Noise is your friend

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Outline

- what this talk is about
 - knowing when to stop
- 2 deterministic partitions
 - idea #1: partition by periodic points
- dynamicist's view of noise
 - idea #2: evolve densities, not noisy trajectories
 - idea #3: for unstable directions, look back
- optimal partition hypothesis



Turbulence?

and what to do about it

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Turbulence: A walk through a repertoire of unstable recurrent patterns?

As a turbulent flow evolves, every so often we catch a glimpse of a familiar pattern:



For any finite spatial resolution, the system follows approximately for a finite time a pattern belonging to a finite alphabet of admissible patterns. The long term dynamics = a walk through the space of such unstable patterns.

Conclusion: Hof is Hope renewed for Hopf's Last Hope for a Theory of Turbulence

Kuramoto-Sivashinsky periodic solutions embody Hopf's vision: They form the repertoire of the recurrent spatio-temporal patterns that is being explored by the turbulent dynamics.

detailed dynamics horrible, but much less so than feared: pieced together from 1-d return maps (!)

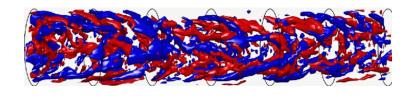
"To do" list:

Q: plane Couette-Taylor shear flow?

Waleffe; Kawahara & Kida: it can be done!

Q: Osborne Reynolds' turbulent pipe flow? need to predict turbulent drag for a given Re...

computation of unstable periodic orbits in high-dimensional state spaces, such as Navier-Stokes,



is at the border of what is feasible numerically, and criteria to identify finite sets of the most important solutions are very much needed.

when are we to stop calculating these solutions?

need the 3D velocity field at every (x, y, z)!

motions of fluids : require ∞ bits?

numerical simulations track 10^2 - 10^6 of computational degrees of freedom; terabytes of data, but how much information is there in all of this?

motions of fluids : require ∞ bits??

that cannot be right...

Science originates from curiosity and bad eyesight.

— Bernard de Fontenelle,

Entretiens sur la Pluralité des Mondes Habités

in practice

every physical problem is coarse partitioned and finite

noise rules the state space

- any physical system experiences (some kind of) noise
- any numerical computation is 'noisy'
- any prediction only needs a desired finite accuracy

mathematician's idealized state space

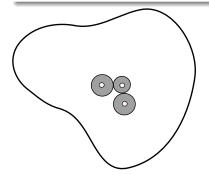
a manifold $\mathcal{M} \in \mathbb{R}^d$: d real numbers determine the state of the system $x \in \mathcal{M}$

noise-limited state space

a 'grid' \mathcal{M}' : N discrete states of the system $a \in \mathcal{M}'$, one for each noise covariance ellipsoid Δ_a

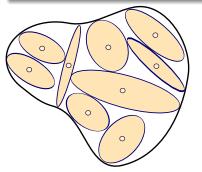
noise limited state space partitions

noise limited cell



a resolvable neighborhood is no smaller than a ball whose radius is the noise amplitude

noise limited partition grid



state space noise-partitioned into neighborhoods indicated by their centers

reasonable to assume that the noise

limits the resolution that can be attained in partitioning the state space

reasonable to assume that the noise

is uniform, leading to a uniform grid partition of the state space

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is uniform, leading to a uniform grid partition of the state space

in dynamics, this is wrong!

noise has memory

noise memory

accumulated noise along dynamical trajectories always coarsens the partition nonuniformly

noise memory

accumulated noise along dynamical trajectories always coarsens the partition nonuniformly

that is good, because

dynamics + noise determine

the **finest attainable** partition

the challenge

turbulence.zip

or 'equation assisted' data compression:

replace the ∞ of turbulent videos by the best possible

small finite set

of videos encoding all physically distinct motions of the turbulent fluid

devil is in the details

fluid dynamics

have equations: can compute the optimal partition

Navier-Stokes

$$\frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla)\mathbf{v} = \frac{1}{B} \nabla^2 \mathbf{v} - \nabla \rho + \mathbf{f}, \qquad \nabla \cdot \mathbf{v} = 0,$$

velocity field $\mathbf{v} \in \mathbb{R}^3$; pressure field p ; driving force \mathbf{f}

dynamical system

state space

a manifold $\mathcal{M} \in \mathbb{R}^d$: d numbers determine the state of the system

representative point

 $x(t) \in \mathcal{M}$

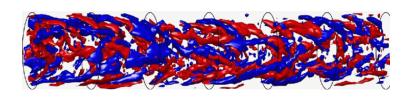
a state of physical system at instant in time

today's experiments

example of a representative point

 $x(t) \in \mathcal{M}, \, d = \infty$ a state of turbulent pipe flow at instant in time

Stereoscopic Particle Image Velocimetry \rightarrow 3-d velocity field over the entire pipe¹

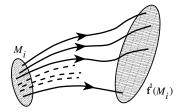


¹Casimir W.H. van Doorne (PhD thesis, Delft 2004)

dynamics

map $f^t(x_0)$ = representative point time t later

evolution in time



 f^t maps a region \mathcal{M}_i of the state space into the region $f^t(\mathcal{M}_i)$

dynamical description of turbulence

dynamical system

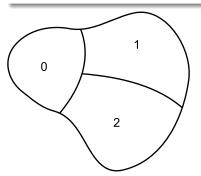
the pair (\mathcal{M}, f)

the problem

enumerate, classify all solutions of (\mathcal{M}, f)

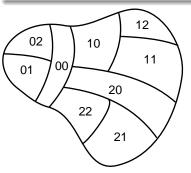
deterministic partition into regions of similar states

1-step memory partition



 $\mathcal{M} = \mathcal{M}_0 \cup \mathcal{M}_1 \cup \mathcal{M}_2$ ternary alphabet $\mathcal{A} = \{1, 2, 3\}.$

2-step memory refinement



 $\mathcal{M}_i = \mathcal{M}_{i0} \cup \mathcal{M}_{i1} \cup \mathcal{M}_{i2}$ labeled by nine 'words' $\{00, 01, 02, \cdots, 21, 22\}.$

deterministic partitions are no good

deterministic dynamics: partitioning can be arbitrarily fine requires exponential # of exponentially small regions

deterministic partitions are no good

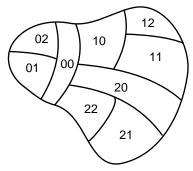
deterministic dynamics: partitioning can be arbitrarily fine requires exponential # of exponentially small regions

yet

in practice

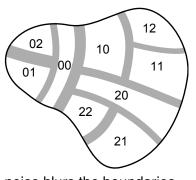
every physical problem must be coarse partitioned

deterministic vs. noisy partitions



deterministic partition

can be refined ad infinitum



noise blurs the boundaries when overlapping, no further refinement of partition

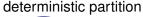
periodic points	instead	of boun	daries

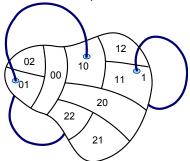
• mhm, do not know how to compute boundaries...

periodic points instead of boundaries

- mhm, do not know how to compute boundaries...
- however, each partition contains a short periodic point

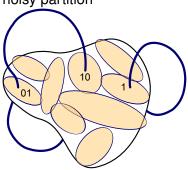
periodic orbit partition





some short periodic points: fixed point $\overline{1} = \{x_1\}$ two-cycle $\overline{01} = \{x_{01}, x_{10}\}\$

noisy partition



periodic points blurred by noise into cigar-shaped densities

periodic points and their cigar	S

 each partition contains a short periodic point smeared into a 'cigar' by noise

periodic points and their cigars

- each partition contains a short periodic point smeared into a 'cigar' by noise
- compute the size of a noisy periodic point neighborhood!

how big is the neighborhood blurred by the accumulated noise?

the (well known) **key formula** that we now derive:

$$Q_{n+1} = M_n Q_n M_n^T + \Delta_n$$

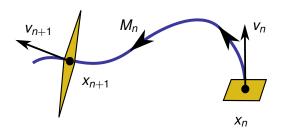
density covariance matrix at time n: Q_n

noise covariance matrix: Δ_n

Jacobian matrix of linearized flow: M_n

Lyapunov equation, doctoral dissertation 1892 Ornstein-Uhlenbeck 1930 Kalman filter 'prediction' 1960

linearized deterministic flow



$$x_{n+1} + z_{n+1} = f(x_n) + M_n z_n$$
, $M_{ij} = \partial f_i / \partial x_j$

in one time step a linearized neighborhood of x_n is

- (1) advected by the flow
- (2) transported by the Jacobian matrix M_n into a neighborhood given by the M eigenvalues and eigenvectors

covariance advection

let the initial density of deviations z from the deterministic center be a Gaussian whose covariance matrix is

$$Q_{jk} = \left\langle z_j z_k^T \right\rangle$$

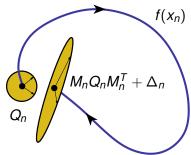
a step later the Gaussian is advected to

$$\left\langle z_j z_k^T \right\rangle \ \rightarrow \ \left\langle (M z)_j (M z)_k^T \right\rangle$$

$$Q \ \rightarrow \ M Q M^T$$

next: add noise

roll your own cigar



in one time step a Gaussian density distribution with covariance matrix Q_n is

- (1) advected by the flow
- (2) smeared with additive noise

into a Gaussian 'cigar' whose widths and orientation are given by the singular values and vectors of Q_{n+1}

covariance evolution

$$Q_{n+1} = M_n Q_n M_n^T + \Delta_n$$

- advect deterministically local density covariance matrix Q → MQM^T
- (2) add noise covariance matrix Δ

covariances add up as sums of squares

cumulative noise along a trajectory

iterate
$$Q_{n+1} = M_n Q_n M_n^T + \Delta_n$$
 along a trajectory

if M is contracting, $|\Lambda_i| < 1$,

the memory of the covariance Q_0 of the starting density is lost, with iteration leading to the limit distribution

$$Q_n = \Delta_n + M_{n-1}\Delta_{n-1}M_{n-1}^T + M_{n-2}^2\Delta_{n-2}(M_{n-2}^2)^T + \cdots$$

example: noise and a single attractive fixed point

if all eigenvalues of M are strictly contracting, all $|\Lambda_j| < 1$ any initial compact measure converges to the unique invariant Gaussian measure $\rho_0(z)$ whose covariance matrix satisfies

Lyapunov equation: time-invariant measure condition

$$Q = MQM^T + \Delta$$

[A. M. Lyapunov doctoral dissertation 1892]

example: Ornstein-Uhlenbeck process

width of the natural measure concentrated at the attractive deterministic fixed point z=0

$$ho_0(z) = rac{1}{\sqrt{2\pi\,Q}}\,\exp\left(-rac{z^2}{2\,Q}
ight)\,, \qquad Q = rac{\Delta}{1-|\Lambda|^2}\,,$$

- is balance between contraction by Λ and noisy smearing by Δ at each time step
- for strongly contracting Λ, the width is due to the noise only
- As $|\Lambda| \to 1$ the width diverges: the trajectories are no longer confined, but diffuse by Brownian motion

214105-7 Nakanishi, Sakaue, and Wakou

J. Chem. Phys. 139, 214105 (2013)

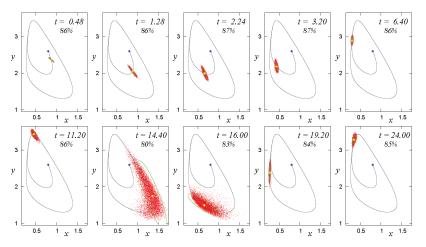


FIG. 2. Time development of distribution for Brusselator. 10 000 samples of Monte Carlo simulations are plotted by the red dots along with the covariance matrix \hat{M} estimated by Eq. (E7); \hat{M} 's are represented by the green ellipses given by δx^2 , \hat{M}^2 - $\delta x = 4/\Omega$, where $\delta x^2 \equiv (x - x^*(t), y - y^*(t))$. The percentages of the samples that fall within the ellipses are shown in each panel. The gray curves represent the trajectory by the rate equation starting from the initial point marked by the blue circles. The system parameters are $k_1 = 0.5$, $k_2 = 1.5$, $k_3 = 1.0$, $k_4 = 1.0$, and $\Omega = 10^4$. The initial point is $(x_3^2, y_4^2) = (0.8, 2.6)$.

remembrance of things past

noisy dynamics of a nonlinear system is fundamentally different from Brownian motion, as the flow ALWAYS induces a local, history dependent effective noise things fall apart, centre cannot hold

but what if *M* has *expanding* eigenvalues?

both deterministic dynamics and noise tend to smear densities away from the fixed point: no peaked Gaussian in your future things fall apart, centre cannot hold

but what if *M* has *expanding* eigenvalues?

look into the past, for initial peaked distribution that spreads to the present state for unstable directions, look back

if *M* has only *expanding* eigenvalues,

balance between the two is attained by iteration from the past, and the evolution of the covariance matrix \tilde{Q} is now given by

$$\tilde{Q}_{n+1} + \Delta_n = M_n \tilde{Q}_n M_n^T \,,$$

[aside to control theorists: reachability and observability Gramians]

solving the Lyapunov equation

iterate $Q_{n+1} = M_n Q_n M_n^T + \Delta_n$ attractive fixed point, $Q = Q_{\infty}$, $M = M_n$, $Q = Q_n$:

$$Q = \Delta + M\Delta M^{\top} + M^{2}\Delta(M^{\top})^{2} + \cdots = \sum_{m,n=0}^{\infty} \delta_{mn}M^{n}\Delta(M^{\top})^{m}$$

bring to resolvent form, $\delta_{mn} = \int_0^{2\pi} \frac{d\theta}{2\pi} e^{i\theta(m-n)}$

for *M* contracting, expanding, or hyperbolic (!)

$$Q = \int_0^{2\pi} \frac{d\theta}{2\pi} \frac{1}{\mathbf{1} - e^{-i\theta}M} \Delta \frac{1}{\mathbf{1} - e^{i\theta}M^{\top}}$$

Cauchy magic

a similarity transformation S separates the expanding and contracting subspaces

$$\Lambda \equiv S^{-1}MS = \left(egin{array}{cc} \Lambda_e & 0 \\ 0 & \Lambda_c \end{array}
ight)$$

transformed noise covariance matrix

$$\hat{\Delta} \equiv \mathcal{S}^{-1} \Delta (\mathcal{S}^{-1})^{ op} = \left(egin{array}{cc} \Delta_{ee} & \Delta_{ec} \ \Delta_{ce} & \Delta_{cc} \end{array}
ight)$$

Cauchy magic

contour integral representation

$$Q = \oint_{\Gamma} \frac{ds}{2\pi} (\mathbf{1} - s^{-1}M)^{-1} \Delta (\mathbf{1} - sM)^{-1}$$

separates Q into expanding and contracting covariances:

$$ilde{Q}_e \equiv S \left(egin{array}{cc} Q_e & 0 \ 0 & 0 \end{array}
ight) S^ op \, , \quad Q_c \equiv S \left(egin{array}{cc} 0 & 0 \ 0 & Q_c \end{array}
ight) S^ op$$

two stationary 'cigars', one in the expanding manifold and the other in the contracting manifold (not orthogonal to each other!)

local problem solved: can compute every cigar

a periodic point of period n is a fixed point of nth iterate of dynamics

global problem solved: can compute all cigars

more algebra: can compute the noisy neighborhoods of all periodic points

optimal partition challenge

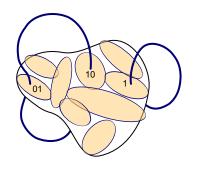
now can address our challenge:

determine the finest possible partition for a given noise

noisy dynamics partitions: strategy

- use periodic orbits to partition state space
- compute local covariances at periodic points to determine their neighborhoods
- done once neighborhoods overlap

optimal partition hypothesis



optimal partition:

the maximal set of resolvable periodic point neighborhoods

how noise frees us from determinism

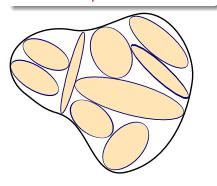
noise memory

accumulated noise along a dynamical trajectory always coarsens the partition

this partition is

- intrinsic to dynamics
- computable

turbulence.zip



example: representative solutions of fluid dynamics

- Professor Zweistein, from the back of the auditorium:
 - (1) she has already done all this in 1969
 - (2) you must be kidding, it cannot be done for turbulence

example: representative solutions of fluid dynamics

- Professor Zweistein, from the back of the auditorium:
 - (1) she has already done all this in 1969
 - (2) you must be kidding, it cannot be done for turbulence
- OK, OK, today we typically have about 10-100 exact invariant solutions populating the state space
 - and we have their Jacobians M (that was hell to get)



• we have not yet tested the method on fluid dynamics

disclosure

- we have not yet tested the method on fluid dynamics
- the brave candidates: step up after the talk

take home message

computation of unstable periodic orbits in high-dimensional state spaces, such as Navier-Stokes,

is at the border of what is feasible numerically, and criteria to identify finite sets of the most important solutions are very much needed

we are to stop calculating these solutions when we attain

take home message

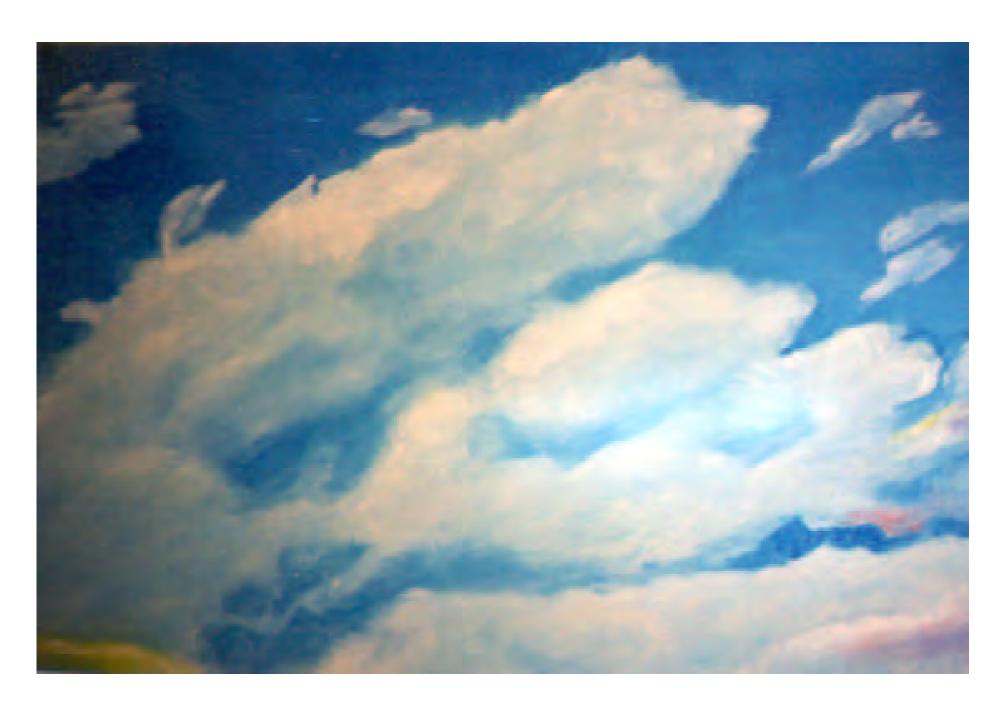
optimal partition hypothesis

take home message

optimal partition hypothesis

- the best of all possible state space partitions
- optimal for the given dynamical system, the given noise

Is this a cloud?



references

- D. Lippolis and P. Cvitanović, How well can one resolve the state space of a chaotic map?, Phys. Rev. Lett. 104, 014101 (2010); arXiv.org:0902.4269
- P. Cvitanović and D. Lippolis, Knowing when to stop: How noise frees us from determinism, in M. Robnik and V.G. Romanovski, eds., Let's Face Chaos through Nonlinear Dynamics (Am. Inst. of Phys., Melville, New York, 2012); arXiv.org:1206.5506