of oscillators. One can transform the parallel part, describing circulation along the orbit, into an oscillator Hamiltonian, but this would make the problem extremelly difficult. Therefore, we carry out the canonical transformations dictated by the Birkhoff procedure only in the orthogonal directions. The $x_{\|}$coordinate plays the role of a parameter. After the tasformation up to order $N$ the Hamiltonian (M.17) is

$$
\begin{equation*}
H\left(x_{\|}, p_{\|}, \tau_{1}, \ldots \tau_{d-1}\right)=H_{0}\left(x_{\|}, p_{\|}, \tau_{1}, \ldots, \tau_{d-1}\right)+\sum_{j=2}^{N} U^{j}\left(x_{\|}, \tau_{1}, \ldots, \tau_{d-1}\right),( \tag{M.24}
\end{equation*}
$$

where $U^{j}$ is a $j$ th order homogeneous polynomial of $\tau$-s with $x_{\|}$dependent coefficients. The orthogonal part can be BS quantized by quantizing the individual oscillators, replacing $\tau$-s as we did in (M.8). This leads to a one dimensional effective potential indexed by $j_{1}, \ldots, j_{d-1}$

$$
\begin{align*}
& H\left(x_{\|}, p_{\|}, j_{1}, \ldots, j_{d-1}\right)=\frac{1}{2} p_{\|}^{2}+U\left(x_{\|}\right)+\sum_{n=1}^{d-1} \hbar s_{n} \omega_{n}\left(j_{n}+1 / 2\right)+\quad \text { (M.25) }  \tag{M.25}\\
& +\sum_{k=2}^{N} U^{k}\left(x_{\|}, \hbar s_{1} \omega_{1}\left(j_{1}+1 / 2\right), \hbar s_{2} \omega_{2}\left(j_{2}+1 / 2\right), \ldots, \hbar s_{d-1} \omega_{d-1}\left(j_{d-1}+1 / 2\right)\right),
\end{align*}
$$

where $j_{n}$ can be any non-negative integer. The term with index $k$ is proportional with $\hbar^{k}$ due to the homogeneity of the polynomials.

The parallel mode now can be BS quantized for any given set of $j$-s

$$
\begin{align*}
& S_{p}\left(E, j_{1}, \ldots, j_{d-1}\right)=\oint p_{\|} d x_{\|}=  \tag{M.26}\\
& =\oint d x_{\|} \sqrt{E-\sum_{n=1}^{d-1} \hbar s_{n} \omega_{n}\left(j_{n}+1 / 2\right)-U\left(x_{\|}, j_{1}, \ldots, j_{d-1}\right)}=2 \pi \hbar\left(m+m_{p} / 2\right)
\end{align*}
$$

where $U$ contains all the $x_{\|}$dependent terms of the Hamiltonian. The spectral determinant becomes

$$
\begin{equation*}
\Delta_{p}(E)=\prod_{j_{1}, \ldots, j_{d-1}}\left(1-e^{i S_{p}\left(E, j_{1}, \ldots, j_{d-1}\right) / \hbar-m_{p} \pi / 2}\right) . \tag{M.27}
\end{equation*}
$$

This expression completes the Sommerfeld method and tells us how to quantize chaotic or general Hamiltonian systems. Unfortunately, quantum mechanics postponed this nice formula until our book.

This formula has been derived with the help of the semiclassical BohrSommerfeld quantization rule and the classical normal form theory. Indeed, if we expand $S_{p}$ in the exponent in the powers of $\hbar$

$$
S_{p}=\sum_{k=0}^{N} \hbar^{k} S_{k}
$$

we get more than just a constant and a linear term. This formula already gives us corrections to the semiclassical zeta function in all powers of $\hbar$. There is a very attracting feature of this semiclassical expansion. $\hbar$ in $S_{p}$ shows up only in the combination $\hbar s_{n} \omega_{n}\left(j_{n}+1 / 2\right)$. A term proportional with $\hbar^{k}$ can only be a homogeneous expression of the oscillator energies $s_{n} \omega_{n}\left(j_{n}+1 / 2\right)$. For example in two dimensions there is only one possibility of the functional form of the order $k$ term

$$
S_{k}=c_{k}(E) \cdot \omega_{n}^{k}(j+1 / 2)^{k},
$$

where $c_{k}(E)$ is the only function to be determined.
The corrections derived sofar are doubly semiclassical, since they give semiclassical corrections to the semiclassical approximation. What can quantum mechanics add to this ? As we have stressed in the previous section, the exact quantum mechanics is not invariant under canonical transformations. In other context, this phenomenon is called the operator ordering problem. Since the operators $\hat{x}$ and $\hat{p}$ do not commute, we
run into problems, when we would like to write down operators for classical quantities like $x^{2} p^{2}$. On the classical level the four possible orderings $x p x p$, $p p x x, p x p x$ and $x x p p$ are equivalent, but they are different in the quantum case. The expression for the energy (M.26) is not exact. We have to go back to the level of the Schrödinger equation if we would like to get the exact expression.

## M. 4 Quantum calculation of $\hbar$ corrections

The Gutzwiller trace formula has originally been derived from the saddle point approximation of the Feynman path integral form of the propagator. The exact trace is a pathsum for all closed paths of the system

$$
\begin{equation*}
\operatorname{Tr} G\left(x, x^{\prime}, t\right)=\int d x G(x, x, t)=\int \mathcal{D} x e^{i S(x, t) / \hbar} \tag{M.28}
\end{equation*}
$$

where $\int \mathcal{D} x$ denotes the discretization and summation for all paths of time length $t$ in the limit of the infinite refination and $S(x, t)$ is the classical action calculated along the path. The trace in the saddle point calculation is a sum for classical periodic orbits and zero length orbits, since these are the extrema of the action $\delta S(x, t)=0$ for closed paths:

$$
\begin{equation*}
\operatorname{Tr} G\left(x, x^{\prime}, t\right)=g_{0}(t)+\sum_{p \in P O} \int \mathcal{D} \xi_{p} e^{i S\left(\xi_{p}+x_{p}(t), t\right) / \hbar} \tag{M.29}
\end{equation*}
$$

where $g_{0}(t)$ is the zero length orbit contribution. We introduced the new coordinate $\xi_{p}$ with respect to the periodic orbit $x_{p}(t), x=\xi_{p}+x_{p}(t)$. Now, each path sum $\int \mathcal{D} \xi_{p}$ is computed in the vicinity of periodic orbits. Since the saddle points are taken in the configuration space, only spatially distinct periodic orbits, the so called prime periodic orbits, appear in the summation. Sofar nothing new has been invented. If we continue the standard textbook calculation scheme, we have to Taylor expand the action in $\xi_{p}$ and keep the quadratic term in the exponent while treating the higher order terms as corrections. Then we can compute the path integrals with the help of Gaussian integrals. The key point here is that we don't compute the path sum directly. We use the correspondence between path integrals and partial differential equations. This idea comes from Maslov [M.5] and a good summary is in ref. [M.6]. We search for that Schrödinger equation, which leads to the path sum

$$
\begin{equation*}
\int \mathcal{D} \xi_{p} e^{i S\left(\xi_{p}+x_{p}(t), t\right) / \hbar} \tag{M.30}
\end{equation*}
$$

where the action around the periodic orbit is in a multi dimensional Taylor expanded form:

$$
\begin{equation*}
S(x, t)=\sum_{\mathbf{n}}^{\infty} s_{\mathbf{n}}(t)\left(x-x_{p}(t)\right)^{\mathbf{n}} / \mathbf{n}!. \tag{M.31}
\end{equation*}
$$

The symbol $\mathbf{n}=\left(n_{1}, n_{2}, \ldots, n_{d}\right)$ denotes the multi index in $d$ dimensions, $\mathbf{n}!=\prod_{i=1}^{d} n_{i}!$ the multi factorial and $\left(x-x_{p}(t)\right)^{\mathbf{n}}=\prod_{i=1}^{d}\left(x_{i}-x_{p, i}(t)\right)^{n_{i}}$, respectively. The expansion coefficients of the action can be determined from the Hamilton-Jacobi equation

$$
\begin{equation*}
\partial_{t} S+\frac{1}{2}(\nabla S)^{2}+U=0 \tag{M.32}
\end{equation*}
$$

in which the potential is expanded in a multidimensional Taylor series around the orbit

$$
\begin{equation*}
U(x)=\sum_{\mathbf{n}} u_{\mathbf{n}}(t)\left(x-x_{p}(t)\right)^{\mathbf{n}} / \mathbf{n}! \tag{М.33}
\end{equation*}
$$

The Schrödinger equation

$$
\begin{equation*}
i \hbar \partial_{t} \psi=\hat{H} \psi=-\frac{\hbar^{2}}{2} \Delta \psi+U \psi \tag{M.34}
\end{equation*}
$$

with this potential also can be expanded around the periodic orbit. Using the WKB ansatz

$$
\begin{equation*}
\psi=\varphi e^{i S / \hbar} \tag{M.35}
\end{equation*}
$$

we can construct a Schrödinger equation corresponding to a given order of the Taylor expansion of the classical action. The Schrödinger equation induces the Hamilton-Jacobi equation (M.32) for the phase and the transport equation of Maslov and Fjedoriuk [M.7] for the amplitude:

$$
\begin{equation*}
\partial_{t} \varphi+\nabla \varphi \nabla S+\frac{1}{2} \varphi \Delta S-\frac{i \hbar}{2} \Delta \varphi=0 . \tag{M.36}
\end{equation*}
$$

This is the partial differential equation, solved in the neighborhood of a periodic orbit with the expanded action (M.31), which belongs to the local pathsum (M.30).

If we know the Green's function $G_{p}\left(\xi, \xi^{\prime}, t\right)$ corresponding to the local equation (M.36), then the local path sum can be converted back into a trace:

$$
\begin{equation*}
\int \mathcal{D} \xi_{p} e^{i / \hbar \sum_{\mathbf{n}} S_{\mathbf{n}}\left(x_{p}(t), t\right) \xi_{p}^{\mathbf{n}} / \mathbf{n !}}=\operatorname{Tr} G_{p}\left(\xi, \xi^{\prime}, t\right) \tag{M.37}
\end{equation*}
$$

The saddle point expansion of the trace in terms of local traces then becomes

$$
\begin{equation*}
\operatorname{Tr} G\left(x, x^{\prime}, t\right)=\operatorname{Tr} G_{W}\left(x, x^{\prime}, t\right)+\sum_{p} \operatorname{Tr} G_{p}\left(\xi, \xi^{\prime}, t\right) \tag{M.38}
\end{equation*}
$$

where $G_{W}\left(x, x^{\prime}, t\right)$ denotes formally the Green's function expanded around zero length (non moving) periodic orbits, known as the Weyl term [M.8]. Each Green's function can be Fourier-Laplace transformed independently and by definition we get in the energy domain:

$$
\begin{equation*}
\operatorname{Tr} G\left(x, x^{\prime}, E\right)=g_{0}(E)+\sum_{p} \operatorname{Tr} G_{p}\left(\xi, \xi^{\prime}, E\right) . \tag{M.39}
\end{equation*}
$$

Notice, that we do not need here to take further saddle points in time, since we are dealing with exact time and energy domain Green's functions. indexGreen's function!energy dependent

The spectral determinant is a function which has zeroes at the eigenenergies $E_{n}$ of the Hamilton operator $\hat{H}$. Formally it is

$$
\Delta(E)=\operatorname{det}(E-\hat{H})=\prod_{n}\left(E-E_{n}\right) .
$$

The logarithmic derivative of the spectral determinant is the trace of the energy domain Green's function:

$$
\begin{equation*}
\operatorname{Tr} G\left(x, x^{\prime}, E\right)=\sum_{n} \frac{1}{E-E_{n}}=\frac{d}{d E} \log \Delta(E) \tag{M.40}
\end{equation*}
$$

We can define the spectral determinant $\Delta_{p}(E)$ also for the local operators and we can write

$$
\begin{equation*}
\operatorname{Tr} G_{p}\left(\xi, \xi^{\prime}, E\right)=\frac{d}{d E} \log \Delta_{p}(E) \tag{M.41}
\end{equation*}
$$

Using (M.39) we can express the full spectral determinant as a product for the sub-determinants

$$
\Delta(E)=e^{W(E)} \prod_{p} \Delta_{p}(E)
$$

where $W(E)=\int^{E} g_{0}\left(E^{\prime}\right) d E^{\prime}$ is the term coming from the Weyl expansion.
The construction of the local spectral determinants can be done easily. We have to consider the stationary eigenvalue problem of the local Schrödinger problem and keep in mind, that we are in a coordinate system moving together with the periodic orbit. If the classical energy of the periodic orbit coincides with an eigenenergy $E$ of the local Schrödinger equation around the periodic orbit, then the corresponding stationary eigenfunction fulfills

$$
\begin{equation*}
\psi_{p}\left(\xi, t+T_{p}\right)=\int d \xi^{\prime} G_{p}\left(\xi, \xi^{\prime}, t+T_{p}\right) \psi_{p}\left(\xi^{\prime}, t\right)=e^{-i E T_{p} / \hbar} \psi_{p}(\xi, t), \tag{M.42}
\end{equation*}
$$

where $T_{p}$ is the period of the prime orbit $p$. If the classical energy of the periodic orbit is not an eigenenergy of the local Schrödinger equation, the non-stationary eigenfunctions fulfill

$$
\begin{equation*}
\psi_{p}^{1}\left(\xi, t+T_{p}\right)=\int d \xi^{\prime} G_{p}\left(\xi, \xi^{\prime}, t+T_{p}\right) \psi_{p}\left(\xi^{\prime}, t\right)=e^{-i E T_{p} / \hbar} \lambda_{p}^{1}(E) \psi_{p}^{1}(t),(\mathrm{M} \tag{M.43}
\end{equation*}
$$

where $\mathbf{l}=\left(l_{1}, l_{2}, \ldots\right)$ is a multi-index of the possible quantum numbers of the local Schrödinger equation. If the eigenvalues $\lambda_{p}^{1}(E)$ are known the local functional determinant can be written as

$$
\begin{equation*}
\Delta_{p}(E)=\prod_{1}\left(1-\lambda_{p}^{1}(E)\right) \tag{M.44}
\end{equation*}
$$

since $\Delta_{p}(E)$ is zero at the eigenenergies of the local Schrödinger problem. We can insert the ansatz (M.35) and reformulate (M.43) as

$$
\begin{equation*}
e^{\frac{i}{\hbar} S\left(t+T_{p}\right)} \varphi_{p}^{1}\left(t+T_{p}\right)=e^{-i E T_{p} / \hbar} \lambda_{p}^{1}(E) e^{\frac{i}{\hbar} S(t)} \varphi_{p}^{1}(t) \tag{M.45}
\end{equation*}
$$

The phase change is given by the action integral for one period $S\left(t+T_{p}\right)-$ $S(t)=\int_{0}^{T_{p}} L(t) d t$. Using this and the identity for the action $S_{p}(E)$ of the periodic orbit

$$
\begin{equation*}
S_{p}(E)=\oint p d q=\int_{0}^{T_{p}} L(t) d t+E T_{p} \tag{M.46}
\end{equation*}
$$

we get

$$
\begin{equation*}
e^{\frac{i}{\hbar} S_{p}(E)} \varphi_{p}^{1}\left(t+T_{p}\right)=\lambda_{p}^{1}(E) \varphi_{p}^{1}(t) . \tag{M.47}
\end{equation*}
$$

Introducing the eigenequation for the amplitude

$$
\begin{equation*}
\varphi_{p}^{1}\left(t+T_{p}\right)=R_{\mathbf{l}, p}(E) \varphi_{p}^{1}(t) \tag{M.48}
\end{equation*}
$$

the local spectral determinant can be expressed as a product for the quantum numbers of the local problem:

$$
\begin{equation*}
\Delta_{p}(E)=\prod_{1}\left(1-R_{\mathbf{l}, p}(E) e^{\frac{i}{\hbar} S_{p}(E)}\right) \tag{M.49}
\end{equation*}
$$

Since $\hbar$ is a small parameter we can develop a perturbation series for the amplitudes $\varphi_{p}^{1}(t)=\sum_{m=0}^{\infty}\left(\frac{i \hbar}{2}\right)^{m} \varphi_{p}^{1(m)}(t)$ which can be inserted into the
equation (M.36) and we get an iterative scheme starting with the semiclassical solution $\varphi^{1(0)}$ :

$$
\begin{align*}
\partial_{t} \varphi^{1(0)}+\nabla \varphi^{1(0)} \nabla S+\frac{1}{2} \varphi^{1(0)} \Delta S & =0,  \tag{M.50}\\
\partial_{t} \varphi^{1(m+1)}+\nabla \varphi^{\mathbf{l}(m+1)} \nabla S+\frac{1}{2} \varphi^{1(m+1)} \Delta S & =\Delta \varphi^{1(m)} .
\end{align*}
$$

The eigenvalue can also be expanded in powers of $i \hbar / 2$ :

$$
\begin{align*}
R_{\mathbf{l}, p}(E) & =\exp \left\{\sum_{m=0}^{\infty}\left(\frac{i \hbar}{2}\right)^{m} C_{\mathbf{l}, p}^{(m)}\right\}  \tag{M.51}\\
=\exp \left(C_{\mathbf{1}, p}^{(0)}\right)\{1 & +\frac{i \hbar}{2} C_{\mathbf{l}, p}^{(1)}+\left(\frac{i \hbar}{2}\right)^{2}\left(\frac{1}{2}\left(C_{\mathbf{1}, p}^{(1)}\right)^{2}+C_{\mathbf{1}, p}^{(2)}\right)+\ldots \tag{M.52}
\end{align*}
$$

The eigenvalue equation (M.48) in $\hbar$ expanded form reads as

$$
\begin{aligned}
\varphi_{p}^{1(0)}\left(t+T_{p}\right) & =\exp \left(C_{1, p}^{(0)}\right) \varphi_{p}^{1(0)}(t), \\
\varphi_{p}^{1(1)}\left(t+T_{p}\right) & =\exp \left(C_{1, p}^{(0)}\right)\left[\varphi_{p}^{1(1)}(t)+C_{1, p}^{(1)} \varphi_{p}^{1(0)}(t)\right], \\
\varphi_{p}^{1(2)}\left(t+T_{p}\right) & =\exp \left(C_{1, p}^{(0)}\right)\left[\varphi_{p}^{1(2)}(t)+C_{1, p}^{(1)} \varphi_{p}^{1(1)}(t)+\left(C_{1, p}^{(2)}+\frac{1}{2}\left(C_{1, p}^{(1)}\right)^{2}\right) \varphi_{p}^{1}(P\rangle(\text { (b) } 3]\right),
\end{aligned}
$$

and so on. These equations are the conditions selecting the eigenvectors and eigenvalues and they hold for all $t$.

It is very convenient to expand the functions $\varphi_{p}^{1(m)}(x, t)$ in Taylor series around the periodic orbit and to solve the equations (M.51) in this basis [M.10], since only a couple of coefficients should be computed to derive the first corrections. This technical part we are going to publish elsewhere [M.9]. One can derive in general the zero order term $C_{1}^{(0)}=$ $i \pi \nu_{p}+\sum_{i=1}^{d-1}\left(l_{i}+\frac{1}{2}\right) u_{p, i}$, where $u_{p, i}=\log \Lambda_{p, i}$ are the logarithms of the eigenvalues of the monodromy matrix $\mathbf{J}_{p}$ and $\nu_{p}$ is the topological index of the periodic orbit. The first correction is given by the integral

$$
C_{1, p}^{(1)}=\int_{0}^{T_{p}} d t \frac{\Delta \varphi_{p}^{1(0)}(t)}{\varphi_{p}^{1(0)}(t)}
$$

When the theory is applied for billiard systems, the wave function should fulfill the Dirichlet boundary condition on hard walls, e.g. it should vanish on the wall. The wave function determined from (M.36) behaves discontinuously when the trajectory $x_{p}(t)$ hits the wall. For the simplicity we consider a two dimensional billiard system here. The wave function on the wall before the bounce $\left(t_{-0}\right)$ is given by

$$
\begin{equation*}
\psi_{i n}(x, y(x), t)=\varphi\left(x, y(x), t_{-0}\right) e^{i S\left(x, y(x), t_{-0}\right) / \hbar} \tag{M.54}
\end{equation*}
$$

where $y(x)=Y_{2} x^{2} / 2!+Y_{3} x^{3} / 3!+Y_{4} x^{4} / 4!+\ldots$ is the parametrization of the wall around the point of reflection (see Fig 1.). The wave function on the wall after the bounce $\left(t_{+0}\right)$ is

$$
\begin{equation*}
\psi_{\text {out }}(x, y(x), t)=\varphi\left(x, y(x), t_{+0}\right) e^{i S\left(x, y(x), t_{+0}\right) / \hbar} \tag{M.55}
\end{equation*}
$$

The sum of these wave functions should vanish on the hard wall. This implies that the incoming and the outgoing amplitudes and the phases are related as

$$
\begin{equation*}
S\left(x, y(x), t_{-0}\right)=S\left(x, y(x), t_{+0}\right), \tag{M.56}
\end{equation*}
$$

and

$$
\begin{equation*}
\varphi\left(x, y(x), t_{-0}\right)=-\varphi\left(x, y(x), t_{+0}\right) . \tag{M.57}
\end{equation*}
$$

The minus sign can be interpreted as the topological phase coming from the hard wall.

Now we can reexpress the spectral determinant with the local eigenvalues:

$$
\begin{equation*}
\Delta(E)=e^{W(E)} \prod_{p} \prod_{1}\left(1-R_{\mathbf{l}, p}(E) e^{\frac{i}{\hbar} S_{p}(E)}\right) . \tag{M.58}
\end{equation*}
$$

This expression is the quantum generalization of the semiclassical Selbergproduct formula [M.11]. A similar decomposition has been found for quantum Baker maps in ref. [M.12]. The functions

$$
\begin{equation*}
\zeta_{\mathbf{l}}^{-1}(E)=\prod_{p}\left(1-R_{\mathbf{l}, p}(E) e^{\frac{i}{\hbar} S_{p}(E)}\right) \tag{M.59}
\end{equation*}
$$

are the generalizations of the Ruelle type [30.24] zeta functions. The trace formula can be recovered from (M.40):

$$
\operatorname{Tr} G(E)=g_{0}(E)+\frac{1}{i \hbar} \sum_{p, \mathbf{1}}\left(T_{p}(E)-i \hbar \frac{d \log R_{\mathbf{l}, p}(E)}{d E}\right) \frac{R_{\mathbf{l}, p}(E) e^{\frac{i}{\hbar} S_{p}(E)}}{1-R_{\mathbf{l}, p}(E) e^{\frac{i}{\hbar} S_{p}(E)}} .(\mathrm{l}
$$

We can rewrite the denominator as a sum of a geometric series and we get

$$
\begin{equation*}
\operatorname{Tr} G(E)=g_{0}(E)+\frac{1}{i \hbar} \sum_{p, r, \mathbf{1}}\left(T_{p}(E)-i \hbar \frac{d \log R_{\mathbf{l}, p}(E)}{d E}\right)\left(R_{\mathbf{l}, p}(E)\right)^{r} e^{\frac{i}{\hbar} r S_{p}(E)} . \tag{M.61}
\end{equation*}
$$

The new index $r$ can be interpreted as the repetition number of the prime orbit $p$. This expression is the generalization of the semiclassical trace
formula for the exact quantum mechanics. We would like to stress here, that the perturbation calculus introduced above is just one way to compute the eigenvalues of the local Schrödinger problems. Non-perturbative methods can be used to calculate the local eigenvalues for stable, unstable and marginal orbits. Therefore, our trace formula is not limited to integrable or hyperbolic systems, it can describe the most general case of systems with mixed phase space.

The semiclassical trace formula can be recovered by dropping the subleading term $-i \hbar d \log R_{\mathrm{l}, p}(E) / d E$ and using the semiclassical eigenvalue $R_{1, p}^{(0)}(E)=e^{C_{p}^{1(0)}}=e^{-i \nu_{p} \pi} e^{-\sum_{i}\left(l_{i}+1 / 2\right) u_{p, i}}$. Summation for the indexes $l_{i}$ yields the celebrated semiclassical amplitude

$$
\begin{equation*}
\sum_{1}\left(R_{1, p}^{(0)}(E)\right)^{r}=\frac{e^{-i r \nu_{p} \pi}}{\left|\operatorname{det}\left(\mathbf{1}-\mathbf{J}_{p}^{r}\right)\right|^{1 / 2}} \tag{М.62}
\end{equation*}
$$

To have an impression about the improvement caused by the quantum corrections we have developed a numerical code [M.13] which calculates the first correction $C_{p, l}^{(1)}$ for general two dimensional billiard systems. The first correction depends only on some basic data of the periodic orbit such as the lengths of the free flights between bounces, the angles of incidence and the first three Taylor expansion coefficients $Y_{2}, Y_{3}, Y_{4}$ of the wall in the point of incidence. To check that our new local method gives the same result as the direct calculation of the Feynman integral, we computed the first $\hbar$ correction $C_{p, 0}^{(1)}$ for the periodic orbits of the 3-disk scattering system [M.14] where the quantum corrections have been We have found agreement up to the fifth decimal digit, while our method generates these numbers with any desired precision. Unfortunately, the $l \neq 0$ coefficients cannot be compared to ref. [M.15], since the $l$ dependence was not realized there due to the lack of general formulas (M.58) and (M.59). However, the $l$ dependence can be checked on the 2 disk scattering system [M.16]. On the standard example [M.14, M.15, M.16, M.18], when the distance of the centers $(R)$ is 6 times the disk radius (a), we got

$$
C_{l}^{(1)}=\frac{1}{\sqrt{2 E}}\left(-0.625 l^{3}-0.3125 l^{2}+1.4375 l+0.625\right)
$$

For $l=0$ and 1 this has been confirmed by A. Wirzba [M.17], who was able to compute $C_{0}^{(1)}$ from his exact quantum calculation. Our method makes it possible to utilize the symmetry reduction of Cvitanović and Eckhardt and to repeat the fundamental domain cycle expansion calculation of ref. [M.18] with the first quantum correction. We computed the correction to the leading 226 prime periodic orbits with 10 or less bounces in the fundamental domain. Table I. shows the numerical values of the exact quantum calculation [M.16], the semiclassical cycle expansion [M.10] and our corrected calculation. One can see, that the error of the corrected calculation vs. the error of the semiclassical calculation decreases with the wavenumber. Besides the improved results, a fast convergence up to six decimal digits can be observed, which is just three decimal digits in the full domain calculation [M.15].

Figure M.2: A typical bounce on a billiard wall. The wall can be characterized by the local expansion $y(x)=Y_{2} x^{2} / 2!+Y_{3} x^{3} / 3!+Y_{4} x^{4} / 4!+\ldots$.

Table M.1: Real part of the resonances ( $\operatorname{Re} k$ ) of the 3-disk scattering system at disk separation 6:1. Semiclassical and first corrected cycle expansion versus exact quantum calculation and the error of the semiclassical $\delta_{S C}$ divided by the error of the first correction $\delta_{\text {Corr }}$. The magnitude of the error in the imaginary part of the resonances remains unchanged.

| Quantum | Semiclassical | First correction | $\delta_{S C} / \delta_{\text {Corr }}$ |
| :---: | :---: | :---: | :---: |
| 0.697995 | 0.758313 | 0.585150 | 0.53 |
| 2.239601 | 2.274278 | 2.222930 | 2.08 |
| 3.762686 | 3.787876 | 3.756594 | 4.13 |
| 5.275666 | 5.296067 | 5.272627 | 6.71 |
| 6.776066 | 6.793636 | 6.774061 | 8.76 |
| $\ldots$ | $\ldots$ | $\ldots$ | $\ldots$ |
| 30.24130 | 30.24555 | 30.24125 | 92.3 |
| 31.72739 | 31.73148 | 31.72734 | 83.8 |
| 32.30110 | 32.30391 | 32.30095 | 20.0 |
| 33.21053 | 33.21446 | 33.21048 | 79.4 |
| 33.85222 | 33.85493 | 33.85211 | 25.2 |
| 34.69157 | 34.69534 | 34.69152 | 77.0 |

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## Appendix N

## Solutions

## Chapter 1

Solution 1.1: 3-disk symbolic dynamics. As explained in sect. 1.4, each orbit segment can be characterized by either of the two symbols 0 and 1, differentiating topologically bouncing back or going onto the third disk. So, there are $2^{n}$ topologically different orbits starting from each disk. Altogether, the 3 -disk pinball has $3 \cdot 2^{n}$ itineraries of length $n$.
Periodic orbits (prime cycles in fundamental domain)

- Of length 2: 12,13,32; or (0).
- Of length 3: 123,321; or (1).
- Of length 4: 1213,2321,3231; or (01).
- Of length 5: $12123,13132,23231,21213,32321,31312$; or (00111).

Some of the cycles are listed in table 11.2 and drawn in figure 22.3.
(Yuheng Lan)
Solution 1.2: Sensitivity to initial conditions. To estimate the pinball sensitivity we consider a narrow beam of point particles bouncing between two disks, figure N.1(a). Or if you find this easier to visualize, think of a narrow ray of light. We


Figure N.1: The 2-disk pinball (a) geometry, (b) defocusing of scattered rays.
assume that the ray of light is focused along the axis between the two points. This is where the least unstable periodic orbit lies, so its stability should give us an upper bound on the number of bounces we can expect to achieve. To estimate the stability we assume that the ray of light has a width $w(t)$ and a "dispersion angle" $\theta(t)$ (we assume both are small), figure N.1(b). Between bounces the dispersion angle stays constant while the width increases as

$$
w(t) \approx w\left(t^{\prime}\right)+\left(t-t^{\prime}\right) \theta
$$

At each bounce the width stays constant while the angle increases by

$$
\theta_{n+1}=\theta_{n}+2 \phi \approx \theta_{n}+w(t) / a
$$

where $\theta_{n}$ denotes the angle after bounce $n$. Denoting the width of the ray at the $n$th bounce by $w_{n}$ then we obtain the pair of coupled equations

$$
\begin{align*}
w_{n+1} & =w_{n}+(R-2 a) \theta_{n}  \tag{N.1}\\
\theta_{n} & =\theta_{n-1}+\frac{w_{n}}{a} \tag{N.2}
\end{align*}
$$

where we ignore corrections of order $w_{n}^{2}$ and $\theta_{n}^{2}$. Solving for $\theta_{n}$ we find

$$
\theta_{n}=\theta_{0}+\frac{1}{a} \sum_{j=1}^{n} w_{n}
$$

Assuming $\theta_{0}=0$ then

$$
w_{n+1}=w_{n}+\frac{R-2 a}{a} \sum_{j=1}^{n} w_{n}
$$

Plugging in the values in the question we find the width at each bounce in Ångstrøms grows as $1,5,29,169,985$, etc. To find the asymptotic behavior for a large number of bounces we try an solution of the form $w_{n}=a x^{n}$. Substituting this into the equation above and ignoring terms that do not grow exponentially we find solutions

$$
w_{n} \approx a w_{n}^{a s y m}=a(3 \pm 2 \sqrt{2})^{n}
$$

The solution with the positive sign will clearly dominate. The constant a we cannot determine by this local analysis although it is clearly proportional to $w_{0}$. However, the asymptotic solution is a good approximation even for quite a small number of bounces. To find an estimate of $a$ we see that $w_{n} / w_{n}^{a s y m}$ very rapidly converges to 0.146447 , thus

$$
w_{n} \approx 0.146447 w_{0}(3+2 \sqrt{2})^{n} \approx 0.1 \times w_{0} \times 5.83^{n}
$$

The outside edges of the ray of light will miss the disk when the width of the ray exceeds 2 cm ; this occurs after 11 bounces.

## (Adam Prügel-Bennett)

Solution 1.2: Sensitivity to initial conditions, another try. Adam's estimate is not very good - do you have a better one? The first problem with it is that the instability is very underestimated. As we shall check in exercise 8.1, the exact formula for the 2-cycle stability is $\Lambda=R-1+R \sqrt{1-2 / R}$. For $R=6, a=1$ this yields $w_{n} / w_{0} \approx(5+2 \sqrt{6})^{n}=9.898979^{n}$, so if that were the whole story, the pinball would be not likely to make it much beyond 8 bounces.

The second problem is that local instability overestimates the escape rate from an enclosure; trajectories are reinjected by scatterers. In the 3-disk pinball the particle leaving a disk can be reinjected by hitting either of other 2 disks, hence $w_{n} / w_{0} \approx$ $(9.9 / 2)^{n}$. This interplay between local instability and global reinjection will be cast into the exact formula involving "Lyapunov exponent" and "Kolmogorov entropy". In order to relate this estimate to our best continuous time escape rate estimate $\gamma=0.4103 \ldots$ (see table 18.2), we will have to also compute the mean free flight time (18.20). As a crude estimate, we take the shortest disk-to-disk distance, $\langle T\rangle=R-2=4$. The continuous time escape rate result implies that $w_{n} / w_{0} \approx e^{(R-2) \gamma n}=(5.16)^{n}$, in the same ballpark as the above expansion-reinjection estimate.
(Predrag Cvitanović)

## Chapter 2

Solution 2.1: Trajectories do not intersect. Suppose that two trajectories $C_{x}$ and $C_{y}$ intersect at some point $z$. We claim that any points $\tilde{x}$ on $C_{x}$ is also a point on $C_{y}$ and vice versa. We only need to prove the first part of the statement.
According to the definition of $C_{x}$, there exist $t_{x}, t_{y}, t_{1} \in \mathbb{R}$ such that $f^{t_{x}}(x)=$ $z, f^{t_{y}}(y)=z, f^{t_{1}}(x)=\tilde{x}$. It is easy to check that $f^{t_{y}-t_{x}+t_{1}}(y)=\tilde{x}$. So, $\tilde{x} \in C_{y}$. Therefore, if two trajectories intersect, then they are the same trajectory.
(Yueheng Lan)
Solution 2.2: Evolution as a group. Let's check the basic defining properties of a group. The members of the set are $f^{t}, t \in \mathbb{R}$ and the "product law" is given by ' o '.

- As $f^{t+s}=f^{t} \circ f^{s}$, the set is closed, i.e., the product of any two members generates another member of the set.
- It is associative, as $\left(f^{t} \circ f^{s}\right) \circ f^{r}=f^{t+s+r}=f^{t} \circ\left(f^{s} \circ f^{r}\right)$.
- $I=f^{0}$ is the identity, as $f^{t} \circ f^{0}=f^{t}$.
- $f^{-t}$ is the inverse of $f^{t}$, as $f^{-t} \circ f^{t}=I$.

So, $\left\{f^{t}, \circ\right\}_{t \in \mathbb{R}}$ forms a group. As $f^{t} \circ f^{s}=f^{t+s}=f^{s} \circ f^{t}$, it is a commutative (Abelian) group.

Any Abelian group can replace the continuous time. For example, $\mathbb{R}$ can be replaced by $\mathbb{Z}_{6}$. To mess things up try a non-commutative group.

Solution 2.3: Almost ODE's. What is an $O D E$ on $\mathbb{R}$ ? An $O D E$ is an equality which reveals explicitly the relation between function $x(t)$ and its time derivatives $\dot{x}, \ddot{x}, \cdots$, i.e., $F(t, x, \dot{x}, \ddot{x}, \cdots)=0$ for some given function $F$. Let's check the equations given in the exercise.
(a) $\dot{x}=\exp (\dot{x})$ is an $O D E$.
(b) $\dot{x}=x(x(t))$ is not an ODE, as $x(x(t))$ is not a known function acting on $x(t)$.
(c) $\dot{x}=x(t+1)$ is not an $O D E$, as $x(t+1)$ is not a value at current time. Actually, it is a difference-differential equation.
(Yueheng Lan)
Solution 2.4: All equilibrium points are fixed points. Given a vector field $v(x)$, the phase space dynamics is defined by

$$
\begin{equation*}
\frac{d}{d t} x(t)=v(x(t)) \tag{N.3}
\end{equation*}
$$

An equilibrium point $a$ of $v$ is defined by $v(a)=0$, so $x(t)=a$ is a constant solution of (N.3). For the flow $f^{t}$ defined by (N.3), this solution satisfies $f^{t}(a)=a, t \in \mathbb{R}$. So, it is a fixed point of the dynamics $f^{t}$.
(Yueheng Lan)

## Solution 2.5: Gradient systems.

1. The directional derivative

$$
\frac{d}{d n} \phi=n \cdot \nabla \phi
$$

produces the increasing rate along the unit vector $n$. So, along the gradient direction $\nabla \phi /|\nabla \phi|, \phi$ has the largest increasing rate. The velocity of the particle has the opposite direction to the gradient, so $\phi$ deceases most rapidly in the velocity direction.
2. An extremum $a$ of $\phi$ satisfies $\nabla \phi(a)=0$. According to exercise 2.4, $a$ is a fixed point of the flow.
3. Two arguments lead to the same conclusion here.

First, near an equilibrium point, the equation is always linearizable. For gradient system, after orthogonal transformation it is even possible to write the linearized equation in diagonal form so that we need only to consider one eigendirection. The corresponding scalar equation is $\dot{x}=\lambda x$. Notice that we moved the origin to the equilibrium point. The solution of this equation is $x(t)=x(0) \exp (\lambda t)$, for $\lambda \neq 0$. if $x(0) \neq 0$, it will take infinite amount of time (positive or negative) for $x(t) \rightarrow 0$. For $\lambda=0$, the approach to zero is even slower as then only higher orders of $x$ take effect.
The second argument seems easier. We know that the solution curve through an equilibrium point is the point itself. According to exercise 2.1, no other solution curve will intersect it, which means that if not starting from the equilibrium point itself, other point can never reach it.
4. On a periodic orbit, the velocity is bounded away from zero. So $\phi$ is always decreasing on a periodic orbit, but in view of the periodicity, we know that this can not happen (at each point, there is only one value of $\phi$.). So, there is no periodic orbit in a gradient system.

Solution 2.10: Classical collinear helium dynamics. An example of a solution are A. Prügel-Bennett's programs, available at ChaosBook.org/extras.

## Solution 2.8: Equilibria of the Rössler system.

1. Solve $\dot{x}=\dot{y}=\dot{z}=0$, to get $x=a z, y=-z$ and $x^{2}-c x+a b=0$. There are two solutions of a quadratic equation, hence there are two equilibrium points:

$$
\begin{equation*}
x^{ \pm}=a z^{ \pm}=-a y^{ \pm}=\left(c \pm \sqrt{c^{2}-4 a b}\right) / 2 \tag{N.4}
\end{equation*}
$$

2. That above expressions are exact. However, it pays to think of $\epsilon=a / c$ as a small parameter in the problem. By subsitution from (2.23),

$$
\begin{equation*}
x^{ \pm}=c p^{ \pm}, \quad y^{ \pm}=-p^{ \pm} / \epsilon, \quad z^{ \pm}=p^{ \pm} / \epsilon \tag{N.5}
\end{equation*}
$$

Expanding $\sqrt{D}$ in $\epsilon$ yields $p^{-}=\epsilon^{2}+o\left(\epsilon^{3}\right)$, and $p^{+}=1-\epsilon^{2}+o\left(\epsilon^{3}\right)$. Hence

$$
\begin{array}{ll}
x^{-}=a^{2} / c+o\left(\epsilon^{3}\right), & x^{+}=c-a^{2} / c+o\left(\epsilon^{3}\right) \\
y^{-}=-a / c+o\left(\epsilon^{2}\right), & z^{+}=c / a+a / c+o\left(\epsilon^{2}\right),  \tag{N.6}\\
z^{-}=a / c+o\left(\epsilon^{2}\right), & z^{+}=c / a-a / c+o\left(\epsilon^{2}\right) .
\end{array}
$$

For $a=b=0.2, c=5.7$ in (2.14), $\epsilon \approx 0.035$, so

$$
\begin{align*}
& \left(x^{-}, y^{-}, z^{-}\right)=(0.0070,-0.0351,0.0351), \\
& \left(x^{+}, y^{+}, z^{+}\right)=(5.6929,-28.464,28.464) . \tag{N.7}
\end{align*}
$$

(Rytis Paškauskas)

## Chapter 3

(No solutions available.)

## Chapter 4

Solution 4.1: Trace-log of a matrix. 1) one method is to first check that this is true for any Hermitian matrix $M$. Then write an arbitrary complex matrix as sum $M=A+z B, A, B$ Hermitian, Taylor expand in $z$ and prove by analytic continuation that the identity applies to arbitrary M. (David Mermin)
2) another method: evaluate $\frac{d}{d t} \operatorname{det}\left(e^{t \ln M}\right)$ by definition of derivative in terms of infinitesimals. (Kasper Juel Eriksen)
3) check appendix K. 1
4) This identity makes sense for a matrix $M \subset \mathbb{C}^{n \times n}$, if $\left|\prod_{i=1}^{n} \lambda_{i}\right|<\infty$ and $\left\{\left|\lambda_{i}\right|>0, \forall i\right\}$, where $\left\{\lambda_{i}\right\}$ is a set of eigenvalues of $M$. Under these conditions there exist a nonsingular $O: M=O D O^{-1}, D=\operatorname{diag}\left[\left\{\lambda_{i}, i=1, \ldots, n\right\}\right]$. If $f(M)$ is a
matrix valued function defined in terms of power series then $f(M)=O f(D) O^{-1}$, and $f(D)=\operatorname{diag}\left[\left\{f\left(\lambda_{i}\right)\right\}\right]$. Using these properties and cyclic property of the trace we obtain

$$
\exp (\operatorname{tr}(\ln M))=\exp \left(\sum_{i} \ln \lambda_{i}\right)=\prod_{i} \lambda_{i}=\operatorname{det}(M)
$$

5) Consider $M=\exp A$.

$$
\operatorname{det} M=\operatorname{det} \lim _{n \rightarrow \infty}\left(\mathbf{1}+\frac{1}{n} A\right)^{n}=\lim _{n \rightarrow \infty}\left(\mathbf{1}+\frac{1}{n} \operatorname{tr} A+\ldots\right)^{n}=\exp (\operatorname{tr}(\ln M))
$$

Solution 4.2: Stability, diagonal case. The relation (4.16) can be verified by noting that the defining product (4.11) can be rewritten as

$$
\begin{align*}
e^{t \mathbf{A}} & =\left(\mathbf{U} \mathbf{U}^{-1}+\frac{t \mathbf{U} \mathbf{A}_{D} \mathbf{U}^{-1}}{m}\right)\left(\mathbf{U U}^{-1}+\frac{t \mathbf{U} \mathbf{A}_{D} \mathbf{U}^{-1}}{m}\right) \cdots \\
& =\mathbf{U}\left(I+\frac{t \mathbf{A}_{D}}{m}\right) \mathbf{U}^{-1} \mathbf{U}\left(I+\frac{t \mathbf{A}_{D}}{m}\right) \mathbf{U}^{-1} \cdots=\mathbf{U} e^{t \mathbf{A}_{D}} \mathbf{U}^{-1} \tag{N.8}
\end{align*}
$$

## Solution 4.3: Topology of the Rössler flow.

1. The characteristic determinant of the matrix of variations that yields the equilibrium point stability (4.24) yields

$$
\begin{aligned}
& \left|\begin{array}{ccc}
-\lambda & -1 & -1 \\
1 & a-\lambda & 0 \\
z^{ \pm} & 0 & x^{ \pm}-c-\lambda
\end{array}\right|=0 \\
& \lambda^{3}+\lambda^{2}\left(-a-x^{ \pm}+c\right)+\lambda\left(a\left(x^{ \pm}-c\right)+1+x^{ \pm} / a\right)+c-2 x^{ \pm}=0 .
\end{aligned}
$$

Equation (4.42) follows after noting that $x^{ \pm}-c=c\left(p^{ \pm}-1\right)=-c p^{\mp}$ and $2 x^{ \pm}-c=c\left(2 p^{ \pm}-1\right)= \pm c \sqrt{D}$, see (2.8).
2. Approximate solutions of (4.42) are obtained by expanding $p^{ \pm}$and $\sqrt{D}$ and substituting into this equation. Namely,

$$
\begin{aligned}
\sqrt{D} & =1-2 \epsilon^{2}-2 \epsilon^{4}-4 \epsilon^{6}-\ldots \\
p^{-} & =\epsilon^{2}+\epsilon^{4}+2 \epsilon^{6}+\ldots \\
p^{+} & =1-\epsilon^{2}-\epsilon^{4}-2 \epsilon^{6}+\ldots
\end{aligned}
$$

In case of the equilibrium " - ", close to the origin expansion of (4.42) results in

$$
\left(\lambda^{2}+1\right)(\lambda+c)=-\epsilon \lambda\left(1-c^{2}-c \lambda\right)+\epsilon^{2} c\left(\lambda^{2}+2\right)+o\left(\epsilon^{2}\right)
$$

The term on the left-hand side suggests the expansion for eigenvalues as

$$
\lambda_{1}=-c+\epsilon a_{1}+\ldots, \quad \lambda_{2}+i \theta_{2}=\epsilon b_{1}+i+\ldots
$$

after some algebra one finds the first order correction coefficients $a_{1}=c /\left(c^{2}+1\right)$ and $b_{1}=\left(c^{3}+i\right) /\left(2\left(c^{2}+1\right)\right)$. Numerical values are $\lambda_{1} \approx-5.694, \lambda_{2}+i \theta_{2} \approx$ $0.0970+i 1.0005$.

In case of $p^{+}$, the leading order term in (4.42) is $1 / \epsilon$. Set $x=\lambda / \epsilon$, then expansion of (4.42) results in

$$
x=c-\epsilon x-\epsilon^{2}(2 c-x)-\epsilon^{3}\left(x^{3}-c x^{2}\right)-\epsilon^{4}\left(2 c-x\left(1+c^{2}\right)+c x^{2}\right)+o\left(\epsilon^{4}\right)
$$

Solve for real eigenvalue first. Set $x=c+\epsilon a_{1}+\epsilon^{2} a_{2}+\epsilon^{3} a_{3}+\epsilon^{4} a_{4}+\ldots$. The subtle point here is that leading order correction term of the real eigenvalue is $\epsilon a_{1}$, but to determine leading order of the real part of complex eigenvalue, one needs all terms $a_{1}$ through $a_{4}$.
Collecting powers of $\epsilon$ results in

$$
\begin{array}{lll}
\epsilon: & a_{1}+c=0 & a_{1}=-c \\
\epsilon^{2}: & c+a_{1}+a_{2}=0 & a_{2}=0 \\
\epsilon^{3}: & a_{1}-a_{2}-a_{3}=0 & a_{3}=-c \\
\epsilon^{4}: & c+c^{2} a_{1}-a_{2}+a_{3}+a_{4}=0 & a_{4}=c^{3} .
\end{array}
$$

hence

$$
\lambda_{1}=\epsilon x=a-a^{2} / c+o\left(\epsilon^{3}\right) \approx 0.192982
$$

To calculate the complex eigenvalue, one can make use of identities $\operatorname{det} A=$ $\Pi \lambda=2 x^{+}-c$, and $\operatorname{tr} A=\sum \lambda=a+x^{+}-c$. Namely,

$$
\begin{aligned}
& \lambda_{2}=\frac{1}{2}\left(a-c p^{-}-\lambda_{1}\right)=-\frac{a^{5}}{2 c^{2}}+o\left(\epsilon^{5}\right) \approx-0.49 \times 10^{-6} \\
& \theta_{2}=\sqrt{\frac{2 x^{+}-c}{\lambda_{1}}-\lambda_{2}^{2}}=\sqrt{\frac{a+c}{a}}(1+o(\epsilon)) \approx 5.431
\end{aligned}
$$

(Rytis Paškauskas)

## Chapter 5

(No solutions available.)

## Chapter 6

Solution 6.1: A pinball simulator. Examples of pretty pinballs are A. PrügelBennett's xpinball.c and W. Benfold's java programs, available at ChaosBook.org/extras

Solution 6.4: Billiard exercises. Korsch and Jodl [1.9] have a whole book of numerical exercises with billiards, including 3-disks.

## Chapter 7

Solution 7.2: Linearization for maps. (difficulty: medium) The first few terms of the map $h$ that conjugates $f$ to $\alpha z$

$$
f(z)=h^{-1}(\alpha h(z)) .
$$

are determined many places, for example in ref. [9.5].

There are conditions on the derivative of $f$ at the origin to assure that the conjugation is always possible. These conditions are formulated in ref. [1.14], among others.

## Chapter 8

(No solutions available.)

## Chapter 9

Solution 9.1: Integrating over Dirac delta functions. (a) Whenever $h(x)$ crosses 0 with a nonzero velocity ( $\operatorname{det} \partial_{x} h(x) \neq 0$ ), the delta function contributes to the integral. Let $x_{0} \in h^{-1}(0)$. Consider a small neighborhood $V_{0}$ of $x_{0}$ so that $h: V_{0} \rightarrow V_{0}$ is a one-to-one map, with the inverse function $x=x(h)$. By changing variable from $x$ to $h$, we have

$$
\begin{aligned}
\int_{V_{0}} d x \delta(h(x)) & =\int_{h\left(V_{0}\right)} d h\left|\operatorname{det} \partial_{h} x\right| \delta(h)=\int_{h\left(V_{0}\right)} d h \frac{1}{\left|\operatorname{det} \partial_{x} h\right|} \delta(h) \\
& =\frac{1}{\left|\operatorname{det} \partial_{x} h\right|_{h=0}} .
\end{aligned}
$$

Here, the absolute value $|\cdot|$ is taken because delta function is always positive and we keep the orientation of the volume when the change of variables is made. Therefore all the contributions from each point in $h^{-1}(0)$ add up to the integral

$$
\int_{\mathbb{R}^{d}} d x \delta(h(x))=\Sigma_{x \in h^{-1}(0)} \frac{1}{\left|\operatorname{det} \partial_{x} h\right|} .
$$

Notice that if $\operatorname{det} \partial_{x} h=0$, then the delta function integral is not well defined.
(b) The formal expression can be written as the limit

$$
F:=\int_{\mathbb{R}} d x \delta\left(x^{2}\right)=\lim _{\sigma \rightarrow 0} \int_{\mathbb{R}} d x \frac{e^{-\frac{x^{4}}{2 \sigma}}}{\sqrt{2 \pi \sigma}}
$$

by invoking the approximation given in the exercise. The change of variable $y=x^{2} / \sqrt{\sigma}$ gives

$$
F=\lim _{\sigma \rightarrow 0} \sigma^{-3 / 4} \int_{\mathbb{R}^{+}} d y \frac{e^{-\frac{y^{2}}{2}}}{\sqrt{2 \pi y}}=\infty
$$

where $\mathbb{R}^{+}$represents the positive part of the real axis. So, the formal expression does not make sense. Notice that $x^{2}$ has a zero derivative at $x=0$, which invalidates the expression in (a).
(Yueheng Lan)
Solution 9.2: Derivatives of Dirac delta functions. We do this problem just by direct evaluation. We denote by $\Omega_{y}$ a sufficiently small neighborhood of $y$.
(a)

$$
\begin{aligned}
\int_{\mathbb{R}} d x \delta^{\prime}(y) & =\Sigma_{x \in y^{-1}(0)} \int_{\Omega_{y}} d y \operatorname{det}\left(\frac{d y}{d x}\right)^{-1} \delta^{\prime}(y) \\
& =\left.\Sigma_{x \in y^{-1}(0)} \frac{\delta(y)}{\left|y^{\prime}\right|}\right|_{-\epsilon} ^{\epsilon}-\int_{\Omega_{y}} d y \frac{\delta(y)}{y^{\prime 2}}\left(-y^{\prime \prime}\right) \frac{1}{y^{\prime}} \\
& =\Sigma_{x \in y^{-1}(0)} \frac{y^{\prime \prime}}{\left|y^{\prime}\right| y^{\prime 2}}
\end{aligned}
$$

where the absolute value is taken to take care of the sign of the volume.
(b)

$$
\begin{aligned}
\int_{\mathbb{R}} d x \delta^{(2)}(y) & =\Sigma_{x \in y^{-1}(0)} \int_{\Omega_{y}} d y \frac{\delta^{(2)}(y)}{y^{\prime}} \\
& =\left.\Sigma_{x \in y^{-1}(0)} \frac{\delta^{\prime}(y)}{\left|y^{\prime}\right|}\right|_{-\epsilon} ^{\epsilon}-\int_{\Omega_{y}} d y \frac{\delta^{\prime}(y)}{y^{2^{2}}}\left(-y^{\prime \prime}\right) \frac{1}{y^{\prime}} \\
& =\left.\Sigma_{x \in y^{-1}(0)} \frac{y^{\prime \prime} \delta(y)}{\left|y^{\prime}\right| y^{\prime 2}}\right|_{-\epsilon} ^{\epsilon}-\int_{\Omega_{y}} d y \delta(y) \frac{d}{d x}\left(\frac{y^{\prime \prime}}{y^{\prime 3}}\right) \frac{1}{y^{\prime}} \\
& =\Sigma_{x \in y^{-1}(0)}-\int_{\Omega_{y}} d y \delta(y)\left(\frac{y^{\prime \prime \prime}}{y^{\prime 3}}-3 \frac{y^{\prime \prime 2}}{y^{\prime 4}}\right) \frac{1}{y^{\prime}} \\
& =\Sigma_{x \in y^{-1}(0)}\left(3 \frac{y^{\prime \prime 2}}{y^{\prime 4}}-\frac{y^{\prime \prime \prime}}{y^{\prime 3}}\right) \frac{1}{\left|y^{\prime}\right|}
\end{aligned}
$$

(c)

$$
\begin{aligned}
\int_{\mathbb{R}} d x b(x) \delta^{(2)}(y) & =\Sigma_{x \in y^{-1}(0)} \int_{\Omega_{y}} d y b(x) \frac{\delta^{(2)}(y)}{y^{\prime}} \\
& =\left.\Sigma_{x \in y^{-1}(0)} \frac{b(x) \delta^{\prime}(y)}{\left|y^{\prime}\right|}\right|_{-\epsilon} ^{\epsilon}-\int_{\Omega_{y}} d y \delta^{\prime}(y) \frac{d}{d x}\left(\frac{b}{y^{\prime}}\right) \frac{1}{y^{\prime}} \\
& =\Sigma_{x \in y^{-1}(0)}-\left.\delta(y) \frac{d}{d x}\left(\frac{b}{y^{\prime}}\right) \frac{1}{y^{\prime}}\right|_{-\epsilon} ^{\epsilon}+\int_{\Omega_{y}} d y \delta(y) \frac{d}{d x}\left(\frac{d}{d x}\left(\frac{b}{y^{\prime}}\right) \frac{1}{y^{\prime}}\right) \frac{1}{y^{\prime}} \\
& \left.=\Sigma_{x \in y^{-1}(0)} \frac{1}{\left|y^{\prime}\right|} \frac{d}{d x}\left(\frac{b^{\prime}}{y^{\prime 2}}-\frac{b y^{\prime \prime}}{y^{\prime 3}}\right)\right) \\
& =\Sigma_{x \in y^{-1}(0)} \frac{1}{\left|y^{\prime}\right|}\left[\frac{b^{\prime \prime}}{y^{\prime 2}}-\frac{b^{\prime} y^{\prime \prime}}{y^{\prime 3}}-2 \frac{b^{\prime} y^{\prime \prime}}{y^{\prime 3}}+b\left(3 \frac{y^{\prime \prime 2}}{y^{\prime 4}}-\frac{y^{\prime \prime \prime}}{y^{\prime 3}}\right)\right] \\
& =\Sigma_{x \in y^{-1}(0)} \frac{1}{\left|y^{\prime}\right|}\left[\frac{b^{\prime \prime}}{y^{\prime 2}}-3 \frac{b^{\prime} y^{\prime \prime}}{y^{\prime 3}}+b\left(3 \frac{y^{\prime \prime 2}}{y^{\prime 4}}-\frac{y^{\prime \prime \prime}}{y^{\prime 3}}\right)\right]
\end{aligned}
$$

Solution 9.3: $\quad \mathcal{L}^{t}$ generates a semigroup. Every "sufficiently good" transformation $f^{t}$ in phase space $\mathcal{M}$ is associated with a Perron-Frobenius operator $\mathcal{L}^{t}$ which is when acting on a function $\rho(x)$ in $\mathcal{M}$

$$
\mathcal{L}^{t} \cdot \rho(x)=\int_{\mathcal{M}} d y \delta\left(x-f^{t}(y)\right) \rho(y)
$$

In some proper function space $\mathcal{F}$ on $\mathcal{M}$, the one parameter family of operators $\left\{\mathcal{L}^{t}\right\}_{t \in \mathbb{R}^{+}}$generate a semigroup. Let's check this statement. For any $t_{1}, t_{2}>0$ and $\rho \in \mathcal{F}$, the product " o " of two operators is defined as usual

$$
\left(\mathcal{L}^{t_{1}} \circ \mathcal{L}^{t_{2}}\right) \cdot \rho(y)=\mathcal{L}^{t_{1}} \cdot\left(\mathcal{L}^{t_{2}} \cdot \rho\right)(y) .
$$

So, we have

$$
\begin{aligned}
\left(\mathcal{L}^{t_{1}} \circ \mathcal{L}^{t_{2}}\right)(y, x) & =\int_{\mathcal{M}} d z \mathcal{L}^{t_{1}}(y, z) \mathcal{L}^{t_{2}}(z, x) \\
& =\int_{\mathcal{M}} d z \delta\left(y-f^{t_{1}}(z)\right) \delta\left(z-f^{t_{2}}(x)\right) \\
& =\delta\left(y-f^{t_{1}}\left(f^{t_{2}}(x)\right)\right) \\
& =\delta\left(y-f^{t_{1}+t_{2}}(x)\right) \\
& =\mathcal{L}^{t_{1}+t_{2}}(y, x),
\end{aligned}
$$

where the semigroup property $f^{t_{1}}\left(f^{t_{2}}(x)\right)=f^{t_{1}+t_{2}}(x)$ of $f^{t}$ has been used. This proves the claim in the title.
(Yueheng Lan)
Solution 9.5: Invariant measure. Hint: Compare the second map to the construction of Exercise 13.6. We do (a),(b),(c),(d) for the first map and (e) for the second.
(a) The partition point is in the middle of $[0,1]$. If the density on the two pieces are two constants $\rho_{0}^{A}$ and $\rho_{0}^{B}$, respectively, the Perron-Frobenius operator still leads to the piecewise constant density

$$
\begin{aligned}
\rho_{1}^{A} & =\frac{1}{2}\left(\rho_{0}^{A}+\rho_{0}^{B}\right), \\
\rho_{1}^{B} & =\frac{1}{2}\left(\rho_{0}^{A}+\rho_{0}^{B}\right) .
\end{aligned}
$$

Notice that in general if a finite Markov partition exists and the map is affine (linear on each partition cell), a finite-dimensional invariant subspace consisting of piecewise constant function can always be identified in the function space.
(b) From the discussion of (a), any constant function on $[0,1]$ is an invariant measure. If we consider the invariant probability measure, then the constant has to be 1 .
(c) As the map is invariant in $[0,1]$ (there is no escaping), the leading eigenvalue of $\mathcal{L}$ is always 1 due to the "mass" conservation.
(d) Take a typical point on $[0,1]$ and record its trajectory under the first map for some time ( $10^{5}$ steps). Plot the histogram...ONLY 0 is left finally!! This happens because of the finite accuracy of the computer arithmetics. A small trick is to change the slope 2 to 1.99999999 . You will find a constant measure on $[0,1]$ which is the natural measure. Still, the finite presicion of the computer will make every point eventually periodic and strictly speaking the measure is defined only on some fine lattice points. But when the resolution improves, the computer-generated measure will steadily approach the natural measue. For the first map, any small deviation from the constant profile will be stretched and smeared out. So, the natural measure has to be constant.
(e) Simple calculation shows that $\alpha$ is the partition point. We may use $A, B$ to mark the left and right part of the partition, respectively. $A$ maps to $B$ and $B$ maps to the whole interval $[0,1]$. As the magnitude of the slope $\Lambda=(\sqrt{5}+1) / 2$ is greater than 1 , we may expect the natural measure is still piecewise constant with eigenvalue 1. The determining equation is

$$
\left(\begin{array}{cc}
0 & 1 / \Lambda \\
1 / \Lambda & 1 / \Lambda
\end{array}\right)\binom{\rho^{A}}{\rho^{B}}=\binom{\rho^{A}}{\rho^{B}},
$$

which gives $\rho^{B} / \rho^{A}=\Lambda$.
(Yueheng Lan)
Solution 9.7: Eigenvalues of the skew Ulam tent map Perron-Frobenius operator. If we have density $\rho_{n}(x)$, the action of the Perron-Frobenius operator associated with $f(x)$ gives a new density

$$
\rho_{n+1}(x)=\frac{1}{\Lambda_{0}} \rho_{n}\left(x / \Lambda_{0}\right)+\frac{1}{\Lambda_{1}} \rho_{n}\left(1-x / \Lambda_{1}\right),
$$

where $\Lambda_{1}=\frac{\Lambda_{0}}{\Lambda_{0}-1}$. The eigenvalue equation is given by

$$
\begin{equation*}
\rho_{n+1}(x)=\lambda \rho_{n}(x) . \tag{N.9}
\end{equation*}
$$

We may solve it by assuming that the eigenfunctions are $N$-th order polynomials $P(N)$ (check it). Indeed, detailed calculation gives the following results:

- $P(0)$ gives $\lambda=1$, corresponding to the expected leading eigenvalue.
- $P(1)$ gives $\lambda=\frac{1}{\Lambda_{0}^{2}}-\frac{1}{\Lambda_{1}^{2}}=\frac{2}{\Lambda_{0}}-1$,
- $P(2)$ gives $\lambda=\frac{1}{\Lambda_{0}^{3}}+\frac{1}{\Lambda_{1}^{3}}$,
- $P(3)$ gives $\lambda=\frac{1}{\Lambda_{0}^{4}}-\frac{1}{\Lambda_{1}^{4}}$,
- The guess is that $P(N)$ gives $\lambda=\frac{1}{\Lambda_{0}^{N+1}}+(-1)^{N} \frac{1}{\Lambda_{1}^{N+1}}$.

The final solution is that the piecewise linear function $\rho^{A}=-\Lambda_{0}, \rho^{B}=\Lambda_{1}$ gives the eigenvalue 0 . If only the continuous functions are considered, this kind of eigenfunction of course should not be included.

Solution 9.7: Eigenvalues of the skew Ulam tent map Perron-Frobenius operator. The first few eigenvalues are

$$
\begin{aligned}
e^{s_{0}} & =1, \quad e^{s_{1}}=\frac{2}{\Lambda_{0}}-1 \\
e^{s_{2}} & =\frac{1}{4}+\frac{3}{4}\left(\frac{2}{\Lambda_{0}}-1\right)^{2}, \quad e^{s_{3}}=\frac{1}{2}\left(\frac{2}{\Lambda_{0}}-1\right)+\frac{1}{2}\left(\frac{2}{\Lambda_{0}}-1\right)^{3} \ldots
\end{aligned}
$$

For eigenvectors (invariant densities for skew tent maps), see for example Invariant densities for skew tent by L. Billings and E.M. Bolt.

Solution 9.10: $\mathcal{A}$ as a generator of translations. If $v$ is a constant in space, Taylor series expansion gives

$$
a(x+t v)=\sum_{k=0}^{\infty} \frac{1}{k!}\left(t v \frac{\partial}{\partial x}\right)^{k} a(x)=e^{t v \frac{\partial}{\partial x}} a(x) .
$$

(Yueheng Lan)

## Chapter 10

## Solution 10.1: How unstable is the Hénon attractor?

1. Evaluate numerically the Lyapunov exponent by iterating the Hénon map; For $a=1.4, b=0.3$ the answer should be close to $\lambda=0.41$. If you have a good estimate and a plot of the convergence of your estimate with $t$, please send us your results for possible inclusion into this text.
2. Both Lyapunov exponents for $a=1.39945219, b=0.3$ are negative, roughly $\lambda_{1}=-0.2712, \lambda_{2}=-0.9328$ (check that these values respect the constant volume contraction condition (4.37) for the Hénon map). Why? Because after a long transient exploration of the Hénon map's non-wandering set, on average after some 11,000 iterates, almost every inital point falls into a stable 13cycle. You can check its existence by starting at one of its periodic points $\left(x_{p}, y_{p}\right)=(-0.2061,-0.3181)$.
If you missed the stable 13-cycle (as all students in one of the courses did), you should treat your computer experiments with great deal of scepticism.

As the product of eigenvalues is the constant $-b$, you need to evaluate only the expanding eigenvalue. There are many ways to implement this calculation - here are a few:

1. The most naive way - take the log of distance of two nearby trajectories, iterate until you run out of accuracy. Tray this many times, estimate an average.
2. Slighly smarter: as above, but keep rescaling the length of the vector connecting neighboring points so it remains small, average over the sum of logs of rescaling factors. You can run this forever.
3. Keep multiplying the [2×2] Jacobian stability matrix (4.36) until you run out of accuracy. Compute the log of the leading eigenvalue (4.18), try this many times, estimate an average.
4. Slighly smarter still: as above, but start with an arbitrary initial tangent space vector, keep multiplying it with the Jacobian stability matrix, and rescaling the length of the vector so it remains small. You can run this forever.
5. There is probably no need to use the $Q R$ decomposition method or any other such numerical method for this 2-dimensional problem.
(Yueheng Lan and P. Cvitanović)

## Chapter 11

Solution 11.1: Binary symbolic dynamics. Read the text.
Solution 11.2: 3-disk fundamental domain symbolic dynamics. Read sect. 1.4.

Solution 11.5: Reduction of 3-disk symbolic dynamics. The answer is given in sect. 22.6. Some remarks concerning part (c):

If an orbit does not have any spatial symmetry, its length in the fundamental domain is equal to that in the full space. One fundamental domain orbit corresponds to six copies of the orbit in the full space related to each other by symmetries. If a periodic orbit does have a spatial symmetry, then its fundamental domain image is a fraction of that in the whole space, and the orbit (and its symmetry pratnenrs) in the full space is tiled by copies of the irreducible segment, corresponding to an orbit in the fundamental domain. The higher symmetry an orbit has, the shorter the irreducible segment.

Another way to construct a fundamental domain orbit: put a periodic orbit and all its spatial symmetry relatives simultaneously in the full space. The segments that fall into a fundamental domain constitute the orbit in the fundamental domain.
(Yueheng Lan)
Solution 11.6: Unimodal map symbolic dynamics. Hint: write down an arbitrary binary number such as $\gamma=.1101001101000 \ldots$ and generate the future itinerary $S^{+}$by checking whether $f^{n}(\gamma)$ is greater or less than $1 / 2$. Then verify that (??) recovers $\gamma$.

## Solution 11.7: "Golden mean" pruned map.

(a) Consider the 3-cycle drawn in the figure. Denote the lengths of the two horizontal intervals by $a$ and $b$. We have

$$
\frac{a}{b}=\frac{b}{a+b},
$$

so the slope is given by the golden mean, $\Lambda=\frac{b}{a}=\frac{1+\sqrt{5}}{2}$, and the piece-wise linear map is given by

$$
f(x)=\left\{\begin{array}{l}
\Lambda x, x \in[0,1 / 2] \\
\Lambda(1-x), x \in[1 / 2,1]
\end{array}\right.
$$

(b) Evaluate

$$
f\left(\frac{1}{2}\right)=\frac{1+\sqrt{5}}{4}, \quad f\left(\frac{1+\sqrt{5}}{4}\right)=\frac{-1+\sqrt{5}}{4}, \quad f\left(\frac{-1+\sqrt{5}}{4}\right)=\frac{1}{2} .
$$

Once a point enters the region covered by the interval $\mathcal{M}$ of length $a+b$, bracketed by the 3-cycle, it will be trapped there forever. Outside $\mathcal{M}$, all points on unit interval will be mapped to $(0,1 / 2]$, except for 0 . The points in the interval $\left(0, \frac{-1+\sqrt{5}}{4}\right)$ approach $\mathcal{M}$ monotonically.
(c) It will be in $\left(\frac{1}{2}, \frac{1+\sqrt{5}}{4}\right)$.
(d) From (b), we know that except for the origin 0, all periodic orbits should be in $\mathcal{M}$. By (c), we cannot have the substring 00 in a periodic orbit (except for the fixed point at 0 ). Hence 00 is the only pruning block, and the symbolic dynamics is a finite subshift, with alphabet $\{0,1\}$ and only one grammar rule: a consecutive repeat of symbol 0 is inadmissible.
(e) Yes. 0 is a periodic orbit with the symbol sequence $\overline{0}$. It is a repeller and no point in its neighborhood will return. So it plays no role in the asymptotic dynamics.
(Yueheng Lan)

## Chapter 12

(No solutions available.)

## Chapter 13

Solution 13.1: A transition matrix for 3-disk pinball. a) As the disk is convex, the transition to itself is forbidden. Therefore, the Markov diagram is

with the corresponding transition matrix

$$
\mathbb{T}=\left(\begin{array}{lll}
0 & 1 & 1 \\
1 & 0 & 1 \\
1 & 1 & 0
\end{array}\right)
$$

Note that $\mathbb{T}^{2}=\mathbb{T}+2$. Suppose that $\mathbb{T}^{n}=a_{n} \mathbb{T}+b_{n}$, then

$$
\mathbb{T}^{n+1}=a_{n} \mathbb{T}^{2}+b_{n} \mathbb{T}=\left(a_{n}+b_{n}\right) \mathbb{T}+2 a_{n}
$$

So $a_{n+1}=a_{n}+b_{n}, b_{n+1}=2 a_{n}$ with $a_{1}=1, b_{1}=0$.
b) From a) we have $a_{n+1}=a_{n}+2 a_{n-1}$. Suppose that $a_{n} \propto \lambda^{n}$. Then $\lambda^{2}=\lambda+2$. Solving this equation and using the initial condition for $n=1$, we obtain the general formula

$$
\begin{aligned}
a_{n} & =\frac{1}{3}\left(2^{n}-(-1)^{n}\right), \\
b_{n} & =\frac{2}{3}\left(2^{n-1}+(-1)^{n}\right) .
\end{aligned}
$$

c) $\mathbb{T}$ has eigenvalue 2 and -1 (degeneracy 2). So the topological entropy is $\ln 2$, the same as in the case of the binary symbolic dynamics.
(Yueheng Lan)
Solution 13.2: Sum of $A_{i j}$ is like a trace. Suppose that $A \phi_{k}=\lambda_{k} \phi_{k}$, where $\lambda_{k}, \phi_{k}$ are eigenvalues and eigenvectors, respectively. Expressing the vector $v=(1,1, \cdots, 1)^{t}$ in terms of the eigenvectors $\phi_{k}$, that is, $v=\Sigma_{k} d_{k} \phi_{k}$, we have

$$
\begin{aligned}
\Gamma_{n} & =\Sigma_{i j}\left[A^{n}\right]_{i j}=v^{t} A^{n} v=\Sigma_{k} v^{t} A^{n} d_{k} \phi_{k}=\Sigma_{k} d_{k} \lambda_{k}^{n}\left(v^{t} \phi_{k}\right) \\
& =\Sigma_{k} c_{k} \lambda_{k}^{n},
\end{aligned}
$$

where $c_{k}=\left(v^{t} \phi_{k}\right) d_{k}$ are constants.
a) As $\operatorname{tr} A^{n}=\Sigma_{k} \lambda_{k}^{n}$, it is easy to see that both $\operatorname{tr} A^{n}$ and $\Gamma_{n}$ are dominated by the largest eigenvalue $\lambda_{0}$. That is

$$
\frac{\ln \left|\operatorname{tr} A^{n}\right|}{\ln \left|\Gamma_{n}\right|}=\frac{n \ln \left|\lambda_{0}\right|+\ln \left|\Sigma_{k}\left(\frac{\lambda_{k}}{\lambda_{0}}\right)^{n}\right|}{n \ln \left|\lambda_{0}\right|+\ln \left|\Sigma_{k} d_{k}\left(\frac{\lambda_{k}}{\lambda_{0}}\right)^{n}\right|} \rightarrow 1 \text { as } n \rightarrow \infty .
$$

b) The nonleading eigenvalues do not need to be distinct, as the ratio in a) is controlled by the largest eigenvalues only.
(Yueheng Lan)
Solution 13.4: Transition matrix and cycle counting. a) According to the definition of $\mathbb{T}_{i j}$, the transition matrix is

$$
\mathbb{T}=\left(\begin{array}{ll}
a & c \\
b & 0
\end{array}\right) .
$$

b) All walks of length three $0000,0001,0010,0100,0101,1000,1001,1010$ (four symbols!) with weights aaa, aac, acb, cba, cbc, baa, bac, bcb. Let's calculate $\mathbb{T}^{3}$,

$$
\mathbb{T}^{3}=\left(\begin{array}{ll}
a^{3}+2 a b c & a^{2} c+b c^{2} \\
a^{2} b+b^{2} c & a b c
\end{array}\right) .
$$

There are altogether 8 terms, corresponding exactly to the terms in all the walks.
c) Let's look at the following equality

$$
\mathbb{T}_{i j}^{n}=\Sigma_{k_{1}, k_{2}, \cdots, k_{n-1}} \mathbb{T}_{i k_{1}} \mathbb{T}_{k_{1} k_{2}} \cdots \mathbb{T}_{k_{n-1} j}
$$

Every term in the sum is a possible path from $i$ to $j$, though the weight could be zero. The summation is over all possible intermediate points ( $n-1$ of them). So, $\mathbb{T}_{i j}^{n}$ gives the total weight (probability or number) of all the walks from $i$ to $j$ in $n$ steps.
d) We take $a=b=c=1$ to just count the number of possible walks in $n$ steps. This is the crudest description of the dynamics. Taking $a, b, c$ as transition probabilities would give a more detailed description. The eigenvlues of $\mathbb{T}$ is $(1 \pm \sqrt{5}) / 2$, so we get $N(n) \propto\left(\frac{1+\sqrt{5}}{2}\right)^{n}$.
e) The topological entropy is then $\ln \frac{1+\sqrt{5}}{2}$.
(Yueheng Lan)
Solution 13.6: "Golden mean" pruned map. It is easy to write the transition matrix $\mathbb{T}$

$$
\mathbb{T}=\left(\begin{array}{ll}
0 & 1 \\
1 & 1
\end{array}\right)
$$

The eigenvalues are $(1 \pm \sqrt{5}) / 2$. The number of periodic orbits of length $n$ is the trace

$$
\mathbb{T}^{n}=\frac{(1+\sqrt{5})^{n}+(1-\sqrt{5})^{n}}{2^{n}}
$$

(Yueheng Lan)
Solution 13.5: 3-disk prime cycle counting. The formula for arbitrary length cycles is derived in sect. 13.4.

Solution 13.45: Alphabet $\{\mathbf{0 , 1}\}$, prune _1000_, _00100_, _01100_.
step 1. _1000_ prunes all cycles with a _000_ subsequence with the exception of the fixed point $\overline{0}$; hence we factor out $\left(1-t_{0}\right)$ explicitly, and prune _000_ from the rest. Physically this means that $x_{0}$ is an isolated fixed point - no cycle stays in its vicinity for more than 2 iterations. In the notation of exercise 13.18, the alphabet is $\{1,2,3 ; \overline{0}\}$, and the remaining pruning rules have to be rewritten in terms of symbols $2=10,3=100$ :
step 2. alphabet $\{1,2,3 ; \overline{0}\}$, prune _33_, _213_,_313_. Physically, the 3-cycle $\overline{3}=\overline{100}$ is pruned and no long cycles stay close enough to it for a single _100_ repeat. As in exercise 13.7, prohibition of _33_ is implemented by dropping the symbol " 3 " and extending the alphabet by the allowed blocks 13, 23:
step 3. alphabet $\{1,2, \underline{13}, \underline{23} ; \overline{0}\}$, prune _213_, _ $\underline{33} \underline{13}-, \underline{13} \underline{13}-$, where $\underline{13}=13$, $\underline{23}=23$ are now used as single letters. Pruning of the repetitions -13 $\underline{13}$ - (the 4-cycle $\overline{13}=\overline{1100}$ is pruned) yields the

Result: alphabet $\{1,2, \underline{23}, \underline{113} ; \overline{0}\}$, unrestricted 4-ary dynamics. The other remaining possible blocks __213_, __2313_ are forbidden by the rules of step 3. The topological zeta function is given by

$$
\begin{equation*}
1 / \zeta=\left(1-t_{0}\right)\left(1-t_{1}-t_{2}-t_{23}-t_{113}\right) \tag{N.10}
\end{equation*}
$$

for unrestricted 4-letter alphabet $\{1,2, \underline{23}, \underline{113}\}$.
Solution 13.8: Spectrum of the "golden mean" pruned map.

1. The idea is that with the redefinition $2=10$, the alphabet $\{1,2\}$ is unrestricted binary, and due to the piecewise linearity of the map, the stability weights factor in a way similar to (14.10).
2. As in (15.9), the spectral determinant for the Perron-Frobenius operator takes form (15.11)

$$
\operatorname{det}(1-z \mathcal{L})=\prod_{k=0}^{\infty} \frac{1}{\zeta_{k}}, \quad \frac{1}{\zeta_{k}}=\prod_{p}\left(1-\frac{z^{n_{p}}}{\left|\Lambda_{p}\right| \Lambda_{p}^{k}}\right)
$$

The mapping is piecewise linear, so the form of the topological zeta function worked out in (13.16) already suggests the form of the answer. The alphabet $\{1,2\}$ is unrestricted binary, so the dynamical zeta functions receive contributions only from the two fixed points, with all other cycle contributions cancelled exactly. The $1 / \zeta_{0}$ is the spectral determinant for the transfer operator like the one in (10.19) with the $T_{00}=0$, and in general

$$
\begin{align*}
\frac{1}{\zeta_{k}} & =\left(1-\frac{z}{\left|\Lambda_{1}\right| \Lambda_{1}^{k}}\right)\left(1-\frac{z^{2}}{\left|\Lambda_{2}\right| \Lambda_{2}^{k}}\right)\left(1-\frac{z^{3}}{\left|\Lambda_{12}\right| \Lambda_{12}^{k}}\right) \cdots \\
& =1-(-1)^{k}\left(\frac{z}{\Lambda^{k+1}}+\frac{z^{2}}{\Lambda^{2 k+2}}\right) \tag{N.11}
\end{align*}
$$

The factor $(-1)^{k}$ arises because both stabilities $\Lambda_{1}$ and $\Lambda_{2}$ include a factor $-\Lambda$ from the right branch of the map.

## Solution 15.2: Dynamical zeta functions

1. Work through section sect. 15.3.2.
2. Generalize the transition matrix (11.15) to a transfer operator.

Solution 13.11: Whence Möbius function? Written out $f(n)$ line-by-line for a few values of $n$, (13.38) yields

$$
\begin{align*}
f(1) & =g(1) \\
f(2) & =g(2)+g(1) \\
f(3) & =g(3)+g(1) \\
f(4) & =g(4)+g(2)+g(1) \\
& \cdots  \tag{N.12}\\
f(6) & =g(6)+g(3)+g(2)+g(1)
\end{align*}
$$

Now invert recursively this infinite tower of equations to obtain

$$
\begin{aligned}
g(1) & =f(1) \\
g(2) & =f(2)-f(1) \\
g(3) & =f(3)-f(1) \\
g(4) & =f(4)-[f(2)-f(1)]-f(1)=f(4)-f(2) \\
& \cdots \\
g(6) & =f(6)-[f(3)-f(1)]-[f(2)-f(1)]-f(1)
\end{aligned}
$$

We see that $f(n)$ contributes with factor -1 if $n$ prime, and not at all if $n$ contains a prime factor to a higher power. This is precisely the raison d'etre for the Möbius function, with whose help the inverse of (13.38) can be written as the Möbius inversion formula [24.29] (13.39).

## Chapter 14

(No solutions available.)

## Chapter 15

Solution 15.1: Numerical estimate of the escape rate for a $1-d$ repeller The logistic map is defined by $x_{n+1}=A x_{n}\left(1-x_{n}\right)$. For $A \leq 4$ any point in the unit interval $[0,1]$ will remain in the interval forever. For $A>4$ almost all points starting in the unit interval will eventually escape towards $-\infty$.

The rate of escape can be easily measured by numerical experiment. We define the fraction of initial conditions that leave the interval after $n$ iterations to be $\Gamma_{n}$. Figure N. 2 shows a plot of $\log \left(\Gamma_{n}\right)$ versus $n$, computed by starting with 10000000 random initial points. Asymptotically the escape rate falls off exponentially as

$$
\Gamma(n)=C \mathrm{e}^{-\gamma n}
$$

Figure N. 2 suggests that this formula is very accurate even for relatively small $n$. We estimate $\gamma$ by measuring the slope of the curve in figure N.2. To avoid errors due to rounding and transients only the points $5 \leq n \leq 10$ were used. A linear regression fit yields the escape rate for $A=6$ :

$$
\gamma=0.8315 \pm 0.0001
$$

where the error is from statistical fluctuations (there may be systematic errors either due to rounding or because we are not in the true asymptotic regime).
(Adam Prügel-Bennet)

## Solution 15.2: Dynamical zeta functions

1. Work through section sect. 15.3.2.
2. Generalize the transition matrix (11.15) to a transfer operator.

Solution 15.5: Dynamical zeta functions as ratios of spectral determinants. Try inserting a factor equal to one in the zeta function and then expanding it. The problem is solved in sect. 15.5.

Solution 15.8: Dynamical zeta functions for Hamiltonian maps. Read sect. 15.7.


Figure N.2: Plot of $\log (\Gamma(n))$ versus $n$ for the logistic map $x_{n+1}=6 x_{n}(1-$ $x_{n}$ ). Error bars show estimated errors in the mean assuming a binomial distribution. 10000000 random initial starting points were used.

## Chapter 16

Solution 16.3: Euler formula. Let

$$
P=\prod_{k=0}^{\infty}\left(1+t u^{k}\right)=\sum_{n=0}^{\infty} P_{n} t^{n}
$$

then

$$
\begin{align*}
P_{n} & =\left.\frac{1}{n!} \frac{\partial^{n} P}{\partial t^{n}}\right|_{t=0}=\frac{1}{n!} \sum_{i_{n} \neq i_{n-1} \neq \cdots \neq i_{1}} u^{i_{n}+i_{n-1}+\cdots+i_{1}} \\
& =\sum_{i_{n}>i_{n-1}>\cdots i_{1} \geq 0} u^{i_{n}+i_{n-1}+\cdots+i_{1}} \tag{N.13}
\end{align*}
$$

Clearly $P_{0}=1$, and

$$
P_{1}=\sum_{i=0} u^{i}
$$

multiplying both sides by $1-u$

$$
(1-u) P_{1}=1+u+u^{2}+\cdots-\left(u+u^{2}+\cdots\right)=1
$$

(since, for $|u|<1, \lim _{n \rightarrow \infty} u^{n}=0$ ). Thus $P_{1}=1 /(1-u)$. Similarly

$$
P_{2}=\sum_{i>j \geq 0} u^{i+j}
$$



Performing the same trick as for $P_{1}$

$$
(1-u) P_{2}=\sum_{i>j \geq 0} u^{i+j}-\sum_{i>j \geq 0} u^{i+(j+1)}
$$

The only terms that survive are those for which $j=i-1$ (that is the top diagonal in the figure) thus

$$
(1-u) P_{2}=u^{-1} \sum_{i=1}^{\infty} u^{2 i}
$$

and

$$
(1-u)\left(1-u^{2}\right) P_{2}=u^{-1}\left(u^{2}+u^{4}+\cdots-\left(u^{4}+u^{6}+\cdots\right)\right)=u
$$

Thus

$$
P_{2}=\frac{u}{(1-u)\left(1-u^{2}\right)}
$$

In general

$$
(1-u) P_{n}=\sum_{i_{n}>i_{n-1}>\cdots i_{1} \geq 0} u^{i_{n}+i_{n-1}+\cdots+i_{1}}-\sum_{i_{n}>i_{n-1}>\cdots i_{1} \geq 0} u^{i_{n}+i_{n-1}+\cdots+\left(i_{1}+1\right)}
$$

$$
\begin{equation*}
=u^{-1} \sum_{i_{n}>i_{n-1}>\cdots i_{2} \geq 1} u^{i_{n}+i_{n-1}+\cdots+2 i_{2}} \tag{N.14}
\end{equation*}
$$

since only the term $i_{1}=i_{2}-1$ survives. Repeating this trick

$$
(1-u)\left(1-u^{2}\right) P_{n}=u^{-1-2} \sum_{i_{n}>i_{n-1}>\cdots i_{3} \geq 2} u^{i_{n}+i_{n-1}+\cdots+3 i_{3}}
$$

and

$$
\prod_{i=1}^{n}\left(1-u^{i}\right) P_{n}=u^{-(1+2+\cdots+n)} u^{n(n-1)}=u^{n(n-1) / 2}
$$

Thus

$$
P_{n}=\frac{u^{n(n-1) / 2}}{\prod_{i=1}^{n}\left(1-u^{i}\right)}
$$

Solution 16.3: Euler formula, 2nd method. The coefficients $Q_{k}$ in (16.4) are given explicitly by the Euler formula

$$
\begin{equation*}
Q_{k}=\frac{1}{1-\Lambda^{-1}} \frac{\Lambda^{-1}}{1-\Lambda^{-2}} \cdots \frac{\Lambda^{-k+1}}{1-\Lambda^{-k}} \tag{N.16}
\end{equation*}
$$

Such a formula is easily proved by considering the finite order product

$$
\mathcal{W}_{j}(z, \gamma)=\prod_{l=0}^{j}\left(1+z \gamma^{l}\right)=\sum_{l=0}^{j+1} \Gamma_{l} z^{l}
$$

Since we have that

$$
\left(1+z \gamma^{j+1}\right) \mathcal{W}_{j}(z, \gamma)=(1+z) \mathcal{W}_{j}(\gamma z, \gamma)
$$

we get the following identity for the coefficients

$$
\Gamma_{m}+\Gamma_{m-1} \gamma^{j+1}=\Gamma_{m} \gamma^{m}+\Gamma_{m-1} \gamma^{m-1} m=1, \ldots
$$

Starting with $\Gamma_{0}=1$, we recursively get

$$
\Gamma_{1}=\frac{1-\gamma^{j+1}}{1-\gamma} \quad \Gamma_{2}=\frac{\left(1-\gamma^{j+1}\right)\left(\gamma-\gamma^{j+1}\right)}{(1-\gamma)\left(1-\gamma^{2}\right)} \ldots
$$

the Euler formula (16.5) follows once we take the $j \rightarrow \infty$ limit for $|\gamma|<1$.

Solution 16.3: Euler formula, 3rd method. First define

$$
\begin{equation*}
f(t, u):=\prod_{k=0}^{\infty}\left(1+t u^{k}\right) \tag{N.17}
\end{equation*}
$$

Note that

$$
\begin{equation*}
f(t, u)=(1+t) f(t u, u) \tag{N.18}
\end{equation*}
$$

by factoring out the first term in the product. Now make the ansatz

$$
\begin{equation*}
f(t, u)=\sum_{n=0}^{\infty} t^{n} g_{n}(u) \tag{N.19}
\end{equation*}
$$

plug it into (N.18), compare the coefficients of $t^{n}$ and get

$$
\begin{equation*}
g_{n}(u)=u^{n} g_{n}(u)+u^{n-1} g_{n-1}(u) . \tag{N.20}
\end{equation*}
$$

Of course $g_{0}(u)=1$. Therefore by solving the recursion (N.20) and by noting that $\sum_{k=1}^{n-1} k=\frac{n(n-1)}{2}$ one finally arrives at

$$
\begin{equation*}
g_{n}(u)=\frac{u^{\frac{n(n-1)}{2}}}{\prod_{k=1}^{n}\left(1-u^{k}\right)} . \tag{N.21}
\end{equation*}
$$

Euler got this formula and he and Jacobi got many nice number theoretical results from it, most prominent the pentagonal number theorem, which says that in the series expansion of $\prod_{k=1}^{\infty}\left(1-q^{k}\right)$ all terms cancel except those which have as an exponent the circumference of a regular pentagon with integer base length.
(Juri Rolf)
Solution 16.4: 2-d product expansion. Now let us try to apply the same trick as above to the two dimensional situation

$$
\begin{equation*}
h(t, u):=\prod_{k=0}^{\infty}\left(1+t u^{k}\right)^{k+1} \tag{N.22}
\end{equation*}
$$

Write down the first terms and note that similar to (N.18)

$$
\begin{equation*}
h(t, u)=f(t, u) h(t u, u), \tag{N.23}
\end{equation*}
$$

where $f$ is the Euler product (N.17). Now make the ansatz

$$
\begin{equation*}
h(t, u)=\sum_{n=0}^{\infty} t^{n} a_{n}(u) \tag{N.24}
\end{equation*}
$$

and use the series expansion for $f$ in (N.23) to get the recursion

$$
\begin{equation*}
a_{n}(u)=\frac{1}{1-u^{n}} \sum_{m=0}^{n-1} u^{m} a_{m}(u) g_{n-m}(u) . \tag{N.25}
\end{equation*}
$$

With this one can at least compute the generalized Euler product effectively, but it would be nice if one could use it for a proof of the general behaviour of the coefficients $a_{n}$.
(Juri Rolf)

## Chapter 17

Solution 17.3: Stability of billiard cycles. The 2-cycle $\overline{0}$ stability (8.1) is the solution to both problems (provided you evaluate correctly the hyperbola curvature on the diagonal).

Solution 17.4: Numerical cycle routines. A number of sample Fortran programs for finding periodic orbits is available on the homepage for this manuscript, www.nbi.dk/ChaosBook/.

Solution 17.11: Inverse iteration method for a Hamiltonian repeller.
For the complete repeller case (all binary sequences are realized), the cycles can be evaluated variationally, as follows. According to (3.12), the coordinates of a periodic orbit of length $n_{p}$ satisfy the equation

$$
\begin{equation*}
x_{p, i+1}+x_{p, i-1}=1-a x_{p, i}^{2}, \quad i=1, \ldots, n_{p} \tag{N.26}
\end{equation*}
$$

with the periodic boundary condition $x_{p, 0}=x_{p, n_{p}}$.
In the complete repeller case, the Hénon map is a realization of the Smale horseshoe, and the symbolic dynamics has a very simple description in terms of the binary alphabet $\epsilon \in\{0,1\}, \epsilon_{p, i}=\left(1+S_{p, i}\right) / 2$, where $S_{p, i}$ are the signs of the corresponding cycle point coordinates, $S_{p, i}=x_{p, i} /\left|x_{p, i}\right|$. We start with a preassigned sign sequence $S_{p, 1}, S_{p, 2}, \ldots, S_{p, n_{p}}$, and a good initial guess for the coordinates $x_{p, i}^{\prime}$. Using the inverse of the equation (17.22)

$$
\begin{equation*}
x_{p, i}^{\prime \prime}=S_{p, i} \sqrt{\frac{1-x_{p, i+1}^{\prime}-x_{p, i-1}^{\prime}}{a}}, i=1, \ldots, n_{p} \tag{N.27}
\end{equation*}
$$

we converge iteratively, at exponential rate, to the desired cycle points $x_{p, i}$. Given the cycle points, the cycle stabilities and periods are easily computed using (4.36). The itineraries and the stabilities of the short periodic orbits for the Hénon repeller (N.26) for $a=6$ are listed in table 17.2; in actual calculations all prime cycles up to topological length $n=20$ have been computed.
(G. Vattay)

## Chapter 18

Solution 18.2: Prime cycles for a 1- $d$ repeller, analytic formulas. For the logistic map the prime cycles, ordered in terms of their symbolic dynamics, are listed in table 11.1

$$
\mathcal{P}=\{0,1,01,001,011,0001,0011,0111, \ldots\}
$$

The position of the prime cycles can be found by iterating the inverse mapping. If we wish to find the position of a prime orbit $p=b_{1} b_{2} \cdots b_{n_{p}}$, where $b_{i} \in\{0,1\}$, then starting from some initial point, $x=1 / 2$ say, we apply one of the inverse mappings

$$
f_{ \pm}^{-1}(x)=\frac{1}{2} \pm \frac{1}{2} \sqrt{1-x / 4 A}
$$



Figure N.3: Periodic orbits and stabilities for the logistics equation $x_{n+1}=6 x_{n}(1-$ $x_{n}$ ).
where we choose $f_{-}^{-1}$ if $b_{1}=0$ or $f_{+}^{-1}$ if $b_{1}=1$. We then apply the inverse mapping again depending on the next element in the prime orbit. Repeating this procedure many times we converge onto the prime cycle. The stability $\Lambda_{p}$ of a prime cycle $p$ is given by the product of slopes of $f$ around the cycle. The first eight prime cycles are shown in figure N.3.

The stabilities of the first five prime orbits can be calculated for arbitrary $A$. We find that $\Lambda_{0}=A, \Lambda_{1}=2-A, \Lambda_{01}=4+2 A-A^{2}$, and

$$
\begin{equation*}
\Lambda_{\substack{001 \\ 011}}=8+2 A-A^{2} \pm A(2-A) \sqrt{A^{2}-2 A-7} \tag{N.28}
\end{equation*}
$$

There is probably a closed form expression for the 4-cycles as well.

For crosschecking purposes: if $A=9 / 2, \Lambda_{0}=9 / 2 \Lambda_{1}=-5 / 2 \Lambda_{01}=-7.25$ $\Lambda_{011}=19.942461 \ldots$

Solution 18.2: Dynamical zeta function for a $1-d$ repeller The escape rate can be estimated from the leading zero in the dynamical zeta function $1 / \zeta(z)$, defined by

$$
1 / \zeta(z)=\prod_{p}\left(1-z^{n_{p}} /\left|\Lambda_{p}\right|\right)
$$

To compute the position of this pole we expand $1 / \zeta(z)$ as a power series (18.5) in $z$

$$
1 / \zeta(z)=1-\sum_{i=1} \hat{c}_{i} z^{i}
$$

where

$$
\begin{aligned}
& \hat{c}_{1}=\left|\Lambda_{0}\right|^{-1}+\left|\Lambda_{1}\right|^{-1}, \quad \hat{c}_{2}=\left|\Lambda_{01}\right|^{-1}-\left|\Lambda_{1} \Lambda_{0}\right|^{-1} \\
& \hat{c}_{3}=\left|\Lambda_{001}\right|^{-1}-\left|\Lambda_{0} \Lambda_{01}\right|^{-1}+\left|\Lambda_{011}\right|^{-1}-\left|\Lambda_{01} \Lambda_{1}\right|^{-1}
\end{aligned}
$$

etc.. Using the cycles up to length 6 we get

$$
\begin{aligned}
1 / \zeta(z)= & 1-0.416667 z-0.00833333 z^{2} \\
& +0.000079446 z^{3}-9.89291 \times 10^{-7} z^{4}+\ldots
\end{aligned}
$$

The leading zero of this Taylor series is an estimate of $\exp (\gamma)$. Using $n=1,2,3$ and 4 we obtain the increasingly accurate estimates for $\gamma: 0.875469,0.830597,0.831519$ and 0.831492 In a hope to improve the convergence we can use the Padé approximates $P_{M}^{N}(z)=\sum_{i=1}^{N} p_{i} z^{i} /\left(1+\sum_{j=1}^{M} q_{j} z^{j}\right)$. Using the Padé approximates $P_{1}^{n-1}(z)$ for $n=2,3$ and 4 we obtain the estimates $0.828585,0.831499$ and 0.831493 .

The above results correspond to $A=6$; in the $A=9 / 2$ case the leading zero is $1 / z=1.43549 \ldots$ and $\gamma=0.36150 \ldots$ (Adam Prügel-Bennet)

Solution 18.2: Spectral determinant for a 1-d repeller We are told the correct expression for the escape rate is also given by the logarithm of the leading zero of the spectral determinant (15.11), expanded as the Taylor series (18.8). The coefficients $c_{i}$ should fall off super-exponentially so that truncating the Taylor series is expected to give a far more accurate estimate of the escape rate than using the dynamical zeta function. How do we compute the $c_{i}$ coefficients in (18.8)? One straightforward method is to first compute the Taylor expansion of $\log (F(z))$

$$
\begin{aligned}
\log (F(z)) & =\sum_{p} \sum_{k=0} \log \left(1-\frac{t_{p}}{\Lambda_{p}^{k}}\right)=-\sum_{p} \sum_{k=0} \sum_{r=1} \frac{t_{p}^{r}}{\Lambda_{p}^{k r}} \\
& =-\sum_{p} \sum_{r=1} \frac{t_{p}^{r}}{1-\Lambda_{p}^{-r}}=-\sum_{p} \sum_{r=1} B_{p}(r) z^{n_{p} r}
\end{aligned}
$$

where $B_{p}(r)=-1 / r\left|\Lambda_{p}^{r}\right|\left(1+\Lambda_{p}^{-r}\right)$. Writing $\log (F(z))$ as a power series

$$
\log (F(z))=-\sum_{i=1} b_{i} z^{i}
$$

we obtain

$$
\begin{align*}
& b_{1}=B_{0}(1)+B_{1}(1) \\
& b_{2}=B_{01}(1)+B_{0}(2)+B_{1}(2) \\
& b_{3}=B_{001}(1)+B_{011}(1)+B_{0}(3)+B_{1}(3) \\
& b_{3}=B_{0001}(1)+B_{0011}(1)+B_{0111}(1)+B_{01}(2)+B_{0}(4)+B_{1}(4) \tag{N.29}
\end{align*}
$$

etc.. To obtain the coefficients for the spectral determinant we solve

$$
F(z)=1-\sum_{i=1} Q_{i} z^{i}=\exp \left(\sum_{i=1} b_{i} z^{i}\right)
$$



Figure N.4: Plot of the Taylor coefficients for the spectral determinant, $c_{i}$, and for the dynamical zeta function, $b_{i}$.
for the $Q_{i}$ 's. This gives

$$
\begin{aligned}
Q_{1} & =b_{1}, \quad Q_{2}=b_{2}+b_{1}^{2} / 2, \quad Q_{3}=b_{3}+b_{1} b_{2}+b_{1}^{3} / 6 \\
Q_{4} & =b_{4}+b_{1} b_{3}+b_{2}^{2} / 2+b_{2} b_{1}^{2} / 2+b_{1}^{4} / 24
\end{aligned}
$$

Using these formulas we find

$$
F(z)=1-0.4 z-0.0152381 z^{2}-0.0000759784 z^{3}+4.5311 \times 10^{-9} z^{4}+\cdots
$$

The logarithm of the leading zero of $F(z)$ again gives the escape rate. Using the $n=1,2,3$, and 4 truncations we find the approximation to $\gamma$ of $0.916291,0.832345$, 0.83149289 and 0.8314929875 . As predicted, the convergence is much faster for the spectral determinant than for the dynamical zeta function.

In figure N. 4 we show a plot of the logarithm of the coefficients for the spectral determinant and for the dynamical zeta function.
(Adam Prügel-Bennet)
The above results correspond to $A=6$; in the $A=9 / 2$ case all cycles up to length 10 yield $\gamma=0.36150966984250926 \ldots$
(Vadim Moroz)
Solution 18.2: Escape rate for a 1- $d$ repeller We can compute an approximate functional dependence of the escape rate on the parameter a using the stabilities of the first five prime orbits computed above, see (N.28). The spectral determinant (for $a>4$ ) is

$$
\begin{align*}
F= & 1-\frac{2 z}{a-1}-\frac{8 z^{2}}{(a-3)(a-1)^{2}(a+1)} \\
& +\left(\frac{2\left(32-18 a+17 a^{2}-16 a^{3}+14 a^{4}-6 a^{5}+a^{6}\right)}{(a-3)(a-1)^{3}(1+a)\left(a^{2}-5 a+7\right)\left(a^{2}+a+1\right)}\right.  \tag{N.30}\\
& \left.-\frac{2 a(a-2) \sqrt{\left(a^{2}-2 a-7\right)}}{\left(a^{2}-5 a+7\right)\left(a^{2}-2 a-7\right)\left(a^{2}+a+1\right)}\right) z^{3}
\end{align*}
$$



Figure N.5: Plot of the escape rate versus $a$ for the logistic map $x_{n+1}=a x_{n}\left(1-x_{n}\right)$ calculated from the first five periodic orbits.

The leading zero is plotted in figure N.5; it always remains real while the other two roots which are large and negative for $a>5.13 \ldots$ become imaginary below this critical value. The accuracy of this truncation is clearly worst for $a \rightarrow 4$, the value at which the hyperbolicity is lost and the escape rate goes to zero.
(Adam Prügel-Bennet)
Solution 18.3: Escape rate for the Ulam map. The answer is worked out in Nonlinearity 3, 325; 3, 361 (1990).

## Chapter 19

## Solution 19.3:

(d) In the $A=9 / 2$ case all cycles up to length 9 yield $\lambda=1.08569 \ldots$.. (Vadim Moroz)

Solution 9.4: The escape rate is the leading zero of the zeta function

$$
0=1 / \zeta(\gamma)=1-e^{\gamma} / 2 a-e^{\gamma} / 2 a=1-e^{\gamma} / a
$$

So, $\gamma=\log (a)$ if $a>a_{c}=1$ and $\gamma=0$ otherwise. For $a \approx a_{c}$ the escape rate behaves like

$$
\gamma(a) \approx\left(a-a_{c}\right)
$$

Solution 19.1: $\quad$ The escape is controlled by the size of the primary hole of the repeller. All subholes in the repeller will be proportional with the main hole. The size of the main hole is $l=\sqrt{1-1 / a}$. Near $a_{c}=1$ the escape rate is

$$
\gamma(a) \sim\left(a-a_{c}\right)^{1 / 2}
$$

We can generalize this and the previous result and conclude that

$$
\gamma(a) \sim\left(a-a_{c}\right)^{1 / z}
$$

where $z$ is the order of the maximum of the single humped map.
Solution 19.2: By direct evaluation we can calculate the zeta functions and the Fredholm determinant of this map. The zeta functions are

$$
1 / \zeta_{k}(z)=\operatorname{det}\left(1-z \mathbf{T}_{k}\right)
$$

where

$$
\mathbf{T}_{k}=\left(\begin{array}{ll}
T_{00}^{k+1} & T_{01}^{k+1} \\
T_{10}^{k+1} & T_{11}^{k+1}
\end{array}\right)
$$

and $T_{00}=1 / a_{1}, T_{01}=\left(b-b / a_{1}\right) /(1-b), T_{11}=\left(1-b-b / a_{2}\right) /(1-b), T_{10}=1 / a_{2}$ are inverses of the slopes of the map. The Fredholm determinant is the product of zeta functions

$$
F(z)=\prod_{k=0}^{\infty} 1 / \zeta_{k}(z)
$$

The leading zeroes of the Fredholm determinant can come from the zeroes of the leading zeta functions.

The zeroes of $1 / \zeta_{0}(z)$ are

$$
\begin{aligned}
& 1 / z_{1}=\frac{T_{00}+T_{11}+\sqrt{\left(T_{00}-T_{11}\right)^{2}+4 T_{01} T_{10}}}{2} \\
& 1 / z_{2}=\frac{T_{00}+T_{11}-\sqrt{\left(T_{00}-T_{11}\right)^{2}+4 T_{01} T_{10}}}{2}
\end{aligned}
$$

The zeroes of $1 / \zeta_{1}(z)$ are

$$
\begin{aligned}
& 1 / z_{3}=\frac{T_{00}^{2}+T_{11}^{2}+\sqrt{\left(T_{00}^{2}-T_{11}^{2}\right)^{2}+4 T_{01}^{2} T_{10}^{2}}}{1 / z_{4}=\frac{T_{00}^{2}+T_{11}^{2}-\sqrt{\left(T_{00}^{2}-T_{11}^{2}\right)^{2}+4 T_{01}^{2} T_{10}^{2}}}{2}} .
\end{aligned}
$$

By substituting the slopes we can show that $z_{1}=1$ is the leading eigenvalue. The next to leading eigenvalue, which is the correlation decay in discrete time, can be $1 / z_{3}$ or $1 / z_{2}$.

## Chapter 20

Solution 20.1: In the higher dimensional case there is no change in the derivation except $\Lambda_{p}$ should be replaced with the product of expanding eigenvalues $\prod_{j}\left|\Lambda_{p, j}\right|$. The logarithm of this product is $\sum_{j} \log \left|\Lambda_{p, j}\right|$. The average of $\log \left|\Lambda_{, j}\right|$ is the $j$-th Lyapunov exponent.

Solution 20.4: The zeta function for the two scale map is

$$
1 / \zeta(z, \beta)=1-z\left(\frac{1}{a^{\beta}}+\frac{1}{b^{\beta}}\right)
$$

The pressure function is

$$
P(\beta)=\log z_{0}(\beta)=-\log \left(\frac{1}{a^{\beta}}+\frac{1}{b^{\beta}}\right)
$$

The escape rate is

$$
\gamma=P(1)=-\log \left(\frac{1}{a}+\frac{1}{b}\right)
$$

The topological entropy is

$$
K_{0}=h_{t o p}=-P(0)=\log 2
$$

The Lyapunov exponent is

$$
\bar{\lambda}=P^{\prime}(1)=\frac{\log a / a+\log b / b}{1 / a+1 / b}
$$

The Kolmogorov entropy is

$$
K_{1}=\bar{\lambda}-\gamma=P^{\prime}(1)-P(1)=\frac{\log a / a+\log b / b}{1 / a+1 / b}+\log \left(\frac{1}{a}+\frac{1}{b}\right)
$$

The Rényi entropies are

$$
K_{\beta}=(P(\beta)-\beta \gamma) /(\beta-1)=\left(\log \left(\frac{1}{a^{\beta}}+\frac{1}{b^{\beta}}\right)+\beta \log \left(\frac{1}{a}+\frac{1}{b}\right)\right) /(1-\beta)
$$

The box counting dimension is the solution of the implicit equation $P\left(D_{0}\right)=0$, which is

$$
1=\frac{1}{a_{0}^{D}}+\frac{1}{b_{0}^{D}}
$$

The information dimension is

$$
D_{1}=1-\gamma / \bar{\lambda}
$$

The rest of the dimensions can be determined from equation $P\left(q-(q-1) D_{q}\right)=\gamma q$. Taking exp of both sides we get

$$
\frac{1}{a^{q-(q-1) D_{q}}}+\frac{1}{b^{q-(q-1) D_{q}}}=\left(\frac{1}{a}+\frac{1}{b}\right)^{q}
$$

For a given $q$ we can find $D_{q}$ from this implicit equation.
Solution 20.5: The zeta function is

$$
1 / \zeta(z, \beta)=\operatorname{det}\left(1-\mathbf{T}_{\beta-1}\right)
$$

where we replaced $k$ with $\beta-1$ in solution $N$. The pressure can be calculated from the leading zero which is (see solution N)

$$
P(\beta)=\log z_{0}(\beta)=-\log \left(\frac{T_{00}^{\beta}+T_{11}^{\beta}+\sqrt{\left(T_{00}^{\beta}-T_{11}^{\beta}\right)^{2}+4 T_{01}^{\beta} T_{10}^{\beta}}}{2}\right)
$$

Solution 20.6: We can easily read off that $b=1 / 2, a_{1}=\arcsin (1 / 2) / 2 \pi$ and $a_{2}=a_{1}$ and do the steps as before.

## Chapter 21

(No solutions available.)

## Chapter 22

(No solutions available.)

(a)
(b)

Figure N.6: (a) (b) A partition of the unit interval into three or five intervals, labeled by the order along the unit interval $\mathcal{A}=\left\{\mathcal{M}_{1}, \mathcal{M}_{2}=\mathcal{M}_{4} \cup\left(\frac{1}{2}\right) \cup \mathcal{M}_{5}, \mathcal{M}_{3}\right\}$. The partition is Markov, as the critical point is also a fixed point. (c) the Markov graph for this Markov partition.

## Chapter 23

Solution 23.1: Diffusion for odd integer $\Lambda$. Consider first the case $\Lambda=3$, illustrated in figure N.6. If $\beta=0$, the dynamics in the elementary cell is simple enough; a partition can be constructed from three intervals, which we label $\left\{\mathcal{M}_{1}, \mathcal{M}_{2}, \mathcal{M}_{3}\right\}$, with the alphabet ordered as the intervals are laid out along the unit interval. The Markov graph is figure N.6(c), and the dynamical zeta function is

$$
1 /\left.\zeta\right|_{\beta=0}=1-\left(t_{1}+t_{2}+t_{3}\right)=1-3 z / \Lambda,
$$

with eigenvalue $z=1$ as required by the flow conservation.
However, description of global diffusion requires more care. As explained in the definition of the map (23.9), we have to split the partition $\mathcal{M}_{2}=\mathcal{M}_{4} \cup\left(\frac{1}{2}\right) \cup \mathcal{M}_{5}$, and exclude the fixed point $f\left(\frac{1}{2}\right)=\frac{1}{2}$, as the map $\hat{f}(\hat{x})$ is not defined at $\hat{f}\left(\frac{1}{2}\right)$. (Are we to jump to the right or to the left at that point?) As we have $f\left(\mathcal{M}_{4}\right)=\mathcal{M}_{1} \cup \mathcal{M}_{4}$, and similarly for $f\left(\mathcal{M}_{5}\right)$, the Markov graph figure N.6(d) is infinite, and so is the dynamical zeta function:

$$
1 / \zeta=1-t_{1}-t_{14}-t_{144}-t_{1444} \cdots-t_{3}-t_{35}-t_{355}-t_{3555} \cdots .
$$

The infinite alphabet $\mathcal{A}=\{1,14,144,1444 \cdots 3,35,355,3555 \cdots\}$ is a consequence of the exclusion of the fixed point(s) $x_{4}, x_{5}$. As is customary in such situations (see exercise 18.10, and chapter 21, inter alia), we deal with this by dividing out the undesired fixed point from the dynamical zeta function. We can factorize and resum the weights using the piecewise linearity of (23.9)

$$
1 / \zeta=1-\frac{t_{1}}{1-t_{4}}-\frac{t_{3}}{1-t_{5}} .
$$

The diffusion constant is now most conveniently evaluated by evaluating the partial derivatives of $1 / \zeta$ as in (18.16)

$$
\begin{align*}
\langle T\rangle_{\zeta} & =-z \frac{\partial}{\partial z} \frac{1}{\zeta}=\left.2\left(\frac{t_{1}}{1-t_{4}}+\frac{t_{1} t_{4}}{\left(1-t_{4}\right)^{2}}\right)\right|_{z=1, \beta=0}=\frac{3}{4} \\
\left.\left\langle\hat{x}^{2}\right\rangle_{\zeta}\right|_{z=1, \beta=0} & =2\left(\frac{\hat{n}_{1}\left(\hat{n}_{1}+\hat{n}_{4}\right) \Lambda^{2}}{(1-1 / \Lambda)^{2}}+2 \frac{\hat{n}_{4}^{2} / \Lambda^{3}}{(1-1 / \Lambda)^{3}}\right)=\frac{1}{2} \tag{N.31}
\end{align*}
$$

yielding $D=1 / 3$, in agreement with in (23.21) for $\Lambda=3$.

## Chapter 24

(No solutions available.)

## Chapter 26

Solution 26.2: Green's function. The Laplace transform of the (time-dependent) quantum propagator

$$
K\left(q, q^{\prime}, t\right)=\sum_{n} \phi_{n}(q) e^{-i E_{n} t / \hbar} \phi_{n}^{*}\left(q^{\prime}\right)
$$

is the (energy-dependent) Green's function

$$
\begin{aligned}
G\left(q, q^{\prime}, E+i \varepsilon\right) & =\frac{1}{i \hbar} \int_{0}^{\infty} d t e^{\frac{i}{\hbar} E t-\frac{\varepsilon}{\hbar} t} \sum_{n} \phi_{n}(q) e^{-i E_{n} t / \hbar} \phi_{n}^{*}\left(q^{\prime}\right) \\
& =\frac{1}{i \hbar} \sum_{n} \phi_{n}(q) \phi_{n}^{*}\left(q^{\prime}\right) \int_{0}^{\infty} d t e^{\frac{i}{\hbar}\left(E-E_{n}+i \varepsilon\right) t} \\
& =-\left.\sum_{n} \phi_{n}(q) \phi_{n}^{*}\left(q^{\prime}\right) \frac{1}{E-E_{n}+i \varepsilon} e^{-\frac{\varepsilon}{\hbar} t} e^{i\left(E-E_{n}\right) t / \hbar}\right|_{t=0} ^{t=\infty} .
\end{aligned}
$$

When $\varepsilon$ is positive, $e^{-\frac{\varepsilon}{\hbar} \infty}=0$, so

$$
G\left(q, q^{\prime}, E+i \varepsilon\right)=\sum_{n} \frac{\phi_{n}(q) \phi_{n}^{*}\left(q^{\prime}\right)}{E-E_{n}+i \varepsilon} .
$$

## Solution 26.1: Lorentzian representation of the Dirac delta function.

General hint: read up on principal parts, positive and negative frequency parts of the Dirac delta function, perhaps the Cauchy theorem, in any good quantum mechanics textbook.

To see that (26.19) satisfies properties of the delta function,

$$
\delta\left(E-E_{n}\right)=-\lim _{\varepsilon \rightarrow 0} \frac{1}{\pi} \operatorname{Im} \frac{1}{E-E_{n}+i \varepsilon},
$$

start by expressing explicitely the imaginary part:

$$
\begin{aligned}
-\operatorname{Im} \frac{1}{E-E_{n}+i \varepsilon} & =-\operatorname{Im} \frac{E-E_{n}-i \varepsilon}{\left(E-E_{n}+i \varepsilon\right)\left(E-E_{n}-i \varepsilon\right)} \\
& =\frac{\varepsilon}{\left(E-E_{n}\right)^{2}+\varepsilon^{2}}
\end{aligned}
$$

This is a Lorenzian of width $\epsilon$, with a peak at $E=E_{n}$. It has the correct normalization for the delta function,

$$
\begin{aligned}
\frac{1}{\pi} \int_{-\infty}^{\infty} d E \frac{\varepsilon}{\left(E-E_{n}\right)^{2}+\varepsilon^{2}} & =\left.\frac{1}{\pi} \frac{\varepsilon}{\varepsilon} \arctan \frac{E-E_{n}}{\varepsilon}\right|_{-\infty} ^{\infty} \\
& =\frac{1}{\pi}(\pi / 2-(-\pi / 2))=1
\end{aligned}
$$

so

$$
\begin{equation*}
\frac{1}{\pi} \int_{-\infty}^{\infty} d E \frac{\epsilon}{\left(E-E_{n}\right)^{2}+\epsilon^{2}}=1 \tag{N.32}
\end{equation*}
$$

independently of the width $\epsilon$.
Next we show that in the $\epsilon \rightarrow \infty$ limit the support of the Lorentzian is concentrated at $E=E_{n}$. When $E=E_{n}$,

$$
\lim _{\varepsilon \rightarrow 0} \frac{1}{\pi}\left(\frac{\varepsilon}{\left(E-E_{n}\right)^{2}+\varepsilon^{2}}\right)=\lim _{\varepsilon \rightarrow 0} \frac{1}{\pi} \frac{1}{\varepsilon}=\infty
$$

and when $E \neq E_{n}$,

$$
\lim _{\varepsilon \rightarrow 0} \frac{1}{\pi} \frac{\varepsilon}{\left(E-E_{n}\right)^{2}+\varepsilon^{2}}=0
$$

Providing that a function convolved with $\delta(s), \int f(E) \delta\left(E-E_{n}\right) d E$ has a continuous first derivative at $E=E_{n}$ and falls of sufficiently rapidly as $E \rightarrow \pm \infty$, this is a representation of the delta function.
(R. Paskauskas, Bo Li)

## Chapter 27

Solution 27.2: Fresnel integral. Start by re-expressing the integral overthe infinite half-line:

$$
\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} d x e^{-\frac{x^{2}}{2 i a}}=\frac{2}{\sqrt{2 \pi}} \int_{0}^{\infty} d x e^{-\frac{x^{2}}{2 i a}}, \quad a \in \mathbb{R}, \quad a \neq 0
$$

When $a>0$, the contour
 vanishes, as it contains no pole:

$$
\begin{align*}
\oint_{C} d z e^{-z^{2} / 2 i a} & =\int_{0}^{\infty} d x e^{-\frac{x^{2}}{2 i a}}+\int_{C^{\prime}}+\int_{\infty}^{0} e^{i \frac{\pi}{4}} e^{-\frac{x^{2}}{2 a}} d x=0 \\
\int_{C^{\prime}} & =\int_{0}^{\frac{\pi}{4}} e^{i R^{2} e^{i 2 \phi} / 2 a} R e^{i \phi} i d \phi=0 \tag{N.33}
\end{align*}
$$

So

$$
\begin{aligned}
& \text { In the } a<0 \text { case take the contour }
\end{aligned}
$$

$$
\begin{aligned}
\oint_{C} d z e^{-z^{2} / 2 i a} & =\int_{0}^{\infty} d x e^{-\frac{x^{2}}{2 i a}}+\int_{C^{\prime}}+\int_{\infty}^{0} e^{-i \frac{\pi}{4}} e^{\frac{x^{2}}{2 a}} d x \\
& =\int_{0}^{\infty} d x e^{-\frac{x^{2}}{2 i a}}-e^{-i \frac{\pi}{4}} \int_{0}^{\infty} d x e^{\frac{x^{2}}{2 a}}=0
\end{aligned}
$$

Again

$$
\frac{2}{\sqrt{2 \pi}} \int_{0}^{\infty} d x e^{-\frac{x^{2}}{2 i a}}=e^{-i \frac{\pi}{4}} \sqrt{|a|}
$$

and, as one should have perhaps intuited by analyticity arguments, for either sign of a we have the same Gaussian integral formula

$$
\frac{1}{\sqrt{2 \pi}} \int_{-\infty}^{\infty} d x e^{-\frac{x^{2}}{2 i a}}=|a|^{1 / 2} e^{i \frac{\pi}{4} \frac{a}{|a|}}=\sqrt{i a}
$$

The vanishing of the $C^{\prime}$ contour segment (N.33) can be proven as follows: Substitute $z=R e^{i \phi}$ into the integral

$$
I_{R}=\int_{0}^{\frac{\pi}{4}} e^{i R^{2} e^{i 2 \phi} / 2 a} R e^{i \phi} i d \phi=\int_{0}^{\frac{\pi}{4}} e^{i R^{2}(\cos 2 \phi+i \sin 2 \phi) / 2 a} R e^{i \phi} i d \phi .
$$

Then

$$
\left|I_{R}\right| \leq R \int_{0}^{\frac{\pi}{4}} e^{-R^{2} \sin 2 \phi / 2 a} d \phi=\frac{R}{2} \int_{0}^{\frac{\pi}{2}} e^{-R^{2} \sin \theta / 2 a} d \theta
$$

In the range $[0, \pi / 2]$ we can replace $\frac{2}{\pi} \theta \leq \sin \theta$, obtain a bound

$$
\left|I_{R}\right| \leq \frac{R}{2} \int_{0}^{\frac{\pi}{2}} e^{-R^{2} \theta / \pi a} d \theta=\frac{R}{2} \frac{1-e^{-R^{2} / 2 a}}{R^{2} / a \pi}
$$

so

$$
\lim _{R \rightarrow \infty}\left|I_{R}\right|=0 .
$$

## Chapter 28

Solution 28.7: Free particle $R$-function. Calculate $R$ by its definition

$$
R\left(q^{\prime}, q, t\right)=\int_{0}^{t} \mathcal{L}\left(\dot{q}\left(t^{\prime}\right), q\left(t^{\prime}\right), t^{\prime}\right) d t^{\prime}
$$

where the solution of Lagrange equations of motion is substituted for $q(t)$.
a) a D-dimensional free particle:

We have
$\mathcal{L}\left(\dot{q}\left(t^{\prime}\right), q\left(t^{\prime}\right), t^{\prime}\right)=\frac{m}{2} \sum_{i=1}^{D}\left[\dot{q}_{i}\left(t^{\prime}\right)\right]^{2}$,
$\dot{q}_{i}(t)=$ const $=\frac{q_{i}^{\prime}-q_{i}}{t}$.
Answer:
$R\left(q^{\prime}, q, t\right)=\frac{m}{2} \sum_{i=1}^{D} \frac{\left[q_{i}^{\prime}-q_{i}\right]^{2}}{t}$.
b) Using symmetric gauge for vector potential and denoting the Larmor frequency by $\omega=\frac{e B}{m c}$, we have

$$
\mathcal{L}=\frac{m}{2}\left(\dot{x}^{2}+\dot{y}^{2}+\dot{z}^{2}+\omega(x \dot{y}-y \dot{x})\right)
$$

The equations of motion are
$\ddot{x}-\omega \dot{y}=0$,
$\ddot{y}+\omega \dot{x}=0$,
$\ddot{z}=0$.
To calculate the expression for the principal function we do integration by parts on $\dot{x}^{2}+\dot{y}^{2}$, and the result is

$$
R=\int \mathcal{L} d t=\frac{m}{2}\left(\left.x \dot{x}\right|_{t_{0}} ^{t}+\left.y \dot{y}\right|_{t_{0}} ^{t}+\frac{\left(z^{\prime}-z\right)^{2}}{t}+\int_{t_{0}}^{t}[x(-\ddot{x}+\omega \dot{y})+y(-\ddot{y}-\omega \dot{x})] d t\right)
$$

however terms inside the integral vanish by equations of motion. Denote $w(t)=$ $x(t)+\iota y(t)$, then the first two equations of motion are equivalent to equation in complex $w(t)$ :

$$
\ddot{w}(t)+\iota \omega \dot{w}(t)=0
$$

Solution to which is

$$
w^{\prime} \equiv w(t)=w+\frac{\dot{w}\left(1-e^{-\iota \omega t}\right)}{\iota \omega}
$$

We must reexpress velocities in $R$ in terms of time and initial and final coordinates. In terms of $\dot{w}$ we have

$$
\begin{aligned}
& \dot{w}_{0}=\frac{\omega}{2} \frac{e^{\frac{\iota \omega t}{2}}\left(w-w_{0}\right)}{\sin \left(\frac{\omega t}{2}\right)} \\
& \dot{w}=\frac{\omega}{2} \frac{e^{\frac{-\iota \omega t}{2}}\left(w-w_{0}\right)}{\sin \left(\frac{\omega t}{2}\right)}
\end{aligned}
$$

Notice that

$$
x \dot{x}+y \dot{y}=\operatorname{Re} w^{*} \dot{w}
$$

$\left.\operatorname{Re} w^{*} \dot{w}\right|_{0} ^{t}=\frac{\omega}{2 \sin \frac{\omega t}{2}}\left(\left(|w|^{2}+\left|w_{0}\right|^{2}\right) \cos \frac{\omega t}{2}-2 \operatorname{Re} w_{0} w^{*} e^{\frac{-\iota \omega}{2}}\right)$
$=\frac{\omega}{2}\left(\cot \left(\frac{\omega t}{2}\right)\left[\left(x-x_{0}\right)^{2}+\left(y-y_{0}\right)^{2}\right]+2\left(x_{0} y-y_{0} x\right)\right)$
$R=\frac{m\left(z-z_{0}\right)^{2}}{2 t}+\frac{m \omega}{4}\left(\cot \left(\frac{\omega t}{2}\right)\left[\left(x-x_{0}\right)^{2}+\left(y-y_{0}\right)^{2}\right]+2\left(x_{0} y-y_{0} x\right)\right)$
c)

Solution 28.2: Dirac delta function, Gaussian representation. To prove that $\delta_{\sigma}$ converges to a dirac delta function, it is enough to show that it has the following properties:

1. $\int_{-\infty}^{\infty} \delta_{\sigma}(x) d x=1$
2. $\lim _{\sigma \rightarrow 0} \int_{-a}^{a} f(x) \delta_{\sigma}(x) d x=f(0)$
for arbitrary $f(x)$ continuous and positive $a$.

First property is satisfied by the choise of normalisation constant.
Second property is verified by the change of variables $y=x / \sqrt{2 \sigma^{2}}$ :

$$
\lim _{\sigma \rightarrow 0} \int_{-a}^{a} f(x) \delta_{\sigma}(x) d x=\lim _{\sigma \rightarrow 0} \frac{1}{\sqrt{\pi}} \int_{\frac{-a}{\sqrt{2 \sigma^{2}}}}^{\frac{a}{\sqrt{\sigma^{2}}}} f\left(\sqrt{2 \sigma^{2}} y\right) e^{-y^{2}} d y=f(0)
$$

## Solution 28.3: d-dimensional Gaussian integrals.

We require that the matrix in the exponent is nondegenerate (i.e. has no zero eigenvalues.) The converse may happen when doing stationary phase approximations which requires going beyond the Gaussian saddlepoint approximation, typically to the Airy-function type stationary points [27.10]. We also assume that $M$ is positivedefinite, otherwise the integral is infinite.

Make a change of variables $y=A x$ such that $A^{T} M^{-1} A=\mathrm{Id}$. Then

$$
I=\frac{1}{(2 \pi)^{d / 2}} \int_{\mathbf{R}^{d}} \exp \left[-\frac{1}{2} \sum_{i}\left(y_{i}^{2}-2(J A)_{i} y_{i}\right)\right]|\operatorname{det} A| d y
$$

Complete each term under in the sum in the exponent to a full square

$$
y_{i}^{2}-2(J A)_{i} y_{i}=\left(y_{i}-(J A)_{i}\right)^{2}-(J A)_{i}^{2}
$$

and shift the origin of integration to $J A / 2$, so that

$$
I=\frac{1}{(2 \pi)^{d / 2}} \exp \left(\frac{1}{2} J^{T} A A^{T} J\right)|\operatorname{det} A| \int_{\mathbf{R}^{d}} \exp \left[-\frac{1}{2} \sum_{i} y_{i}^{2}\right] d y
$$

Note that $A A^{T} M^{-1} A A^{T}=A A^{T}$, therefore $A A^{T}=M$ and $|\operatorname{det} A|=\sqrt{\operatorname{det} M}$. The remaining integral is equal to a Poisson integral raised to the d-th power, i.e. $(2 \pi)^{d / 2}$. Answer:

$$
I=\sqrt{\operatorname{det} M} \exp \left[\frac{1}{2} J^{T} M J\right]
$$

## Solution 28.4: Stationary phase approximation.

Main contribution to this integral come from critical points of $\Phi(x)$. Suppose that $p$ is such a nondegenerate critical point, $p: D \Phi(p)=0$, and $D^{2} \Phi(p)$ has full rank. Then there is a local coordinate system $y$ in the neighbourhood of $p$ such that $\Phi(p+y)=\Phi(p)-\sum_{i=1}^{\lambda} y_{i}^{2}+\sum_{i=\lambda+1}^{d} y_{i}^{2}$, where $\lambda$ is the number of negative eigenvalues of $D^{2} \Phi(p)$. Indeed, if we set $x-p=A y$, then $\Phi(x) \approx \Phi(p)+\frac{1}{2} y A^{T} D^{2} \Phi(p) A y$. There exist such $A$ that $\frac{1}{2} A^{T} D^{2} \Phi(p) A=\operatorname{diag}[\underbrace{-1, \ldots-1}_{\lambda}, \underbrace{+1, \ldots}_{d-\lambda}]$. With this change of variables in mind, we have

$$
I=e^{\frac{\iota \Phi(p)}{\hbar}} \int_{\mathrm{R}^{d}} e^{\frac{\iota}{\hbar}\left(-\sum_{i=1}^{\lambda} y_{i}^{2}+\sum_{i=\lambda+1}^{d} y_{i}^{2}\right)}|\operatorname{det} A| d y=e^{\frac{\iota \Phi(p)}{\hbar}}(\pi \hbar)^{d / 2} e^{\frac{\iota \pi}{4}(-2 \lambda+d)}|\operatorname{det} A|
$$

Furthermore, $(\operatorname{det} A)^{2} \operatorname{det} D^{2} \Phi(p)=2^{d} \exp \iota \pi \lambda$, therefore

$$
|\operatorname{det} A|=\frac{2^{d / 2} \exp \frac{\iota \pi \lambda}{2}}{\sqrt{\operatorname{det} D^{2} \Phi(p)}} .
$$

Phase factors $\exp \iota \pi \lambda / 2$ and $\exp -\iota \pi \lambda / 2$ cancel out. Substitute $\exp \iota \pi d / 2=\iota^{d / 2}$.
The result:

$$
I=\frac{(2 \iota \pi \hbar)^{d / 2} e^{\frac{\iota \Phi(p)}{\hbar}}}{\sqrt{\operatorname{det} D^{2} \Phi(p)}}
$$

Critical nondegenerate points are isolated. Therefore if $\Phi$ has more than one critical point, then equivalent local approximation can be made in the neighbourhoods of each critical point and the complete approximation to the integral made by adding contributions of all critical points.

Answer:

$$
I=\sum_{p: D \Phi(p)=0} \frac{(2 \iota \pi \hbar)^{d / 2} e^{\frac{\iota \Phi(p)}{h}} A(p)}{\sqrt{\operatorname{det} D^{2} \Phi(p)}}
$$

Rytis Paškauskas

## Solution 28.4: Stationary phase approximation.

values of $x$ of stationary phase, the points for which the gradient of the phase vanishes

$$
\frac{\partial}{\partial x} \Phi(x)=0
$$

Intuitively, these are the important contributions as for $\hbar \rightarrow 0$ the phase $\Phi(x) / \hbar$ grows large and the function $e^{i \Phi(x) / \hbar}$ oscillates rapidly as a function of $x$, with the negative
and positive parts cancelling each other. More precisely, if the stationary points are well separated local extrema of $\Phi(x)$, we can deform the integration contour and approximate $\Phi(x) / \hbar$ up to the second order in $x$ by

$$
I \approx \sum_{n} A\left(x_{n}\right) e^{i \Phi\left(x_{n}\right) / \hbar} \int d^{d} x e^{\frac{i}{2 \hbar}\left(x-x_{n}\right)^{T} \mathbf{D}^{2} \Phi\left(x_{n}\right)\left(x-x_{n}\right)} .
$$

The second derivative matrix is a real symmetric matrix, so we can transform it to a diagonal matrix by a similarity transformation

$$
\operatorname{Diag}\left(\lambda_{1}, \ldots, \lambda_{d}\right)=\mathbf{O D}^{2} \Phi \mathbf{O}^{+}
$$

where $\mathbf{O}$ is a matrix of an orthogonal transfomation. In the rotated coordinate system $u=\mathbf{O}\left(x-x_{n}\right)$ and the integral takes form

$$
I \approx \sum_{n} A\left(x_{n}\right) e^{i \Phi\left(x_{n}\right) / \hbar} \int d^{d} u e^{\sum_{k=1}^{d} i \lambda_{k} u_{k}^{2} / 2 \hbar}
$$

where we used the fact that the Jacobi determinant of an orthogonal transformation is $\operatorname{det} \mathbf{O}=1$. Carrying out the Gauss integrals

$$
\begin{equation*}
\int d u e^{i \lambda u^{2} / 2 \hbar}=\frac{(2 \pi i \hbar)^{1 / 2}}{\sqrt{\lambda}} \tag{N.34}
\end{equation*}
$$

and using $\operatorname{det} \mathbf{D}^{2} \Phi\left(x_{n}\right)=\prod_{k=1}^{d} \lambda_{k}$ we obtain the stationary phase estimate of (28.54).
A nice exposition of the subject is given in ref. [27.10].
Solution 30.2: Stationary phase approximation in higher dimensions. In this case $1 / \hbar$ is assumed to be a very large number. parameter. The idea of this method is that we only evaluate part of integral $I$ where $e^{i \varphi}$ is stationary that is, $\varphi \approx$ const. That means we need extrema (saddle points) of manifold $\Phi$. In this case

$$
\frac{\partial \Phi}{\partial x_{s p, \mu}}=0
$$

Introduce a new d-dimensional variable $s$ such, that

$$
i \Phi(x)=i \Phi\left(x_{s p, \mu}\right)-s^{2}
$$

Integral I in terms of new variables is

$$
I=\sum_{n} e^{i \Phi\left(x_{n}\right) / \hbar} \int e^{-s^{2} / \hbar} A\left(x_{n}(s)\right)\left|\frac{D x}{D s}\right| d^{d} s
$$

Here $n$ sums all stationary phase points which the path of integration (in complex plane!) meets. next, we need to calculate the Jacobian J:

$$
J=1 /\left|\frac{\partial s_{i}}{\partial x_{k}}\right|
$$

where

$$
\frac{\partial s_{i}}{\partial x_{k}}=\frac{1}{2 i s_{i}} \frac{\partial \Phi}{\partial x_{k}}
$$

This expression is undetermined at stationary phase points, because its right hand side becomes division zero by zero. However, by the chain rule

$$
\frac{\partial s_{i}}{\partial x_{k}}=\frac{1}{2 i \frac{\partial s_{i}}{\partial x_{m}}} \frac{\partial \Phi^{2}}{\partial x_{k} \partial x_{m}}
$$

where $x=x_{s p}$ are evaluated at the stationary phase point. From this expression we obtain that

$$
\left[\left(\frac{\partial s}{\partial x}\right)^{2}\right]_{i, k}=\frac{1}{2 i} \frac{\partial \Phi^{2}}{\partial x_{i} \partial x_{k}}
$$

So the Jacobian is (employing a standard notation for a second derivative)

$$
J=\frac{(2 i)^{d / 2}}{\sqrt{\operatorname{det} D^{2} \Phi}}
$$

Since the exponential factor $e^{-s^{2} / \hbar}$ cuts integration sharply because of a very large parameter $1 / \hbar$, the function is evaluated only at the stationary point $s=0$, and the integral is approximately

$$
I \approx \sum_{n} e^{i \Phi\left(x_{n}\right) / \hbar} A\left(x_{n}\right) \frac{(2 i)^{d / 2}}{\sqrt{\operatorname{det} D^{2} \Phi\left(x_{n}\right)}} \int e^{-s^{2} / \hbar} d^{d} s
$$

Limits of integration may depend on particular situation. If limits are infinite, then

$$
\int e^{-s^{2} / \hbar} d^{d} s=\left(\int_{-\infty}^{\infty} e^{-s^{2} / \hbar} d s\right)=(\pi \hbar)^{d / 2}
$$

We substitute this into $I$ and get the answer.
(Rytis Paškauskas)
Solution 28.12: $D$-dimensional free particle propagator. A free particle reachs $q$ from $q^{\prime}$ by only one trajectory. Taking this into account the semiclassical Van Vleck propagator is

$$
K_{s c}\left(q, q^{\prime}, t\right)=\frac{e^{\frac{i R}{\hbar}}}{(2 \pi i \hbar)^{d / 2}}\left|\operatorname{det} \frac{\partial^{2} R}{\partial q_{i} \partial q_{j}^{\prime}}\right|^{1 / 2}
$$

The principal function of free motion in D-dimensions is

$$
R\left(q, q^{\prime}, t\right)=\frac{m}{2 t} \sum_{\mu=1}^{D}\left(q_{\mu}-q_{\mu}^{\prime}\right)^{2}
$$

The derivative is

$$
\frac{\partial^{2} R}{\partial q_{i} \partial q_{j}^{\prime}}=-\delta_{i, j} \frac{m}{t}
$$

According to that determinant is

$$
\left|\operatorname{det} \frac{\partial^{2} R}{\partial q_{i} \partial q_{j}^{\prime}}\right|^{1 / 2}=e^{i \pi D / 2}\left(\frac{m}{t}\right)^{D / 2}
$$

and the Van Vleck propagator is

$$
K_{s c}\left(q, q^{\prime}, t\right)=e^{i \pi D / 4}\left(\frac{m}{2 \pi \hbar t}\right)^{D / 2} \prod_{\mu=1}^{D} \exp \left[\frac{i m}{2 \hbar t}\left(q_{\mu}-q_{\mu}^{\prime}\right)^{2}\right]
$$

The next step is to calculate the exact quantum propagator:

$$
K\left(q, q^{\prime}, t\right)=\sum_{n} \phi_{n}(q) e^{-i E_{n} t / \hbar} \phi_{n}^{*}\left(q^{\prime}\right)
$$

Taking that particle wave function in free space is

$$
\phi_{p}(q)=\frac{1}{(2 \pi \hbar)^{D / 2}} e^{i p q / \hbar}
$$

we derive that propagator $K$ is

$$
\frac{1}{(2 \pi \hbar)^{D}} \int e^{-\frac{i t}{2 m \hbar} p^{2}+i p\left(q-q^{\prime}\right)} d^{D} p
$$

We can split multi-dimensional integral that stands here into a product of one dimensional integrals. Then we should change variables for purpose of reduction to Poisson-type integrals. We have omitted some straightforward algebra. The result is that the semiclassical Van Vleck propagator and the exact quantum propagator are identical:

$$
K\left(q, q^{\prime}, t\right)=e^{i \pi D / 4}\left(\frac{m}{2 \pi \hbar t}\right)^{D / 2} \prod_{\mu=1}^{D} \exp \left[\frac{i m}{2 \hbar t}\left(q_{\mu}-q_{\mu}^{\prime}\right)^{2}\right]=K_{s c}\left(q, q^{\prime}, t\right)
$$

This result could have been anticipated because approximate formula (??.37) becomes exact for the free particle Lagrangian.

## Chapter 30

Solution 30.1: Monodromy matrix from second variations of the action. If we take two points in the configuration space $q$ and $q^{\prime}$ connected with a trajectory with energy $E$ and vary them in such a way that the variation of their initial and final points are transverse to the velocity of the orbit in that point, we can write the variations of the initial and final momenta as

$$
\begin{equation*}
\delta p_{\perp i}=\frac{\partial^{2} S\left(q, q^{\prime}, E\right)}{\partial q_{\perp i} \partial q_{\perp k}} \delta q_{\perp k}+\frac{\partial^{2} S\left(q, q^{\prime}, E\right)}{\partial q_{\perp i} \partial q_{\perp k}^{\prime}} \delta q_{\perp k}^{\prime} \tag{N.35}
\end{equation*}
$$

and

$$
\begin{equation*}
\delta p_{\perp i}^{\prime}=-\frac{\partial^{2} S\left(q, q^{\prime}, E\right)}{\partial q_{\perp i}^{\prime} \partial q_{\perp k}} \delta q_{\perp k}-\frac{\partial^{2} S\left(q, q^{\prime}, E\right)}{\partial q_{\perp i}^{\prime} \partial q_{\perp k}^{\prime}} \delta q_{\perp k}^{\prime} \tag{N.36}
\end{equation*}
$$

Next we express the variations of the final momenta and coordinates in terms of the initial ones. In the obvious shorthand we can write (N.36) as

$$
\delta q_{\perp}=-S_{q^{\prime} q}^{-1} S_{q^{\prime} q^{\prime}} \delta q_{\perp}^{\prime}-S_{q^{\prime} q}^{-1} \delta p_{\perp}^{\prime}
$$

From (N.35) it then follows that

$$
\begin{equation*}
\delta p_{\perp}=\left(S_{q q^{\prime}}-S_{q q} S_{q^{\prime} q}^{-1} S_{q^{\prime} q^{\prime}}\right) \delta q_{\perp}^{\prime}-S_{q q} S_{q^{\prime} q}^{-1} \delta p_{\perp}^{\prime} \tag{N.37}
\end{equation*}
$$

These relations remain valid in the $q^{\prime} \rightarrow q$ limit, with $q$ on the periodic orbit, and can also be expressed in terms of the monodromy matrix of the periodic orbit. The monodromy matrix for a surface of section transverse to the orbit within the constant energy $E=H(q, p)$ shell is

$$
\begin{align*}
\delta q_{\perp} & =\mathbf{J}_{q q} \delta q_{\perp}^{\prime}+\mathbf{J}_{q p} \delta p_{\perp}^{\prime} \\
\delta p_{\perp} & =\mathbf{J}_{p q} \delta q_{\perp}^{\prime}+\mathbf{J}_{p p} \delta p_{\perp}^{\prime} \tag{N.38}
\end{align*}
$$

In terms of the second derivatives of the action the monodromy matrix is

$$
\begin{aligned}
\mathbf{J}_{q q}=-S_{q^{\prime} q}^{-1} S_{q^{\prime} q^{\prime}}, & \mathbf{J}_{q p}=-S_{q^{\prime} q}^{-1} \\
\mathbf{J}_{p q}=\left(S_{q q^{\prime}}-S_{q q} S_{q^{\prime} q}^{-1} S_{q^{\prime} q^{\prime}}\right), & \mathbf{J}_{p p}=-S_{q q} S_{q^{\prime} q}^{-1}
\end{aligned}
$$

and vice versa

$$
\begin{aligned}
S_{q q}=\mathbf{J}_{p p} \mathbf{J}_{q p}^{-1}, & S_{q q^{\prime}}=\mathbf{J}_{p q}-\mathbf{J}_{p p} \mathbf{J}_{q p}^{-1} \mathbf{J}_{q q} \\
S_{q^{\prime} q}=-\mathbf{J}_{q p}^{-1}, & S_{q^{\prime} q^{\prime}}=-\mathbf{J}_{q p}^{-1} \mathbf{J}_{q q}
\end{aligned}
$$

Now do exercise 30.2.

Solution 30.2: Jacobi gymnastics. We express the Jacobi matrix elements in $\operatorname{det}(\mathbf{1}-\mathbf{J})$ with the derivative matrices of $S$

$$
\operatorname{det}(1-J)=\operatorname{det}\left(\begin{array}{cc}
I+S_{q^{\prime} q}^{-1} S_{q^{\prime} q^{\prime}} & S_{q^{\prime} q}^{-1} \\
-S_{q q^{\prime}}+S_{q q} S_{q^{\prime} q}^{-1} S_{q^{\prime} q^{\prime}} & I+S_{q q} S_{q^{\prime} q}^{-1}
\end{array}\right) .
$$

We can multiply the second column with $S_{q^{\prime} q^{\prime}}$ from the and substract from the first column, leaving the determinant unchanged

$$
\operatorname{det}(1-J)=\operatorname{det}\left(\begin{array}{cc}
I & S_{q^{\prime} q}^{-1} \\
-S_{q q^{\prime}}-S_{q^{\prime} q^{\prime}} & I+S_{q q} S_{q^{\prime} q}^{-1}
\end{array}\right) .
$$

Then, we multiply the second column with $S_{q^{\prime} q}$ from the right and compensate this by dividing the determinant with $\operatorname{det} S_{q^{\prime} q}$

$$
\operatorname{det}(1-J)=\operatorname{det}\left(\begin{array}{cc}
I & I \\
-S_{q q^{\prime}}-S_{q^{\prime} q^{\prime}} & S_{q^{\prime} q}+S_{q q}
\end{array}\right) / \operatorname{det} S_{q^{\prime} q} .
$$

Finally we subtract the first column from the second one

$$
\left.\operatorname{det}\left(1-J_{j}\right)\right)=\operatorname{det}\left(\begin{array}{cc}
I & 0 \\
S_{q q^{\prime}}+S_{q^{\prime} q^{\prime}} & S_{q q^{\prime}}+S_{q^{\prime} q^{\prime}}+S_{q^{\prime} q}+S_{q q}
\end{array}\right) / \operatorname{det} S_{q^{\prime} q}
$$

The last determinant can now be evaluated and yields the desired result (30.2)

$$
\operatorname{det}\left(1-J_{j}\right)=\operatorname{det}\left(S_{q q^{\prime}}+S_{q^{\prime} q^{\prime}}+S_{q^{\prime} q}+S_{q q}\right) / \operatorname{det} S_{q^{\prime} q}
$$

## Chapter 31

Solution 31.1: Evaluation of cycles by minimization. To start with a guess path where each bounce is given some arbitrary position on the correct disk and then iteratively improve on the guess. To accomplish this an improvement cycle is constructed whereby each bouncing point in the orbit is taken in turn and placed in a new position so that it minimizes the path. Since the positions of all the other bounces are kept constant this involves choosing the new bounce position which minimizes the path from the previous bounce to the next bounce. This problem is schematically represented in figure N. 7

Finding the point $B$ involves a one dimensional minimization. We define the vectors $\vec{A}=\overrightarrow{O A}, \vec{B}=\overrightarrow{O B}$ and $\vec{C}=\overrightarrow{O C}$. We wish to minimize the length $L_{A B C}$ by varying $\vec{B}$ subject to the constraint that $|\vec{B}|=a$. Clearly

$$
\begin{aligned}
L_{A B C} & =|\vec{A}-\vec{B}|+|\vec{C}-\vec{B}| \\
& =\sqrt{\vec{A}^{2}+\vec{B}^{2}-2 \vec{A} \cdot \vec{B}}+\sqrt{\vec{C}^{2}+\vec{B}^{2}-2 \vec{C} \cdot \vec{B}}
\end{aligned}
$$



Figure N.7: Minimizing the path from the previous bounce to the next bounce.
writing

$$
\vec{B}(\theta)=a(\cos \theta, \sin \theta)
$$

then the minima is given by

$$
\frac{d L_{A B C}}{d \theta}=-\left(\frac{\vec{A}}{\sqrt{\vec{A}^{2}+\vec{B}^{2}-2 \vec{A} \cdot \vec{B}}}+\frac{\vec{C}}{\sqrt{\vec{C}^{2}+\vec{B}^{2}-2 \vec{C} \cdot \vec{B}}}\right) \cdot \vec{B}^{\prime}(\theta)=0 .
$$

The minima can then be found using a bisection algorithm or using Newton-Raphson. A simpler way is to observe that $\vec{B}^{\prime}(\theta)$ is orthogonal to $\vec{B}(\theta)$ so that the vector

$$
\vec{D}=\frac{\vec{A}}{\sqrt{\vec{A}^{2}+\vec{B}^{2}-2 \vec{A} \cdot \vec{B}}}+\frac{\vec{C}}{\sqrt{\vec{C}^{2}+\vec{B}^{2}-2 \vec{C} \cdot \vec{B}}}
$$

will be proportional to $\vec{B}$. This then provides an iterative sequence for finding $\vec{B}$

- Starting from your current guess for $\vec{B}$ calculate $\vec{D}$
- Put $\vec{B}=a \vec{D} /|\vec{D}|$
- Repeat the first step until you converge.

At each iteration of the improvement cycle the total length of the orbit is measured. The minimization is complete when the path length stops improving. Although this algorithm is not as fast as the Newton-Raphson method, it nevertheless converges very rapidly.
(Adam Prügel-Bennet)

## Chapter 32

## Solution 32.2: The one-disk scattering wave function.

$$
\begin{equation*}
\psi(\vec{r})=\frac{1}{2} \sum_{m=-\infty}^{\infty}\left(H_{m}^{(2)}(k r)-\frac{H_{m}^{(2)}(k a)}{H_{m}^{(1)}(k a)} H_{m}^{(1)}(k r)\right) e^{i m\left(\Phi_{r}-\Phi_{k}\right)} \tag{N.39}
\end{equation*}
$$

(For $r<a, \psi(\vec{r})=0$ of course.)

## Chapter 34

(No solutions available.)

## Chapter 35

(No solutions available.)

## Chapter D

(No solutions available.)

## Chapter E

(No solutions available.)

## Chapter F

Solution F.1: Lefschetz zeta function. Starting with dynamical zeta function ref. [13.11] develops the Atiyah-Bott-Lefschetz fixed point formula and relates is to Weyl characters. Might be worth learning.

## Chapter H

Solution H.1: Using the multiplicative property of the Jacobi matrix we can write

$$
\Lambda^{t^{\prime}+t}\left(x_{0}, \mathbf{u}_{0}\right)=\left\|\mathbf{J}^{t^{\prime}+t}\left(x_{0}\right) \mathbf{u}_{0}\right\|=\left\|\mathbf{J}^{t^{\prime}}(x(t)) \mathbf{J}^{t}\left(x_{0}\right) \mathbf{u}_{0}\right\|
$$

We can introduce the time evolved unit vector

$$
\mathbf{u}(t)=\mathbf{J}^{t}\left(x_{0}\right) \mathbf{u}_{0} /\left\|\boldsymbol{J}^{t}\left(x_{0}\right) \mathbf{u}_{0}\right\| .
$$

Then

$$
\left\|\mathbf{J}^{t^{\prime}}(x(t)) \mathbf{J}^{t}\left(x_{0}\right) \mathbf{u}_{0}\right\|=\left\|\mathbf{J}^{t^{\prime}}(x(t)) \mathbf{u}(t)\right\|\left\|\mathbf{J}^{t}\left(x_{0}\right) \mathbf{u}_{0}\right\|
$$

which is the desired result.
We have to adjoin the tangent space, since the stretching factor depends on $\mathbf{u}$ and not just on $x$. The stretching factor is multiplicative along the entire trajectory $(x(t), \mathbf{u}(t))$. However, it is not multiplicative along the phase space trajectory $x(t)$ with a fixed $\mathbf{u}$.

Solution H.2: $\quad$ If $b=a^{2}$ and $T_{b}=2 T_{a}$ we can introduce the variable $y=e^{s T_{a}}$. The dynamo rate equation then reads

$$
0=1-x+x^{2}
$$

The solutions of this are $x_{ \pm}=(1 \pm i \sqrt{3}) / 2$. The dynamo rate is then a complex cojugate pair $\nu=\log x_{ \pm} / T_{a}$.

The escape rate equation is

$$
0=1-x / a-x^{2} / a^{2}
$$

The solutions are $x_{ \pm}=a(-1 \pm \sqrt{5}) / 2$. The escape rate is $\gamma=\log \left(x_{+}\right) / T_{a}$.
In the reverse case the escape rate remains unchanged, while the dynamo rate becomes $\nu=\log ((\sqrt{5}+1) / 2) / T_{a}$. In this case the advected field grows with an exponential rate. In the previous case it shows oscillations in addition to the exponential growth due to the imaginary part of the rate.

## Chapter L

(No solutions available.)

## Appendix O

## Projects

You are urged to try to work through the essential steps in a project that combines the techniques learned in the course with some application of interest to you for other reasons. It is OK to share computer programs and such, but otherwise each project should be distinct, not a group project. The essential steps are:

## - Dynamics

1. construct a symbolic dynamics
2. count prime cycles
3. prune inadmissible itineraries, construct Markov graphs if appropriate
4. implement a numerical simulator for your problem
5. compute a set of the shortest periodic orbits
6. compute cycle stabilities

## - Averaging, numerical

1. estimate by numerical simulation some observable quantity, like the escape rate,
2. or check the flow conservation, compute something like the Lyapunov exponent

- Averaging, periodic orbits

1. implement the appropriate cycle expansions
2. check flow conservation as function of cycle length truncation, if the system is closed
3. implement desymmetrization, factorization of zeta functions, if dynamics possesses a discrete symmetry
4. compute a quantity like the escape rate as a leading zero of a spectral determinant or a dynamical zeta function.
5. or evaluate a sequence of truncated cycle expansions for averages, such as the Lyapunov exponent or/and diffusion coefficients
6. compute a physically intersting quantity, such as the conductance
7. compute some number of the classical and/or quantum eigenvalues, if appropriate

## O. 1 Deterministic diffusion, zig-zag map

To illustrate the main idea of chapter 23 , tracking of a globally diffusing orbit by the associated confined orbit restricted to the fundamental cell, we consider a class of simple 1-d dynamical systems, chains of piecewise linear maps, where all transport coefficients can be evaluated analytically. The translational symmetry (23.10) relates the unbounded dynamics on the real line to the dynamics restricted to a "fundamental cell" - in the present example the unit interval curled up into a circle. An example of such map is the sawtooth map

$$
\hat{f}(x)= \begin{cases}\Lambda x & x \in[0,1 / 4+1 / 4 \Lambda]  \tag{O.1}\\ -\Lambda x+(\Lambda+1) / 2 & x \in[1 / 4+1 / 4 \Lambda, 3 / 4-1 / 4 \Lambda] \\ \Lambda x+(1-\Lambda) & x \in[3 / 4-1 / 4 \Lambda, 1]\end{cases}
$$

The corresponding circle map $f(x)$ is obtained by modulo the integer part. The elementary cell map $f(x)$ is sketched in figure O.1. The map has the symmetry property

$$
\begin{equation*}
\hat{f}(\hat{x})=-\hat{f}(-\hat{x}), \tag{O.2}
\end{equation*}
$$

so that the dynamics has no drift, and all odd derivatives of the generating function (23.3) with respect to $\beta$ evaluated at $\beta=0$ vanish.

The cycle weights are given by

$$
\begin{equation*}
t_{p}=z^{n_{p}} \frac{e^{\beta \hat{n}_{p}}}{\left|\Lambda_{p}\right|} . \tag{O.3}
\end{equation*}
$$

The diffusion constant formula for $1-d$ maps is

$$
\begin{equation*}
D=\frac{1}{2} \frac{\left\langle\hat{n}^{2}\right\rangle_{\zeta}}{\langle n\rangle_{\zeta}} \tag{O.4}
\end{equation*}
$$

where the "mean cycle time" is given by

$$
\begin{equation*}
\langle n\rangle_{\zeta}=\left.z \frac{\partial}{\partial z} \frac{1}{\zeta(0, z)}\right|_{z=1}=-\sum^{\prime}(-1)^{k} \frac{n_{p_{1}}+\cdots+n_{p_{k}}}{\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right|} \tag{O.5}
\end{equation*}
$$

the mean cycle displacement squared by

$$
\begin{equation*}
\left\langle\hat{n}^{2}\right\rangle_{\zeta}=\left.\frac{\partial^{2}}{\partial \beta^{2}} \frac{1}{\zeta(\beta, 1)}\right|_{\beta=0}=-\sum^{\prime}(-1)^{k} \frac{\left(\hat{n}_{p_{1}}+\cdots+\hat{n}_{p_{k}}\right)^{2}}{\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right|} \tag{O.6}
\end{equation*}
$$



Figure O.1: (a)-(f) The sawtooth map (0.1) for the 6 values of parameter $a$ for which the folding point of the map aligns with the endpoint of one of the 7 intervals and yields a finite Markov partition (from ref. [0.1]). The corresponding Markov graphs are given in figure O.2.
and the sum is over all distinct non-repeating combinations of prime cycles. Most of results expected in this projects require no more than pencil and paper computations.

Implementing the symmetry factorization (23.35) is convenient, but not essential for this project, so if you find sect. 22.1.2 too long a read, skip the symmetrization.

## O.1.1 The full shift

Take the map (O.1) and extend it to the real line. As in example of figure 23.3, denote by $a$ the critical value of the map (the maximum height in the unit cell)

$$
\begin{equation*}
a=\hat{f}\left(\frac{1}{4}+\frac{1}{4 \Lambda}\right)=\frac{\Lambda+1}{4} . \tag{O.7}
\end{equation*}
$$

Describe the symbolic dynamics that you obtain when $a$ is an integer, and derive the formula for the diffusion constant:

$$
\begin{equation*}
D=\frac{\left(\Lambda^{2}-1\right)(\Lambda-3)}{96 \Lambda} \quad \text { for } \Lambda=4 a-1, a \in \mathbb{Z} \tag{0.8}
\end{equation*}
$$

If you are going strong, derive also the fromula for the half-integer $a=$ $(2 k+1) / 2, \Lambda=4 a+1$ case and email it to DasBuch@nbi.dk. You will need to partition $\mathcal{M}_{2}$ into the left and right half, $\mathcal{M}_{2}=\mathcal{M}_{8} \cup \mathcal{M}_{9}$, as in the derivation of (23.21).

## O.1.2 Subshifts of finite type

We now work out an example when the partition is Markov, although the slope is not an integer number. The key step is that of having a partition where intervals are mapped onto unions of intervals. Consider for example the case in which $\Lambda=4 a-1$, where $1 \leq a \leq 2$. A first partition is constructed from seven intervals, which we label $\left\{\mathcal{M}_{1}, \mathcal{M}_{4}, \mathcal{M}_{5}, \mathcal{M}_{2}, \mathcal{M}_{6}, \mathcal{M}_{7}, \mathcal{M}_{3}\right\}$, with the alphabet ordered as the intervals are laid out along the unit interval. In general the critical value $a$ will not correspond to an interval border, but now we choose $a$ such that the critical point is mapped onto the right border of $\mathcal{M}_{1}$, as in figure O.1(a). The critical value of $f()$ is $f\left(\frac{\Lambda+1}{4 \Lambda}\right)=a-1=(\Lambda-3) / 4$. Equating this with the right border of $\mathcal{M}_{1}$, $x=1 / \Lambda$, we obtain a quadratic equation with the expanding solution $\Lambda=4$. We have that $f\left(\mathcal{M}_{4}\right)=f\left(\mathcal{M}_{5}\right)=\mathcal{M}_{1}$, so the transition matrix (11.2) is


Figure O.2: (a) The sawtooth map (0.1) partition tree for figure O.1(a); while intervals $\mathcal{M}_{1}, \mathcal{M}_{2}, \mathcal{M}_{3}$ map onto the whole unit interval, $f\left(\mathcal{M}_{1}\right)=f\left(\mathcal{M}_{2}\right)=f\left(\mathcal{M}_{3}\right)=$ $\mathcal{M}$, intervals $\mathcal{M}_{4}, \mathcal{M}_{5}$ map onto $\mathcal{M}_{1}$ only, $f\left(\mathcal{M}_{4}\right)=f\left(\mathcal{M}_{5}\right)=\mathcal{M}_{1}$, and similarly for intervals $\mathcal{M}_{6}, \mathcal{M}_{7}$. An initial point starting out in the interval $\mathcal{M}_{1}, \mathcal{M}_{2}$ or $\mathcal{M}_{3}$ can land anywhere on the unit interval, so the subtrees originating from the corresponding nodes on the partition three are similar to the whole tree and can be identified (as, for example, in figure 11.14), yielding (b) the Markov graph for the Markov partition of figure O.1(a). (c) the Markov graph in the compact notation of (23.26).
given by

$$
\phi^{\prime}=T \phi=\left(\begin{array}{ccccccc}
1 & 1 & 1 & 1 & 0 & 0 & 1  \tag{O.9}\\
1 & 0 & 0 & 1 & 0 & 0 & 1 \\
1 & 0 & 0 & 1 & 0 & 0 & 1 \\
1 & 0 & 0 & 1 & 0 & 0 & 1 \\
1 & 0 & 0 & 1 & 0 & 0 & 1 \\
1 & 0 & 0 & 1 & 1 & 1 & 1
\end{array}\right)\left(\begin{array}{l}
\phi_{1} \\
\phi_{4} \\
\phi_{5} \\
\phi_{2} \\
\phi_{6} \\
\phi_{7} \\
\phi_{3}
\end{array}\right)
$$

and the dynamics is unrestricted in the alphabet

$$
\{1, \underline{41}, \underline{51}, 2, \underline{63}, \underline{73}, 3,\} .
$$

One could diagonalize (O.9) on the computer, but, as we saw in sect. 11.5, the Markov graph figure O.2(b) corresponding to figure O.1(a) offers more insight into the dynamics. The dynamical zeta function

$$
\begin{align*}
1 / \zeta & =1-\left(t_{1}+t_{2}+t_{3}\right)-2\left(t_{14}+t_{37}\right) \\
1 / \zeta & =1-3 \frac{z}{\Lambda}-4 \cosh \beta \frac{z^{2}}{\Lambda^{2}} \tag{O.10}
\end{align*}
$$

follows from the loop expansion (13.13) of sect. 13.3.
The material flow conservation sect. 19.2 and the symmetry factorization (23.35) yield

$$
0=\frac{1}{\zeta(0,1)}=\left(1+\frac{1}{\Lambda}\right)\left(1-\frac{4}{\Lambda}\right)
$$

| figure O.1 | $\Lambda$ | $D$ |
| :---: | :---: | :---: |
| (a) | 3 | 0 |
| (b) | $\sqrt{5}+2$ | $\frac{1}{10}$ |
| (c) | $\frac{1}{2}(\sqrt{17}+5)$ | $\frac{1}{2 \sqrt{5}}$ |
| (c') | 5 | $\frac{2}{5}$ |
| (d) | $\frac{1}{2}(\sqrt{33}+5)$ | $\frac{1}{8}+\frac{5}{88} \sqrt{33}$ |
| (e) | $2 \sqrt{2}+3$ | $\frac{1}{2 \sqrt{2}}$ |
| (f) | $\frac{1}{2}(\sqrt{33}+7)$ | $\frac{1}{4}+\frac{1}{4 \sqrt{33}}$ |
|  | 7 | $\frac{2}{7}$ |

Table O.1: The diffusion constant as function of the slope $\Lambda$ for the $a=1,2$ values of ( 0.8 ) and the 6 Markov partitions of figure 0.1
which indeed is satisfied by the given value of $\Lambda$. Conversely, we can use the desired Markov partition topology to write down the corresponding dynamical zeta function, and use the $1 / \zeta(0,1)=0$ condition to fix $\Lambda$. For more complicated transition matrices the factorization (23.35) is very helpful in reducing the order of the polynomial condition that fixes $\Lambda$.

The diffusion constant follows from (23.36) and (O.4)

$$
\begin{aligned}
& \langle n\rangle_{\zeta}=-\left(1+\frac{1}{\Lambda}\right)\left(-\frac{4}{\Lambda}\right), \quad\left\langle\hat{n}^{2}\right\rangle_{\zeta}=\frac{4}{\Lambda^{2}} \\
& D=\frac{1}{2} \frac{1}{\Lambda+1}=\frac{1}{10}
\end{aligned}
$$

Think up other non-integer values of the parameter for which the symbolic dynamics is given in terms of Markov partitions: in particular consider the cases illustrated in figure O.1 and determine for what value of the parameter $a$ each of them is realized. Work out the Markov graph, symmetrization factorization and the diffusion constant, and check the material flow conservation for each case. Derive the diffusion constants listed in table O.1. It is not clear why the final answers tend to be so simple. Numerically, the case of figure O.1(c) appears to yield the maximal diffusion constant. Does it? Is there an argument that it should be so?

The seven cases considered here (see table O.1, figure O.1 and (O.8)) are the 7 simplest complete Markov partitions, the criterion being that the critical points map onto partition boundary points. This is, for example, what happens for unimodal tent map; if the critical point is preperiodic to an unstable cycle, the grammar is complete. The simplest example is the case in which the tent map critical point is preperiodic to a unimodal map 3-cycle, in which case the grammar is of golden mean type, with _00_ substring prohibited (see figure 11.14). In case at hand, the "critical" point is the junction of branches 4 and 5 (symmetry automatically takes care of the other critical point, at the junction of branches 6 and 7), and for the
cases considered the critical point maps into the endpoint of each of the seven branches.

One can fill out parameter $a$ axis arbitrarily densely with such points each of the 7 primary intervals can be subdivided into 7 intervals obtained by 2 -nd iterate of the map, and for the critical point mapping into any of those in 2 steps the grammar (and the corresponding cycle expansion) is finite, and so on.

## O.1.3 Diffusion coefficient, numerically

## (optional:)

Attempt a numerical evaluation of

$$
\begin{equation*}
D=\frac{1}{2} \lim _{n \rightarrow \infty} \frac{1}{n}\left\langle\hat{x}_{n}^{2}\right\rangle . \tag{O.11}
\end{equation*}
$$

Study the convergence by comparing your numerical results to the exact answers derived above. Is it better to use few initial $\hat{x}$ and average for long times, or to use many initial $\hat{x}$ for shorter times? Or should one fit the distribution of $\hat{x}^{2}$ with a Gaussian and get the $D$ this way? Try to plot dependence of $D$ on $\Lambda$; perhaps blow up a small region to show that the dependance of $D$ on the parameter $\Lambda$ is fractal. Compare with figure 23.5 and figures in refs. [O.1, O.2, 23.7, 23.8].

## O.1.4 $D$ is a nonuniform function of the parameters

## (optional:)

The dependence of $D$ on the map parameter $\Lambda$ is rather unexpected - even though for larger $\Lambda$ more points are mapped outside the unit cell in one iteration, the diffusion constant does not necessarily grow. An interpretation of this lack of monotonicity would be interesting.

You can also try applying periodic orbit theory to the sawtooth map (O.1) for a random "generic" value of the parameter $\Lambda$, for example $\Lambda=6$. The idea is to bracket this value of $\Lambda$ by the nearby ones, for which higher and higher iterates of the critical value $a=(\Lambda+1) / 4$ fall onto the partition boundaries, compute the exact diffusion constant for each such approximate Markov partition, and study their convergence toward the value of $D$ for $\Lambda=6$. Judging how difficult such problem is already for a tent map (see sect. 13.6 and appendix E.1), this is too ambitious for a week-long exam.

## References

[O.1] H.-C. Tseng, H.-J. Chen, P.-C. Li, W.-Y. Lai, C.-H. Chou and H.-W. Chen, "Some exact results for the diffusion coefficients of maps with pruned cycles", Phys. Lett. A 195, 74 (1994).
[O.2] C.-C. Chen, "Diffusion Coefficient of Piecewise Linear Maps", Phys. Rev. E51, 2815 (1995).
[O.3] H.-C. Tseng and H.-J. Chen, "Analytic results for the diffusion coefficient of a piecewise linear map", Int. J. Mod. Phys.B 10, 1913 (1996).

## O. 2 Deterministic diffusion, sawtooth map

To illustrate the main idea of chapter 23 , tracking of a globally diffusing orbit by the associated confined orbit restricted to the fundamental cell, we consider in more detail the class of simple 1-d dynamical systems, chains of piecewise linear maps (23.9). The translational symmetry (23.10) relates the unbounded dynamics on the real line to the dynamics restricted to a "fundamental cell" - in the present example the unit interval curled up into a circle. The corresponding circle map $f(x)$ is obtained by modulo the integer part. The elementary cell map $f(x)$ is sketched in figure 23.3. The map has the symmetry property

$$
\begin{equation*}
\hat{f}(\hat{x})=-\hat{f}(-\hat{x}) \tag{O.12}
\end{equation*}
$$

so that the dynamics has no drift, and all odd derivatives of the generating function (23.3) with respect to $\beta$ evaluated at $\beta=0$ vanish.

The cycle weights are given by

$$
\begin{equation*}
t_{p}=z^{n_{p}} \frac{e^{\beta \hat{n}_{p}}}{\left|\Lambda_{p}\right|} \tag{O.13}
\end{equation*}
$$

The diffusion constant formula for 1-d maps is

$$
\begin{equation*}
D=\frac{1}{2} \frac{\left\langle\hat{n}^{2}\right\rangle_{\zeta}}{\langle n\rangle_{\zeta}} \tag{O.14}
\end{equation*}
$$

where the "mean cycle time" is given by

$$
\begin{equation*}
\langle n\rangle_{\zeta}=\left.z \frac{\partial}{\partial z} \frac{1}{\zeta(0, z)}\right|_{z=1}=-\sum^{\prime}(-1)^{k} \frac{n_{p_{1}}+\cdots+n_{p_{k}}}{\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right|} \tag{O.15}
\end{equation*}
$$

the mean cycle displacement squared by

$$
\begin{equation*}
\left\langle\hat{n}^{2}\right\rangle_{\zeta}=\left.\frac{\partial^{2}}{\partial \beta^{2}} \frac{1}{\zeta(\beta, 1)}\right|_{\beta=0}=-\sum^{\prime}(-1)^{k} \frac{\left(\hat{n}_{p_{1}}+\cdots+\hat{n}_{p_{k}}\right)^{2}}{\left|\Lambda_{p_{1}} \cdots \Lambda_{p_{k}}\right|} \tag{O.16}
\end{equation*}
$$

and the sum is over all distinct non-repeating combinations of prime cycles. Most of results expected in this projects require no more than pencil and paper computations.

## O.2.1 The full shift

Reproduce the formulas of sect. 23.2.1 for the diffusion constant $D$ for $\Lambda$ both even and odd integer.

| figure 23.4 | $\Lambda$ | $D$ |
| :---: | :---: | :---: |
|  | 4 | $\frac{1}{4}$ |
| (a) | $2+\sqrt{6}$ | $1-\frac{3}{4} \sqrt{6}$ |
| (b) | $2 \sqrt{2}+2$ | $\frac{15+2 \sqrt{2}}{16+4 \sqrt{2}}$ |
| (c) | 5 | 1 |
| (d) | $3+\sqrt{5}$ | $\frac{5}{2} \frac{\Lambda-1}{3 \Lambda-4}$ |
| (e) | $3+\sqrt{7}$ | $\frac{5 \Lambda-4}{3 \Lambda-2}$ |
|  | 6 | $\frac{5}{6}$ |

Table O.2: The diffusion constant as function of the slope $\Lambda$ for the $\Lambda=4,6$ values of (23.20) and the 5 Markov partitions like the one indicated in figure 23.4.

## O.2.2 Subshifts of finite type

We now work out examples when the partition is Markov, although the slope is not an integer number. The key step is that of having a partition where intervals are mapped onto unions of intervals.

Start by reproducing the formula (23.28) of sect. 23.2 .3 for the diffusion constant $D$ for the Markov partition, the case where the critical point is mapped onto the right border of $I_{1_{+}}$.

Think up other non-integer values of the parameter $\Lambda$ for which the symbolic dynamics is given in terms of Markov partitions: in particular consider the remaing four cases for which the critical point is mapped onto a border of a partion in one iteration. Work out the Markov graph symmetrization factorization and the diffusion constant, and check the material flow conservation for each case. Fill in the diffusion constants missing in table O.2. It is not clear why the final answers tend to be so simple. What value of $\Lambda$ appears to yield the maximal diffusion constant?

The 7 cases considered here (see table 0.2 and figure 23.4) are the 7 simplest complete Markov partitions in the $4 \leq \Lambda \leq 6$ interval, the criterion being that the critical points map onto partition boundary points. In case at hand, the "critical" point is the highest point of the left branch of the map (symmetry automatically takes care of the other critical point, the lowest point of the left branch), and for the cases considered the critical point maps into the endpoint of each of the seven branches.

One can fill out parameter $a$ axis arbitrarily densely with such points each of the 6 primary intervals can be subdivided into 6 intervals obtained by 2 -nd iterate of the map, and for the critical point mapping into any of those in 2 steps the grammar (and the corresponding cycle expansion) is finite, and so on.

## O.2.3 Diffusion coefficient, numerically

## (optional:)

Attempt a numerical evaluation of

$$
\begin{equation*}
D=\frac{1}{2} \lim _{n \rightarrow \infty} \frac{1}{n}\left\langle\hat{x}_{n}^{2}\right\rangle . \tag{O.17}
\end{equation*}
$$

Study the convergence by comparing your numerical results to the exact answers derived above. Is it better to use few initial $\hat{x}$ and average for long times, or to use many initial $\hat{x}$ for shorter times? Or should one fit the distribution of $\hat{x}^{2}$ with a Gaussian and get the $D$ this way? Try to plot dependence of $D$ on $\Lambda$; perhaps blow up a small region to show that the dependance of $D$ on the parameter $\Lambda$ is fractal. Compare with figure 23.5 and figures in refs. [O.1, O.2, 23.7, 23.8].

## O.2.4 $D$ is a nonuniform function of the parameters

## (optional:)

The dependence of $D$ on the map parameter $\Lambda$ is rather unexpected even though for larger $\Lambda$ more points are mapped outside the unit cell in one iteration, the diffusion constant does not necessarily grow. Figure 23.5 taken from ref. [23.7] illustrates the fractal dependence of diffusion constant on the map parameter. An interpretation of this lack of monotonicity would be interesting.

You can also try applying periodic orbit theory to the sawtooth map (23.9) for a random "generic" value of the parameter $\Lambda$, for example $\Lambda=$ 4.5. The idea is to bracket this value of $\Lambda$ by the nearby ones, for which higher and higher iterates of the critical value $a=\Lambda / 2$ fall onto the partition boundaries, compute the exact diffusion constant for each such approximate Markov partition, and study their convergence toward the value of $D$ for $\Lambda=4.5$. Judging how difficult such problem is already for a tent map (see sect. 13.6 and appendix E.1), this is too ambitious for a week-long exam.


[^0]:    1.2
    page 30

[^1]:    [运 sect. 9.1

[^2]:    11.8
    page 185

[^3]:    ${ }^{1}$ Actually, these are double-letter symbols as $\sigma_{i}$ and $l_{i}$ are only counted as a product.

[^4]:    ${ }^{3}$ It is really a pity, that in 1926 Schrödinger introduced the wave mechanics and blocked the development of Sommerfeld's concept.

